

Assignment 10 Solutions

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Question 1:

a) See your lecture notes for details!

- ① The central force is spherically symmetric, and so is the associated potential,

$$V(r) = -a/r^2$$

- ② The symmetry of the force/potential is inherited by the Lagrangian describing the particle's motion.

Spherical symmetry \Rightarrow conservation of total angular momentum.

- ③ By definition, our motion must be in a plane perpendicular to the angular momentum vector \vec{L} ,

posn vector of particle \vec{r}

$$\begin{aligned}\vec{r} \cdot \vec{L} &= m \vec{r} \cdot (\vec{r} \times \dot{\vec{r}}) \\ &= m \dot{\vec{r}} \cdot (\vec{r} \times \vec{r}) = 0 \quad \therefore \vec{r} \perp \vec{L}\end{aligned}$$

b) Using polar co-ordinates to describe motion in the plane, the Lagrangian is: (2)

$$L = \frac{m}{2} (\dot{r}^2 + r^2 \dot{\varphi}^2) - V(r)$$

The equations of motion are:

i) $\frac{d}{dt} (mr^2 \dot{\varphi}) = 0 \Rightarrow l = mr^2 \dot{\varphi}$ is conserved magnitude of angular momentum.

ii) $m \ddot{r} - mr \dot{\varphi}^2 + \frac{\partial V}{\partial r} = 0$

↗ replace w/ $\dot{\varphi} = l/mr^2$

↓

$$m \ddot{r} - \frac{l^2}{mr^3} + \frac{\partial V}{\partial r} = 0$$

or

$$m \ddot{r} = F_{\text{eff}}(r) \quad \text{w/} \quad F_{\text{eff}} = \frac{l^2}{mr^3} - \frac{\partial V}{\partial r}$$

$$\text{and } V_{\text{eff}} = \frac{l^2}{2mr^2} + V(r)$$

$$\text{such that } F_{\text{eff}} = -\frac{\partial V_{\text{eff}}}{\partial r}!$$

c)

$$V_{\text{eff}} = \frac{\frac{L^2}{2m} - a}{r^2}$$

$$\propto \frac{1}{r^2}$$

sign depends on
relative magnitude of
 $\frac{L^2}{2m} + a$.

i)

$$\frac{L^2}{2m} - a > 0$$



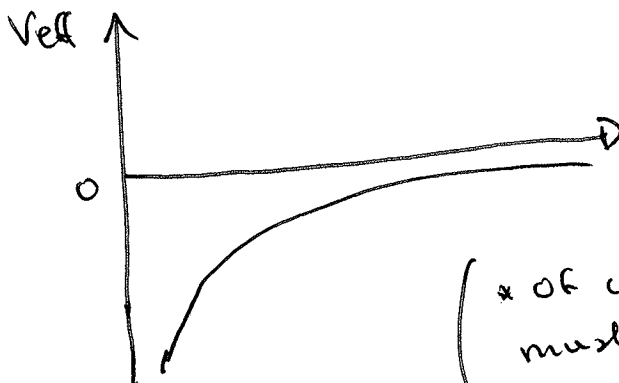
scattered motion
for any $E_0 > 0$

ii)

$$\frac{L^2}{2m} - a < 0$$

\Rightarrow divergent attractive
potential as $r \rightarrow 0$

\therefore particle is attracted
to $r=0$ for any
initial energy.



(* of course, physically our model
must break down as $r \rightarrow 0$...)

iii) A circular orbit requires $F_{\text{eff}} = 0$.

$$\Rightarrow \text{Solve for } \frac{\partial V_{\text{eff}}}{\partial r} = 0$$

\Rightarrow Only solution is trivial case w/

$$\frac{l^2}{2m} - a = 0 \quad \text{or} \quad V_{\text{eff}} = 0 !$$

d) We have that total mechanical energy can be written as,

$$E = T + V = \frac{m\dot{r}^2}{2} + V_{\text{eff}}(r)$$

and it is conserved.

Hence, rearranging we can obtain,

$$\dot{r} = \pm \sqrt{\frac{2}{m}(E - V_{\text{eff}})}$$

Separately,

$$\frac{d\varphi}{dt} = \frac{d\varphi}{dr} \cdot \frac{dr}{dt}$$

or,

$$\frac{d\varphi}{dr} = \frac{\dot{\varphi}}{\dot{r}}$$

$$= \pm \frac{L}{mr^2} \cdot \frac{1}{\sqrt{\frac{2}{m}(E - V_{eff})}}$$

$$\Rightarrow \varphi - \varphi_0 = \pm \int_{r_0}^r dr' \frac{L}{mr'^2} \frac{1}{\sqrt{\frac{2}{m}(E - V_{eff}(r'))}}$$

Question 2:

a) See your lecture notes!

b) This system of equations is known for exhibiting an Andronov-Hopf bifurcation.

Let's write down the equations in polar co-ordinates:

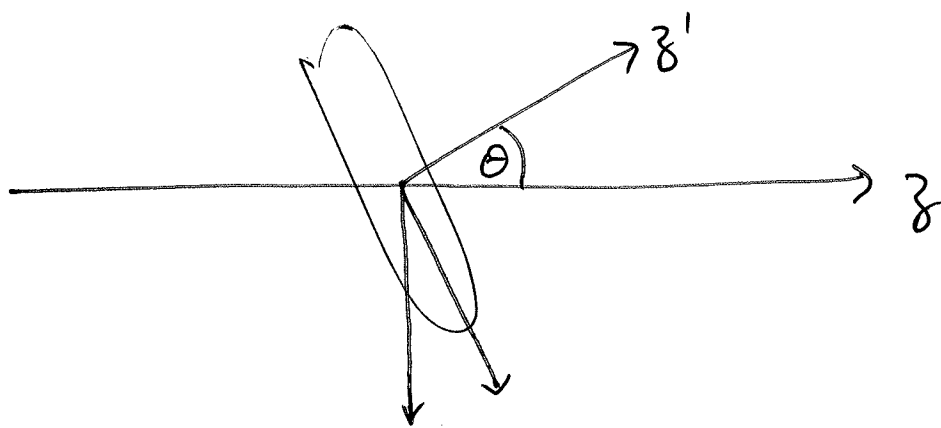
$$\dot{r} = ar - cr^3 = ar \left(1 - \frac{c}{a} r^2\right)$$

$$\dot{\theta} = b$$

Question 3:

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a)



Working w/ body-fixed z' along the normal to the disc, we compute the principal moment of inertia $I_{z'}$.

$$I_z = \int dV \rho_0 (x^2 + y^2) \quad (= I_{zz})$$

\nwarrow $\rho_0 = \frac{M}{\pi R^2} = \text{mass density}$

$$= \rho_0 \int_0^R dr \int_0^{2\pi} d\theta \, r^2 \, r dr \quad \left[\text{disc is of negligible thickness along } z' \right]$$

$$= 2\pi \rho_0 \frac{R^4}{4} = \frac{m R^2}{2}$$

The remaining moments are equal, $I_1 = I_2$. There are ⑨ two ways to obtain them.

i) $I_1 = I_2 = \int dV \cancel{y^2 + z^2} \overset{I_1}{\nearrow} 0$ as disc is negligibly thin

$$= \rho_0 \int_0^{2\pi} d\theta \int_0^R dr \, r^3 \sin\theta \quad \sim \text{evaluate using known integral or formula sheet.}$$

$$= \frac{mR^2}{4}$$

ii) $I_1 + I_2 = \int dV \cancel{x^2 + y^2 + 2z^2} \nearrow 0$

$$= I_3$$

$$\therefore I_1 = I_2 = I_3/2 = \frac{mR^2}{4}.$$

b) This is the same as the barbell problem we studied in Assignment 7. The key part to realize is that θ is fixed, & there is no rotation about body-fixed \hat{z} (the ~~the~~ disc is rigidly attached). Depending on how you choose your axes you should obtain something resembling,

$$\vec{\omega} = \dot{\phi} \begin{pmatrix} \sin\phi \sin\theta \\ \cos\phi \sin\theta \\ \cos\theta \end{pmatrix} \quad \text{w/ the } \sin\phi \text{ \& } \cos\phi \text{ terms depending on axes choice.}$$

c) The kinetic energy is solely due to rotation (as the origin of our body-fixed axes is at the com). Thus, we evaluate it using body-fixed quantities,

$$T = T_{\text{rot}} = \frac{1}{2} [I_1 \omega_1^2 + I_2 \omega_2^2 + I_3 \omega_3^2]$$

↙ $\vec{\omega}$ in body-fixed axes.

Using that $I_1 = I_2 = I_{\perp}$ we have:

$$T = \frac{I_{\perp}}{2} \dot{\phi}^2 \sin^2 \theta + \frac{I_3}{2} \dot{\phi}^2 \cos^2 \theta$$

d) We want $\dot{\vec{\omega}} = 0$ for the motion to be preserved (this is the same as the barbell in A7!).

Here w/ a fixed rotation frequency Ω ,

$$\vec{\omega} \rightarrow \Omega \begin{pmatrix} \sin \phi \sin \theta \\ \cos \phi \sin \theta \\ \cos \theta \end{pmatrix} \Rightarrow \sqrt{\omega_1^2 + \omega_2^2} = \cancel{\cos \theta} \Omega \sin \theta$$

Euler's equations give us:

$$\begin{aligned} (I_3 - I_2) \omega_2 \omega_3 &= N_1, & (I_2 - I_1) \omega_1 \omega_2 &= N_3 \\ (I_1 - I_3) \omega_1 \omega_3 &= N_2 \end{aligned}$$

$$I_1 = I_2 \Rightarrow N_3 = 0$$

(11)

$$|\vec{N}|^2 = \sqrt{N_1^2 + N_2^2} = \sqrt{\left(\frac{mR^2}{4} \omega_2 \omega_3\right)^2 + \left(\frac{mR^2}{4} \omega_1 \omega_3\right)^2}$$

$$= \frac{mR^2}{4} \sqrt{\omega_1^2 + \omega_2^2} |\omega_3|$$

$$= \frac{mR^2}{4} L^2 \cos \theta \sin \theta \quad (\text{for } 0 < \theta < \pi/2)$$