Density Matri	×	DI
See Sec. 3.4:	Density operators and pure versus mix	ed
	ensembles	
Situation 1:	Consider spin- & partile:	
	oven beam	
	Let's say: Each particle is prepared	
	in a superposition of spin-up and	sp
	sorpægenster down: we have a fixe	d
	phase I = C + 1 + > + C - 1 - > (coher super spin-up state ket of	nt mail
	Spin-up Spin-d	own
	state ket of	
	one partile: we have multiple	
	particles and they	
	the state het 17 >	
	Let spin point in direction û.	
	Angle with z-axis described by (0, 4).	
	$= \frac{c_{+}}{c_{-}} = \frac{\cos(\theta/2)}{\sin(\theta/2)}$	

Of course: |c+|2 + |c-|2 = 1 with

|c+12 = cos2(\$) and |c-12 = sin2(\$).

The description we have developed so far cannot describe a collection of atoms or electrons with

random spin orientation.

Situation 2: Particles coming and of a hot oven with random spin orientation, ise., no preferred direction.

> We expect: 50% of the members of the ensemble are characterized by 1+> and the remaining 50% by

1->. p = 0.5) fractional I populations . Jos spin-up p_ = 0.5 and spin-down, respectively.

populations

= real numbers

 \rightarrow 10, $\pm |c_{+}|^{2}$ and $p_{-} \pm |c_{-}|^{2}$

(pr and pr are more closely related to classical porobability theory)

DM-3 The discussion on the previous two pages shows that we are in need for a formalism that can treat ensembles. La density matrices loperators

Density matrices

We will start with pure state.

De fine $\hat{g} = 14 \times 241$ looking ahead:

here, just one

some state state and p=1(will be different for mixed state)

Let 14> = 2 cn 19n>

complete set of states

=> p = \(\sum_{n} \) \(\text{c}_{n} \) \(\text{c}_{n} \) \(\text{c}_{n} \) \(\text{c}_{m} \) \(\text{c}_{m} \) \(\text{c}_{m} \) \(\text{c}_{m} \)

The matrix elements of & in the d19023 basis

Per = < /2 | 9 | 40>

= < 9/2 | 5 5 cm cm 1 9, > < 9m 1 9 > = Ch Ce

A few properties that can be verified readily: $(i) \hat{g}^2 = \hat{g}$ (ini) (Â) = Tr (Â g)

(ii) Tr g = 1

Expectation value

Let's look at each of these three properties: ρ² = | Ψ><Ψ| Ψ><Ψ| = |Ψ><Ψ| = ρ̂. plugging in def.
of \hat{g} , using $\hat{g}^2 = \hat{g}\hat{g}$. Tr g = \(\frac{1}{k} \) \quad \qq \quad \ (in) evaluate use vesult in some basis from above def. of (A) expand 4 in basis = Z S Pnm Amn = E (S Sum Amn)

nse earlier nesult

rewrite in way! rewrite in = \(\sum \) \(\sum \) Amn \(\sum \) $= T_r \left(\hat{g} \hat{A} \right) = T_r \left(\hat{A} \hat{g} \right)$ Importantly: Properties (i), (iii), (iii) are independent of the busis chosen! Also: 9+=9. Eigenvalues of § are I and O. all others once

We will now look at mixed state.

For a mixed state, we define:

- IF; > are energy eigenstates

y Hamiltonian A

E pi =1

As before: I am be expanded in terms of complete set of states:

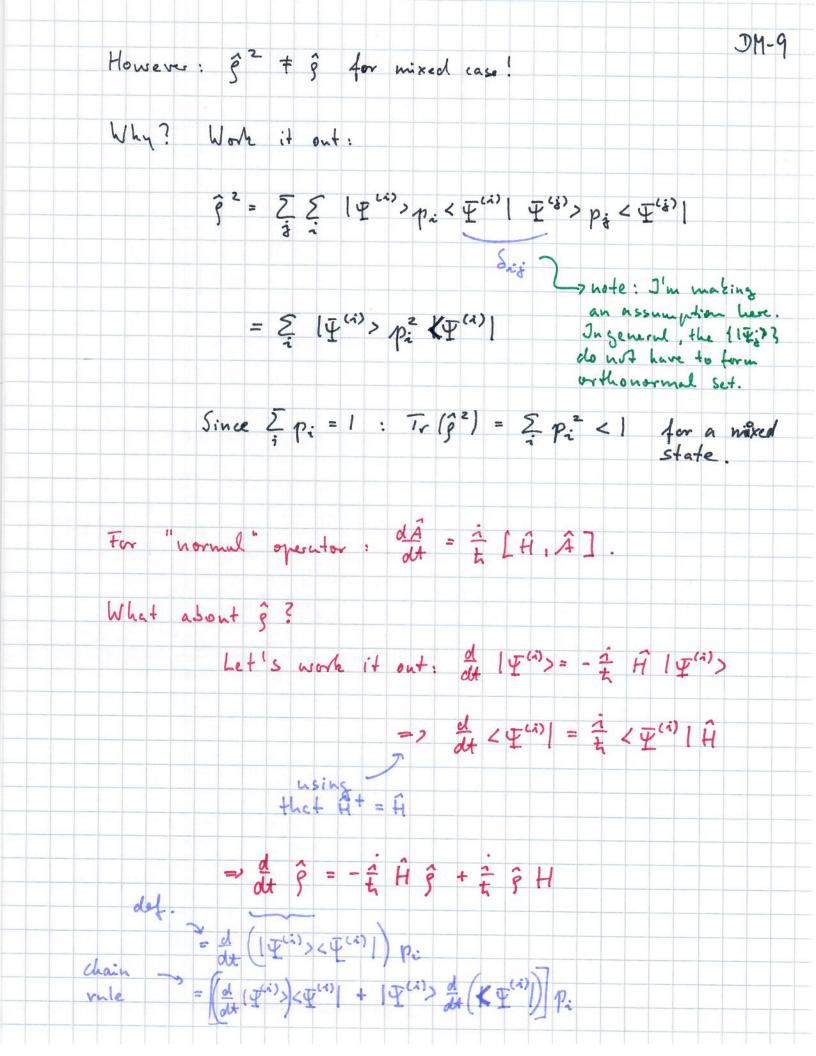
14; > = 5 ch (4)

the expansion coefficients for each of the energy expension tes are distinct

As in the pure case, we can evaluate \(\hat{g} \) in any given basis:

Pre = < 9/2 1 9 1 4e> = = pi ch (ce) +

We can read off: ghe = get => & hernitian.



So: dt g = [fi.g]

opprosite to normal" cace!

or it = = - [g(t), fi]

This equation of motion can be regarded as the quantum mechanical analog of Liouville's theorem in classical statistical mechanics:

Felassical = - [Pelassical , H]

Palassical: density of representative points in phase space.

The analogy with classical statistical mechanics can be pushed further: ensemble average

Quantum mechanically: <A>= tr (&A)

Example: Consider particle with magnetic moment et subjected to uniform magnetic field in the Z-direction. $\hat{H} = -\frac{e}{mc} \hat{S} \cdot \vec{B} = -\frac{eB}{mc} \hat{S}_{2}$ eigenvalues are # = E = 7 2mc (or E, = - tw/2 and E = tw/2) Let w = 2 B => E = 7 tw/2 let p = e + B tw/z and p = e - B tw/z $\hat{f} = \hat{p}_{+} |+> < + |+| + \hat{p}_{-}|-> < -|$ with $\hat{p}_{+} + \hat{p}_{-} = |+| -> |+| + |\hat{p}_{+}| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+| + |+|$ then: 3+ follows: $\begin{array}{c|c}
\hline
\text{in } 1+2, 1-2
\end{array}$ $\begin{array}{c|c}
P_{+} & 0 \\
\hline
\text{o} & \overline{P}_{-}
\end{array}$ $\begin{array}{c|c}
\hline
P_{+} & 0 \\
\hline
\text{o} & \overline{P}_{-}
\end{array}$ basis Tr g2 = p+2 +p=

High temperature himit: The canonical ensemble becomes a completely random ensemble in which all energy eigenstates are equally populated. Low temperature limit: Canonical ensemble becomes a pure ensemble where only the ground state is populated. Let us contrast the above with a quantum state It > of the form 14> = c+ (+> + c- 1-> => 0= 1平>< 平1 = (c+1+> + c-1->)(c+<+1+c+<-1) = 1412 1+><+1 + 14-12 1-><-1 + c_ c+ 1-><+1 + c+ c- 1+><-1 "coherences"

