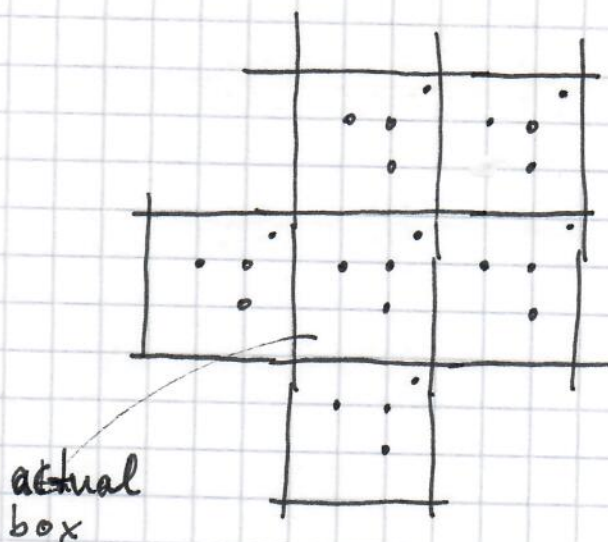


periodic boundary conditions:



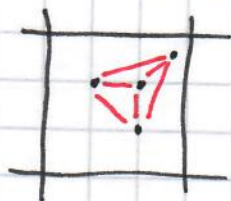
relevant parameter

$$\rho = \frac{N}{V}$$

With periodic boundary conditions, the system has infinite extent \rightarrow it is infinitely large

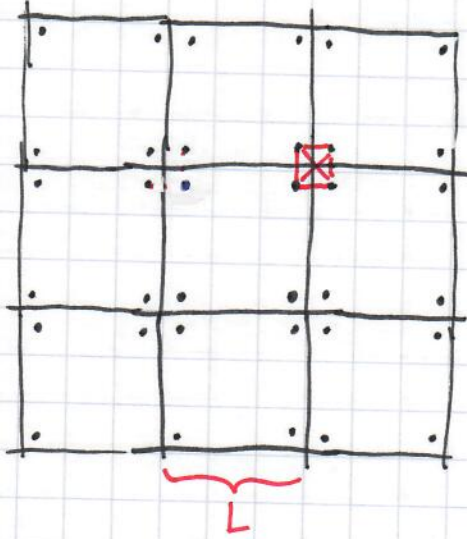
How many two-body interactions do we have in the case shown?

4 particles: $\frac{4 \cdot 3}{2} = 6$ two-body pairs



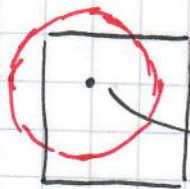
6 two-body pairs

Let's look at a slightly different arrangement



again, 6 two-body pairs

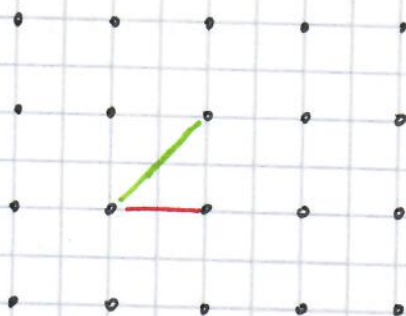
In principle, one would have an infinite number of pairs \rightarrow in practice (i.e., in computer simulations), one needs to choose a cut-off. Typically, we look for and account for all pairs that have a distance smaller than $\frac{L}{2}$ to all other particles.



this particle interacts w/ all particles located in the circle w/ radius $\frac{L}{2}$

Note: in a lattice, we talk about nearest neighbor and next-to-nearest neighbor

Square lattice :



— nearest neighbor interaction
— next-to-nearest neighbor interaction

A bit of review from undergrad quantum:

1144

Consider a quantum particle of mass m in a two-dimensional periodic box of area L^2 .

↑ note: we are not given the "statistical nature" of the quantum particle, i.e., we don't know whether it's a boson or a fermion.

→ does not matter for now. Eventually: Degeneracy factor.

- (a) What is the number of single-particle quantum states with energy less than E ?
- (b) What is the density of the single-particle states as a fct. of E ?

(a) Since we are considering a quantum particle, we need to solve the 2D Schrödinger equation:

$$-\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \psi(x, y) = E \psi(x, y)$$

We know: $\psi(x, y) = \varphi(x) \varphi(y)$

$$\frac{1}{\sqrt{L}} e^{ik_x x} \quad \frac{1}{\sqrt{L}} e^{ik_y y}$$

We need to make sure that our eigenstates fulfill the periodic boundary conditions:

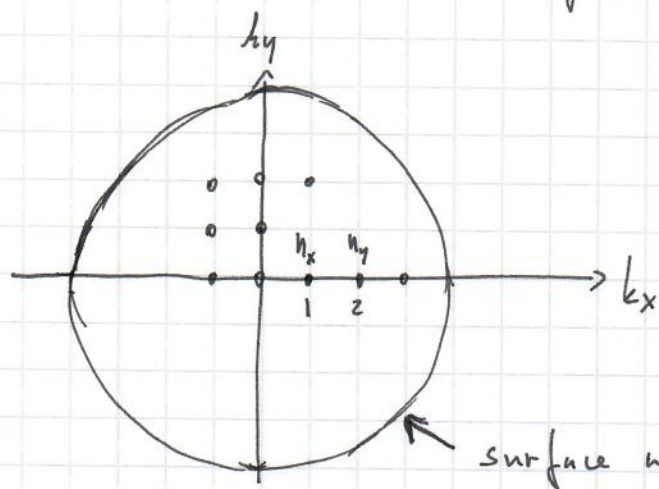
$$\begin{aligned} \varphi(x+L) &= \varphi(x) \\ \Rightarrow e^{ik_x(x+L)} &= e^{ik_x x} \Rightarrow e^{ik_x L} = 1 \end{aligned}$$

It follows: $k_x = \frac{2\pi}{L} n_x$, where $n_x = 0, \pm 1, \pm 2, \dots$

Similarly $k_y = \frac{2\pi}{L} n_y$, where $n_y = 0, \pm 1, \pm 2, \dots$

$$\begin{aligned} E \text{ is then given by } E &= \frac{\hbar^2 k_x^2}{2m} + \frac{\hbar^2 k_y^2}{2m} \\ &= \frac{\hbar^2}{2m} (k_x^2 + k_y^2) = \frac{(2\pi\hbar)^2}{2mL^2} (n_x^2 + n_y^2) \end{aligned}$$

We are looking for the number of states with $\varepsilon \leq E$, where E is fixed.



$$k_x^2 + k_y^2 = k^2$$

$$k = \sqrt{\frac{2mE}{\hbar^2}}$$

surface with $|\vec{k}| = k = \sqrt{\frac{2mE}{\hbar^2}}$

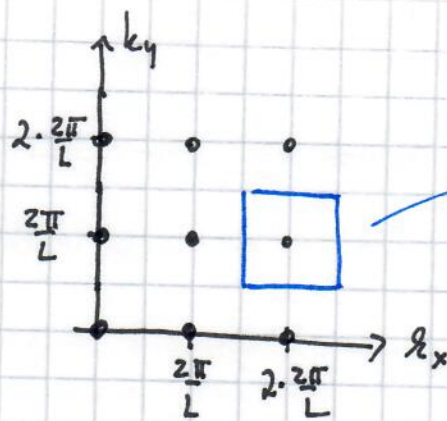
\Rightarrow we need to count all the states that fall inside this circle, which has an area of πk^2 .

We have exactly one state per $\left(\frac{2\pi}{L}\right)^2$

$$\Rightarrow \text{number of states } N(E) = \frac{\pi k^2}{\left(\frac{2\pi}{L}\right)^2} = \frac{L^2 \pi 2mE}{(2\pi \hbar)^2}$$

$$\Rightarrow N(E) = \frac{mEL^2}{2\pi \hbar^2}$$

(b) Density of states: $\frac{dN(E)}{dE} = \frac{mL^2}{2\pi \hbar^2}$



one state in each
area $(\frac{2\pi}{L})^2$

this is an area
in k -space and
thus has units
of length^{-2}

Let's rewrite the density of states expression:

$$\frac{dN(E)}{dE} = \frac{\frac{1}{2\pi}}{\frac{\hbar^2}{mL^2}}$$

this is for a two-
dimensional system

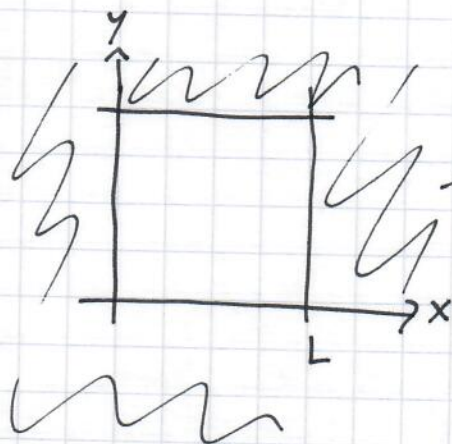
↑
scales as E^0

For a three-dimensional system:

$$\frac{dN(E)}{dE} = \frac{\frac{1}{\sqrt{2}} \frac{1}{\pi^2}}{\frac{\hbar^2}{mL^2}} \left(\frac{E}{\frac{\hbar^2}{mL^2}} \right)^{1/2}$$

↑
this scales as \sqrt{E}

Now: Let us consider the 2D system again but let us consider ~~like~~ hard wall boundary conditions instead.



infinite potential

$$V(x,y) = 0 \text{ for}$$

$$0 < x < L \text{ and } 0 < y < L$$

Again, we start off by looking at the single-particle SE:

$$-\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} \right) \psi_{\vec{n}}(x,y) = E_{\vec{n}} \psi_{\vec{n}}(x,y)$$

$$\psi_{\vec{n}}(0,0) = 0 = \psi_{\vec{n}}(L,L) = \psi_{\vec{n}}(x,0) = \psi_{\vec{n}}(0,y) =$$

$$\psi_{\vec{n}}(x,L) = \psi_{\vec{n}}(L,y)$$

$$\psi_{\vec{n}}(x,y) = \frac{2}{L} \sin\left(n_x \frac{\pi x}{L}\right) \sin\left(n_y \frac{\pi y}{L}\right)$$

note: $e^{ikx}, e^{-ikx} \rightarrow \sin(kx), \cos(kx)$
 two linearly indep. solutions two linearly indep. solutions

\sin/\cos are easier to work with in this case. Since $\cos(0) \neq 0$, the \cos solution needs to be discarded.

left with $\sin(k_x x)$ and $\sin(k_y y)$

need $\sin(k_x x) = 0$ for $x=0$ and $x=L$
 naturally fulfilled

$$\Rightarrow \sin(k_x L) = 0 \quad \text{or} \quad k_x L = n_x \pi$$

What values can n_x take?

To start w/ : $n_x = 0, \pm 1, \pm 2, \dots$

problem: if $n_x = 0$, then

$$\sin(k_x x) = 0$$

\Rightarrow we don't find the particle anywhere



What about $n_x = \pm 1$?

$$n_x = +1: \sin\left(\pi \frac{L}{x}\right)$$

$$n_x = -1: \sin\left(-\pi \frac{L}{x}\right) = -\sin\left(\pi \frac{L}{x}\right)$$

the \pm solutions are
not linearly independent

$\Rightarrow n_x = -1$ needs to
be excluded!

So:

$$\begin{aligned} n_x &= 1, 2, 3, \dots \\ n_y &= 1, 2, 3, \dots \end{aligned}$$

$$\epsilon_{\vec{n}} = \frac{\hbar^2 k_{\vec{n}}^2}{2m} = \frac{\hbar^2}{2m} \frac{\pi^2}{L^2} (n_x^2 + n_y^2)$$

these differ!

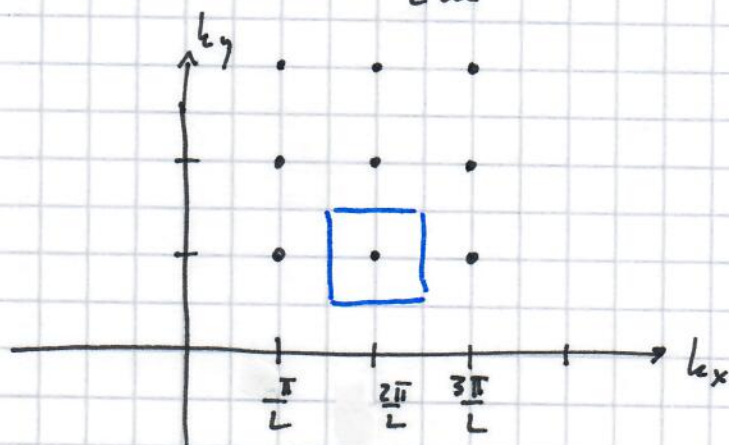
single-particle energy for hard wall
boundary conditions

Compare: Periodic BC: $\epsilon_{\vec{n}} = \frac{\hbar^2}{2m} \frac{\pi^2}{L^2} 4 (n_x^2 + n_y^2)$

$$n_x, n_y = 0, \pm 1, \pm 2, \dots$$

How many states do we have w/ energy less than or equal to E ?

look at $E = \frac{\hbar^2 k^2}{2m}$ or $k = \sqrt{\frac{2mE}{\hbar^2}}$



these are the allowed states in k -space

→ note: no allowed states at negative k_x or negative k_y

different from periodic BC case

→ the area in k -space is $(\frac{\pi}{L})^2$

If we are looking at states with $\leq E$ or $\leq k$, then we see that area taking up in k -space is $\frac{1}{4} \pi k^2$ (quarter of area of circle)

we are only looking at positive k since this is what is being imposed by the hard wall boundary conditions

$$\Rightarrow N(E) = \frac{\frac{1}{4} \pi \hbar^2}{\left(\frac{\pi}{L}\right)^2}$$

number of states
for 2D case - hard
wall BCs

L152

$$= \frac{\frac{1}{4} \pi \frac{2m}{\hbar^2} E}{\frac{\pi^2}{L^2}}$$

$$= \frac{\frac{1}{2\pi} E}{\frac{\hbar^2}{mL^2}} = \frac{1}{2\pi} E \frac{1}{\frac{\hbar^2}{mL^2}}$$

Density of states for
hard wall case:

$$\frac{dN(E)}{dE} = \frac{1}{2\pi} \frac{1}{\frac{\hbar^2}{mL^2}}$$



same as in case
of periodic BCs



same as in case of
periodic BCs

8 Quantum Statistical Mechanics

153

What did we do for the canonical ensemble?

↙
physical situation:
energy exchange

classical!

Partition function $Q(V, T, N)$: — or $Q_N(V, T)$

$$Q(V, T, N) = \frac{1}{N! h^{3N}} \int e^{-\beta \mathcal{H}(\vec{p}, \vec{q})} d^{3N} \vec{p} d^{3N} \vec{q}$$

$$= \text{Tr} \left(\exp(-\beta \mathcal{H}(\vec{p}, \vec{q})) \right)$$

trace implies
prefactors
and that we
are integrating
over the entire
phase space

↗
view this equation as a
definition of the trace (Tr)
in classical statistical mechanics
when working in the canonical
ensemble

$$-\beta A = \log(Q(V, T, N))$$

We had defined

$$\langle f \rangle = \frac{\int \exp(-\beta \mathcal{H}) f \, d^{3N} \vec{p} \, d^{3N} \vec{q}}{\int \exp(-\beta \mathcal{H}) \, d^{3N} \vec{p} \, d^{3N} \vec{q}}$$

↗
mean value of
observable f (or ensemble average)

Let's rewrite this:

$$\langle f \rangle = \frac{\frac{1}{N! h^{3N}} \int \exp(-\beta \mathcal{H}) f \, d^{3N} \vec{p} \, d^{3N} \vec{q}}{Q(V, T, N)}$$

Let's define a density (matrix) $\rho_{\text{can}}(\vec{p}, \vec{q})$:

$$\rho_{\text{can}}(\vec{p}, \vec{q}) = \frac{e^{-\beta \mathcal{H}(\vec{p}, \vec{q})}}{Q(V, T, N)}$$

$$\Rightarrow \langle f \rangle = \frac{1}{N! h^{3N}} \int \rho_{\text{can}}(\vec{p}, \vec{q}) f \, d^{3N} \vec{p} \, d^{3N} \vec{q}$$

$$= \text{Tr}(\rho_{\text{can}} f)$$

We haven't done anything new!

All we did on pages 153/154 is to rewrite the known classical statistical^{mechanics} expressions for the canonical ensemble in a "fancy" compact notation: Tr

ρ_{can}

Why? "Trace" and "density matrix" are concepts well established in quantum mechanics.

Idea: Use quantum mechanical density matrix $\hat{\rho}_{\text{can}}$ and quantum mechanical "Tr operation" and adopt all formalism / equations from classical statistical mechanics.

Let's try to make this more explicit.

What does $Q_N(V, T) = \text{Tr} (e^{-\beta \mathcal{H}})$ mean in quantum statistical mechanics?

$\mathcal{H} \rightarrow \hat{\mathcal{H}}$
 \uparrow
 classically,
 $\mathcal{H} = \mathcal{H}(\vec{p}, \vec{q})$

\nwarrow
 this becomes an operator

$$\Rightarrow \text{Tr} (e^{-\beta \hat{\mathcal{H}}}) = \dots ?$$

if we have an operator, then we need to act on "s.th."...

Moreover, if we have an operator in the exponent, then we need to do a Taylor expansion of the exponential, then act with each term in the infinite sum onto "s.th.", then hopefully collect or resum infinite set of terms to get compact expression.

Let's do it: $e^{-x} = \sum_{n=0}^{\infty} \frac{(-1)^n}{n!} x^n$

$$\Rightarrow e^{-\beta \hat{\mathcal{H}}} = \sum_{n=0}^{\infty} \frac{(-1)^n}{n!} (\beta \hat{\mathcal{H}})^n$$

$\underbrace{\beta^n \hat{\mathcal{H}}^n}_{\hat{\mathcal{H}} \cdot \hat{\mathcal{H}} \cdots \hat{\mathcal{H}}}$
 $n \text{ factors}$

What next? Let $\{\psi_n\}$ be a complete set of eigenstates of $\hat{\mathcal{H}}$.

Note: In general, we don't know the ψ_n , especially not if we are dealing with an interacting N -particle system!

$$\hat{\mathcal{H}} |\psi_n\rangle = E_n |\psi_n\rangle$$

$\underbrace{\quad}_{\text{eigenenergy}}$ — $\underbrace{\quad}_{\text{eigenstate}}$

Closure relation: $\sum_n |\psi_n\rangle \langle \psi_n| = 1$ the sum goes over a complete set of states!

if we have sc. states,
the sum turns into
integral

Importantly: the sum extends
over all eigenstates
of the complete
set!

$$\text{So: } \exp(-\beta \hat{\mathcal{H}}) = \exp(-\beta \hat{\mathcal{H}}) \left(\sum_m |\psi_m\rangle \langle \psi_m| \right) = \hat{1}$$

$$= \sum_n \frac{(-1)^n}{n!} \beta^n \hat{\mathcal{H}}^n \sum_m |\psi_m\rangle \langle \psi_m|$$

$$= \sum_m \sum_n \frac{(-1)^n}{n!} \beta^n \hat{\mathcal{H}}^n |\psi_m\rangle \langle \psi_m|$$

$$\underbrace{\hat{\mathcal{H}} \hat{\mathcal{H}} \dots \hat{\mathcal{H}} |\psi_m\rangle}_{E_m |\psi_m\rangle} = (E_m)^n |\psi_m\rangle$$

$$= \sum_m \sum_n \underbrace{\frac{(-1)^n}{n!} \beta^n (E_m)^n}_{\exp(-\beta E_m)} |\psi_m\rangle \langle \psi_m|$$

$$= \sum_m \exp(-\beta E_m) |\psi_m\rangle \langle \psi_m|$$

$$\text{So: } e^{-\beta \hat{\mathcal{H}}} = \sum_m \exp(-\beta E_m) |\psi_m\rangle \langle \psi_m|$$