

# Rb group workshop : Intro to turbulence for opticians

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# Summary

- Intro
- The language of turbulence : Navier-Stokes Equations
- From determinism to chaos
- Notion of scale : cascade and Kolmogorov Law
- 2D Turbulence : what changes ?
- Quantum turbulence

## Appendices :

- How to compute turbulent flows
- Beyond Kholmogorov theory

Intro : Where/Why/What

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# Where is turbulence ?



**everywhere...**

astrophysics,  
atmospheric physics,  
weather predictions,  
geophysics,  
engineering,  
aviation,  
industry, ...

# Where is turbulence ?



- Galaxies  
 $10^{23}$  m



- Nebula  
 $10^{18}$  m



- Solar Wind  
 $10^{11}$  m



- Stellar convection  
 $10^9$  m



- Atmosphere  
 $10^7$  m



- Ocean  
 $10^5$  m



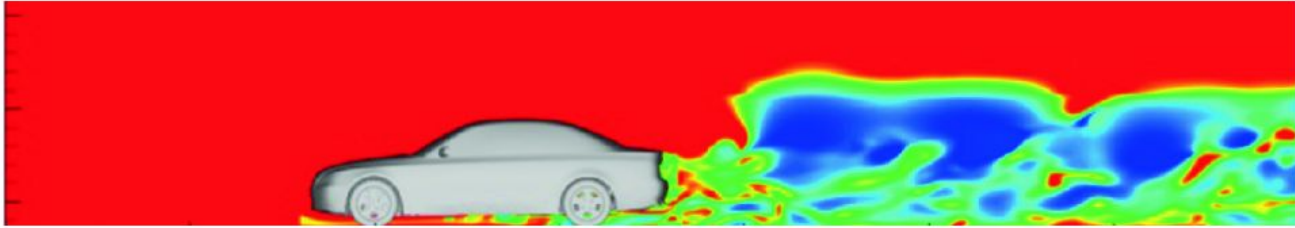
- Navy  
 $10^2$  m



- Coffee  
 $10^{-1}$  m

Can you think of other physical phenomena that exist in such a large diversity of scales ?

# Why is turbulence important ?



- In the absence of turbulence

$$U \simeq 72 \text{ km/h}, \quad L \simeq 2 \text{ m}, \quad \nu = 1.5 \cdot 10^{-5} \text{ m}^2/\text{s}, \quad \rho = 1.2 \text{ kg/m}^3$$

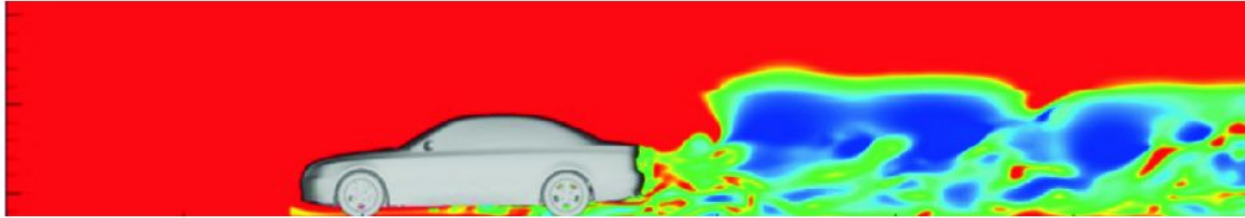
rate of energy dissipation (drag)

$$\epsilon \sim (\rho L^3)(\nu U^2/L^2) = 0.014 \text{ kgm}^2/\text{s}^3$$

- In the presence of turbulence rate of energy dissipation (drag)

$$\epsilon \sim (\rho L^3)(U^3/L) = 38000 \text{ kgm}^2/\text{s}^3$$

# Why is turbulence important ?



$$\begin{aligned}\frac{\text{Turbulent dissipation rate}}{\text{Laminar dissipation rate}} &= \frac{(\rho L^3)(U^3/L)}{(\rho L^3)(\nu U^2/L^2)} \\ &= \frac{UL}{\nu} \\ &= \mathbf{Re} \\ &= 2.6 \cdot 10^6\end{aligned}$$

$$Re = \frac{UL}{\nu} = \frac{L^2/\nu}{L/U} = \frac{\text{viscous time scale}}{\text{advection time scale}}$$

Laminar flow :

$$Re < 2300$$

Turbulent flow :

$$Re > 2900$$



# In the end... **what** is turbulence ?

**No clear commonly accepted definition of turbulence exists!**

Some observations

- Erratic (random) motions
- Unpredictability
- A large range of scales excited from large to small
- Structures exists but are **not** periodic in space or time
- Strong dissipation



- ← Turbulence
- ← First Bifurcations
- ← Laminar





Turbulence is the most important  
unsolved problem of classical  
physics.

— Richard P. Feynman —

AZ QUOTES

*"I am an old man now, and when I die and  
go to Heaven there are two matters on which  
I hope enlightenment. One is quantum  
electro-dynamics and the other is turbulence.*

*About the former, I am really rather  
optimistic"*



Sir Horace Lamb (1932)

From S. Goldstein, Ann. Rev. Fluid Mech, 1, 23 (1969)

*"What is turbulence ?  
Turbulence is like pornography.  
It is hard to define,  
but if you see it, you recognize it  
immediately"*

G. K. Vallis, 1999

# The language of turbulence : Navier-Stokes equations

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The Navier-Stokes (NS) equations :

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla P + \nu \nabla^2 \mathbf{u} + \mathbf{f}$$

Incompressibility condition

$$\nabla \cdot \mathbf{u} = 0$$

$$\nu = \mu / \rho \equiv \text{kinematic viscosity}$$

To study turbulence, we almost always assume **incompressibility**.  
(pseudo-incomp. / low Mach number)

# Breaking down Navier-Stokes



$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla P + \nu \nabla^2 \mathbf{u} + \mathbf{f}$$

$$\nabla \cdot \mathbf{u} = 0$$

- time derivative
- **Non-linear terms**: advection and pressure
- **Linear terms**: viscous terms responsible for energy dissipation
- **Non-homogeneous terms**: Forcing - energy injection
- **Divergence-free constraint**

# Other forms of Navier-Stokes

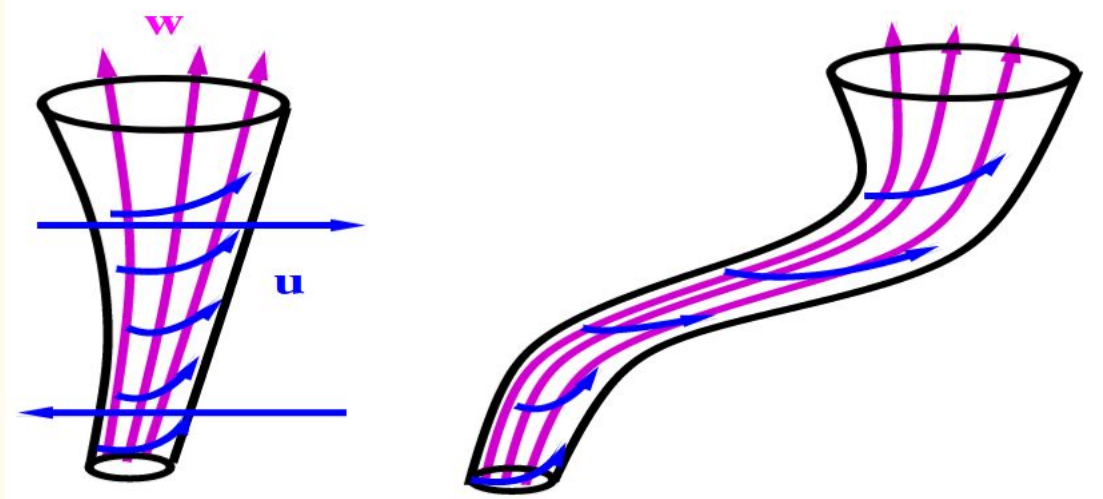
The **vorticity** equation :

$$\partial_t \mathbf{w} = \nabla \times (\mathbf{u} \times \mathbf{w}) + \nu \nabla^2 \mathbf{w} + \nabla \times \mathbf{f}$$

$$\text{where } \mathbf{w} = \nabla \times \mathbf{u}$$

We can visualize vorticity in a “tube” of current :

Vorticity field lines move with the flow.



# Other forms of Navier-Stokes

Navier-Stokes in Fourier space :

$$\partial_t \tilde{\mathbf{u}}_{\mathbf{k}} = - \sum_{\mathbf{p}+\mathbf{q}=\mathbf{k}} i \left( (\mathbf{p} \cdot \tilde{\mathbf{u}}_{\mathbf{q}}) \tilde{\mathbf{u}}_{\mathbf{p}} - \mathbf{k} \frac{(\mathbf{k} \cdot \tilde{\mathbf{u}}_{\mathbf{p}})(\mathbf{p} \cdot \tilde{\mathbf{u}}_{\mathbf{q}})}{|\mathbf{k}|^2} \right) - \nu |\mathbf{k}|^2 \tilde{\mathbf{u}}_{\mathbf{k}} + \tilde{\mathbf{f}}_{\mathbf{k}}$$

and

$$\mathbf{k} \cdot \mathbf{u}_{\mathbf{k}} = 0$$



Ugly ! BUT very convenient for calculations (spectral methods for resolution of NS)

# Conserved quantities

$$\langle f(\mathbf{x}) \rangle = \frac{1}{L_x L_y L_z} \int f(\mathbf{x}) dV$$

## Navier-Stokes

$$\partial_t \mathbf{u} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla P + \nu \nabla^2 \mathbf{u} + \mathbf{f}$$

Spatial  
average

$$\partial_t \langle \mathbf{u} \rangle = 0$$

Scalar product  
with  $\mathbf{u}$  and average

$$\frac{1}{2} \frac{d}{dt} \langle |\mathbf{u}|^2 \rangle = -\nu \langle \nabla_j u_i \nabla_j u_i \rangle + \langle \mathbf{u} \cdot \mathbf{f} \rangle$$

$$\frac{d}{dt} \mathcal{E} = -\epsilon + \mathcal{I}$$

Dissipation

Injection

Momentum and energy are  
**conserved** by the NS equations

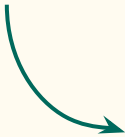


# Conserved quantities

## Vorticity equation

$$\partial_t \mathbf{w} = \nabla \times (\mathbf{u} \times \mathbf{w}) + \nu \nabla^2 \mathbf{w} + \nabla \times \mathbf{f}$$

Scalar product  
with  $\mathbf{u}$  and average



$$\frac{1}{2} \frac{d}{dt} \langle \mathbf{u} \cdot \mathbf{w} \rangle = -\nu \langle \mathbf{w} \cdot \nabla \times \mathbf{w} \rangle + \langle \mathbf{w} \cdot \mathbf{f} \rangle$$

We discover a new conserved quantity :

Helicity

$$\mathcal{H} = \frac{1}{2} \langle \mathbf{u} \cdot \mathbf{w} \rangle$$

$$\frac{d}{dt} \mathcal{H} = -\epsilon_{\mathcal{H}} + \mathcal{I}_{\mathcal{H}}$$



Dissipation

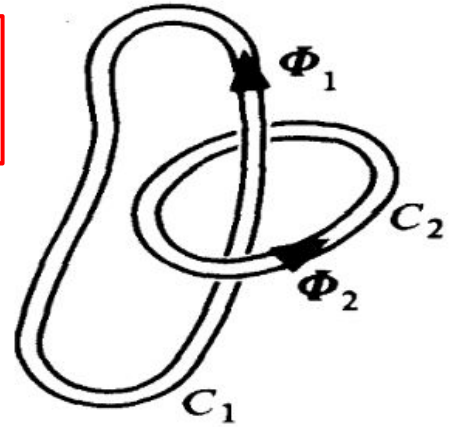


Injection

# Understanding Helicity

$$\mathcal{H} = \frac{1}{2} \langle \mathbf{u} \cdot \mathbf{w} \rangle$$

$$\begin{aligned}\mathcal{H} &= \frac{1}{2} \int_{V_1} \mathbf{u} \cdot \mathbf{w} dx^3 + \frac{1}{2} \int_{V_2} \mathbf{u} \cdot \mathbf{w} dx^3 \\ &= \frac{1}{2} \oint_{C_1} \int_{S_1} \mathbf{u} \cdot \mathbf{w} ds d\ell + \frac{1}{2} \oint_{C_2} \int_{S_2} \mathbf{u} \cdot \mathbf{w} ds d\ell \\ &= \frac{1}{2} \Phi_1 \oint_{C_1} \mathbf{u} \cdot d\ell + \frac{1}{2} \Phi_2 \oint_{C_2} \mathbf{u} \cdot d\ell \\ &= \frac{1}{2} \Phi_1 \oint_{A_1} \mathbf{w} \cdot d\mathbf{s} + \frac{1}{2} \Phi_2 \oint_{A_2} \mathbf{w} \cdot d\mathbf{s} \\ &= \Phi_1 \Phi_2\end{aligned}$$



Helicity is a measure of  
how much vortices  
**interact** !

Useful for turbulence...

# Symmetries of Navier-Stokes

- Space translations:  $x' = x + \ell$
- Time translations:  $t' = t + T$
- Galilean transformations:  $\mathbf{x}' = \mathbf{x} - \mathbf{c}t$ ,  $t' = t$ ,  $\mathbf{u}' = \mathbf{u} + \mathbf{c}$
- Rotations  $\mathbf{u}' = \mathcal{R}\mathbf{u}$ ,  $\mathbf{x}' = \mathcal{R}^{-1}\mathbf{x}$
- Parity (reflections):  $\mathbf{u}' = -\mathbf{u}$ ,  $\mathbf{x}' = -\mathbf{x}$
- Scaling  $\mathbf{x}' = \mathbf{x}/\lambda$ ,  $t' = t/\lambda^\alpha$ ,  $\mathbf{u}' = \lambda^\beta \mathbf{u}$   
 $\beta = -1$  and  $\alpha = 2$

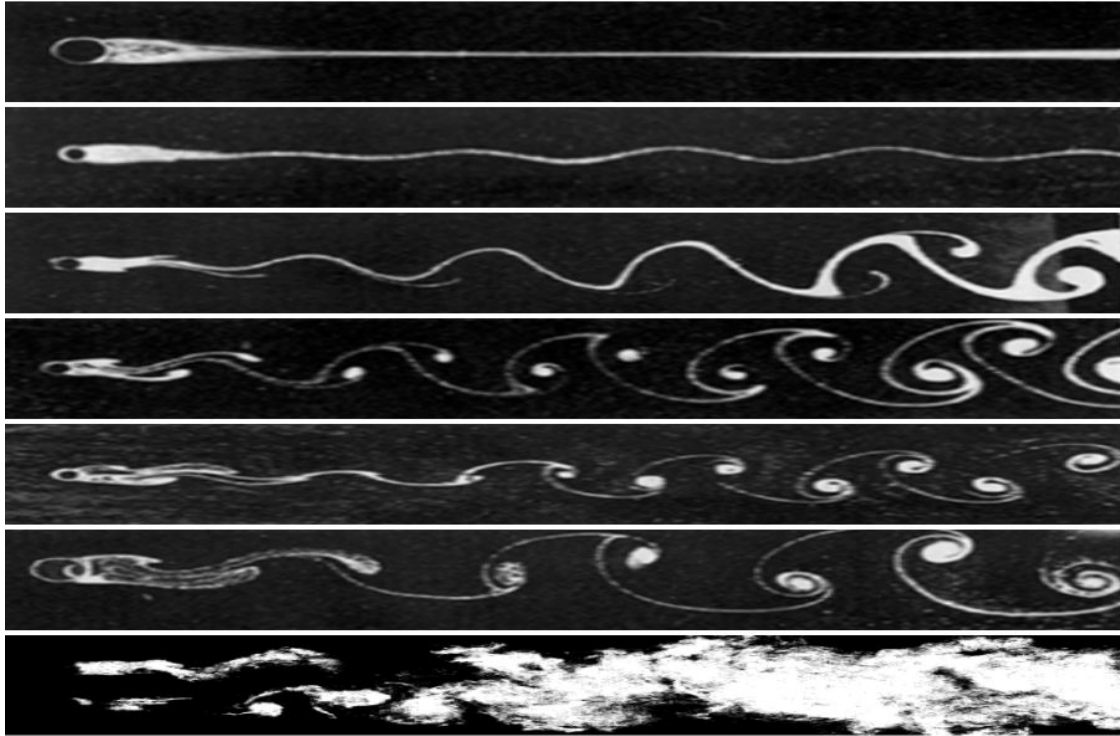
But how do we explain the apparent “**chaotic**” behavior of turbulent flow ?

From determinism to chaos

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# Breaking of symmetry

$$Re = \frac{UL}{\nu}$$



Re= 32

Re= 55

Re= 60

Re= 73

Re=102

Re=161

Re=1800

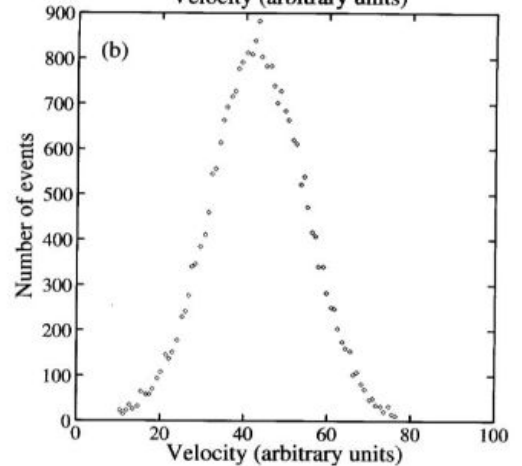
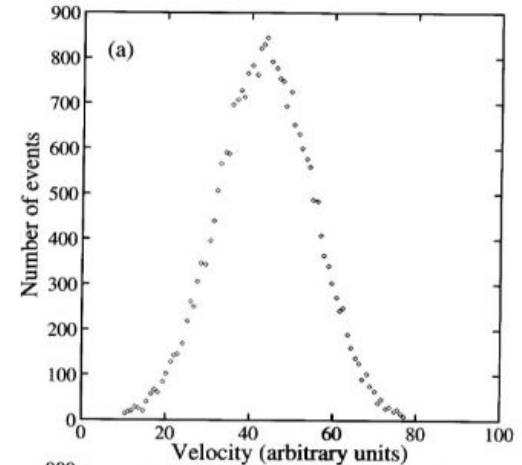
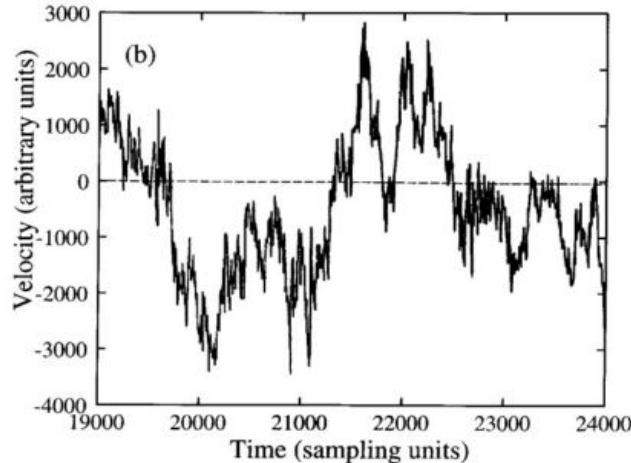
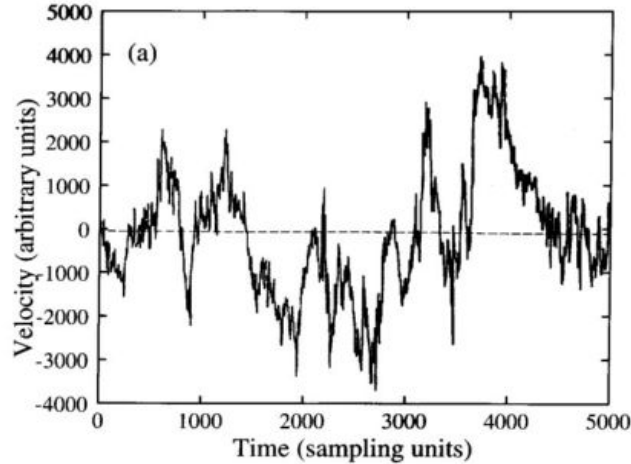
Flow is observed in a wind tunnel.

The flow first loses symmetry and then predictability...

# Statistical study

Here are velocity measurements through time in the wind tunnel.

Raw signals are not reproducible, but histograms (related to statistics) are.



# What causes chaos ?

We choose two initial

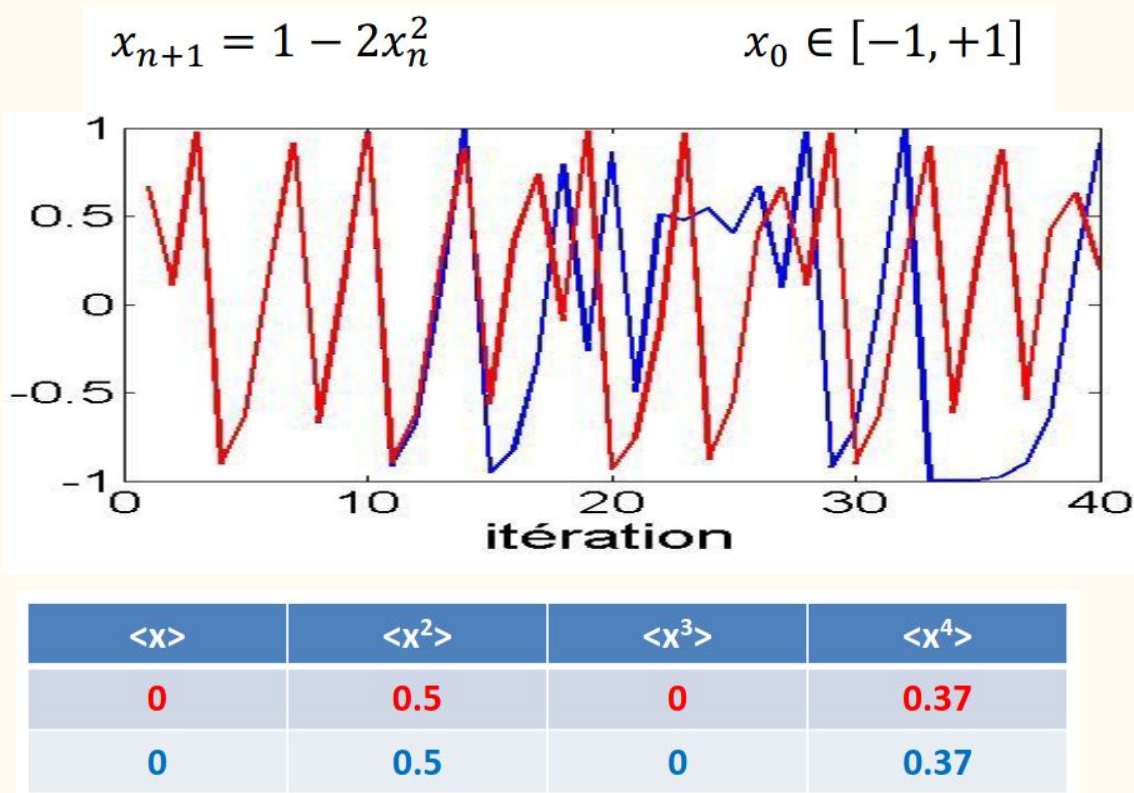
conditions :

$$x_0 = 0.6667$$

$$x_0 = 0.66667$$

Non linear term makes the system extremely sensitive to variations in the initial conditions : **chaos !**

Statistics however are **the same** :





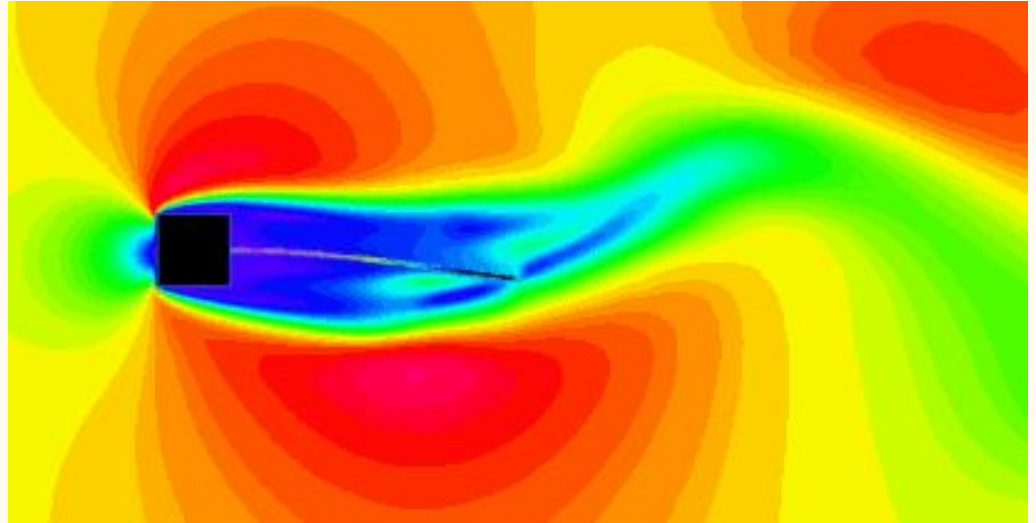
# Summary

- At very small  $r$ , ( $Re$ ) symmetric solutions are observed
- When  $r$ , ( $Re$ ) is increased symmetries are broken: ie observed individual solutions do not satisfy them
- Symmetries are recovered in a statistical sense
- Individual solution convey little information
- The system needs to be described in a statistical way:  $P(\mathbf{u})$

# Quizz !

Can there be turbulence in the well known Stokes equation ?

$$\partial_t \mathbf{u} = \nu \nabla^2 \mathbf{u} + \mathbf{f}$$



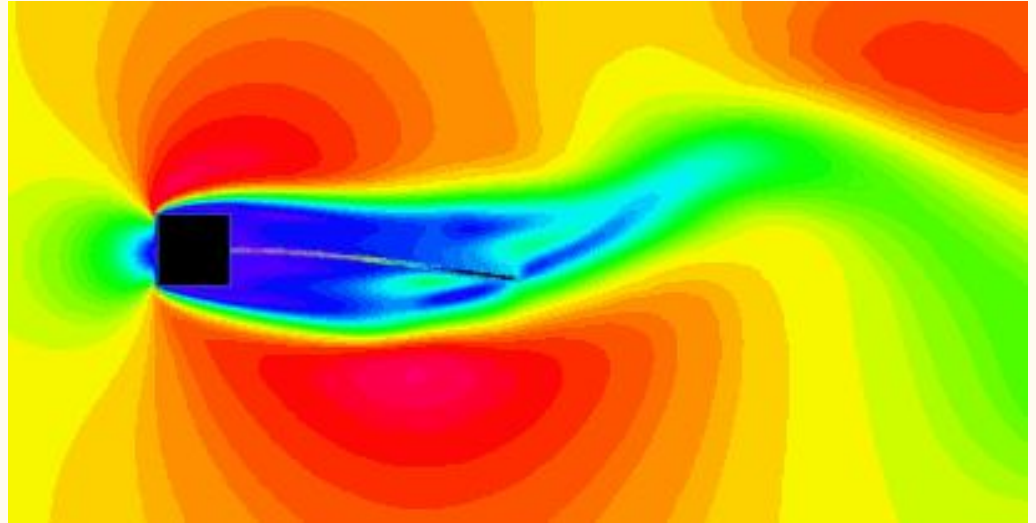
# Quizz !

Can there be turbulence in the well known Stokes equation ?

$$\partial_t \mathbf{u} = \nu \nabla^2 \mathbf{u} + \mathbf{f}$$

**Answer :**

**NO**, there is only linear terms in the Stokes equation, it describes flows with small velocity, small length-scale or high viscosity.



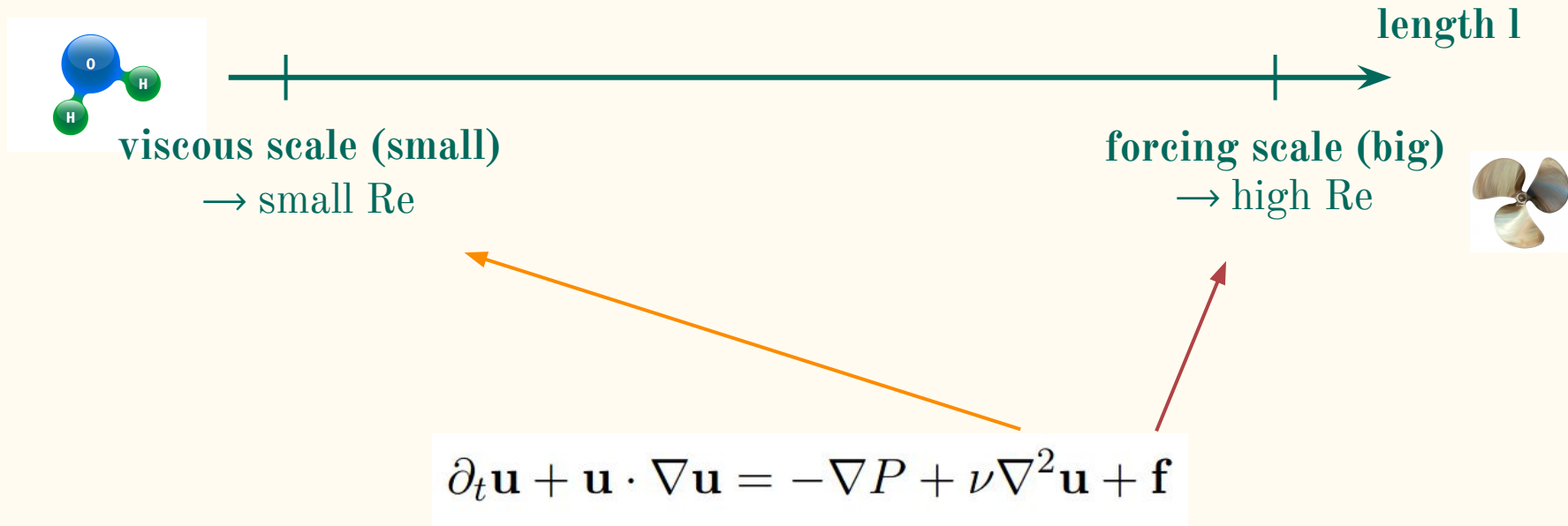
Notion of scale :  
Cascade and Kolmogorov law

—

# Intuition about scales

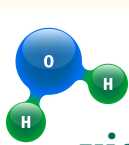
$$Re = \frac{UL}{\nu}$$

Turbulence covers many scales, what are they ?

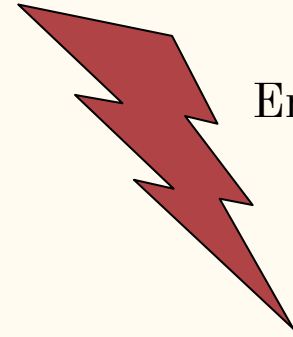


# Intuition about scales

Energy Dissipation



viscous scale (small)



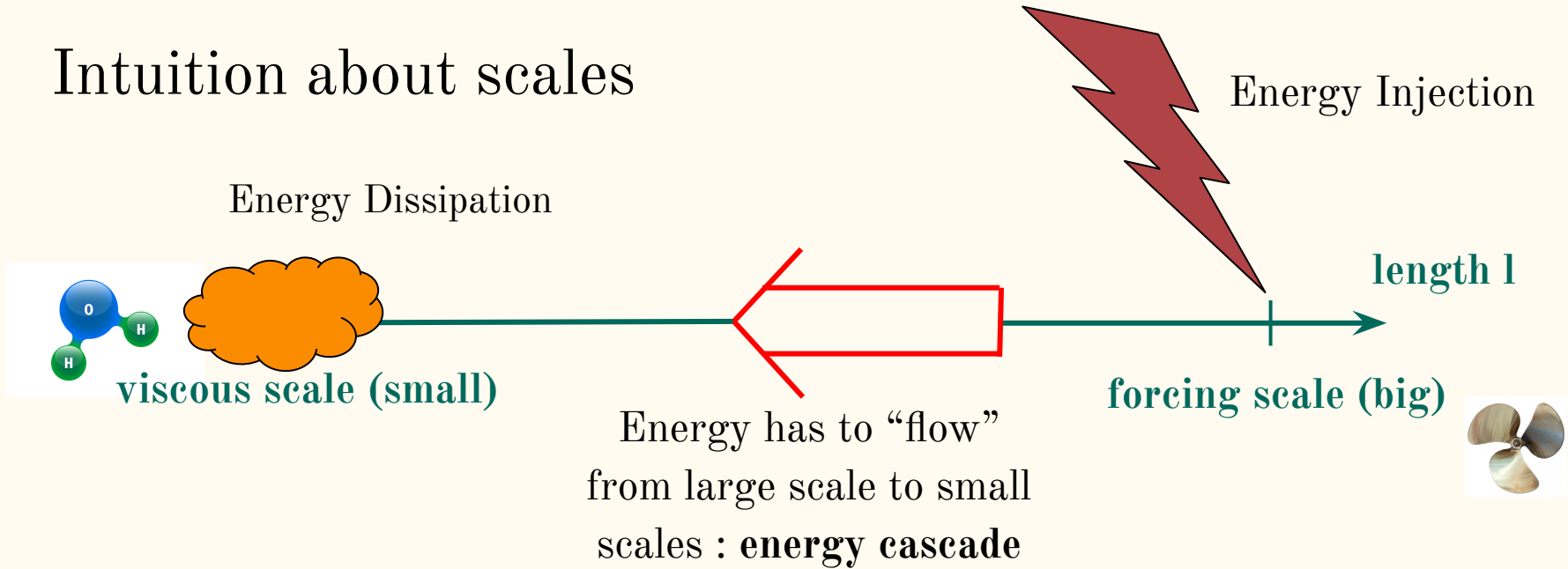
Energy Injection

length  $l$

forcing scale (big)



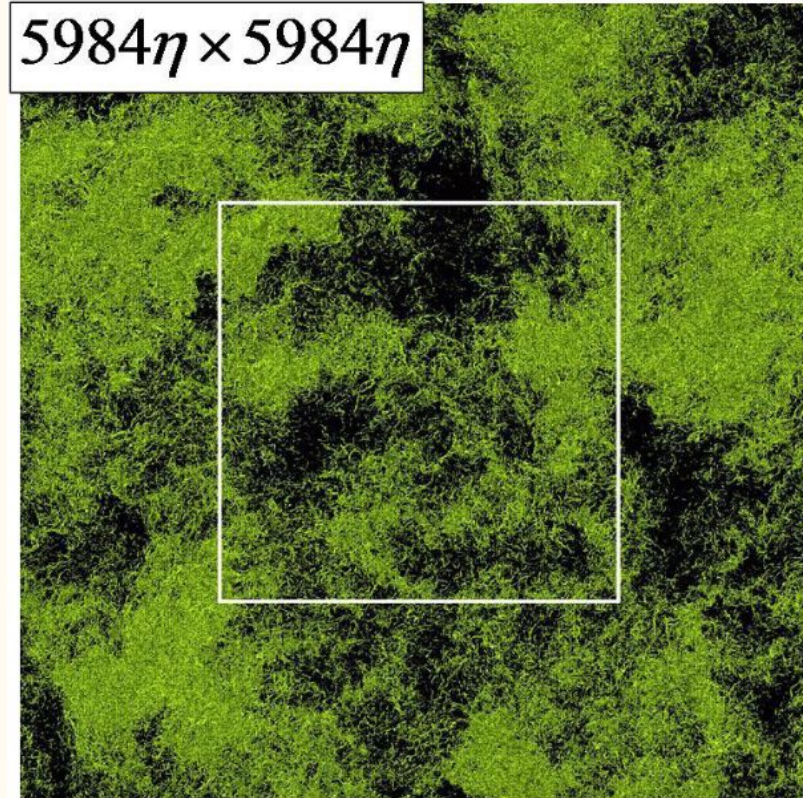
# Intuition about scales



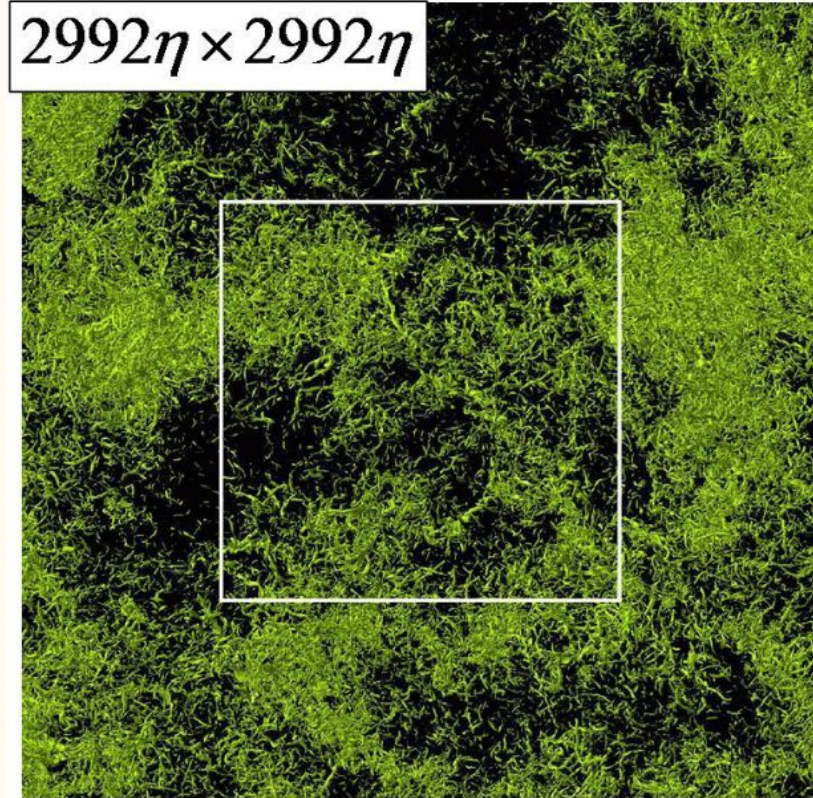


# What happens between scales ?

Vorticity plot for  
3D turbulent flow :

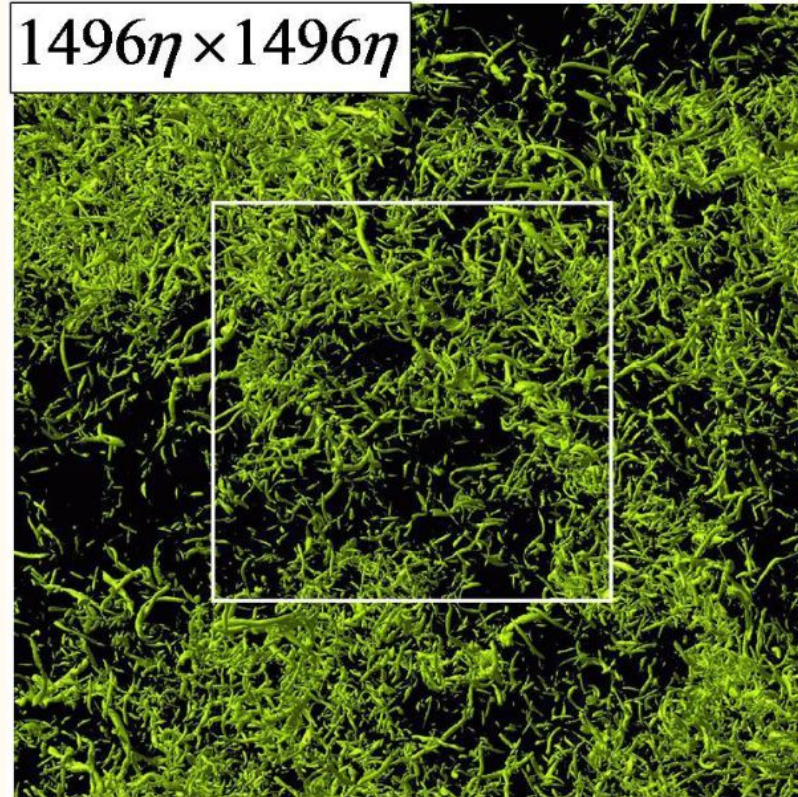


What happens between scales ?

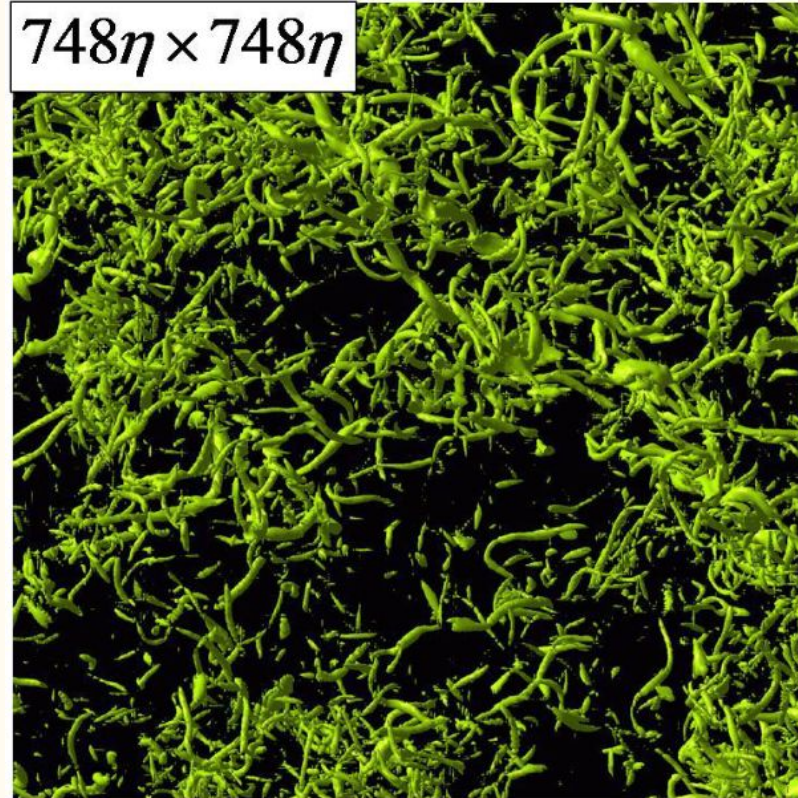




What happens between scales ?



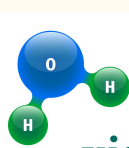
# What happens between scales ?



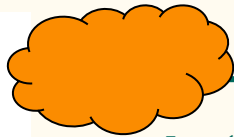
We observe some sort  
of **scale invariance...**

# Kolmogorov's assumption

Energy Dissipation



viscous scale (small)



The rate at which energy flows is **constant across scales**

Energy Injection

length  $l$

forcing scale (big)



$$\mathcal{I} = \Pi_\ell = \epsilon$$

Energy injection = energy flux = Energy dissipation

# Scale formulation of Navier-Stokes

We can use Fourier to define the scales : **scale**  $\ell \sim 1/k$

$$\mathbf{u}(\mathbf{x}, t) = \sum \tilde{\mathbf{u}}_{\mathbf{k}}(t) e^{i\mathbf{k} \cdot \mathbf{x}}, \quad \tilde{\mathbf{u}}_{\mathbf{k}}(t) = \langle \mathbf{u} e^{-i\mathbf{k} \cdot \mathbf{x}} \rangle$$

**Energy Spectrum**

$$E(k) = \frac{1}{2\delta k} \sum_{k \leq |\mathbf{k}| < k + \delta k} |\tilde{\mathbf{u}}_{\mathbf{k}}|^2 \sim \frac{\text{Energy}}{\text{per unit wavenumber}}$$

**Energy at scale**  $\ell \sim 1/k$

$$u_{\ell}^2 \propto k E(k)$$

# Calculating the spectrum

We know that the rate of energy dissipation scales as :

$$\epsilon \propto \frac{\text{energy}}{\text{time}} = \frac{u^2}{\tau}$$

And because :

$\mathcal{I}$	=	$\Pi_\ell$	=	$\epsilon$
Energy injection	=	energy flux	=	Energy dissipation

We have :

$$\Pi \propto \frac{u_\ell^2}{\tau_\ell} = \frac{u_\ell^3}{\ell}$$



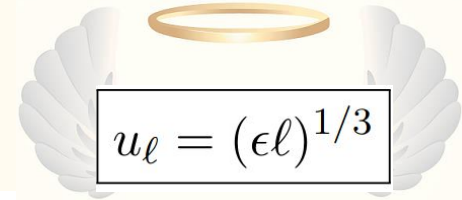
$u_\ell = (\epsilon \ell)^{1/3}$
----------------------------------

**Kolmogorov  
K41 law !!**

(1941)



# Calculating the spectrum


$$u_\ell = (\epsilon \ell)^{1/3}$$

The spectrum scales as :

$$E(k) \propto \frac{u_\ell^2}{k} = \frac{(\epsilon \ell)^{2/3}}{1/\ell} = \frac{(\epsilon/k)^{2/3}}{k}$$

$$E(k) \propto \epsilon^{2/3} k^{-5/3}$$

We can also find the scale at which dissipation takes control :

**From N.S.**

$$\nu \nabla^2 \mathbf{u}$$



$$\epsilon \propto \nu \frac{u_\nu^2}{\ell_\nu^2} \propto \nu \frac{(\epsilon^{1/3} \ell_\nu^{1/3})^2}{\ell_\nu^2} \propto \nu \frac{\epsilon^{2/3}}{\ell_\nu^{4/3}}$$

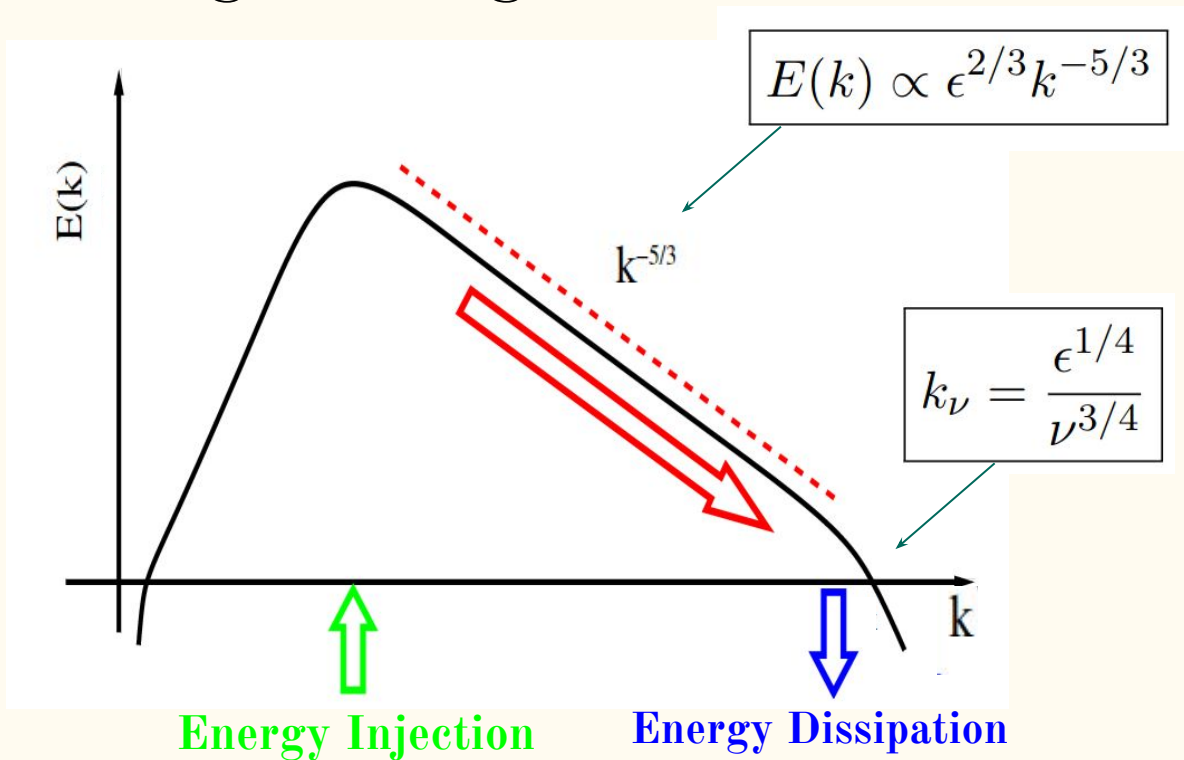


$$\ell_\nu = \frac{\nu^{3/4}}{\epsilon^{1/4}}$$

or

$$k_\nu = \frac{\epsilon^{1/4}}{\nu^{3/4}}$$

# Putting it all together



The turbulence spectrum is **UNIVERSAL** in the **inertial range**.

It only depends on  $\nu$  and  $\epsilon$  and we have :

$$\epsilon = \mathcal{I} = \frac{u_L^3}{L}$$

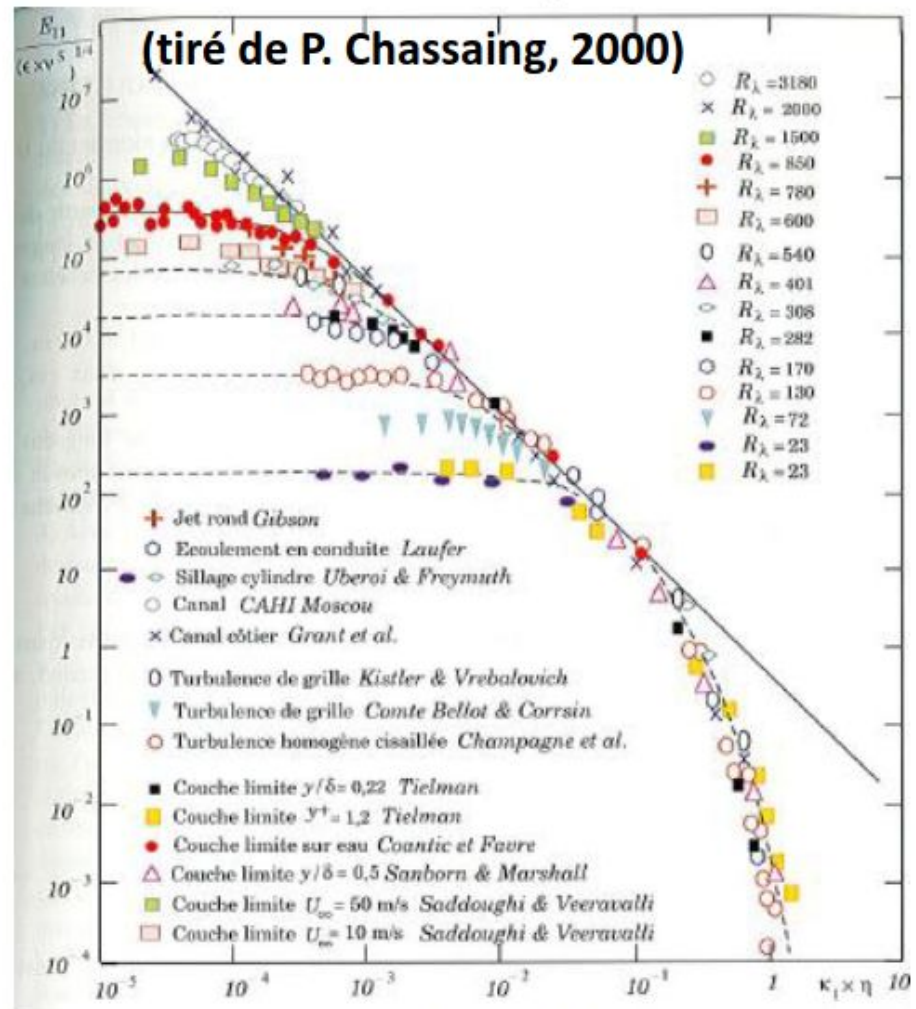
→ dissipation does not depend on  $\nu$  !

# Universality

We indeed observe the  $-5/3$  spectrum in a wide variety of real flows :

→ Kolmogorov hypothesis of self similarity **was correct** !

→ There is an energy cascade from large scales to small scale in turbulent flows !



In the end :

Turbulence is a process that **transports energy** injected into the system **at large scales to the small scales** where it is dissipated by viscosity.

But which term in Navier-Stokes is responsible for this “flux” :

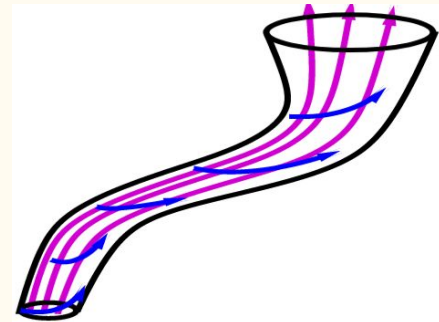
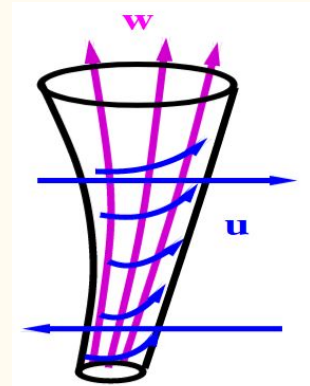
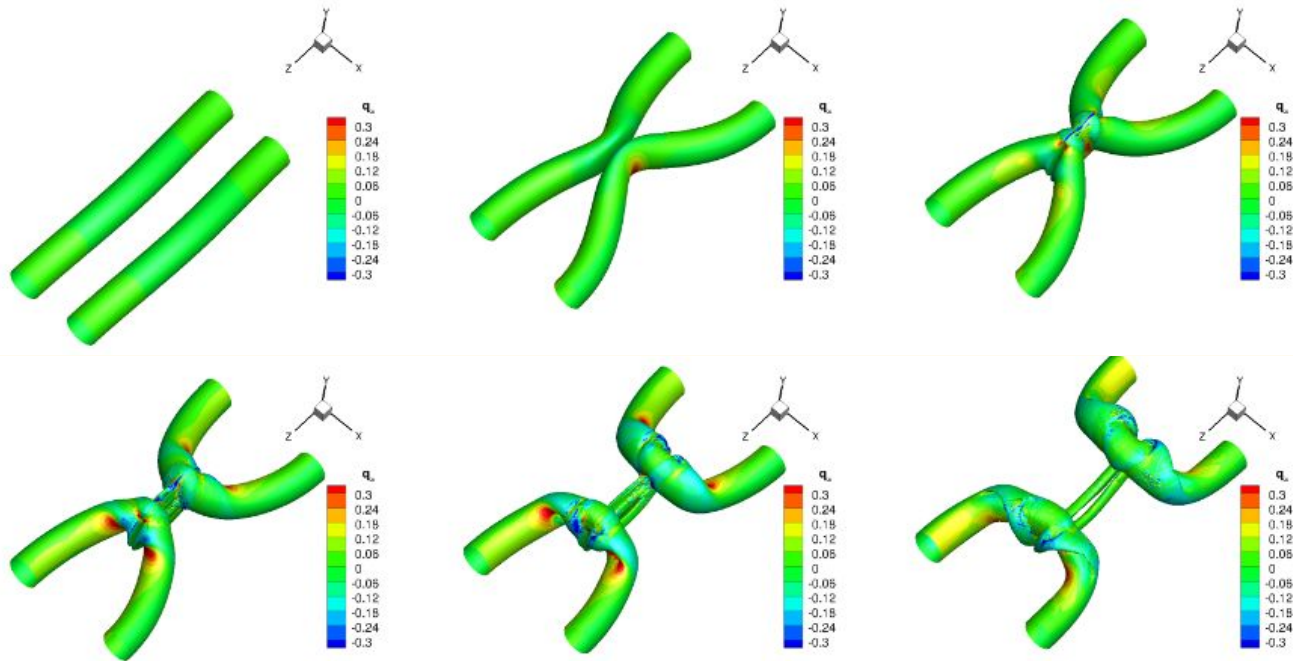
$$\partial_t \mathbf{u} + \overset{?}{\mathbf{u} \cdot \nabla \mathbf{u}} = \overset{?}{-\nabla P} + \nu \nabla^2 \mathbf{u} + \mathbf{f}$$

Turns out the **nonlinear advection term** is responsible for the energy cascade :

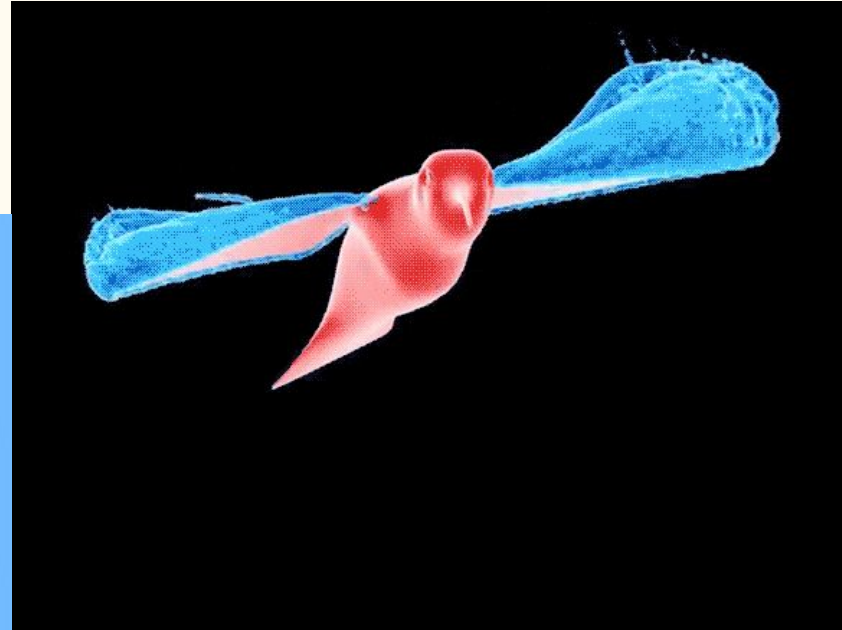
$$\mathbf{u} \cdot \nabla \mathbf{u}$$

But HOW is the energy transferred exactly ?

# Vortex stretching



# Vortex stretching



# Videos

[https://youtu.be/59LL\\_IRs1MQ](https://youtu.be/59LL_IRs1MQ)

<https://www.youtube.com/watch?v=J4MIiAJNAMO>

<https://www.youtube.com/watch?app=desktop&v=JaABLY6E8HE>



# Take a step back : 1st law of Thermodynamics

Mechanical work is transformed into heat :

$$\Delta U = Q - W$$

1st observation : Joule's experiment (1845)

→ Joule measured the rise in temperature of a fluid after stirring it with a propeller.

He observed that **fluids are good converters of mechanical energy into heat**

→ now we know why : **energy cascade**





# Quizz !

Why don't we observe turbulence in honey ?  
(10 000 x more viscous than water)



# Quizz !

Why don't we observe turbulence in honey ?  
(10 000 x more viscous than water)

**Answer 1 (boring) :**

Because for high viscosity  $\nu$  we  
have a **low Reynolds number**  
and thus a laminar flow.

$$Re = \frac{UL}{\nu}$$



# Quizz !

Why don't we observe turbulence in honey ?  
(10 000 x more viscous than water)

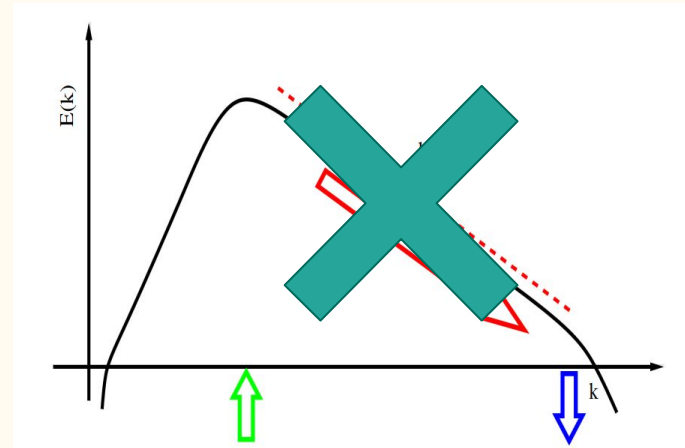
**Answer 2 :**

Because  $\nu$  is high, we have a **large viscous scale** and no energy cascade can happen

$$\ell_\nu = \frac{\nu^{3/4}}{\epsilon^{1/4}}$$

I calculated  $\ell_\nu \simeq 10 \text{ cm} !!$

→ dissipation occurs at human scale, no need for turbulence...



# Back to the definition

We will say that a 3D flow is turbulent if

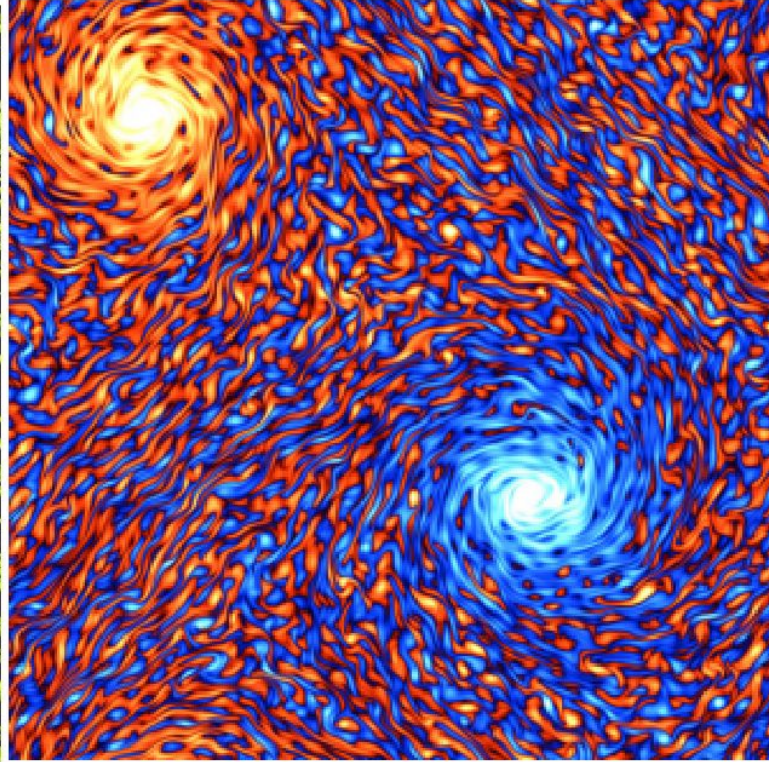
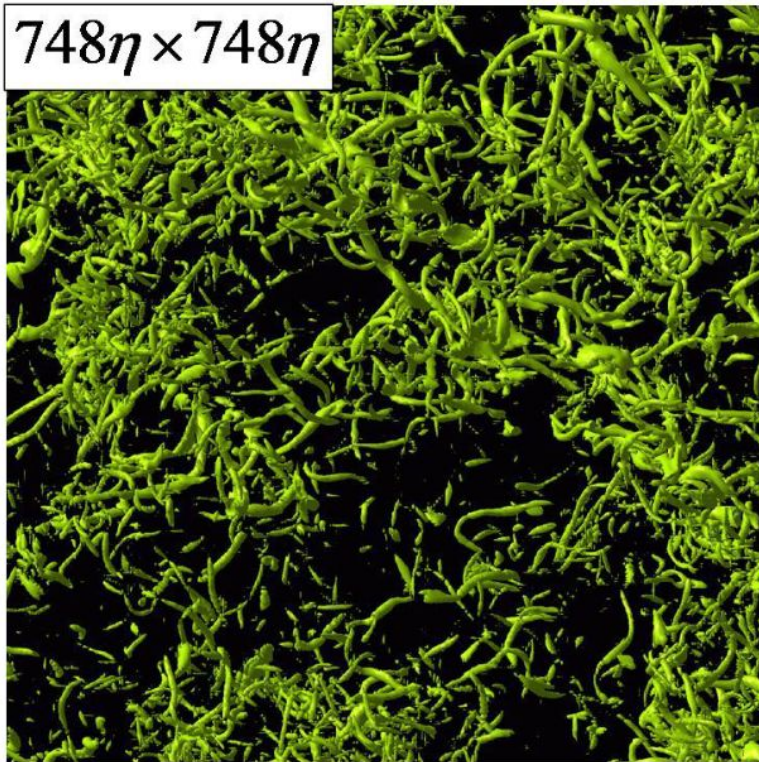
- **It is chaotic** (phase space trajectories are unstable: the slightest perturbation will lead to a different state thus leading to unpredictability)
- **It involves a large range of scales** ( many degrees of freedom )
- **Energy is dissipated at a scale much different than the scale it is injected** (ie there is a Cascade )
- **It is a strongly out of equilibrium process**

2D turbulence : what changes ?

—

## 2D vs 3D

$748\eta \times 748\eta$





## 2D vs 3D

Mathematically we have :  $u_z = 0, \quad \partial_z \mathbf{u} = 0, \quad \Rightarrow w_x = w_y = 0$

Helicity is now irrelevant ( $= 0$ ) the new conserved quantity is **Enstrophy** :

And we still have :

$$\frac{d}{dt} \mathcal{E} = -\epsilon + \mathcal{I}_{\mathcal{E}}$$

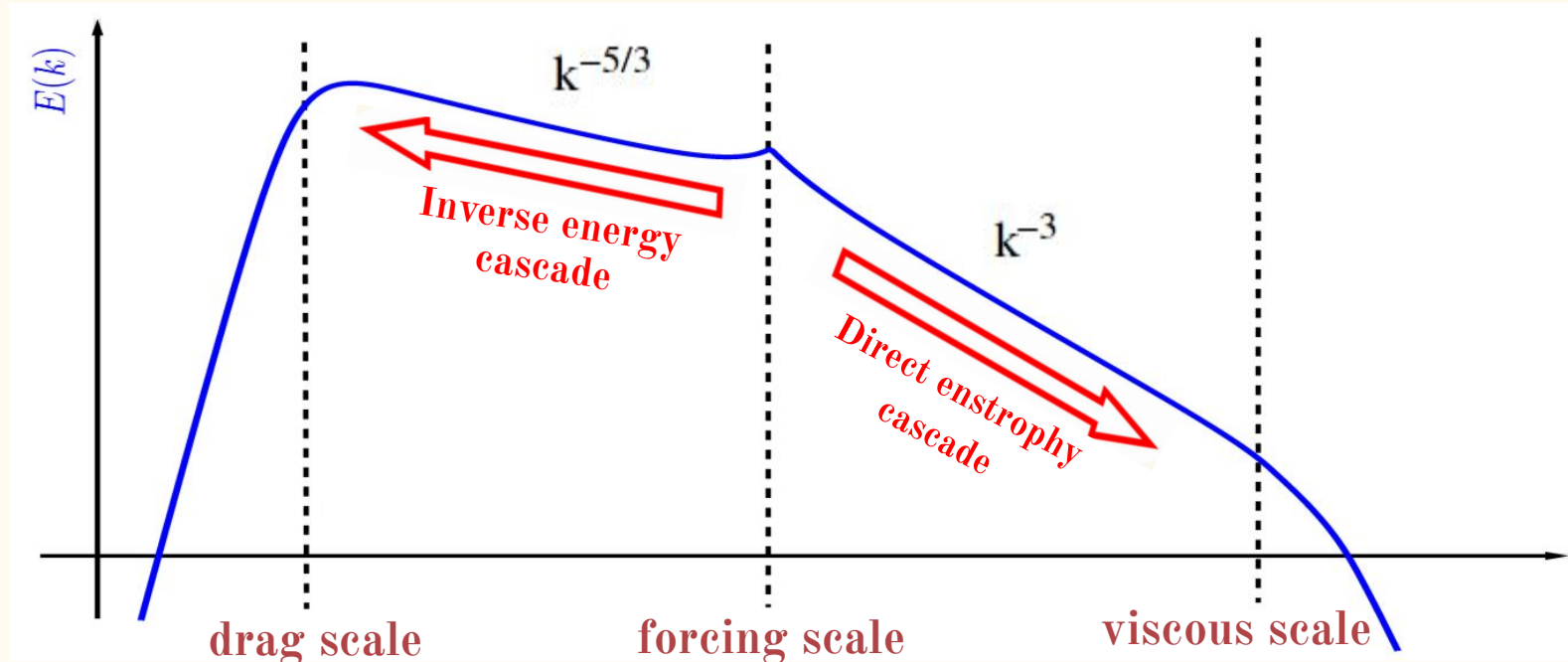
$$\frac{d}{dt} \Omega = -\eta + \mathcal{I}_{\Omega}$$

$$\Omega = \frac{1}{2} \langle w_z^2 \rangle$$

En **ergy** : with **work**  
En **strophy** : with **twist**

# Energy and Enstrophy cascade :

In 2D the cascade is **reversed** : energy flows from smaller scales to larger scales





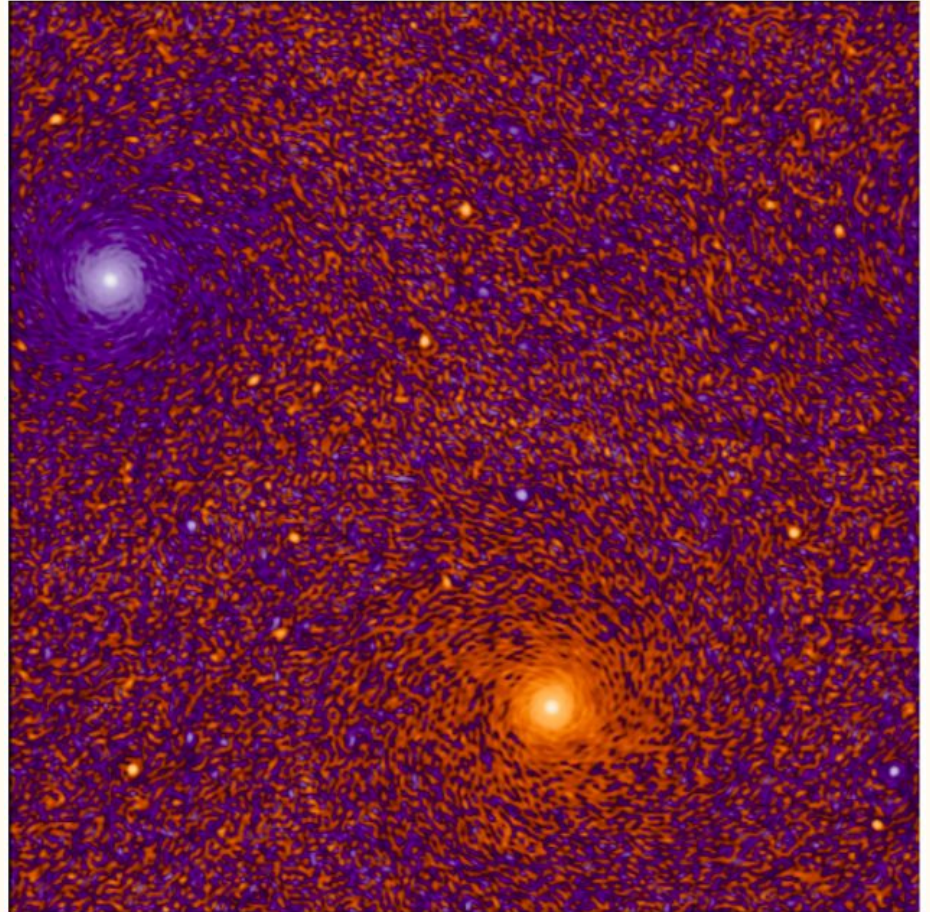
# Condensation

If the size of the system is finite,  
energy moves upscale and  
condensation occurs :

<https://youtu.be/Ax688IYmxf8>

<https://youtu.be/LiarsYToJUw>

<https://youtu.be/OzUCPAD4YDQ>



# Quantum turbulence

—

# Classical vs Quantum turbulence

Navier-Stokes equation on  $\omega$  :

$$\partial_t \omega + \{\omega, \phi\} = \nu \nabla^2 \omega - \alpha \omega + f$$

Where  $\Phi$  is the stream function :

$$\omega(\mathbf{r}, t) = -\nabla^2 \phi$$

$$(u, v) = (\partial_y \phi, -\partial_x \phi)$$

We assume **incompressibility**.

Gross-Pitaevskii equation on  $\psi$  :

$$i\partial_t \psi = \frac{c}{\sqrt{2}\xi} \left( -\xi^2 \nabla^2 \psi + \frac{|\psi|^2}{n_0} \psi - \psi \right)$$

Where :

$$c = \sqrt{gn_0/m}$$
$$\xi = \hbar / \sqrt{2mgn_0}$$

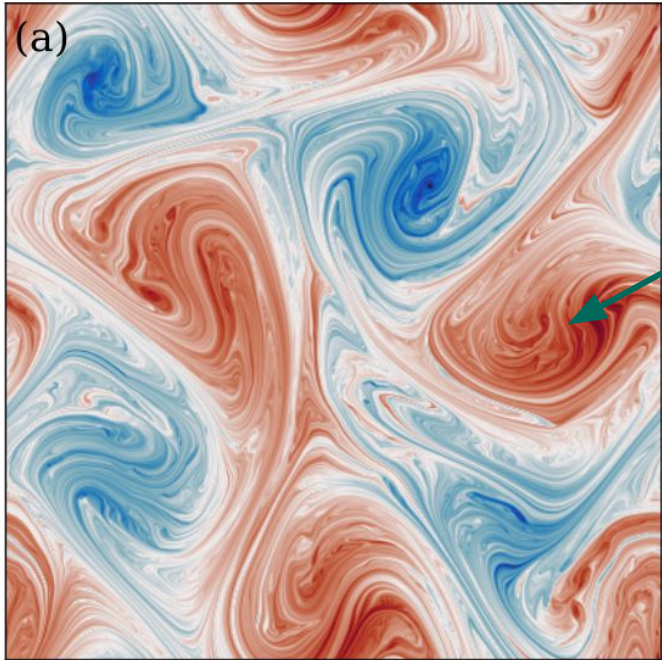
We can define the velocity field :

$$\mathbf{v} = -\sqrt{2}c\xi \text{Im}(\psi \nabla \psi^*) / \rho$$

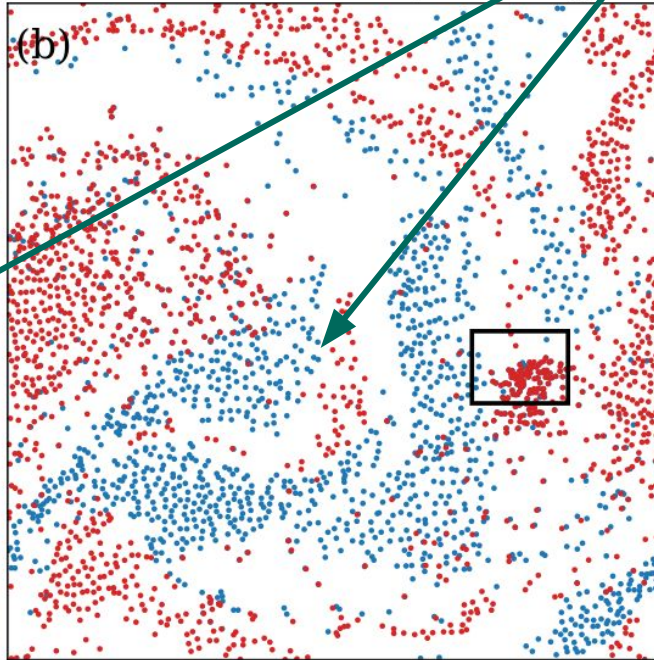
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Same **large scale structures** !

Navier-Stokes



Gross-Pitaevskii



For Quantum turbulence, vortices are **quantized**

$\Rightarrow$  the vorticity field is a superposition of  $\delta$ -Dirac terms

$\Rightarrow \Omega$  is ill-defined

One can also have annihilation of vortices of opposite charge that reduce turbulence...



# Mach number's influence

Strong density gradients can lead to “quasi-shocks”

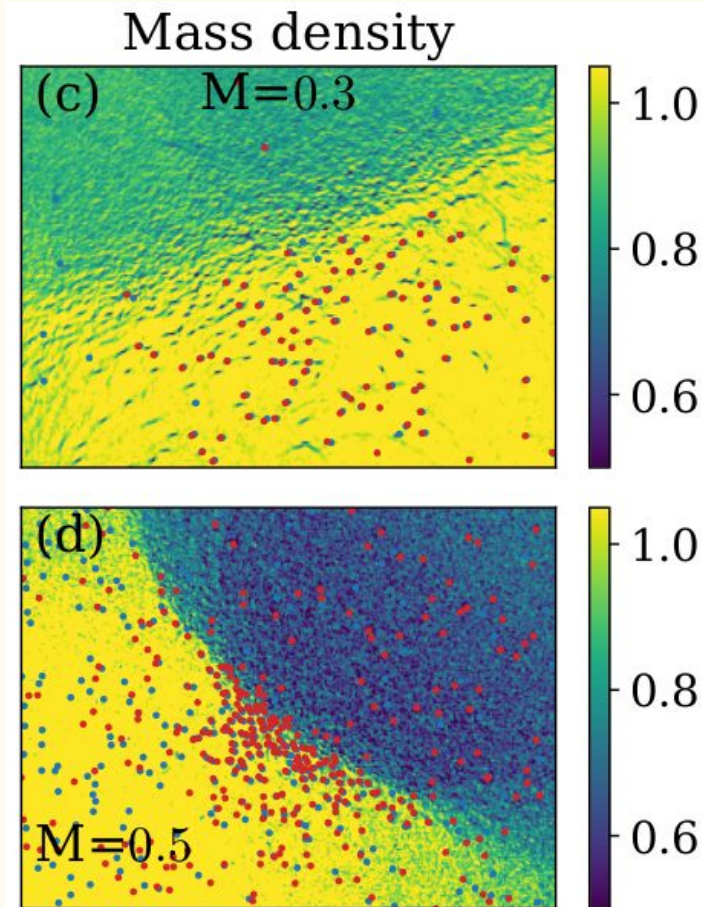
⇒ spontaneous generation of vortices

⇒ change in the flux statistics

This is quantified by the Mach number :

$$M = v_{\text{rms}}/c$$

Two simulations are done with  $M = 0.3$  &  $0.5$  :



# Thanks !

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*Big whorls have little whorls  
Which feed on their velocity  
And little whorls  
have lesser whorls,  
And so on to viscosity.*

*[Concerning atmospheric turbulence.]*

*Lewis Fry Richardson*

