

## Question 1:

### Assignment 3 solutions

①

Let's begin by defining the position of the centerpoint of each wheel (e.g., where the axles connect to the wheels) by the vectors  $\vec{r}$  &  $\vec{r}'$  respectively. Then, the co-ordinates of the midpoint of the axle are:

$$(x, y) = \frac{1}{2} (r_x + r'_x, r_y + r'_y) \quad (i)$$

Now, per the definitions given in the question, when one wheel rotates by  $d\phi$  ( $d\phi'$ ) the vector  $\vec{r}$  ( $\vec{r}'$ ) is shifted by a distance  $Rd\phi$  ( $Rd\phi'$ ) in ~~the~~ a direction defined by  $\theta$  (see Fig. 1). Then, we can equivalently write:

$$d\vec{r} = Rd\phi [\sin\theta \hat{x} - \cos\theta \hat{y}]$$

$$d\vec{r}' = Rd\phi' [\sin\theta \hat{x} - \cos\theta \hat{y}]$$

see Fig. 1 to understand  
this dependence.

From the eqns for  $d\vec{r}$  &  $d\vec{r}'$  & Eq (i) we then have: ②

$$dx = \frac{R}{2} \sin \theta (d\phi + d\phi') \quad (ii)$$

$$dy = -\frac{R}{2} \cos \theta (d\phi + d\phi') \quad (iii)$$

Multiply Eqs (ii) & (iii) by  $\sin \theta$  &  $\cos \theta$  & then adding / subtracting the resulting equations we can then also write:

$$\sin \theta dx - \cos \theta dy = \frac{R}{2} (\sin^2 \theta + \cos^2 \theta) (d\phi + d\phi')$$

$$\text{+ } \cos \theta dx + \sin \theta dy = 0$$

Which is the constraint Eq. 1. ~~for~~ 2

Now to obtain the holonomic constraint (Eq. 2 of Q1), consider the relative vector:

$$\vec{r}_{12} = \vec{r} - \vec{r}'$$

By definition, the angle  $\theta$  can be written as,

$$\tan \theta = \frac{r_{12,y}}{r_{12,x}}$$

and the differential of this ~~vector~~ result is, ③

$$\sec^2 \theta d\theta = -\frac{r_{12,y}}{r_{12,x}^2} dr_{12,x} + \frac{dr_{12,y}}{r_{12,x}}$$

Plugging  $d\vec{r}$  &  $d\vec{r}'$  into the definition of  $\vec{r}_{12}$  & computing differentials then let's us write:

$$\begin{aligned} \sec^2 \theta d\theta &= R(d\phi - d\phi') \left[ -\frac{r_{12,y}}{r_{12,x}^2} \sin \theta - \frac{\cos \theta}{r_{12,x}} \right] \\ &\equiv -R \frac{(d\phi - d\phi')}{r_{12,x}} [\tan \theta \sin \theta + \cos \theta] \end{aligned}$$

$$\begin{aligned} \downarrow \\ d\theta &= -\frac{R(d\phi - d\phi')}{r_{12,x}} [\sin^2 \theta \cos \theta + \cos^3 \theta] \\ &= -\frac{R(d\phi - d\phi')}{r_{12,x}} \cos \theta \end{aligned}$$

↙ trig identities,  
 $\sin^2 \theta + \cos^2 \theta = 1$

Again, we consider the definition of  $\theta$  to realize,

$$\cos \theta = \frac{r_{12,x}}{r}$$

So that,

(4)

$$d\theta = -\frac{R}{L}(d\phi - d\phi')$$

Integrating both sides then yields the result  
for Eq. 2 (~~Eq. 2~~):

$$\theta = \text{const.} - \frac{R}{L}(\phi - \phi')$$

$$\hookrightarrow \theta + \frac{R}{L}(\phi - \phi') = \text{const.}$$

So in summary, the condition of no slipping  
leads to two nonholonomic constraint eqns:

~~Eq. 1~~

$$\sin\theta dx - \cos\theta dy = \frac{R}{2}(d\phi + d\phi')$$

$$\cos\theta dx + \sin\theta dy = 0$$

(nonholonomic  $\rightarrow$  they can be equivalently written w/

$$\frac{dx}{dt}, \frac{dy}{dt}, \frac{d\phi}{dt} + \frac{d\phi'}{dt} \rightarrow \text{depend on velocities!})$$

and one holonomic constraint:

$$\theta + \frac{R}{L}(\phi - \phi') = \text{const.} \rightarrow \text{depends only on co-ordinates } \theta, \phi, \phi'$$

## Question 2:

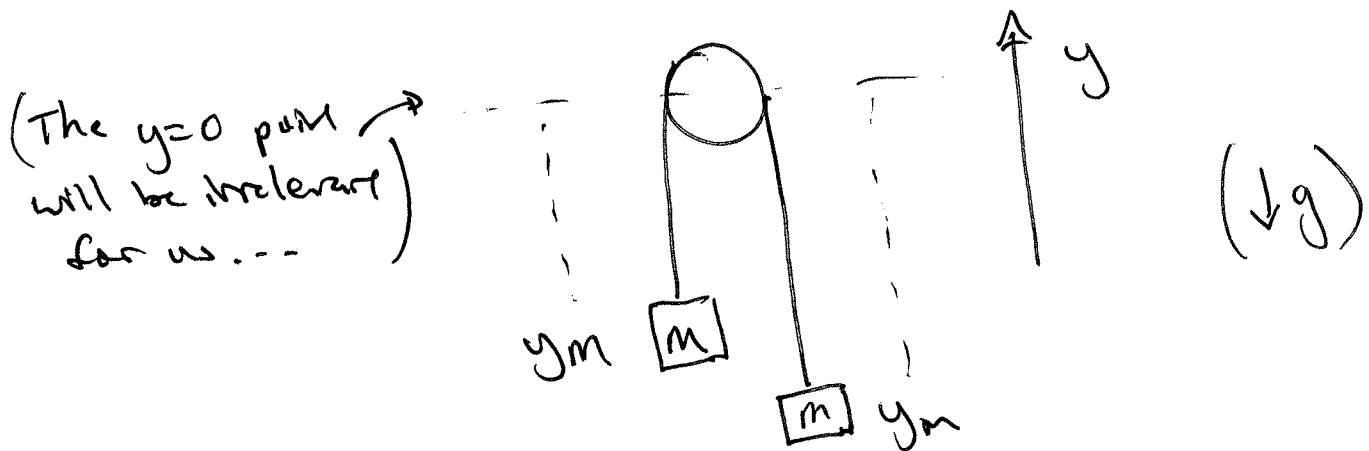
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### Part 1

a) D'Alembert's principle states that,

$$\sum_i (F_i^{(a)} - \dot{p}_i) \cdot \delta r_i = 0$$

Let us define the co-ordinates of the system by:



Then for mass  $m$ ,

$$F_m^{(a)} = -mg \quad \text{and} \quad \dot{p}_m = (m\dot{y}_m) = m\ddot{y}_m$$

and for mass  $M$ ,

$$F_M^{(a)} = -Mg \quad \dot{p}_M = M\ddot{y}_M$$

Then d'Alembert's principle gives:

(6)

$$-m(g + \ddot{y}_m) \cdot \delta y_m - M(g + \ddot{y}_m) \cdot \delta y_m = 0 \quad (i)$$

b)  $y_m$  &  $y_M$  are not independent, because the masses are connected by the fixed length of rope. In particular,

$$L + y_m + y_M = 0 \quad (\text{assuming } y=0 \text{ at pulley ---})$$

Then,

$$\ddot{y}_m = -\ddot{y}_M \quad + \quad \delta y_m = -\delta y_M$$

and thus,

$$(i) \rightarrow -g(m-M) + (m+M)\ddot{y}_m = 0$$

$$\downarrow$$
$$\ddot{y}_m = \frac{m-M}{m+M} g$$

$$\text{or } \ddot{y}_m = \frac{M-m}{m+M} g$$

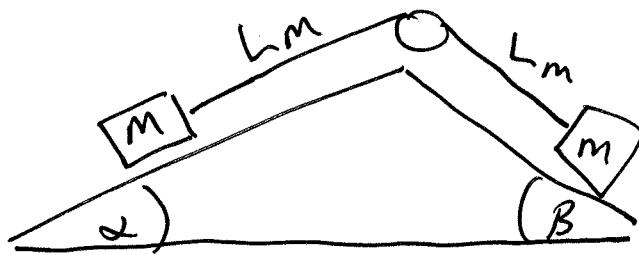
c) We can again use d'Alembert's principle ⑦  
 to compute the relevant equations of motion.  
 Appropriately compute the component of gravity  
 along the surfaces of the wedge:

$$F_m^{(a)} = +mg \sin \beta$$

$$p_m = +m \ddot{L}_m$$

$$\cancel{F}_M^{(a)} = +Mg \sin \alpha$$

$$p_M = +M \ddot{L}_m$$



↓ g

Then,  $\sum_i (F_i^{(a)} - p_i) \cdot \delta r_i = 0$ , yields:

$$\delta L_m (mg \sin \beta - m \ddot{L}_m) + \cancel{\delta L_m} (Mg \sin \alpha - M \ddot{L}_m) = 0$$

As in (b), we have that  $L_M + L_m$  are not independent, i.e.,

$$\ddot{L}_M = -\ddot{L}_m \quad \& \quad \delta L_M = -\delta L_m$$

Thus, d'Alembert's principle then implies,

(8)

$$+ mgs \sin \beta - mL \ddot{\alpha} - Mgs \sin \alpha - ML \ddot{\alpha} = 0$$

↓

~~mg sin β~~

$$g(m \sin \beta - M \sin \alpha) - (m+M)L \ddot{\alpha} = 0$$

↓

$$L \ddot{\alpha} = g \frac{m \sin \beta - M \sin \alpha}{m+M}$$

## Part 2

d)

In principle we might think at least 3 co-ordinates are needed to describe the positions of the three weights. However, given that the two ropes (length  $L$  each) form a pair of constraints, we will find only 2 generalized co-ordinates are required.



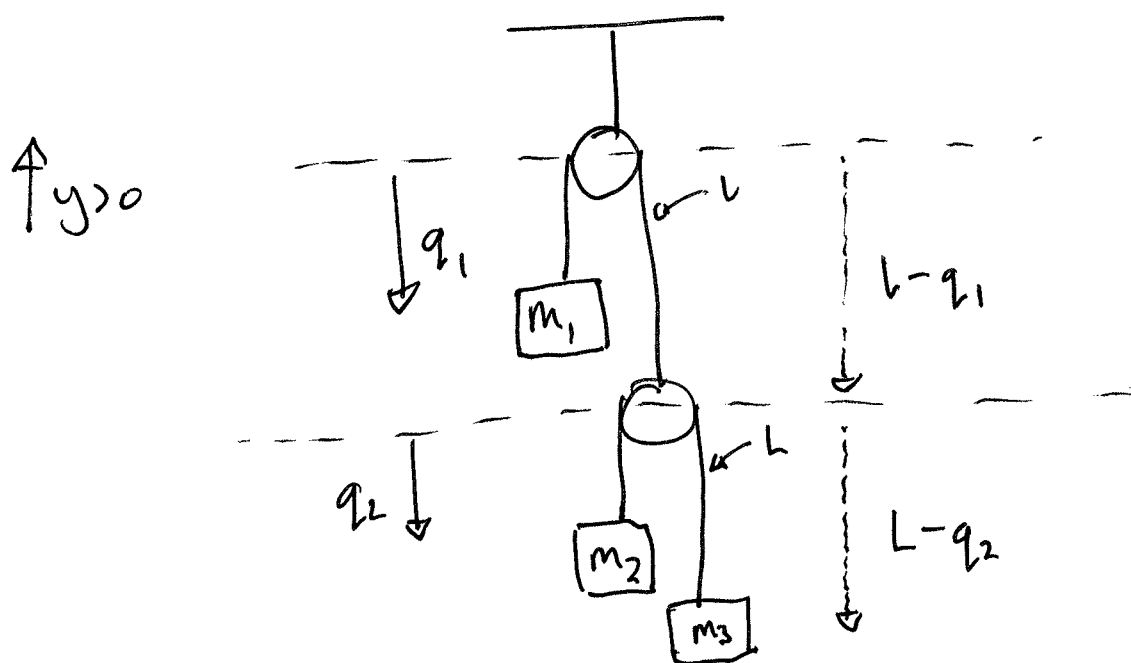
Why 2? In general,

$$n_{\text{dof}} = DN - m$$

$\uparrow$  # of d.o.f. for generalized co-ordinates  
 $\uparrow$   $D$  dimensional motion  
 $\nwarrow$   $N$  moving particles  
 $\nwarrow$   $m$  constraints (holonomic)

We have  $N = 4$  moving objects (3 blocks + 1 moving pulley) moving in  $D=1$ . The two rope lengths give us  $m=2$  constraints. Thus  $n_{\text{dof}} = 2$ !

Let's next draw a diagram to set up our relevant co-ordinates:



Defining  $y_i$  as the original cartesian co-ordinates of each block w/ mass  $m_i$ , we can define everything in terms of the two generalized co-ordinates  $q_1 + q_2$ .

[  $q_1 \Rightarrow$  distance of  $m_1$  from top pulley height,  
 $q_2 \Rightarrow$  distance of  $m_2$  from bottom pulley height. ]

We have:

$$y_1 = -q_1, \quad y_2 = -l + q_1 - q_2$$

$$\text{e } y_3 = -(l+L) + q_1 + q_2.$$

Also, for completeness,

$$\delta y_1 = -\delta q_1, \quad \delta y_2 = \delta q_1 - \delta q_2$$

$$\text{+ } \delta y_3 = \delta q_1 + \delta q_2.$$

e) We can obtain the relevant equations of motion for  $q_1 + q_2$  from d'Alembert's principle.

$$\sum_i (F_i^{(n)} - \dot{p}_i) \cdot \delta r_i = 0$$

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~~$$m_1 \ddot{q}_1 + m_2 \ddot{q}_1$$~~

pls

$$\Rightarrow 0 = (-m_1 g - m_1 \ddot{y}_1) \delta y_1 + (-m_2 g - m_2 \ddot{y}_2) \delta y_2 \quad (1)$$

$$+ (-m_3 g - m_3 \ddot{y}_3) \delta y_3$$

$$= (m_1 g - m_1 \ddot{q}_1) \delta q_1 + (-m_2 g + m_2 \ddot{q}_2 - m_2 \ddot{q}_1) (\delta q_1 - \delta q_2)$$

$$+ (-m_3 g - m_3 \ddot{q}_1 - m_3 \ddot{q}_2) (\delta q_1 + \delta q_2)$$

As both  $\delta q_1$  &  $\delta q_2$  are independent variations,

$$\delta q_1 \rightarrow \cancel{m_1 g} (m_1 g - m_2 g - m_3 g) + \ddot{q}_1 (-m_1 - m_2 - m_3) \quad (1)$$

$$0 = + \ddot{q}_2 (m_2 - m_3)$$

$$\delta q_2 \rightarrow 0 = (m_2 g - m_3 g) + \ddot{q}_1 (m_2 - m_3) \quad (2)$$

$$+ \ddot{q}_2 (-m_2 - m_3)$$

Eqs ① + ② are a linear system w/ two unknowns (e.g.  $\ddot{q}_1$  &  $\ddot{q}_2$ ) that can be solved. In particular, using the shorthand,

$$\textcircled{1} \rightarrow A_1 \ddot{q}_1 + B_1 \ddot{q}_2 + C_1 = 0$$

$$\textcircled{2} \rightarrow A_2 \ddot{q}_1 + B_2 \ddot{q}_2 + C_2 = 0$$

then, for example,

(12)

$$\ddot{q}_1 = \frac{B_2 C_1 - B_1 C_2}{A_2 B_1 - A_1 B_2}$$

$$+ \ddot{q}_2 = \frac{A_2 C_1 - A_1 C_2}{A_1 B_2 - A_2 B_1} \quad \square$$

f) By the definitions established in (d), we have:

$$\ddot{y}_1 = -\ddot{q}_1 = \frac{B_1 C_2 - B_2 C_1}{A_2 B_1 - A_1 B_2}$$

and similarly,

$$\ddot{y}_2 = \ddot{q}_1 - \ddot{q}_2 = \dots$$

$$\ddot{y}_3 = \ddot{q}_1 + \ddot{q}_2 = \dots$$

For  $m_1$  to be stationary, we want to establish when:

$$\ddot{y}_1 = \frac{B_1 C_2 - B_2 C_1}{A_2 B_1 - A_1 B_2} = 0$$

ie., want to solve:

$$B_1 C_2 - B_2 C_1 = 0$$

(assuming  $A_2 B_1 - A_1 B_2 \neq 0$ )

↓

$$(m_2 - m_3)^2 g + (m_2 + m_3)(m_1 - (m_2 + m_3))g = 0$$

↓

$$m_1 = m_2 + m_3 - \frac{(m_2 - m_3)^2}{m_2 + m_3}$$

An example is for  $m_1 = m_2 + m_3$  ~~and~~  $m_2 = m_3$ .  
 Then  $m_1$  is stationary (which makes sense because the mass of the lower pulley system balances the mass  $m_1$ ) and it is also straightforward to show that  $m_2 = m_3$  implies  $\ddot{q}_2 = 0$  & thus  $\ddot{y}_2 = 0$ , so the lower pulley system is also stationary.

Question 3:

a) First, let us assume the length of the rope is such that  $L=h$  (see Fig 3), such that the gravitational potential energy for the block can be written as,

$$V = Mgd$$

↖ height of mass  $M$  above floor.

Further, we can use polar co-ordinates to characterize the motion of the block of mass  $m$  on the table,

$\theta \rightarrow$  angular co-ordinate

$r \rightarrow$  radial co-ord. (defined relative to hole in table)

With this in hand, the kinetic energy of both masses can be ~~was~~ written as,

$$T_m = \cancel{\frac{m}{2} \dot{x}^2} + \frac{m}{2} \dot{r}^2 + \frac{m}{2} r^2 \dot{\theta}^2 \quad (\Rightarrow \text{from } T = \frac{m}{2} \dot{x}^2 + \frac{m}{2} \dot{y}^2)$$

$$T_m = \frac{M}{2} \dot{d}^2$$

However, note that the fixed length ~~holding the~~ of the rope holding both masses together implies, (15)

$$\dot{r} = \dot{d}$$

& thus:

$$T_m = \frac{m}{2} \dot{d}^2 + \frac{m}{2} d^2 \dot{\theta}^2 \quad \left[ d \text{ \& } \theta \text{ are thus our two generalized co-ordinates} \right]$$

Hence, the Lagrangian of the total system is then,

$$L = \left( \frac{m+M}{2} \right) \dot{d}^2 + \frac{m}{2} d^2 \dot{\theta}^2 - Mgd$$

Lagrange's eqns of motion then yield,

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_j} \right) - \frac{\partial L}{\partial q_j} = 0 \quad \text{for } q_j = \theta, d$$

↳

$$\text{i) } \frac{d}{dt} (m d^2 \dot{\theta}) = 0$$

$$\text{ii) } \frac{d}{dt} ((m+M) \dot{d}) = m d \dot{\theta}^2 - Mg$$

The former equation illustrates that angular momentum is conserved,

$$m d^2 \dot{\Theta} = \text{const.} = \dot{C}$$

when enables us to rewrite (ii) as,

$$(m+M)\ddot{d} = -Mg + \frac{\dot{C}^2}{md^3} \quad (*)$$

Formally solving for  $d$  from this equation implicitly allows us to subsequently solve for  $\Theta$  from  $\dot{\Theta}(d)$ .

b) Clearly we have identified that angular momentum is conserved:

$$L = I \omega = \text{const.}$$

$$\begin{array}{cc} \nearrow & \nearrow \\ I = md^2 & \omega = \dot{\Theta} \\ \text{(moment of inertia)} & \text{(angular frequency)} \end{array}$$

Moreover inspection of (\*) indicates that:

$$\ddot{d} = 0 \rightarrow \frac{\dot{C}^2}{md^3} = Mg$$

e.g., the lower block is stationary when the force due to rotation of upper block balances gravity.



## Question 4

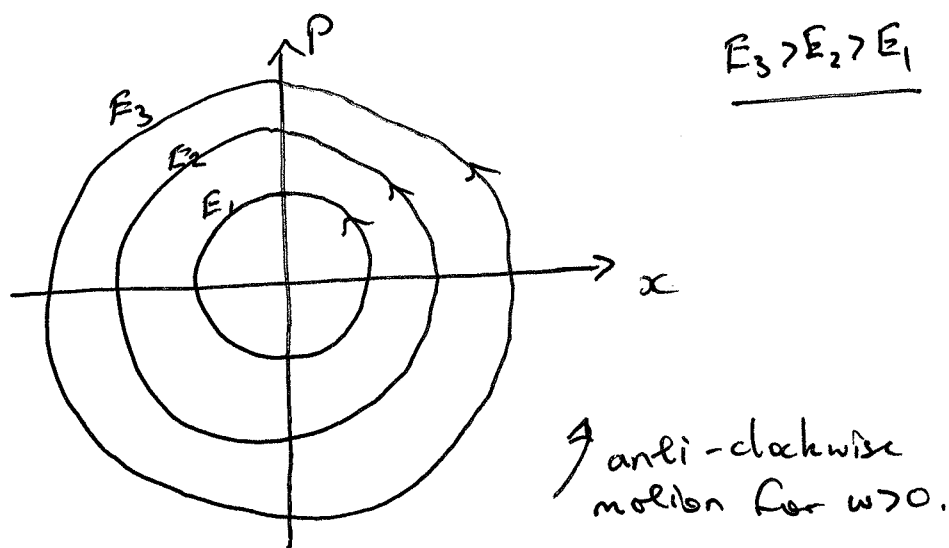
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a) The harmonic oscillator in 1D traces out circles in ~~the~~ phase space (for appropriately normalized position + momentum co-ordinates).

Here we take  $m=\omega=1$  for simplicity such that for any trajectory we have that,

$$E = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 x^2 = \frac{p^2}{2} + \frac{x^2}{2} = \text{constant}.$$

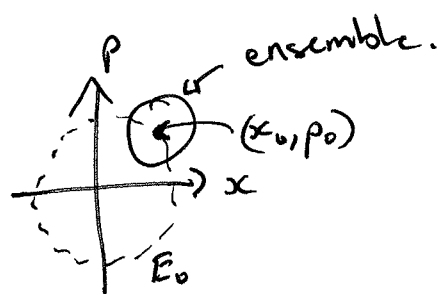
Phase portrait:



b) ~~We~~ We consider a swarm of points w/in an area defined by,

$$(x-x_0)^2 + (p-p_0)^2 = a$$

e.g., a circle centred @  $(x_0, p_0)$  w/ initial energy  $\bar{E}_0 = \frac{x_0^2}{2} + \frac{p_0^2}{2}$



Now, by definition of simple harmonic motion, each point w/in the ensemble will orbit about  $(x, p) = (0, 0)$  maintaining a fixed radius (due to conservation of energy) and at a fixed angular frequency ( $\omega = 1$ ). Hence, the circular area itself orbits w/ freq  $\omega = 1$  about the origin w/out being deformed. ~~there~~.

$\Rightarrow$  Phase-space area occupied by ensemble is conserved under evolution.

c) See Fig 4 of assignment.

Area of ensemble

$$A = \int_{p_1}^{p_2} dp \int_{q(p)} dq \cdot 1$$

$q \equiv$  spatial co-ordinate.

where  $q(p)$  is defined by  $E' + E''$  & the relation,

$$p = \pm \sqrt{2m(E - mgq)}$$

$E = T + V = \text{const.}$

$V(q)$  due to gravity

Then,

$$q = \frac{(E - p^2/2m)}{mg} \quad (18)$$

✓

$$A = \int_{p_1}^{p_2} dp \left[ (E'' - p^2/2m) - (E' - p^2/2m) \right] \times \frac{1}{mg}$$
$$= \frac{(E'' - E')(p_2 - p_1)}{mg} \quad (\sim \Delta E \Delta p)$$

d) Conservation of energy implies that,

$$p(t) = \pm \sqrt{2m(E - mgq(t))}$$

or

$$q(t) = \frac{1}{mg} \left( E - \frac{p(t)^2}{2m} \right)$$

Now, for a particle in 1D subject to gravity,

$$\ddot{q} = -g \rightarrow q(t) = -\frac{gt^2}{2} + \frac{p(0)t}{m} + q(0)$$

and thus,  $p(t) = -mgt + p(0)$

The latter solution for  $p(t)$  implies:

$$p_2(t) - p_1(t) = p_2(0) - p_1(0) \text{ is conserved.}$$

Thus the boundaries of the integrated region of the ensemble evolve in a well defined way:

$$\begin{aligned}
A(t) &= \int_{p_1(t)}^{p_2(t)} \int_{q(p)} dq \\
&= \int_{p_1(t)}^{p_2(t)} dp \left[ (E'' - p^2/2m) - (E' - p^2/2m) \right] \frac{1}{mg} \\
&= \frac{(E'' - E')(p_2(t) - p_1(t))}{mg} \\
&= \frac{(E'' - E')(p_2(0) - p_1(0))}{mg} = A(0)
\end{aligned}$$

Thus the area of the ensemble in phase-space is conserved, consistent w/ Liouville's theorem.