

Problem 1  $P_{\text{speed}}(v) = 4\pi \left(\frac{m}{2\pi kT}\right)^{3/2} v^2 e^{-v^2 m/2kT} = \alpha v^2 e^{-v^2 \delta}$

a.)  $\frac{dP_{\text{speed}}}{dv} = 2\alpha v e^{-v^2 \delta} - 2\alpha \delta v^3 e^{-v^2 \delta} = 0$

$$2\alpha v e^{-v^2 \delta} = 2\alpha \delta v^3 e^{-v^2 \delta} \quad \therefore 1 = \delta v^2 \quad \Rightarrow \quad v^2 = \frac{1}{\delta}$$

$$v^2 = \frac{2kT}{m} \quad \therefore \quad v = \pm \sqrt{\frac{2kT}{m}}$$

$$v = \pm \sqrt{\frac{2kT}{m}}$$

b.)  $\bar{v} = \sum v_i P(v_i) = \int_0^\infty v P(v) dv = \alpha \int_0^\infty v^3 e^{-v^2 \delta} dv$

$$\bar{v} = \alpha \int_0^\infty v^3 e^{-v^2 \delta} dv : \int_0^\infty x^n e^{-x^2 \alpha} = \frac{n!}{2(x^{n+1})} : n = 2k+1 \quad \therefore \quad k = \frac{n-1}{2} \quad \Rightarrow \quad k = \frac{3-1}{2} = 1$$

$$\bar{v} = \alpha \cdot \frac{1!}{2 \cdot 3^2} = 4\pi \left(\frac{m}{2\pi kT}\right)^{3/2} \cdot \frac{1}{2} \left(\frac{4\pi k^2 T^2}{m^2}\right) = 4\pi \left(\frac{m^3}{8\pi^3 k^3 T^3}\right)^{1/2} \cdot \frac{1}{2} \left(\frac{16\pi^4 T^4}{m^4}\right)^{1/2}$$

$$= 2\pi \sqrt{\frac{m^3}{8\pi^3 k^3 T^3}} \cdot \frac{16\pi^4 T^4}{m^4} = 2\pi \sqrt{\frac{2kT}{m\pi^3}} = \sqrt{4\pi^2} \sqrt{\frac{2kT}{m\pi^3}} = \sqrt{\frac{8\pi^2 kT}{m\pi^3}} = \sqrt{\frac{8kT}{m\pi}}$$

$$\bar{v} = \sqrt{\frac{8kT}{m\pi}}, \quad v_{\text{mp}} = \pm \sqrt{\frac{2kT}{m}} : \quad \frac{\bar{v}}{v_{\text{mp}}} = \sqrt{\frac{8kT}{m\pi}} / \sqrt{\frac{2kT}{m}} = \sqrt{\frac{\frac{8kT}{m\pi}}{\frac{2kT}{m}}} = \sqrt{\frac{8kT}{2kT\pi}} = \sqrt{\frac{4}{\pi}}$$

$$\frac{\bar{v}}{v_{\text{mp}}} = \sqrt{\frac{4}{\pi}} = \frac{2}{\sqrt{\pi}}$$

$$\bar{v} = \sqrt{\frac{8kT}{m\pi}}, \quad \bar{v} \text{ differs } \frac{2}{\sqrt{\pi}} \text{ from the most probable speed}$$

c.) As  $T$  increases, the most likely and mean speed both increase

d.) Average kinetic energy is  $\bar{v}^2$

$$\bar{v}^2 = \int_0^\infty v^2 P(v) dv = \int_0^\infty v^2 \cdot \alpha \cdot v^2 e^{-v^2 \delta} dv : \int_0^\infty x^n e^{-x^2 \alpha} : n = 2k+1 \quad \therefore \quad k = \frac{n-1}{2}$$

$$\bar{v}^2 = \alpha \frac{(\frac{3}{2})!}{2 \cdot 3^{5/2}} = 4\pi \left(\frac{m}{2\pi kT}\right)^{3/2} \frac{(\frac{3}{2})!}{2} \left(\frac{2kT}{m}\right)^{5/2} = 2\pi \left(\frac{m}{2\pi kT}\right)^{3/2} \left(\frac{3}{2}\right)! \left(\frac{2kT}{m}\right)^{5/2}$$

$$\bar{v}^2 = 2\pi \left(\frac{3}{2}\right)! \sqrt{\left(\frac{m}{2\pi kT}\right)^3 \cdot \left(\frac{2kT}{m}\right)^5} = 2\pi \left(\frac{3}{2}\right)! \sqrt{\frac{(2kT)^2}{\pi^3 m^2}} : \left(\frac{3}{2}\right)! = \frac{3}{4} \sqrt{\pi}$$

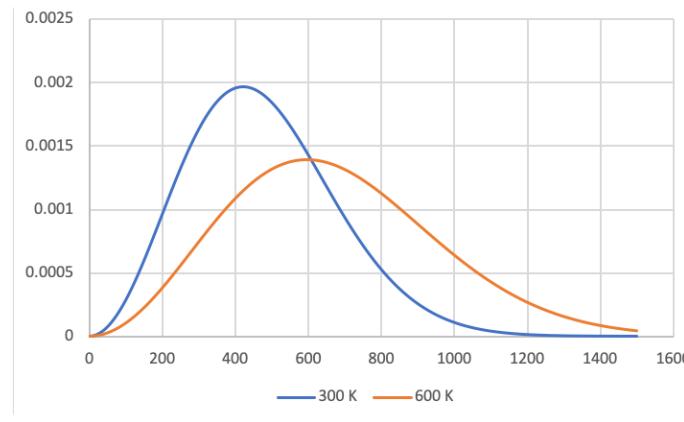
Problem 1 Continued

$$\bar{v}^2 = 2\pi \cdot \frac{3}{4} \sqrt{\pi k} \cdot \sqrt{\frac{4k^2 T^2}{9m^2}} = \frac{3}{2} \cdot \sqrt{\pi k} \cdot \frac{2kT}{m} \cdot \frac{1}{2\sqrt{2}} = \frac{3kT}{m} \quad : \quad \bar{v}^2 = \frac{3kT}{m}$$

$$E = \frac{1}{2}mv^2 = \frac{1}{2}m\left(\frac{3kT}{m}\right) = \frac{3}{2}kT$$

$$\boxed{\bar{v}^2 = \frac{3kT}{m}, E = \frac{3}{2}kT}$$

c.)



There is a higher peak for the 300 K line compared to that of the 600 K line. As the speed's increase, the 300 K line dives towards zero faster than the 600 K line.

## Problem 2

$$a.) V_{mp} = \sqrt{\frac{3kT}{m}}$$

at peak for gas A :  $V_{mp} = 400 \text{ m/s}$ , at peak for gas B :  $V_{mp} = 500 \text{ m/s}$

$$V_{mpA} = \sqrt{\frac{3kT_A}{m}} \quad \therefore \quad T_A = \frac{V_{mpA}^2 \cdot m}{3k} \quad ; \quad V_{mpB} = \sqrt{\frac{3kT_B}{m}} \quad \therefore \quad T_B = \frac{V_{mpB}^2 \cdot m}{3k}$$

$$T_A = 160,000 (\text{m/s})^2 \cdot \frac{m}{3k} \quad T_B = 640,000 (\text{m/s})^2 \quad \therefore$$

$$\frac{T_B}{T_A} = \frac{640,000 (\text{m/s})^2}{160,000 (\text{m/s})^2} \cdot \frac{m}{3k} \cdot \frac{3k}{m} = \frac{4}{1}$$

Gas B's temp is 4 times greater than that of Gas A's

$$b.) V_{mp} = \sqrt{\frac{3kT}{m}}$$

at peak for gas A :  $V_{mp} = 600 \text{ m/s}$ , at peak for gas B :  $V_{mp} = 500 \text{ m/s}$

$$V_{mpA} = \sqrt{\frac{3kT}{m_A}} \quad \therefore \quad m_A = \frac{3kT}{V_{mpA}^2} \quad ; \quad V_{mpB} = \sqrt{\frac{3kT}{m_B}} \quad \therefore \quad m_B = \frac{3kT}{V_{mpB}^2}$$

$$m_A = \frac{3kT}{360,000 (\text{m/s})^2}, \quad m_B = \frac{3kT}{640,000 (\text{m/s})^2}$$

$$\frac{m_B}{m_A} = \frac{3kT}{640,000 (\text{m/s})^2} \cdot \frac{360,000 (\text{m/s})^2}{3kT} = \frac{36}{64} = \frac{9}{16} = \frac{9}{16}$$

$m_B$  is  $\frac{9}{16}$  that of  $m_A$

### Problem 3

$$a.) P_{\text{spread}}(v) = 4\pi \left(\frac{m}{2kT}\right)^{3/2} v^2 e^{-v^2 m/2kT} : P(v) = \int_0^\infty 4\pi \left(\frac{m}{2kT}\right)^{3/2} v^2 e^{-v^2 m/2kT} dv$$

$$u = \frac{v}{\sqrt{2kT/m}} : v = u\sqrt{2kT/m} : \frac{du}{dv} = \frac{1}{\sqrt{2kT/m}} \rightarrow dv = \sqrt{2kT/m} du$$

$$P(v) = 4 \int_0^\infty \left(\frac{m^3 \pi^2}{8\pi^5 k^3 T^3}\right)^{1/2} u^2 \cdot \frac{2kT}{m} e^{-u^2 (2kT/m)} \cdot \cancel{\frac{m}{2kT}} du$$

$$P(v) = \frac{4}{\sqrt{\pi}} \int_0^\infty \left(\frac{m^3}{8k^3 T^3}\right)^{1/2} \left(\frac{4k^2 T^2}{m^2}\right)^{1/2} u^2 e^{-u^2} du = \frac{4}{\sqrt{\pi}} \int_0^\infty \left(\frac{m^3}{2.8k^3 T^3} \cdot \frac{4k^2 T^2}{m^2}\right)^{1/2} u^2 e^{-u^2} du$$

$$P(v) = \frac{4}{\sqrt{\pi}} \int_0^\infty \left(\frac{m}{2kT}\right)^{1/2} u^2 e^{-u^2} du = \frac{4}{\sqrt{\pi}} \int_{v_{\min}/\sqrt{2kT/m}}^\infty \left(\frac{m}{2kT}\right)^{1/2} u^2 e^{-u^2} du$$

$$P(v) = \frac{4}{\sqrt{\pi}} \int_{v_{\min}/\sqrt{2kT/m}}^\infty u^2 e^{-u^2} du$$

$$b.) m_{He} = 6.646 \times 10^{-27} \text{ kg}, v_{\min} = 11,000 \text{ m/s}, T = 300 \text{ K}, k = 1.38 \times 10^{-23} \text{ J/K}$$

$$\frac{v_{\min}}{\sqrt{2kT/m}} = \frac{11,000 \text{ m/s}}{\sqrt{2 \cdot 1.38 \times 10^{-23} \text{ J/K} / 300 \text{ K} / 6.646 \times 10^{-27} \text{ kg}}} \approx 9.85 \rightarrow \alpha$$

$$P(v) = \frac{4}{\sqrt{\pi}} \int_\alpha^\infty u^2 e^{-u^2} du = 7.396 \times 10^{-42} \rightarrow \text{TI Inspire}$$

$$P = 7.396 \times 10^{-42}$$

$$c.) m_{N_2} = 4.665 \times 10^{-26} \text{ kg}, v_{\min} = 11,008 \text{ m/s}, T = 300 \text{ K}, k = 1.38 \times 10^{-23} \text{ J/K}$$

$$\frac{v_{\min}}{\sqrt{2kT/m}} = \frac{11,000 \text{ m/s}}{\sqrt{2 \cdot 1.38 \times 10^{-23} \text{ J/K} / 300 \text{ K} / 4.665 \times 10^{-26} \text{ kg}}} \approx 26.07 \rightarrow \alpha$$

$$P(v) = \frac{4}{\sqrt{\pi}} \int_\alpha^\infty u^2 e^{-u^2} du = 2.256 \times 10^{-294} \rightarrow \text{TI Inspire}$$

$$\frac{P_{He}}{P_{N_2}} = \frac{7.396 \times 10^{-42}}{2.256 \times 10^{-294}} = 3.278 \times 10^{252}$$

$$P = 3.278 \times 10^{252}$$

A Helium molecule is  $3.278 \times 10^{252}$  more likely to escape Earth's atmosphere than Nitrogen. This tells us that there should be more Nitrogen than Helium in our atmosphere.

### Problem 4

a.)  $E = \frac{p_x^2}{2m} + \frac{1}{2} m \omega^2 r^2$ ,  $Z_{\text{single}} = \sum e^{-\frac{E_i \beta}{kT}} = \alpha \int dx \int dy \int dz \int dp_x \int dp_y \int dp_z e^{-(p_x^2 + p_y^2 + p_z^2) \beta / 2m}$

$$Z_{\text{single}} = \alpha \int_{-\infty}^{\infty} e^{-\frac{m\omega^2 x^2 \beta}{2m}} dx \int_{-\infty}^{\infty} e^{-\frac{p_x^2 \beta}{2m}} dp_x \int_{-\infty}^{\infty} e^{-\frac{m\omega^2 y^2 \beta}{2m}} dy \int_{-\infty}^{\infty} e^{-\frac{p_y^2 \beta}{2m}} dp_y \int_{-\infty}^{\infty} e^{-\frac{m\omega^2 z^2 \beta}{2m}} dz \int_{-\infty}^{\infty} e^{-\frac{p_z^2 \beta}{2m}} dp_z$$

Using  $\int_{-\infty}^{\infty} e^{-\alpha x^2} dx = \sqrt{\frac{\pi}{\alpha}}$

$$Z_{\text{single}} = \alpha \sqrt{\frac{2\pi}{m\omega^2 \beta}} \sqrt{\frac{2\pi}{m\omega^2 \beta}} \sqrt{\frac{2\pi}{m\omega^2 \beta}} \sqrt{\frac{2m\pi}{\beta}} \sqrt{\frac{2m\pi}{\beta}} \sqrt{\frac{2m\pi}{\beta}} = \left(\frac{2\pi}{m\omega^2 \beta}\right)^{3/2} \left(\frac{2m\pi}{\beta}\right)^{3/2}$$

$$Z_{\text{single}} = \alpha \left(\frac{2\pi}{m\omega^2 \beta} \cdot \frac{2m\pi}{\beta}\right)^{3/2} = \alpha \left(\frac{4\pi^2}{m\omega^2 \beta}\right)^{3/2} = \alpha \left(\frac{2\pi}{m\omega \beta}\right)^3 : Z_{\text{single}} = \alpha \left(\frac{2\pi}{m\omega \beta}\right)^3$$

$$Z_{\text{ensemble}} = (Z_{\text{single}})^N \quad \therefore \quad Z_{\text{ensemble}} = \left(\alpha^N \left(\frac{2\pi}{m\omega \beta}\right)^3\right)$$

b.)  $\bar{E} = -\frac{\partial}{\partial \beta} \ln(Z) = -\frac{\partial}{\partial \beta} \ln \left( \left( \alpha \left( \frac{2\pi}{m\omega \beta}\right)^3 \right)^N \right)$

$$\begin{aligned} \bar{E} &= -\frac{\partial}{\partial \beta} \cdot N \ln \left( \alpha \left( \frac{2\pi}{m\omega \beta}\right)^3 \right) = N \cdot \frac{\partial}{\partial \beta} \ln \left( \alpha \left( \frac{m\omega \beta}{2\pi}\right)^3 \right) = N \cdot \frac{\partial}{\partial \beta} \left( \ln \left( \alpha \right) + \ln \left( \frac{m\omega \beta}{2\pi}\right)^3 \right) \\ &= N \cdot \frac{\partial}{\partial \beta} \left( \ln(\alpha) \cdot \ln(\alpha) + 3 \ln \left( \frac{m\omega \beta}{2\pi}\right) \right) = N \frac{\partial}{\partial \beta} \left( 3 \ln \left( \frac{m\omega \beta}{2\pi}\right) - \ln(\alpha) \right) \end{aligned}$$

$$\bar{E} = N \cdot 3 \cdot \frac{4\pi}{2\pi k} \cdot \frac{2\pi}{m\omega \beta} = 3N \cdot \frac{1}{\beta} = 3N \cdot \frac{1}{\frac{1}{kT}} = 3NkT$$

$$\bar{E} = 3NkT$$

$$C = \frac{\partial \bar{E}}{\partial T} = 3Nk$$

$$C = 3Nk$$

c.) Dulong and Petit  $C = 3 \cdot 6.02 \times 10^{23} \cdot 1.38 \times 10^{-23} \frac{J}{\text{mol} \cdot \text{K}} = 24.92 \frac{J}{\text{mol} \cdot \text{K}}$  :  $C_{\text{DP}} = 24.92 \frac{J}{\text{mol} \cdot \text{K}}$

$$C_{\text{Cu}} = 385.1 \frac{J}{\text{kg} \cdot \text{K}} \cdot \frac{0.06355 \text{ kg}}{\text{mol}} = 24.47 \frac{J}{\text{mol} \cdot \text{K}} : \text{PRE} = \frac{|24.92 - 24.47|}{24.92} = 1.8\%$$

$$C_{\text{Al}} = 897.1 \frac{J}{\text{kg} \cdot \text{K}} \cdot \frac{0.09698 \text{ kg}}{\text{mol}} = 24.20 \frac{J}{\text{mol} \cdot \text{K}} : \text{PRE} = \frac{|24.92 - 24.20|}{24.92} = 2.9\%$$

$$C_{\text{Fe}} = 449.3 \frac{J}{\text{kg} \cdot \text{K}} \cdot \frac{0.05585 \text{ kg}}{\text{mol}} = 25.09 \frac{J}{\text{mol} \cdot \text{K}} : \text{PRE} = \frac{|24.92 - 25.09|}{24.92} = 0.63\%$$

As we can see, the Dulong Petit law does reasonably well for predicting heat capacity.

Problem 5

a.)

State	Occupancy
1	—●—●—●—
2	— —●—●—●—

b.)

State	Occupancy
1	— — — — ● — — — — ● — — — — ● — —
2	— — — ● — — — — ● — — — — ● — — — — ● — —
3	— — ● — — — — ● — — — — ● — — — — ● — — — — ● — —
4	— ● — — — — ● — — — — ● — — — — ● — — — — ● — — — — ● — —

### Problem 6

a.)  $\bar{n}_k = \frac{1}{\beta} \frac{\partial}{\partial \mu} \ln [Z_{bk}] , Z_{bk} = \sum e^{-\epsilon_k(\epsilon_k - \mu)\beta} , \mu = 0$

$$Z_{bk} = \sum e^{-\epsilon_k(\epsilon_k - \mu)\beta} = \frac{1}{1 - e^{-(\epsilon_k - \mu)\beta}}$$

$$\therefore \bar{n}_k = \frac{1}{\beta} \frac{\partial}{\partial \mu} \ln \left[ \frac{1}{1 - e^{-(\epsilon_k - \mu)\beta}} \right] = -\frac{1}{\beta} \frac{\partial}{\partial \mu} \ln (1 - e^{-(\epsilon_k - \mu)\beta})$$

$$\bar{n}_k = \cancel{\frac{1}{\beta}} \left( \frac{-\beta e^{-(\epsilon_k - \mu)\beta}}{1 - e^{-(\epsilon_k - \mu)\beta}} \right) = \frac{e^{-(\epsilon_k - \mu)\beta}}{1 - e^{-(\epsilon_k - \mu)\beta}} : \mu = 0$$

$$\bar{n}_k = \frac{e^{\epsilon_k \beta}}{1 - e^{\epsilon_k \beta}} \cdot \frac{e^{\epsilon_k \beta}}{e^{\epsilon_k \beta} - 1} : \bar{n}_k = \frac{1}{e^{\epsilon_k \beta} - 1} = \frac{1}{e^{\epsilon_k / kT} - 1}$$

$$\lim_{T \rightarrow 0} (\bar{n}_k) = \frac{1}{\infty - 1} = 0 \quad \boxed{T \rightarrow 0, \bar{n}_k = 0}$$

$$\bar{E} = \sum \epsilon_k \bar{n}_k = \frac{\epsilon_k}{e^{\epsilon_k / kT} - 1} : \lim_{T \rightarrow 0} (\bar{E}) = \frac{\epsilon_k}{\infty - 1} = 0$$

$$\boxed{T \rightarrow 0, \bar{E} = 0}$$

$$C = \frac{\partial \bar{E}}{\partial T} = \frac{\partial}{\partial T} (\epsilon_k (e^{\epsilon_k / kT} - 1)^{-1}) = \frac{\epsilon^2}{4k \sinh(\epsilon_k / kT)^2 t^2}$$

$$\lim_{T \rightarrow 0} C = \frac{\epsilon^2}{4k \sinh(\infty) \cdot 0} = \infty \quad \boxed{T \rightarrow 0, C = \infty}$$

b.)  $\bar{n}_k = \frac{1}{e^{\epsilon_k \beta} - 1} = \frac{1}{e^{\epsilon_k / kT} - 1}$

$$\lim_{T \rightarrow \infty} (\bar{n}_k) = \frac{1}{1 - 1} = \frac{1}{0} = \infty \quad \boxed{T \rightarrow \infty, \bar{n}_k = \infty}$$

$$\bar{E} = \frac{\epsilon_k}{e^{\epsilon_k / kT} - 1}$$

$$\lim_{T \rightarrow \infty} \bar{E} = \frac{\epsilon_k}{1 - 1} = \frac{\epsilon_k}{0} = \infty \quad \boxed{T \rightarrow \infty, \bar{E} = \infty}$$

$$C = \frac{\epsilon^2}{4k \sinh(\epsilon_k / kT)^2 t^2}$$

$$\lim_{T \rightarrow \infty} C = \frac{\epsilon^2}{4k \sinh(0) \cdot \infty} = 0 \quad \boxed{T \rightarrow \infty, C = 0}$$

Problem 6 continued

c.)  $\bar{E} = E$

$$\bar{E} = \frac{E_k}{e^{E_k/kT} - 1} \rightarrow \lambda\bar{E} = \frac{\bar{E}}{e^{\bar{E}/kT} - 1} \rightarrow 1 = \frac{1}{e^{\bar{E}/kT} - 1} \rightarrow e^{\bar{E}/kT} - 1 = 1$$

$$e^{\bar{E}/kT} = 2 \quad \ln(2) = \frac{\bar{E}}{kT} \quad \therefore T = \frac{\bar{E}}{k \ln(2)}$$

$$T = \frac{\bar{E}}{k \ln(2)}$$

Problem 7

a.)  $\bar{n}_k = \frac{1}{\beta} \frac{\partial}{\partial \mu} \ln [Z_{Bk}]$ ,  $Z_{Bk} = \sum e^{-\eta_k(\epsilon_k - \mu)\beta}$ ,  $Z_{Bk} = \sum_i e^{-\eta_k(\epsilon_k - \mu)\beta} = 1 + e^{-(\epsilon - \mu)\beta}$

$$Z_B = 1 + e^{-(\epsilon - \mu)\beta} \quad : \quad \ln(Z_B) = \ln(1 + e^{-(\epsilon - \mu)\beta})$$

$$\frac{\partial}{\partial \mu} \ln(Z_B) = \frac{\beta e^{-(\epsilon - \mu)\beta}}{1 + e^{-(\epsilon - \mu)\beta}} \quad \therefore \bar{n} = \frac{1}{\beta} \cdot \frac{\beta e^{-(\epsilon - \mu)\beta}}{1 + e^{-(\epsilon - \mu)\beta}} \quad : \quad \bar{n} = \frac{e^{-(\epsilon - \mu)\beta}}{1 + e^{-(\epsilon - \mu)\beta}}$$

$$\bar{n} = \frac{e^{-(\epsilon - \mu)\beta}}{1 + e^{-(\epsilon - \mu)\beta}} \cdot \frac{e^{(\epsilon - \mu)\beta}}{e^{(\epsilon - \mu)\beta} + 1} = \frac{1}{e^{(\epsilon - \mu)\beta} + 1} \quad \text{lim}_{T \rightarrow \infty} \bar{n} = \frac{1}{e^{(\epsilon - \mu)/kT} + 1} = \frac{1}{e^0 + 1} = \frac{1}{2}$$

As  $T \rightarrow \infty$ ,  $\bar{n} = \frac{1}{2}$

As  $T \rightarrow 0$ ,  $\bar{n}$  does not depend upon  $\mu$ . This tells us that the likelihood of a Fermion occupying a state is about 50%.

b.)

$$\text{if } \epsilon > \mu, T \rightarrow 0 \quad \bar{n} = \frac{1}{e^{\infty} + 1} = \frac{1}{\infty} = 0 \quad \boxed{\bar{n}=0 \text{ as } T \rightarrow 0 \text{ for } \epsilon > \mu}$$

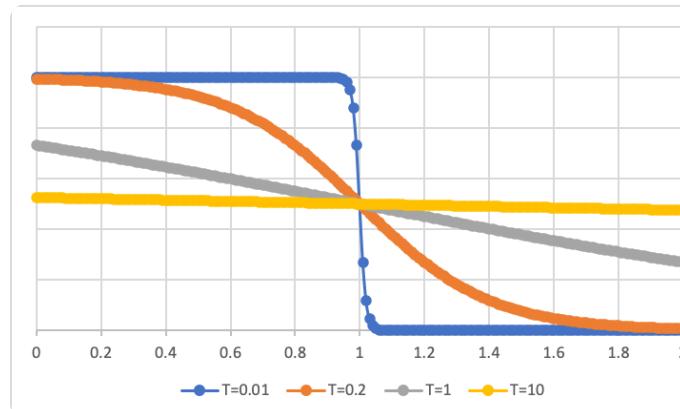
$$\text{if } \epsilon < \mu, T \rightarrow 0 \quad \bar{n} = \frac{1}{e^{\infty} + 1} = 1 \quad \boxed{\bar{n}=1 \text{ as } T \rightarrow 0 \text{ for } \epsilon < \mu}$$

This implies the likelihood of Fermion occupying a state to be 0 if  $\epsilon > \mu$  and 1 if  $\epsilon < \mu$  for small temperatures.

c.) If  $\epsilon = \mu$

$$\bar{n} = \frac{1}{e^0 + 1} = \frac{1}{1+1} = \frac{1}{2} \quad \boxed{\bar{n} = \frac{1}{2}}$$

d.)



e.) At very low temperatures the states will be either very populated or not populated. If  $\epsilon' > 1 \rightarrow \bar{n} = 0$ ,  $\epsilon' < 1 \rightarrow \bar{n} = 1$ . And at very high temperatures  $\bar{n} = \frac{1}{2}$  regardless of  $\epsilon'$  value.