

Importantly: The results are independent of how exactly ~~partitioned~~ the energy levels of the particles are being grouped into cells.

In fact, the results can be derived without this "grouping into cells".

(This is the reason why the grouping is legitimate.)

8.6 The ideal gases: grand canonical ensemble

plural since we will be looking at different statistics

Our starting point is the partition function $Q_N(T, V)$ in the canonical ensemble:

$$Q_N(T, V) = \sum_{\{n_k\}} g\{n_k\} e^{-\beta E\{n_k\}}$$

Constraint on occupation numbers: $\sum_k n_k = N$

Also: $E\{n_k\} = \sum_k n_k \epsilon_k$

$g\{n_k\}$: statistical weight factor associated with the distribution set $\{n_k\}$

$\sum_{\{n_k\}}$... goes over all distribution sets consistent with the particle number constraint equation

For identical bosons: $g\{n_k\} = 1$

For identical fermions: $g\{n_k\} = 1$ provided all n_k are zero or one
 $= 0$ if n_k are not zero or one

For Boltzmann particles: $g\{n_k\} = \prod_k \frac{1}{n_k!}$

After a bit of work, one finds:

$$\langle n_k \rangle = \frac{1}{z^{-1} e^{\beta \epsilon_k} + a}$$

mean or average occupation number of single-particle energy level ϵ_k

$$a = \begin{cases} -1 & \text{identical bosons} \\ +1 & \text{identical fermions} \\ 0 & \text{Boltzmann} \end{cases}$$

$$N = \sum_{\vec{k}} \langle n_{\vec{k}} \rangle$$

$$U = \sum_{\vec{k}} \epsilon_{\vec{k}} \langle n_{\vec{k}} \rangle$$

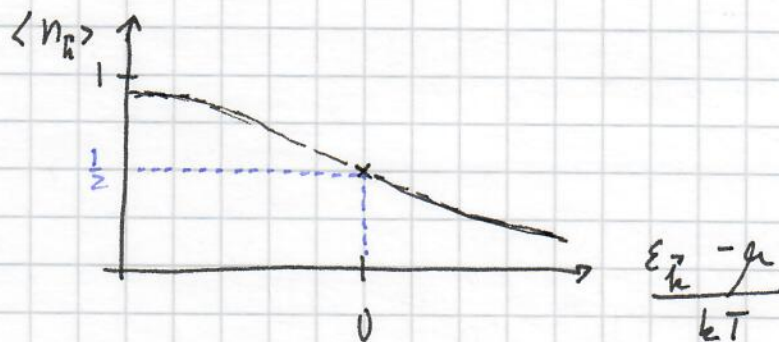
Let's look at the average or mean occupation numbers:

$$\langle n_{\vec{k}} \rangle = \frac{1}{e^{(\epsilon_{\vec{k}} - \mu)/kT} + a}$$

$a = 1$ (Fermi-Dirac distribution): $n_{\vec{k}}$ can only take the values 0 or 1 $\Rightarrow \langle n_{\vec{k}} \rangle \leq 1$

$\epsilon_{\vec{k}} < \mu \Rightarrow e^{(\epsilon_{\vec{k}} - \mu)/kT}$ is exponentially decaying.

If $|\epsilon_{\vec{k}} - \mu| \gg kT \Rightarrow \langle n_{\vec{k}} \rangle \rightarrow 1$



$a = -1$ (Bose-Einstein distribution):

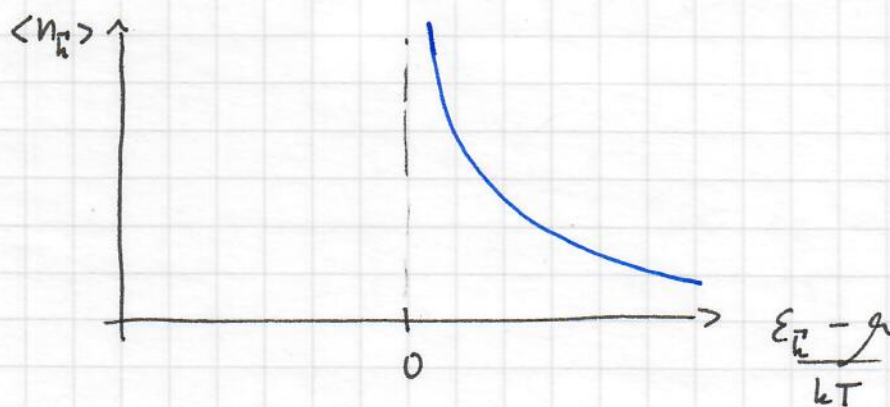
• requires $\sum e^{-\beta \epsilon_k} < 1$

$$e^{\beta(\mu - \epsilon_k)} < 1$$

$$\Rightarrow \mu < \epsilon_k \text{ for all } \epsilon_k$$

say, the lowest possible ϵ_k is ϵ_0 :

as $\mu \rightarrow \epsilon_0$, the occupation of that level becomes macroscopically large

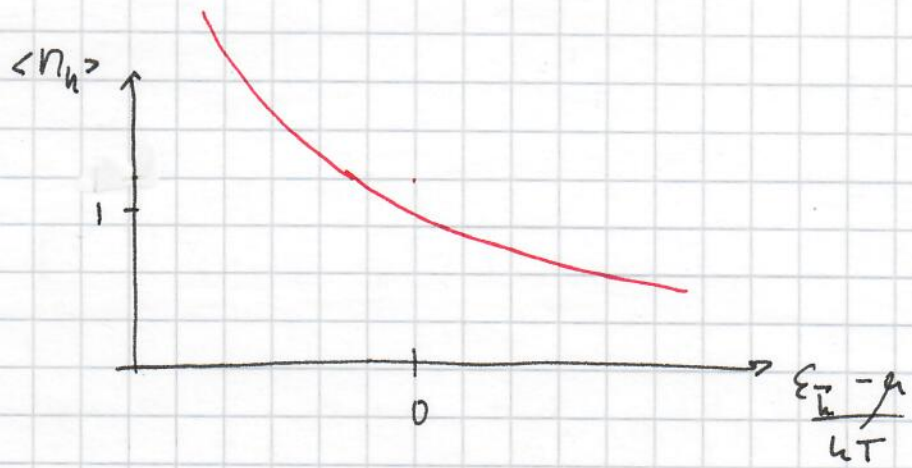


$a = 0$ (Maxwell-Boltzmann distribution):

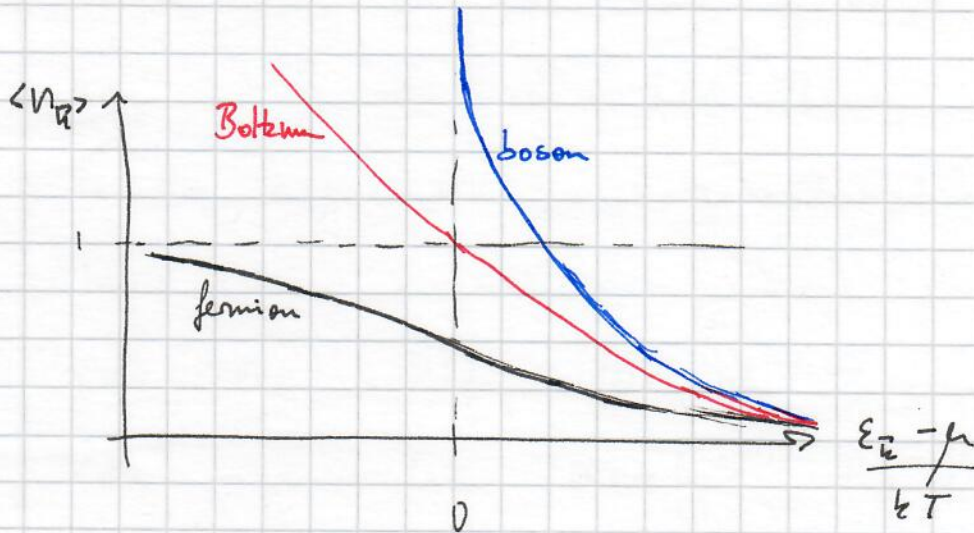
$$\langle n_k \rangle = e^{\mu/kT} e^{-\epsilon_k/kT}$$



recall: this is a quantum result
but without the quantum statistics



Let's put the three plots together:



large $(\epsilon_k - \mu)/kT$ limit:
large e limit:

$$\langle n_k \rangle = \frac{1}{e^{(\epsilon_k - \mu)/kT} + a} \xrightarrow{\text{neglect } a = \pm 1 \text{ (small compared to } e^{\epsilon_k - \mu/kT})} \frac{1}{e^{(\epsilon_k - \mu)/kT}}$$

$$= e^{\mu/kT} e^{-\epsilon_k/kT} \quad (\text{Boltzmann})$$

Physical picture:

large $\exp[(\epsilon_k - \mu)/kT]$ implies $\langle n_k \rangle \ll 1$

If $\langle n_k \rangle \ll 1$, then the $g\{n_k\}$ factors for the Boltzmann gas approach 1, i.e., $g\{n_k\} = \prod_k \frac{1}{n_k!} \rightarrow 1$.

Alternatively, we can think about what happens to the Bose and Fermi occupations or the associated exchange statistics: if we have very small occupations, then the particles hardly feel the existence of other particles. Therefore, the exchange statistics kicks in less and less as the average or mean occupation numbers go down.

It is also instructive to look at the relative mean-square fluctuations:

$$\frac{\langle n_k^2 \rangle - \langle n_k \rangle^2}{\langle n_k \rangle^2} = \frac{1}{\langle n_k \rangle} - a$$

without derivation
(see exercise 8.4 of
text)

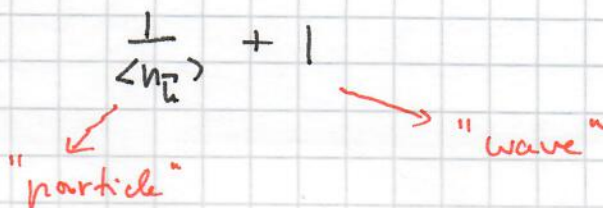
Boltzmann gas: $a = 0 \implies$ we refer to the $\frac{1}{\langle n_k \rangle}$ behavior as "normal"

$$\sqrt{\frac{\langle n_k^2 \rangle - \langle n_k \rangle^2}{\langle n_k \rangle^2}} = \frac{1}{\sqrt{\langle n_k \rangle}}$$

Fermi gas: $a = +1 \implies$ "below normal"

relative mean square fluctuations
tend to vanish as $\langle n_k \rangle \rightarrow 1$

Bose gas: $a = -1 \implies$ "above normal"

$\frac{1}{\langle n_k \rangle} + 1$

 "particle" "wave"

Let's look at statistics of occupation numbers:

$p_{\epsilon_k}(n)$: probability that there are n particles
in a state with energy ϵ_k

One finds:

Bose-Einstein :
$$p_{\epsilon_k}(n) = \frac{(\langle n_k \rangle)^n}{(\langle n_k \rangle + 1)^{n+1}}$$

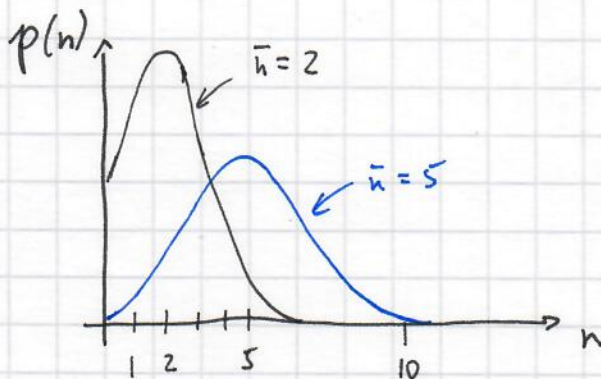
Fermi-Dirac :
$$p_{\epsilon_k}(n) = \begin{cases} 1 - \langle n_k \rangle & \text{for } n=0 \\ \langle n_k \rangle & \text{for } n=1 \end{cases}$$

Maxwell-Boltzmann:
$$p_{\epsilon_k}(n) = \frac{(\langle n_k \rangle)^n}{n!} e^{-\langle n_k \rangle}$$

↑
this is a
Poisson distri-
bution

Let's look at $p_{\varepsilon_k}(n) = \frac{(\langle n_k \rangle)^n}{n!} e^{-\langle n_k \rangle}$

Let $\langle \vec{n}_k \rangle = \bar{n} \rightarrow$ simplifies notation



$\langle n \rangle = \langle n_k \rangle$ of course, this has to be true by definition

$$\langle n^2 \rangle = \langle n_k \rangle^2 + \langle n_k \rangle$$

$$\Rightarrow \Delta n^2 = \langle n^2 \rangle - (\langle n \rangle)^2 = \langle n_k \rangle$$

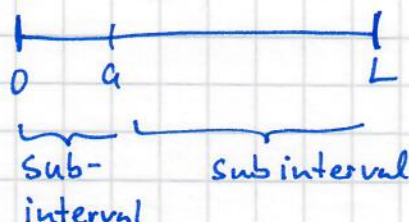
$$\begin{aligned} \Rightarrow \sqrt{\frac{\Delta n^2}{\langle n \rangle^2}} &= \sqrt{\frac{\langle n_k \rangle - \langle n_k \rangle^2}{\langle n_k \rangle^2}} \\ &= \frac{1}{\sqrt{\langle n_k \rangle}} \end{aligned}$$

this is just saying that
the statistics of the n_k
(Poisson distr.) is consistent
with our result from p. 201

of course, this is what we
had before

Detour on binomial and Poisson distribution:

Let's consider interval $[0, L]$



Want to distribute N particles completely randomly so that the probability that a particle be found in first subinterval is $\frac{a}{L}$ and that a particle be found in second subinterval is $\frac{L-a}{L}$.

Probability that n particles are in interval $[0, a]$ is

$$P_n = \left(\frac{a}{L}\right)^n \left(\frac{L-a}{L}\right)^{N-n} \binom{N}{n}$$

probability to have
 n particles in $[0, a]$
→ probability for
one particular
arrangement

↑
number of ways of
choosing n objects
from set of N

Want to rewrite P_n ...

$$\binom{N}{n} = \frac{N(N-1)(N-2)\cdots(N-n+1)}{n!}$$

$$\xrightarrow{\substack{n \text{ terms} \\ \text{in numerator}}} = N^n \frac{1(1-\frac{1}{N})(1-\frac{2}{N})\cdots(1-\frac{n-1}{N})}{n!}$$

$$\Rightarrow p_n = \left(\frac{aN}{L}\right)^n \left(1 - \frac{a}{L}\right)^{N-n} \frac{1}{n!} (1-\frac{1}{N})(1-\frac{2}{N})\cdots(1-\frac{n-1}{N})$$

$$= \bar{n}^n \frac{1}{n!} \underbrace{\left(1 - \frac{\bar{n}}{N}\right)^N}_{\text{using}} \frac{(1-\frac{1}{N})(1-\frac{2}{N})\cdots(1-\frac{n-1}{N})}{\left(1 - \frac{a}{L}\right)^n}$$

$$\text{using } \sum_{n=0}^N p_n \cdot n = \underbrace{\langle n \rangle}_{\bar{n}} = \frac{a}{L} N$$

$$\sum_{n=0}^N p_n = 1$$

$$\underbrace{\sum_{n=0}^N p_n^2 - \langle n \rangle^2}_{\substack{\langle n^2 \rangle \\ (\Delta n)^2}} = \frac{a}{L} \left(1 - \frac{a}{L}\right) N$$

$$p_n \xrightarrow[N \rightarrow \infty]{\frac{a}{L} \rightarrow 0} \bar{n}^n \frac{1}{n!} \underbrace{e^{-\bar{n}}}_{\text{Poissonian distribution}} \cdot \underbrace{1}_{\text{Poissonian distribution}} = \frac{\bar{n}^n}{n!} e^{-\bar{n}}$$

Poissonian
distribution

Boltzmann particles \leadsto distinguishable particles
 \leadsto no correlations

(Poissonian distribution for randomly selected "events")

Compare to Bose-Einstein:

$$p_{E_k}(n) = \frac{(\langle n_k \rangle)^n}{(\langle n_k \rangle + 1)^{n+1}}$$

$$\Rightarrow \frac{p_{E_k}(n)}{p_{E_k}(n-1)} = \frac{\langle n_k \rangle}{\langle n_k \rangle + 1}$$

\nwarrow independent of n , i.e., independent of the number of particles already in the state

\Rightarrow by comparison to Boltzmann case, where we had $\langle n_k \rangle / n$, bosons have a tendency to bunch together (i.e., we have positive statistical correlations among bosons)

\swarrow in contrast, fermions exhibit negative statistical correlations