

11.3 ff.: Magnetic behavior of an ideal Fermi

gas

Consider : gas of non-interacting fermions in the presence of an external magnetic field \vec{B}

Comment: We previously looked at system of N magnetic moments

• • • • •

magnetic moments
are fixed in space
→ can point up
or down

$$\mathcal{H} = \sum_{i=1}^N h_i$$

$$\mu = N \bar{\mu}$$

/ could be \vec{s}
gyromagnetic
factor

Partition function:
(single magnetic moment)

$$e^{\beta \mu B} + e^{-\beta \mu B}$$

$$= 2 \cosh(\beta \mu B)$$

$$p_{\pm} = e^{\pm \beta \mu B} \frac{1}{2 \cosh(\beta \mu B)}$$

$$\Rightarrow M = N \langle \mu \rangle = N \mu \tanh\left(\frac{\mu B}{kT}\right)$$

magnetization $\langle \mu \rangle = \mu(p_+ - p_-)$

- In this chapter we are considering a different situation :
- the magnetic moments are not fixed in space instead
 - the fermions form a non-interacting gas ($\hat{=}$ are treated as a non-interacting gas) and placed into an external magnetic field

M : Average induced magnetic moment per unit volume of the system along the direction of an external magnetic field B

Note: the text uses H instead of B for B -field

$$M = \frac{1}{V} \left\langle - \frac{\partial \mathcal{H}}{\partial B} \right\rangle \quad \text{like a "force equation"}$$

$$\Rightarrow \boxed{M = kT \left(\frac{\partial}{\partial B} \frac{\log Q_N}{V} \right)_{T,V,N}}$$

derived on next page

for canonical ensemble

$$M = kT \left(\frac{\partial}{\partial B} \left(\frac{\log \Omega}{V} \right) \right)_{T,V,z}$$

for grand canonical ensemble

z to be eliminated in terms of N

Math detail (derivation of $H = kT \frac{\partial}{\partial \beta} (\ln Q_N)$):

Start with: this eq. defines H

$$H = \frac{1}{V} \left\langle - \frac{\partial \mathcal{Z}}{\partial \beta} \right\rangle$$

by def.: classical SM,
canonical
ensemble

$$= \frac{-1}{V} \frac{\int \frac{\partial \mathcal{Z}}{\partial \beta} e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}{\int e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}$$

$$= -\frac{1}{V\beta} \frac{\int \left(\frac{\partial}{\partial \beta} (\beta \mathcal{E}) \right) e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}{\int e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}$$

$$\beta = \frac{1}{kT}$$

using \star $\rightarrow = \frac{kT}{V} \frac{\int \frac{\partial}{\partial \beta} (e^{-\beta \mathcal{E}}) d^{3N} p d^{3N} q}{\int e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}$

Since β is a constant, the derivative can be taken outside the integral

$$\text{Now: } \frac{\partial}{\partial \beta} (e^{-\beta \mathcal{E}}) = e^{-\beta \mathcal{E}} \left(\frac{\partial}{\partial \beta} (-\beta \mathcal{E}) \right) \star$$

$$= \frac{kT}{V} \underbrace{\frac{\int \frac{\partial}{\partial \beta} e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}{\int e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}}$$

of the form $\frac{\frac{\partial}{\partial \beta} f}{f}$

rewrite as $\frac{\partial}{\partial \beta} \log f$

$$= \frac{kT}{V} \frac{\partial}{\partial \beta} \left(\log \left(\underbrace{\int e^{-\beta \mathcal{E}} d^{3N} p d^{3N} q}_{Q_N} \right) \right)$$

$$H = kT \left(\frac{\partial}{\partial \beta} \left(\frac{\log Q_N}{V} \right) \right)$$

We can do the same thing for quantum SM.

χ : magnetic susceptibility per unit volume

$$\chi = \frac{\partial M}{\partial B}$$

Back to classical system discussed on p. 254:

using
 $M = N\mu \tanh\left(\frac{\mu B}{kT}\right)$
from page 254

at high T : $M = \frac{N\mu^2 B}{kT}$ (use $\tanh x \approx x$)

$$\chi = \frac{N\mu^2}{kT}$$
 Curie law

$$\begin{aligned} & \frac{\partial}{\partial B} \tanh\left(\frac{\mu B}{kT}\right) \\ &= \frac{k}{kT} \underbrace{\text{Sech}^2\left(\frac{\mu B}{kT}\right)}_{\rightarrow 1 \text{ for } B \rightarrow 0} \end{aligned}$$

Susceptibility at zero magnetic field at any temperature:

$$\chi_0 = \lim_{B \rightarrow 0} \frac{\partial M}{\partial B} = \frac{N\mu^2}{kT}$$

$\chi > 0$: paramagnet

at low T : "energy dominates" \rightarrow smaller E
at higher T : "entropy dominates"
 \hookrightarrow larger S

system wants to minimize free energy
 $A = U - TS$ Helmholtz

In general :

$\chi > 0$: paramagnetism

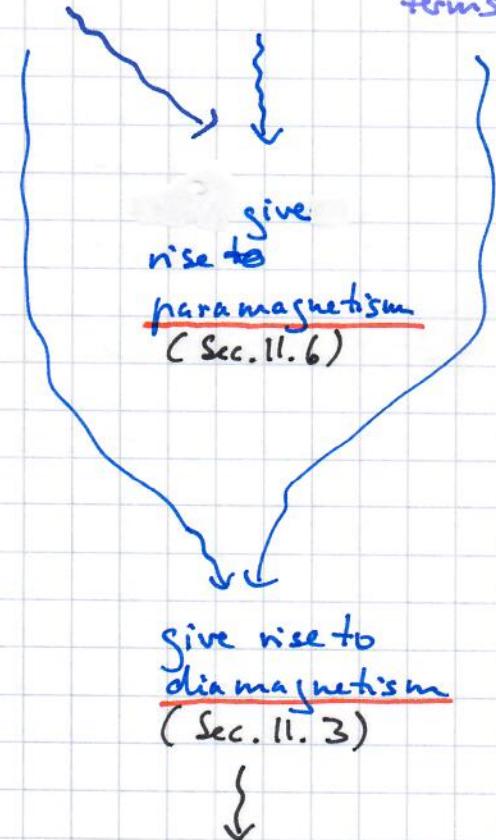
$\chi < 0$: diamagnetism

We want to treat non-interacting Fermi gas in B -field quantum mechanically :

single-particle energy will be $\frac{p^2}{2m} - \vec{\mu} \cdot \vec{B} + \vec{A}$ -dep. terms

$\vec{\mu}$: intrinsic magnetic moment of particle

\vec{A} : vector potential



due to quantization of the orbits
of charged particles in the presence
of external magnetic field

might alternatively say:

due to the quantization of the (kinetic) energy of charged particles associated with their motion perpendicular to the direction of the field

In general, we have a "competition" of paramagnetic and diamagnetic tendency

→ spin-orbit coupling and electron-electron (or fermion-fermion) interactions also

need to be taken into account → they lead to a coupling of the two tendencies

note: in condensed matter systems, spin-orbit coupling is $\vec{p} \cdot \vec{s}$ term

(in atomic and nuclear physics, spin-orbit coupling refers to $3 \cdot \ell$ coupling term → plays a key role in explaining shell structure of nuclei!!!)

Let's start with Pauli paramagnetism (Sec. 11.6)

Single-particle Hamiltonian:

$$\hat{H} = \frac{\hat{P}^2}{2m} - \hat{\vec{\mu}} \cdot \hat{\vec{B}}$$

↑
intrinsic (spin)
magnetic
moment

Say, we focus on spin- $\frac{1}{2}$ particle for simplicity.
specifically, electron

spin
Intrinsic magnetic moment of electron:

$$\hat{\vec{\mu}} = -g_s \mu_B \frac{1}{\hbar} \hat{\vec{S}}$$

↑
magnetic moment

is anti-parallel to

the spin angular momentum

in SI units

g_s : Bohr magneton ($\mu_B = \frac{e\hbar}{2m_e}$)

g_s : spin g-factor ($g_s \approx 2.002\ 319\dots$)

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note: g_s is 2 according to
Dirac equation \rightarrow the small
deviation from 2 is due to
quantum electrodynamics

numerical value of g_B :

$$\mu_B = 5.788 \cdot 10^{-5} \frac{\text{eV}}{\text{T}}$$

$$1 \text{ Tesla} = 10,000 \text{ G} = 10^4 \text{ G}$$

$$\text{So: } g_B = 5.788 \cdot 10^{-1} \frac{\text{eV}}{\text{G}}$$

In laboratory settings, B-field strengths up to 1000 G can be generated without too much effort

Let magnetic field lie along +z-axis:

$$\vec{B} = B \hat{e}_z \Rightarrow -\hat{\mu} \cdot \hat{B} = -B \hat{\mu}_z$$

↗
 B is strength
 of B-field

↑
 z-component

$$\frac{1}{\hbar} \hat{S}_z = \frac{1}{2} \hat{b}_z = \frac{1}{2} (| \uparrow \rangle \langle \uparrow | - | \downarrow \rangle \langle \downarrow |)$$

$$\text{So, } \hat{H} = \frac{\hat{p}^2}{2m} + g_B B \hat{c}_z$$

\downarrow
 $- \hat{\mu} \cdot \vec{B} = - \underbrace{\left(-2g_B \frac{1}{2} \hat{c}_z \right) B}_{+}$

using $g_s = 2$

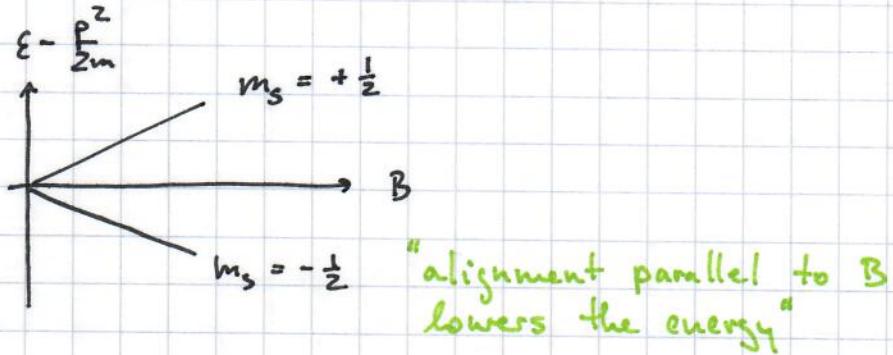
(note: the sign of
the second term changed!)

Eigenvalues of single-particle Hamiltonian:

$$\epsilon = \frac{\hat{p}^2}{2m} \pm g_B B$$

$m_s = \frac{1}{2}$ ↗ + : spin ↑ ($\hat{\mu}$ anti-parallel to \hat{B})

$m_s = -\frac{1}{2}$ ↙ - : spin ↓ ($\hat{\mu}$ parallel to \hat{B})



Our goal is now to obtain the susceptibility χ as a function of temperature.

$$E_N = \sum_k \sum_{m_s} E_{k,m_s} n_{k,m_s}$$

↑
energy of
 N -particle
system

the subscript
labels the N -particle
energies



Occupation number of the
single-particle energy level E_{k,m_s}

$$\rightarrow = \sum_k \left[\underbrace{\left(\frac{\hbar^2 k^2}{2m} - \mu_B B \right) n_k^+}_{m_s = -\frac{1}{2}} + \underbrace{\left(\frac{\hbar^2 k^2}{2m} + \mu_B B \right) n_k^-}_{m_s = +\frac{1}{2}} \right]$$

"Resolve" the
sum over m_s



Superscript +:
magnetic moment
parallel to field
($m_s = -\frac{1}{2}$)

$$n_k^+ \text{ and } n_k^- : 0 \text{ or } 1$$

$$\sum_k n_k^+ = N^+$$

$$\sum_k n_k^- = N^- = N - N^+$$

} definitions
(equations that
define N^+ and N^-)

$$\Rightarrow E_N = \sum_k (n_k^+ + n_k^-) \frac{\hbar^2 k^2}{2m} - \mu_B B (N^+ - N^-)$$

Since we have the N -particle energies (at least
in principle), we can calculate the partition function.

$$Q_N = \sum' e^{-\beta E_h}$$

$\{n_h^+\}, \{n_h^-\}$

the prime on the sum reminds us that we have restrictions

Sum over all E_h subject to the conditions / restrictions:

$$n_h^+, n_h^- = 0 \text{ or } 1$$

$$\sum_h n_h^+ + \sum_h n_h^- = N$$

The evaluation of the sum is a bit tricky due to the fact that we have constraints.

Let's pick one specific N^+ value and sum over all sets $\{n_h^+\}, \{n_h^-\}$ such that

$$\sum_h n_h^+ = N^+ \quad \text{restriction 1}$$

$$\text{and } \sum_h n_h^- = N - N^+ \quad \text{restriction 2}$$

Then we sum over all possible N_+ ($N_+ = 0, 1, \dots, N$).

$$\Rightarrow Q_N = \sum_{N_+=0}^N \left[e^{\beta g_B B (2N^+ - N)} \left\{ \sum'' e^{-\beta \sum_h \frac{t_h^2 h^2}{2m} n_h^+} \left| \begin{array}{l} \sum''' e^{-\beta \sum_h \frac{t_h^2 h^2}{2m} n_h^-} \\ \{n_h^-\} \end{array} \right. \right\} \right]$$

↑
Subject to restriction 1

↑
Subject to restriction 2

this is "just"
 $e^{\beta g_B B (N^+ - N^-)}$

$$\text{Define: } Q_N^{(0)} = \sum_{\sum n_i^z = N} e^{-\beta \sum_i \frac{\hbar^2 \omega_i}{2m} n_i^z} = e^{-\beta A_0(N)}$$

earlier, this type of sum was written as

$$\sum_{\sum n_i^z}$$

this is the partition fct. of the spinless ideal Fermi gas consisting of N particles

$$\Rightarrow Q_N = e^{-\beta g_B B N} \sum_{N_+ = 0}^N e^{2\beta g_B B N_+} Q_{N_+}^{(0)} Q_N^{(0)}$$

or $Q_{N-N_+}^{(0)}$

taking logarithm
=>

$$\frac{1}{N} \log Q_N = -\beta g_B B + \frac{1}{N} \log \left(\sum_{N_+ = 0}^N e^{2\beta g_B B N_+ - \beta A_0(N_+) - \beta A_0(N_-)} \right)$$

$N+1$ positive terms

$\log(\text{largest term in sum}) + \frac{1}{N} \log N$
type terms
correction

To find largest contribution: look at exponent and take derivative with respect to N_+

$$\rightarrow 2g_B B - \frac{\partial A_0(N_+)}{\partial N_+} - \frac{\partial A_0(N-N_+)}{\partial N_+} = 0$$

we're forcing this to be zero

NR:

$$\frac{\partial A_0(N-N_+)}{\partial N_+}$$

$$= \frac{\partial A_0(N-N_+)}{\partial N_-} \frac{\partial N_-}{\partial N_+}$$

$$= - \frac{\partial A_0(N-N_+)}{\partial N_-}$$

$$\underbrace{g_0(N_+)}_{\text{---}}$$

$$\underbrace{-g_0(N-N_+)}_{\text{---}}$$

this equation yields \bar{N}_+

(the \bar{N}_+ found in this way gives the largest contribution to the sum inside the logarithm)

$$\rightsquigarrow 2g_B B = g_0(\bar{N}_+) - g_0(N - \bar{N}_+)$$

1 /

recall, these are chemical potentials of our fictitious spinless Fermi gas

While not overly explicit, we have found a general solution!

More detail on: How to get $M = \frac{1}{V} g_B N_r$?

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By definition : $\mathcal{M} = kT \left(\frac{\partial}{\partial B} \left(\frac{\log Q_N}{V} \right) \right)_{N, T, V}$ see page 255

What do we have for $\log Q_N$? Look at bottom of page 264.

$$\log Q_N = -\beta N \mu_B B + \log ("(N+1) \text{ positive terms}")$$

We found that the largest contribution was given for $\sum \mu_B B = \mu_0 \left(\frac{1+r}{2} N \right) - \mu_0 \left(\frac{1-r}{2} N \right)$

result given on page 266, \star

$$\Rightarrow \log \left(\sum_{N_+=0}^N e^{2\beta \mu_B B N_+ - \beta A_0(N_+) - \beta A_0(N_-)} \right)$$

$$\xrightarrow{\text{keeping just one term}} 2\beta \mu_B \bar{N}_+ - \beta A_0(\bar{N}_+) - \beta A_0(N - \bar{N}_+)$$

$$\Rightarrow \log Q_N \approx \beta \mu_B B \underbrace{(2\bar{N}_+ - N)}_{N_r} - \beta A_0 \left(\frac{r+1}{2} N \right) - \beta A_0 \left(\frac{1-r}{2} N \right)$$

these are the Helmholtz free energies of the fictitious spinless Fermi gas

→ these terms are independent of B !!!

$$\Rightarrow \mathcal{M} = kT \left(\frac{\partial}{\partial B} \left(\frac{\log Q_N}{V} \right) \right)_{N, T, V} = \frac{1}{V} \mu_B N_r = \frac{1}{V} \mu_B (\bar{N}_+ - \bar{N}_-) \xrightarrow{\text{by def. of } r}$$

Let's define : $\bar{N}_+ - \bar{N}_- = Nr$, where $0 \leq r \leq 1$.

Using r , we can write: $2\bar{N}_+ - N \rightarrow 2\bar{N}_+ = N(r+1) \Rightarrow \bar{N}_+ = \frac{1+r}{2} N$

$$2g_B B = \mu_0 \left(\frac{1+r}{2} N \right) - \mu_0 \left(\frac{1-r}{2} N \right) \quad (*)$$

on previous page: \bar{N}_+ $N - \bar{N}_+$

$$\left. \begin{aligned} B=0 \text{ implies} \\ \mu_0 \left(\frac{1+r}{2} N \right) \\ = \mu_0 \left(\frac{1-r}{2} N \right) \\ \Rightarrow r=0 \end{aligned} \right\}$$

if $B \rightarrow 0 \Rightarrow r \rightarrow 0$ and $\bar{N}_+ \rightarrow \bar{N}_-$

So, we can think of small r as corresponding to small B

$$\text{Note: } \boxed{V\mathcal{M} = \mu_B (\bar{N}_+ - \bar{N}_-) = \mu_B N r}$$

See next page
for derivation of
 $V\mathcal{M} = \mu_B N r$

to get small- r expression,

Taylor expand r.h.s. of $(*)$

around $r=0$ (leading terms cancel,
first order terms give $\frac{1}{2} + \frac{1}{2} = 1$):

$$\sim 2g_B B \approx \left. \frac{\partial \mu_0(xN)}{\partial x} \right|_{x=\frac{1}{2}} r$$

$$\Rightarrow r = \left. \frac{2g_B B}{\partial \mu_0(xN)} \right|_{x=\frac{1}{2}}$$

Susceptibility per unit volume $\Rightarrow \chi = \frac{\partial M}{\partial B}$

From $M = \frac{1}{V} \mu_B N r$

we get

$$M = \frac{1}{V} \mu_B^2 N \frac{2B}{\left. \frac{\partial \mu_0(xN)}{\partial x} \right|_{x=1}}$$

using result for r

$$= \frac{\partial}{\partial B} \left(\frac{2\mu_B^2 B N}{\sqrt{\left. \frac{\partial \mu_0(xN)}{\partial x} \right|_{x=\frac{1}{2}}}} \right)$$

$$= \frac{1}{V} \frac{2\mu_B^2 N}{\left. \frac{\partial \mu_0(xN)}{\partial x} \right|_{x=\frac{1}{2}}} = \frac{2n\mu_B^2}{\left. \frac{\partial \mu_0(xN)}{\partial x} \right|_{x=\frac{1}{2}}}$$

density $n = \frac{N}{V}$

What is $\mu_0(xN)$?

$$\mu_0(xN) = \left(\frac{3xN}{4\pi V} \right)^{2/3} \frac{h^2}{Z_m} \quad \text{(this is } h - \text{not } h!)$$

So: χ is positive \rightsquigarrow paramagnetic effect !!!

Next: Landau diamagnetism (see Sec. 11.3)

Note: the discussion here is fairly brief → I want to bring the idea across.

↓
orbital motion
feels B field

e positive
(charge of electron - e)

Classically: $\mathcal{H}_{cl} = \frac{1}{2m} \left(\vec{p} + \frac{e}{c} \vec{A} \right)^2$ (for electron)

↑
single-particle Hamiltonian

↓
vector potential

Let's choose the following gauge: $A_x = -By$

$A_y = 0$	magnetic field along z-direction
$A_z = 0$	

$$\Rightarrow \mathcal{H}_d = \frac{1}{2m} \left(p_x - \frac{e}{c} By \right)^2 + \frac{p_y^2}{2m} + \frac{p_z^2}{2m}$$

Now quantize:

$$\hat{\mathcal{H}} = \frac{1}{2m} \left(\hat{p}_x - \frac{e}{c} B \hat{y} \right)^2 + \frac{\hat{p}_y^2}{2m} + \frac{\hat{p}_z^2}{2m}$$

Strategy:

- Solve single-particle Hamiltonian
- Construct grand partition function
- Calculate susceptibility

Make educated guess:

$$\psi(x, y, z) = e^{i(k_x x + k_z z)} f(y)$$

plane waves
in x and z

We find: $\left[\frac{\hat{p}_y^2}{2m} + \frac{1}{2} m \omega_0^2 (y - y_0)^2 \right] f(y) = \bar{\Sigma} f(y)$

where $\bar{\Sigma} = \Sigma - \frac{\hbar^2 k_z^2}{2m}$

$$\omega_0 = \frac{eB}{mc}$$

cyclotron frequency

$$y_0 = \frac{\hbar c}{eB} k_x$$

$$\Rightarrow \Sigma = \underbrace{\frac{e\hbar B}{mc}}_{\hbar \omega_0} \left(j + \frac{1}{2} \right) + \frac{\hbar^2 k_z^2}{2m}$$

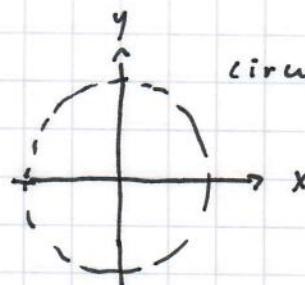
↙

$$j = 0, 1, 2, \dots$$

→ no dependence
on k_x
yes and no...

there is a somewhat
hidden dependence due to
degeneracy of energy levels

Classical trajectory:



circular motion \rightarrow
angular velocity ω_0



helical path
(constant linear
velocity)

Quantum mechanical energy level structure:

xy-plane: quantization in units of two.

recall $\gamma_0 = \frac{hc}{eB} k_x \rightarrow$ allowed values of k_x :

$$k_x = \frac{2\pi}{L} n_x ; n_x = 0, \pm 1, \pm 2, \dots$$

Assume system is inside a cube of size L^3

$\Rightarrow \gamma_0$ must lie between 0 and L

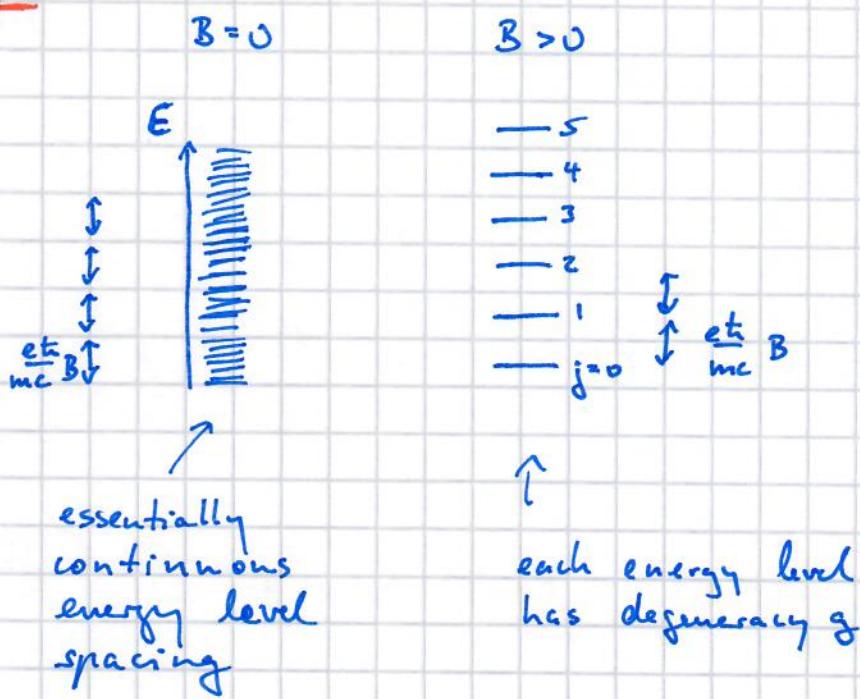
$\Rightarrow k_x = \frac{eB}{hc} \gamma_0$ must lie between 0 and $\frac{eB}{hc} L$

$\Rightarrow n_x \geq 0$ and $n_x \leq \frac{L k_x}{2\pi} \leq \frac{eB}{hc} L^2$

$g = \frac{eB}{hc} L^2$ is degeneracy
of each Landau
level

The degeneracy or multiplicity factor g is a quantum mechanical measure of the freedom available to the particle (for the center of its orbit to be located anywhere in the area L^3).

Pictorially:



Grand partition function:

$$\log \Omega = \sum_{\varepsilon} \log (1 + z e^{-\beta \varepsilon})$$

↑
Sum over all
single-particle
states

equilibrium number \bar{N} of particles:

$$\bar{N} = \left(-\frac{\partial}{\partial z} \log \Omega \right)_{B, V, T}$$

magnetic moment:

$$M = \mu T \left(\frac{\partial}{\partial B} \left(\frac{\log \Omega}{V} \right) \right)_{T, V, z}$$

After a bit of work: (let $x = \sqrt{\frac{e h}{4 \pi m c}} B = \beta g_B B$)

$$\bar{N} = \frac{z V}{\lambda^3} \frac{x}{\sinh x}$$

high temperature limit:

$$M = \frac{z}{\lambda^3} \frac{e h}{4 \pi m c} \left(\frac{1}{\sinh x} - \frac{x \cosh x}{\sinh^2 x} \right)$$

$z \ll 1$

$$\Rightarrow M = -\frac{\bar{N}}{V} \frac{e h}{4 \pi m c} \left(\coth x - \frac{1}{x} \right)$$

$$\begin{cases} g_B = \frac{e h}{2 m c} \\ \text{in CGS units} \end{cases}$$

$$= -\frac{\bar{N}}{V} g_B \left(\coth x - \frac{1}{x} \right)$$

Langevin fit.

$$X = -\frac{1}{3 k T_r} g_B^2$$

↑ diamagnetism

$\frac{1}{T}$ behavior (as in Curie's law)