Although the problem in principle involves a 3D sphen, we can simplify our analysis by restricting molion to a 2D x-y plane (gravely less along-y). Our solution will be valid for any other plane generaled by a robation through 3.

a) For a point particle (ie., w/ zero radius), we have a constraint of the general form:

 $x^2+y^2 > R^2$ 

where (x,y) dendes the location of the partitle of R is the sphere radius.

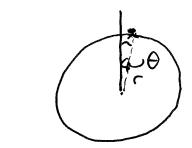
In part (b) we will be Interested in determining how the height at which the particle falls of the sphere. Hence for our solution we can assume we only need a description where the constraint is holonomic,

## $52+y^2-R^2=0$

and the inequality will only become velevarl once the particle has left the sphere.

Next, given that we are looking at motion on a spherical surface, let's more to polar co-ordinalis,

 $SC = rSin \Theta$  [chosen s.l.  $\Theta = 0$  7 north  $y = r\cos \Theta$  [chosen s.l.  $\Theta = 0$  7 north sphere.



In these co-ordinales, the constraint simplifies to n-R=O. We could just plug this straight into our définition of TeV (ie. set c=R inocey) but we are going to want the forces of condraint, so let's leave r as a free variable for now. To get the Corner of constraint & Eom, lets stare w the Lagrangians

$$T = \frac{m}{2} \left( \dot{x}^2 + \dot{y}^2 \right) = \frac{m}{2} \dot{r}^2 + \frac{m}{2} \dot{r}^2 \dot{o}^2 \quad \begin{cases} \text{see class} \\ \text{or do the} \end{cases}$$

$$\frac{1}{2} \left( \dot{x}^2 + \dot{y}^2 \right) = \frac{m}{2} \dot{r}^2 + \frac{m}{2} \dot{r}^2 \dot{o}^2 \quad \begin{cases} \text{see class} \\ \text{or do the} \end{cases}$$

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radial argular contribution.

Then, 
$$L = \frac{m}{2} \left( \dot{r}^2 + r^2 \dot{\theta}^2 \right) - mgrees \theta$$
.

Next, ue use the generalté Lagrange equalions:

$$\begin{pmatrix}
\frac{dr}{dr} \begin{pmatrix} \frac{10}{3r} \\ \frac{3r}{3r} \end{pmatrix} - \frac{90}{9r} = \frac{90}{3r}$$

$$\begin{pmatrix}
\frac{dr}{dr} \begin{pmatrix} \frac{9!}{3r} \\ \frac{3r}{3r} \end{pmatrix} - \frac{9c}{3r} = \frac{9c}{3r}$$

where x is an unknown lagrange multiplier strains from our single constraint,  $f(r) = r - R \quad (= 0)$ 

From I we the obtain:

$$m\ddot{c} - mr\dot{\theta}^2 + mgcos\theta = \lambda$$
 ①  $\left(\frac{\partial f}{\partial r} = 1\right)$ 
 $mr^2\ddot{\theta} + 2mr\dot{r}\dot{\theta} - mgrsin\theta = 0$  ②  $\left(\frac{\partial f}{\partial \theta} = 0\right)$ 

Now, of supplies the 3rd egn to enable us to eliminate & in paraiple, ie. by sulling r=K & i=i=0.

$$(0 =) -mR \dot{\theta}^2 + mg \cos \theta = \lambda$$
 (2a)

Exe Do can be rearranged,  $\Theta = 9/R \sin \theta$ , solve solve for  $\lambda$  in  $(\omega)$ . To solve our then solve for  $\lambda$  in  $(\omega)$ . To solve our equalion for  $\dot{\omega}$  lets use a trick,

We could formally integrate (X),

Plugging in  $\dot{\Theta} = g/g \sin \theta$ ,

$$\int_{0}^{\theta_{2}} \frac{9}{R} \sin \theta d\theta = \int_{0}^{\theta_{2}} \dot{\theta} d\dot{\theta}$$

$$\frac{1}{-g_{k}}\cos\theta\Big|_{\theta_{1}}^{\theta_{2}} = \frac{\mathring{\theta}^{2}}{2}\Big|_{\dot{\theta}_{1}}^{\dot{\theta}_{2}}$$

Our pareide  $\Xi$  is indially at rest,  $\dot{\theta}_1 = 0$ , to on the north pole,  $\dot{\theta}_1 = 0$ , so:

$$-\frac{9}{8}\left(\cos\theta_{2}-1\right)=\frac{1}{2}\dot{\theta}_{2}^{2}$$
or  $\theta_{2}\rightarrow\theta_{1}$ 

$$\dot{\Theta} = \frac{29}{R} \left( 1 - \cos \theta \right)$$

Our solve for  $\lambda$ ,

$$\dot{\theta} \Rightarrow (a) \Rightarrow \lambda = -2mg + 3mg\cos\theta$$

The force of constraint associated w/ r is then,

$$Q_r = + \lambda \frac{\partial f}{\partial r} = \frac{1}{2} 2 mg^{\frac{1}{2}} 3 mg \cos \theta$$

b) When does the ball point particle fall off the sphen? & Precisely when the force of constraint vanishes (this is when we instead more) to  $x^2ty^2 > R^2$ 

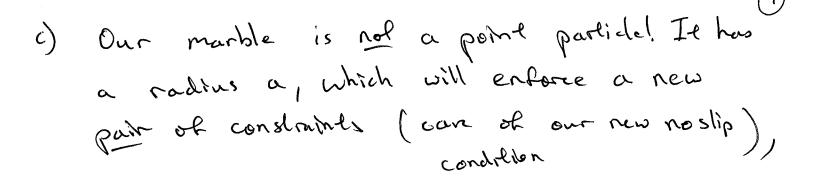
 $\cos \theta = 2/3$  is the angle at which it fulls off.

Converling back to our carlesian co-ordholis,

$$y = R \cos \theta = \frac{2K}{3}$$

: The thing height at which the particle falls of the sphere is  $y = \frac{2k}{3}$ .

Notice > this is above  $\frac{2k}{3}$ !



ii) 
$$\alpha(\phi-\theta)=R\theta$$
 [no slip]

Where does (ii) come from?!? Flost, seep set up co-ordhdes/dingram.



A marble when rolling wont elipping traces out on arclongth,

A S = a Ø

radius of angle marble rolls through.
marble

But also, in terms of the angle O above the large sphere,

Combining both:  $a(\phi-\theta) = R\theta = 2(i)$ 

As holonomiz constraints:

(i) 
$$\Rightarrow$$
  $f_{1}(r, \theta, \phi) = r - \alpha - R$ 

(ii) 
$$\rightarrow f_2(r, \theta, \phi) = R\theta - \alpha(\phi - \theta)$$

Building from

from a), our Lagrangian is:

$$\tau = \frac{m}{2} \left( \dot{r}^2 + r^2 \dot{\theta}^2 \right) + \frac{m}{2} a^2 \frac{\dot{\theta}}{2} \dot{\theta}^2$$

new rolational contaibution.

V = mgrcos 0

$$\exists L = \frac{m}{2} \left( \dot{r}^2 + r^2 \dot{\theta}^2 + \alpha^2 \dot{\phi}^2 \right) - mgrcos \theta$$

Our Eom in this case become:

$$\frac{d}{dt} \left( \frac{\partial L}{\partial q_{N}} \right) - \frac{\partial L}{\partial q_{N}} = \frac{3}{j=1} \lambda_{j} \frac{\partial f_{j}}{\partial q_{N}}$$

$$q_{N} = 7, 0, 0$$

A

$$\begin{cases} m\ddot{r} - mr\dot{\Theta}^2 + mycos\theta = \lambda_1 \\ mr^2\dot{\Theta} + 2mr\dot{\Theta}\dot{r} + myrsin\theta = \lambda_2(Rta) & 2 \end{cases}$$

$$m r^2\ddot{\Theta} = -\lambda_2 a \qquad 3$$

W/ 2 constraint egns:

$$f_1 \rightarrow r = Rta$$
,  $r = \dot{r} = 0$ 

$$f_2 \rightarrow \phi = \frac{1}{\alpha} \Rightarrow \frac{$$

The insight from the constraint egns can be used to solve 3:

Similarly, From D:

Lo 
$$\theta = \frac{\lambda_2 - mgsh\theta}{m(R+a)}$$
 (2a)

Selling & Ba = Qu, we obtain:

$$\lambda_2 = \frac{\text{masin9}}{2}$$

We then phy this back into (30) to solve for,

$$\theta = -\frac{9}{2(Rta)}$$
 Sin  $\theta$ 

Using the relation OdO = Odo from b), and following similar working,

$$\dot{\theta}^2 = \frac{9}{\text{Res}} - \frac{9}{\text{Res}^{\text{ta}}} \cos \theta$$

Finally, this goes into O to yield,

$$\lambda_1 = mg \left(2\cos\theta - 1\right)$$

From  $\lambda_1 \propto Q_T$  we know that the condition for the marble to full off the sphere is equivalent to,

$$\lambda_{1}=0 \Rightarrow 2\cos\theta-1=0$$

$$\downarrow \qquad \qquad \downarrow \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \qquad \downarrow \qquad \qquad \qquad \downarrow \qquad \qquad \qquad$$

This converts to a height:

$$y = R + a$$

Question 2 (Noh m + M are interchanged)

a) Let us slare by seeling 3=0 lo correspond to the table surface (3<0 under the table).

We obtain our lagrangian as,

 $L = \frac{M}{2} \left( i^2 + i^2 \dot{\varphi}^2 \right) + \frac{m}{2} \dot{\hat{s}}^2 - \frac{mg}{2}$ 

The rope provides a constraint:

f(n,418) = L+ 3-r =0 (360)

Then we have equalities of moliton in the the form,

de (3h) - 3h = > 3f de (3qn) - 3qn

for qu= 1,4,8.

$$\begin{pmatrix}
M^{2}\dot{\phi} - M^{2}\dot{\phi} = -\lambda & 0 \\
\frac{d}{de}(M^{2}\dot{\phi}) = 0 & 0 \\
m^{2}\dot{\phi} + my = \lambda & 0
\end{pmatrix}$$

② says that  $p_{\psi} = Mr^2 \mathring{\psi}$  is a constant of moldon!

Additionally, our constraint let's us set,

f -> 3 = "

and using this wol Mx3%

 $mM\ddot{r} + mMg = M\lambda$  (3a)

Similiarly, m x D yields,

 $m M \ddot{r} - m M \dot{q}^2 r = -m \lambda$  (a)

Then taking in To - Ba gives us,

 $\lambda = \frac{mM}{m+M} \dot{\varphi}^2 r + \frac{mM}{m+M} g$ 

Defining the reduced news,  $u = \frac{mM}{m+m}$ , the The forces of contraint one then oblained one

$$Q_{r} = -ug - u\dot{q}^{2}r$$

$$Q_{\varphi} = 0$$

$$Q_{\varphi} = u\dot{q}^{2}r + ug$$

$$Q_{\varphi} = u\dot{q}^{2}r + ug$$

the hole, we want

ie.,

$$ug + u\mathring{p}^{2}r - mg = 0$$

$$LD r\mathring{q}^{2} = mg - ug$$

$$= m - \frac{mM}{m+M}g = \frac{m}{m}g$$

$$\frac{mM}{m+M}g = \frac{m}{m}g$$

or  $\hat{\varphi} = \sqrt{\frac{mg}{Mr}}$  (taking the squarrood)

Reset by indial condition

(a) In principle, we know that,

 $E = T + V = \frac{M}{2} \left( i^2 + c^2 \dot{\varphi}^2 \right) + \frac{m}{2} \dot{s}^2 + mg_3 = const.$ 

We can use this to obtain an equalton for t(r).

First, we manipulate our expression for E by using that our constant implies i=3,

 $E = \frac{m+m}{2}i^2 + \frac{pq}{2mr^2} + mg(r-1)$ 

This can be rearranged to solve for i,

 $r = \pm \sqrt{\frac{2}{m+m}} \sqrt{\frac{E - p_{\ell}^2}{2Mr^2} - mg(r-l)}$ 

15

Worlding i = dr we then got,

$$dt = \pm \sqrt{\frac{m+m}{2}} \frac{3 dr}{\sqrt{E - p_u^2 - m_y(r-1)}}$$

e integrating yields,

$$t - t_o = \pm \int \frac{m+m'}{2} \int \frac{dr'}{\left(\frac{1}{2} - \frac{1}{2} - \frac{1}{2}$$

Question 3

a)

From Lagrange's equalien of molion,

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{q}}\right) - \frac{\partial L}{\partial q} = 0$$

Lo 9+89 =0

This egn looks jud like that of a damped harmonic oscillator.

makes sense because of terms  $\alpha q^2 + q^2$  in brackets. The e<sup>82</sup> factor factor in L is thus connected to the exponential damping we expect for a danped SHO.

P

 $\frac{\partial L}{\partial t}$  to care of the ext prefactor.

Hence the energy Function,

$$h = \frac{\dot{q}}{\partial \dot{q}} - L$$
 is not conserved.

of molion. -.

$$Q = q e^{\sigma t/2}$$

nolice,

$$e^{\gamma t}q^2 = e^{\gamma t}Q^2e^{-\gamma t} = Q^2$$

and,
$$e^{8t}\dot{q}^{2} = e^{8t}\left(\dot{Q}e^{-8t/2} + -\frac{\gamma}{2}Qe^{-8t/2}\right)^{2}$$

$$= \left(\dot{Q} - \frac{\gamma}{2}Q\right)^{2}$$

Thus,

$$L = \frac{m}{2} \left( \dot{Q} - \frac{\chi}{2} Q \right)^2 - \frac{k}{2} Q^2$$

W equalion of motion,

d) Our new Lagrangian features a conserved G energy funct,  $h = \frac{\partial JL}{\partial \dot{a}} - L$  as  $\frac{\partial L}{\partial t} = 0$ .

Explicitly,

h= ma(a-3a) - m(a-3a) + k a

 $= \frac{m}{2}\dot{Q}^2 + \frac{1}{2}Q^2$ 

where,  $\overline{k} = k - m(\overline{x})^2$ 

& Clearly, h looks to like the energy of an simple harmonte oscillator. We can also

 $m\omega^2 - m\left(\frac{\sigma}{2}\right)^2$  $= m \left[ \omega^2 - \left(\frac{\zeta}{2}\right)^2 \right]$ 

This is equivalent to the effective frequency,  $\bar{w} = \sqrt{w^2 - (\frac{\pi}{2})^2}$  we deduced for the danped harmonic Oscillator.

In faul our eqn for h thes nicely into our discussion of under artical over-damped systems: For our system to be an effective SHO -> 5/2 > 0 required! ie., w2 > (2)2 !

a) The problem is best tackled by using cylindrical co-ordinates,

$$x = r \cos \theta$$

$$y = r \sin \theta$$

$$\delta = \delta$$

The constraint for the particle to remaine on the surface of the cylinder can be wrotten via:

$$x^2 + y^2 - R^2 = 0$$
or  $r - R = 0$ 

It is thus a holonomic constraint.

with this in hand, we are left with two generalized co-ordinales: 3, 9

b) To compute the Lagrangian we need to first find the potential associate w/ the force. Clearly,

exhibits an indepense of O (ie. strength depends only on 171). It is simplest to write the associated potential into terms of r, w  $r^2 = R^2 + 3^2$ 

$$V(c) = \frac{1}{2}kc^2 \Rightarrow \hat{F}(c) = -\frac{3V}{3}\hat{c}$$

$$= -kc\hat{c}$$

$$\hat{c} = \frac{3}{2}kc^2 \Rightarrow \hat{F}(c) = -\frac{3V}{3}\hat{c}$$

Note you could have also done this in carlesian co-ordinales of a Few more lines of working.

Our kinelie contribution is,

$$T = \frac{m}{2} \left( \dot{x}^2 + \dot{y}^2 + \dot{z}^2 \right) = \frac{m}{2} \left( R^2 \dot{\theta}^2 + \dot{z}^2 \right)$$
Exist a sign or identify
as a sign or argular motion

e thus,

$$L = T - V = \frac{m}{2} \left( R^2 \dot{\Theta}^2 + \dot{z}^2 \right) - \frac{1}{2} k \left( R^2 + z^2 \right)$$

Nole: the contribution of the R2 could be ignored as a (23) constant shift in the potential

Next, we obtain the Eom, from Lagrange's egn:

$$\theta \rightarrow \frac{d}{dt} (mR^2 \dot{0}) = 0$$

The former enables us to identify mp20 as a constant of mother. It corresponds to a conserved angular momentum.

- The equalions describe a particle that:
  - () rotates at fixed role about the 3-axis
  - ii) undergoes harmonie motion about in  $\mathcal{F} = \frac{k}{m} \mathcal{F}$

because the force is independent of 0!

\$ \$ \$ \$