

Problem 1:

(a) We have a frequency distribution function $D(\omega)$:

$$D(\omega) = \frac{A}{\omega} \quad \text{for } \omega_a \leq \omega \leq \omega_b$$

If we know the energy of an oscillator for a given temperature, then we can average that energy over the frequency distribution fct.

For 1D osc. with frequency ω at temperature T , we have

$$U_\omega = \frac{\hbar\omega}{2} \coth\left(\frac{\hbar\omega}{2kT}\right)$$

we can use
our result from
class for canonical
ensemble (see class notes p. 168)

$$\Rightarrow U = \int_{\omega_a}^{\omega_b} D(\omega) U_\omega d\omega$$

$$= A \frac{\hbar}{2} \int_{\omega_a}^{\omega_b} \coth\left(\frac{\hbar\omega}{2kT}\right) d\omega$$

$$\int \coth(x\omega) d\omega = \frac{1}{x} \log(\sinh(x\omega))$$

$$\rightarrow = \frac{A}{\beta} \log\left(\frac{\sinh\left(\frac{\hbar\omega_b}{2kT}\right)}{\sinh\left(\frac{\hbar\omega_a}{2kT}\right)}\right)$$

But: we also need to make sure that we are accounting for all the particles.

$$N = \int_{\omega_a}^{\omega_b} \mathcal{D}(\omega) d\omega = A \int_{\omega_a}^{\omega_b} \frac{1}{\omega} d\omega$$

$$= A \log\left(\frac{\omega_b}{\omega_a}\right)$$

$$\Rightarrow A = N \log\left(\frac{\omega_b}{\omega_a}\right)$$

↑

"A" can really be interpreted as a normalization constant

$$\frac{U}{N} = \frac{kT}{\log\left(\frac{\omega_b}{\omega_a}\right)} \log\left(\frac{\sinh\left(\frac{k\omega_b}{2kT}\right)}{\sinh\left(\frac{k\omega_a}{kT}\right)}\right)$$

specific heat:
per particle

$$\frac{d\left(\frac{U}{N}\right)}{dT} \stackrel{\text{chain rule}}{=} \frac{k}{\log\left(\frac{\omega_b}{\omega_a}\right)} \log\left(\frac{\sinh\left(\frac{k\omega_b}{2kT}\right)}{\sinh\left(\frac{k\omega_a}{kT}\right)}\right)$$

$$+ \frac{k\omega_a}{2T} \frac{\coth\left(\frac{k\omega_a}{2kT}\right)}{\log\left(\frac{\omega_b}{\omega_a}\right)}$$

$$- \frac{k\omega_b}{2T} \frac{\coth\left(\frac{k\omega_b}{2kT}\right)}{\log(\omega_b/\omega_a)}$$

$$\frac{d}{dT} \left[\log\left(\sinh\left(\frac{x}{T}\right)\right) \right]$$

$$= - \frac{x \coth\left(\frac{x}{T}\right)}{T^2}$$

(b)

for large T : $2kT \gg \hbar\omega_a$ and $2kT \gg \hbar\omega_b$

Note: the low T regime would imply

$2kT \ll \hbar\omega_a$ and $2kT \ll \hbar\omega_b$ (of course, the factor of 2 can be dropped).

Back to large T : $\sinh x \approx x$ for small $\frac{\hbar\omega}{2kT} = x$

$\coth x \approx x^{-1}$ for small x

$$\Rightarrow \frac{d(\frac{E}{N})}{dT} \longrightarrow \frac{2}{\log(\frac{\omega_b}{\omega_a})} \log\left(\frac{\omega_b}{\omega_a}\right) = k$$

↑
specific heat
per particle

note: the second and third terms cancel to a good approximation

So: specific heat per particle $\rightarrow k$ in

high T limit \rightarrow the frequency distribution becomes irrelevant

(c) ^{internal} energy per particle for single-frequency case:

$$U_{\omega} = \frac{\hbar\omega}{2} \coth\left(\frac{\hbar\omega}{2kT}\right)$$

$$\Rightarrow \frac{dU_{\omega}}{dT} = \frac{\hbar\omega}{2} \frac{\hbar\omega}{2k} \frac{1}{T^2} \frac{1}{\sinh^2\left(\frac{\hbar\omega}{2kT}\right)}$$

$$= \frac{(\hbar\omega)^2}{4(kT)^2} k \frac{1}{\sinh^2\left(\frac{\hbar\omega}{2kT}\right)}$$

$$\frac{d}{dT} \coth\left(\frac{x}{T}\right)$$

$$= \frac{x}{T^2 (\sinh(\frac{x}{T}))^2}$$

in the notebook : $\hbar\omega \rightarrow \hbar\omega$

$\hbar\omega_a \rightarrow \hbar\omega_a$

$\hbar\omega_b \rightarrow \hbar\omega_b$

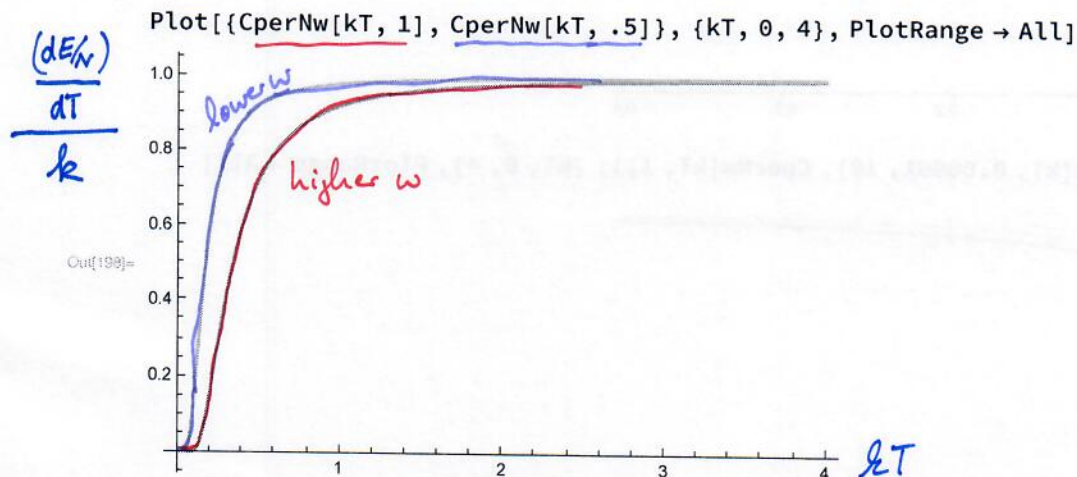
$kT \rightarrow kT$

$$\text{In}[190]:= \text{CperNw}[kT_, hw_] := \frac{\frac{hw^2}{4 kT^2}}{\text{Sinh}\left[\frac{hw}{2 kT}\right]^2}$$

Out[190]=

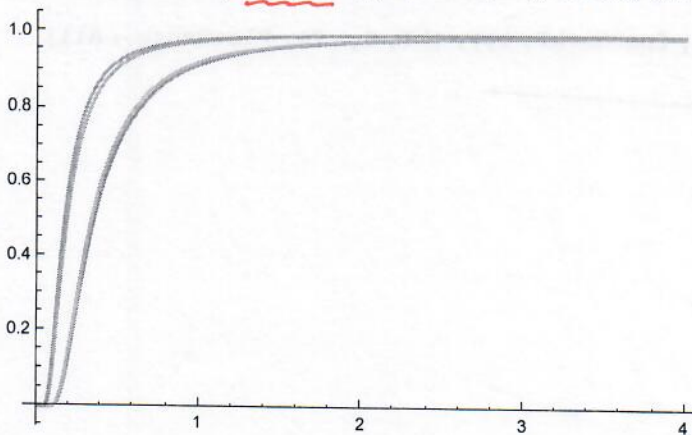
$$\text{CperNwave}[kT_, hwa_, hwb_] := \frac{1}{\text{Log}\left[\frac{hwb}{hwa}\right]} \text{Log}\left[\frac{\text{Sinh}\left[\frac{hwb}{2 kT}\right]}{\text{Sinh}\left[\frac{hwa}{2 kT}\right]}\right] + \frac{\frac{hwa}{2 kT} \text{Coth}\left[\frac{hwa}{2 kT}\right]}{\text{Log}\left[\frac{hwb}{hwa}\right]} - \frac{\frac{hwb}{2 kT} \text{Coth}\left[\frac{hwb}{2 kT}\right]}{\text{Log}\left[\frac{hwb}{hwa}\right]}$$

Out[195]=



Out[209]=

Plot[{CperNwave[kT, 1., 1.1], CperNw[kT, 1],
CperNwave[kT, 0.5, 0.6], CperNw[kT, 0.5]}, {kT, 0, 4}, PlotRange -> All]

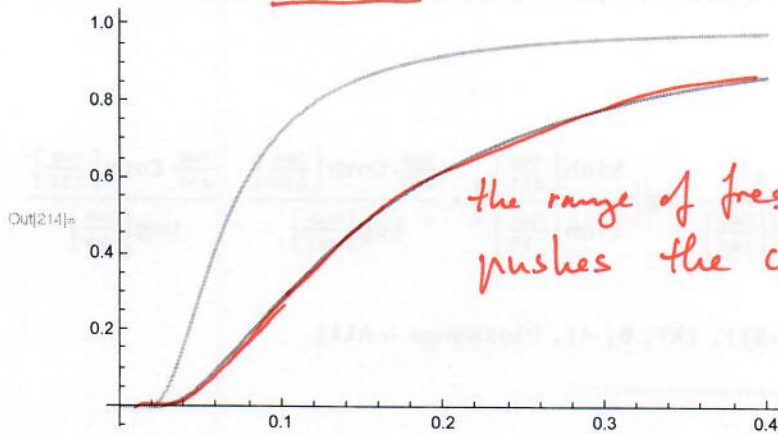


super small range
for w_a, w_b

$\rightarrow w_a \approx w_b$

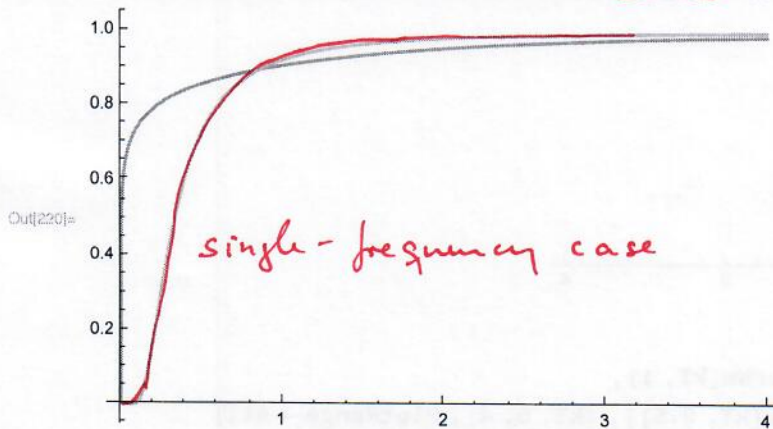
agreement with
single frequency result
(almost...)

In[214] = Plot[{CperNwave[kT, 0.2, 1], CperNw[kT, .2]}, {kT, 0, .4}, PlotRange -> All]



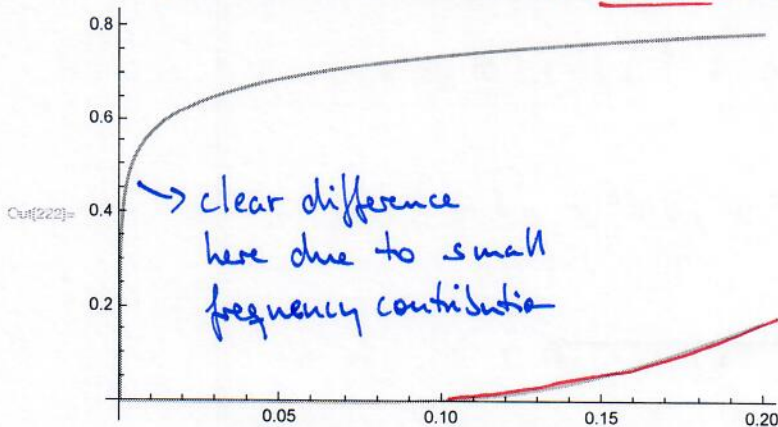
the range of frequencies $> \omega$
pushes the curve to the right

In[220] = Plot[{CperNwave[kT, 0.00001, 10], CperNw[kT, 1]}, {kT, 0, 4}, PlotRange -> All]



single-frequency case

In[222] = Plot[{CperNwave[kT, 0.00001, 10], CperNw[kT, 1]}, {kT, 0, .2}, PlotRange -> All]



→ clear difference
here due to small
frequency contribution

single-frequency case

same plot
but
blown
up

Problem 2:

(a) $3 |\vec{p}|^4 \rightarrow \text{energy}$

↓

$p^2 \rightarrow \text{mass} \cdot \text{energy}$

$\Rightarrow p^4 \rightarrow \text{mass}^2 \text{ energy}^2$

$\Rightarrow B = (\text{mass}^2 \text{ energy})^{-1} = \left(\text{mass}^2 \text{ mass} \frac{\text{meter}^2}{\text{time}^2} \right)^{-1}$

(b)

two dimensions $dx dy \rightarrow 2\pi \hbar dk$

$N = 2 \sum_{\vec{k}} \langle n_{\vec{k}} \rangle_{T=0}$

↑

from spin degree of freedom

assume system is in "box" of area L^2
square

$N = 2 \frac{1}{(2\pi)^2} \iint \langle n_{\vec{k}} \rangle_{T=0} d^2 k$

$= 2 \frac{L^2}{(2\pi)^2} 2\pi \int_0^{k_F} \langle n_{\vec{k}} \rangle_{T=0} \hbar dk$

now: $\langle n_{\vec{k}} \rangle$ at $T=0$ is either 1 or 0

$$\text{for } k < k_F \rightarrow \langle n_k \rangle = 1$$

$$\Rightarrow N = \frac{L^2}{2\pi} \underbrace{\int_0^{k_F} k \, dk}_{\frac{1}{2} k_F^2}$$

$$\Rightarrow N = \frac{L^2}{2\pi} k_F^2$$

in 2D, the particle density is $n = \frac{N}{L^2}$

$$\Rightarrow k_F^2 = \frac{2\pi N}{L^2} = 2\pi n$$

We want the Fermi energy $\leadsto \hbar k_F = p_F$

$$\text{and } E = B|\vec{p}|^4$$

$$\Rightarrow E_F = B p_F^4 = B (\hbar k_F)^4$$

$$\text{So } \boxed{E_F = B \hbar^4 (2\pi n)^2}$$

(c)

$$\begin{array}{c} \swarrow \\ \left(\frac{1}{\text{mass}^2} \frac{1}{\text{energy}} \right) \quad \left(\text{energy}^4 \text{ time}^4 \right) \quad \frac{1}{\text{length}^4} \end{array}$$

energy³

$$\frac{\text{time}^4}{\text{mass}^2 \text{ length}^4}$$

$$\frac{1}{\text{energy}^2}$$

$$E = \frac{1}{2} m v^2$$

$$\Rightarrow \text{mass} \frac{\text{length}^2}{\text{time}^2}$$

→ energy

so, units work out!

Preliminary remarks for Problem 3:

Energy between ^{spin} magnetic moment $\vec{\mu}_s$ and uniform magnetic field \vec{B} :

$$\hat{V} = - \hat{\vec{\mu}}_s \cdot \vec{B}$$

$$\Rightarrow \hat{H} = - \hat{\vec{\mu}}_s \cdot \vec{B}$$

For the electron: $\hat{\vec{\mu}}_s = - \frac{2\mu_B}{\hbar} \hat{\vec{S}}$ and $\hat{\vec{S}} = \frac{\hbar}{2} \hat{\vec{\sigma}}$

Spin and spin magnetic moment of the electron are anti-parallel!

more accurately:

$$\hat{\vec{\mu}}_s = - g_s \frac{\mu_B}{\hbar} \hat{\vec{S}}$$

$$g_s = 2.0023 \dots$$

small deviation from 2 due to quantum electrodynamics

$$\begin{aligned} \Rightarrow \hat{H} &= - \hat{\vec{\mu}}_s \cdot \vec{B} = - \left(- \frac{2\mu_B}{\hbar} \right) \hat{\vec{S}} \cdot \vec{B} \\ &= \frac{2\mu_B}{\hbar} \frac{\hbar}{2} \hat{\vec{\sigma}} \cdot \vec{B} = \mu_B \hat{\vec{\sigma}} \cdot \vec{B} \end{aligned}$$

Let $\vec{B} = B_z \hat{e}_z$

$$= \mu_B B_z \hat{\sigma}_z$$

We are now ready to tackle the problem.

We want: $\hat{\rho} = \frac{e^{-\beta \hat{H}}}{\text{Tr}(e^{-\beta \hat{H}})}$

in matrix form:

$$\hat{\rho} = \begin{pmatrix} \langle \uparrow | \hat{\rho} | \uparrow \rangle & \langle \uparrow | \hat{\rho} | \downarrow \rangle \\ \langle \downarrow | \hat{\rho} | \uparrow \rangle & \langle \downarrow | \hat{\rho} | \downarrow \rangle \end{pmatrix}$$

↑
matrix

Let's start w/ $e^{-\beta \hat{H}}$:

$$e^{-\beta \hat{H}} = e^{-\beta \hat{H}} \left(\underbrace{|\uparrow\rangle\langle\uparrow| + |\downarrow\rangle\langle\downarrow|}_{\text{identity}} \right)$$
$$= e^{-\beta \hat{H}} |\uparrow\rangle\langle\uparrow| + e^{-\beta \hat{H}} |\downarrow\rangle\langle\downarrow|$$

look at $e^{-\beta \hat{H}} |\uparrow\rangle = e^{+\beta \mu_B B_z \hat{G}_z} |\uparrow\rangle$

→ $= \sum_{n=0}^{\infty} \frac{1}{n!} (\beta \mu_B B_z \hat{G}_z)^n |\uparrow\rangle$

Taylor expansion

$$= \sum_{n=0}^{\infty} \frac{1}{n!} (\beta \mu_B B_z)^n |\uparrow\rangle$$

$$= e^{\beta \mu_B B_z} |\uparrow\rangle$$

Similarly:

$$e^{-\beta \hat{x}} |\downarrow\rangle = e^{\beta \mu_B B_z \hat{\sigma}_z} |\downarrow\rangle$$

$$= \sum_{n=0}^{\infty} \frac{1}{n!} (\beta \mu_B B_z \hat{\sigma}_z)^n |\downarrow\rangle$$

$$\begin{aligned} \hat{\sigma}_z^n |\downarrow\rangle &= (-1)^n |\downarrow\rangle \\ &= e^{-\beta \mu_B B_z} |\downarrow\rangle \end{aligned}$$

$$\Rightarrow e^{-\beta \hat{x}} = e^{\beta \mu_B B_z} |\uparrow\rangle\langle\uparrow| + e^{-\beta \mu_B B_z} |\downarrow\rangle\langle\downarrow|$$

In matrix form:

$$\begin{pmatrix} \langle\uparrow| e^{-\beta \hat{x}} |\uparrow\rangle & \langle\uparrow| e^{-\beta \hat{x}} |\downarrow\rangle \\ \langle\downarrow| e^{-\beta \hat{x}} |\uparrow\rangle & \langle\downarrow| e^{-\beta \hat{x}} |\downarrow\rangle \end{pmatrix} = \begin{pmatrix} e^{\beta \mu_B B_z} & 0 \\ 0 & e^{-\beta \mu_B B_z} \end{pmatrix}$$

What about $\text{Tr } e^{-\beta \hat{x}}$?

$$\text{Tr } e^{-\beta \hat{x}} = e^{\beta \mu_B B_z} + e^{-\beta \mu_B B_z}$$

Summing up the diagonal elements

$$\Rightarrow \hat{\rho} = \frac{1}{e^{\beta \mu_B B_z} + e^{-\beta \mu_B B_z}} \begin{pmatrix} e^{\beta \mu_B B_z} & 0 \\ 0 & e^{-\beta \mu_B B_z} \end{pmatrix}$$

We see: $\text{Tr}(\hat{\rho}) = 1.$

Thermal expectation value of $\langle \hat{G}_z \rangle$:

$$\langle \hat{G}_z \rangle = \frac{\text{Tr}(\hat{\rho} \hat{G}_z)}{\text{Tr}(\hat{\rho})} = \text{Tr}(\hat{\rho} \hat{G}_z)$$

$$= \text{Tr} \left(\begin{pmatrix} e^{\beta \mu_B B_z} & 0 \\ 0 & e^{-\beta \mu_B B_z} \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \right) \cdot \frac{1}{e^{\beta \mu_B B_z} + e^{-\beta \mu_B B_z}}$$

$$= \frac{e^{\beta \mu_B B_z} - e^{-\beta \mu_B B_z}}{e^{\beta \mu_B B_z} + e^{-\beta \mu_B B_z}}$$

$$= \tanh(\beta \mu_B B_z)$$

Problem 4:

(a) Free particle in a box w/ periodic BCs.

$$E = \frac{\hbar^2 k^2}{2m}$$

$$\phi_{\vec{k}}(\vec{r}) = \frac{1}{L^{3/2}} e^{i\vec{k} \cdot \vec{r}} \quad (\text{"box normalization"})$$

$$\vec{k} = \frac{2\pi}{L} (n_x, n_y, n_z) \quad \text{where } n_x = 0, \pm 1, \pm 2, \dots$$

$$n_y = \dots$$

$$n_z = \dots$$

So far, we have just collected information from previous homework.

We want the density matrix $\hat{\rho}$ in the canonical ensemble: $\hat{\rho} = \frac{e^{-\beta \hat{H}}}{\text{Tr}(e^{-\beta \hat{H}})}$

We need to find an expression for $e^{-\beta \hat{H}}$.

As before, insert a complete set.

We have:

$$\sum_{\vec{k}} |\phi_{\vec{k}}\rangle \langle \phi_{\vec{k}}| = 1$$

$$\Rightarrow e^{-\beta \hat{H}} = \sum_{\vec{k}} e^{-\beta \hat{H}} \sum_{\vec{k}} |\phi_{\vec{k}}\rangle \langle \phi_{\vec{k}}|$$

$$= \sum_{\vec{k}} e^{-\beta \hat{H}} |\phi_{\vec{k}}\rangle \langle \phi_{\vec{k}}|$$

to evaluate this,
we need to perform
a Taylor expansion,
then act w/ \hat{H}^n onto
 $|\phi_{\vec{k}}\rangle$, and then collect
terms

$$\text{This yields } e^{-\beta \hat{H}} |\phi_{\vec{k}}\rangle$$

$$= e^{-\beta \frac{\hbar^2 k^2}{2m}} |\phi_{\vec{k}}\rangle$$

$$= \sum_{\vec{k}} e^{-\beta \frac{\hbar^2 k^2}{2m}} |\phi_{\vec{k}}\rangle \langle \phi_{\vec{k}}|$$

$$\Rightarrow e^{-\beta \hat{H}} = \sum_{\vec{k}} |\phi_{\vec{k}}\rangle \exp\left(-\beta \frac{\hbar^2 k^2}{2m}\right) \langle \phi_{\vec{k}}|$$

$$\Rightarrow \langle \vec{r} | e^{-\beta \hat{H}} | \vec{r}' \rangle = \frac{1}{L^3} \sum_{\vec{k}} e^{+i\vec{k} \cdot \vec{r}} e^{-\beta \frac{\hbar^2 k^2}{2m}} e^{-i\vec{k} \cdot \vec{r}'}$$

$$\text{where we used } \langle \vec{r} | \phi_{\vec{k}} \rangle = \frac{1}{L^3} e^{i\vec{k} \cdot \vec{r}} \quad \&$$

$$\langle \phi_{\vec{h}} | \vec{r}' \rangle = \frac{1}{L^{3/2}} e^{-i\vec{h} \cdot \vec{r}'}$$

$\frac{1}{\text{Tr}(e^{-\beta \hat{H}})} \langle \vec{r} | e^{-\beta \hat{H}} | \vec{r}' \rangle$ is referred to as density matrix in the coordinate representation.

Now, we want to simplify this \rightarrow the infinite sum is very inconvenient.

Let's rewrite the sum as an integral over $dk_x dk_y dk_z$

$$k_x = \frac{2\pi}{L} n_x \quad \text{where } n_x = 0, \pm 1, \pm 2, \dots$$

$$\left\{ \begin{array}{l} \Delta n_x = 1 \end{array} \right.$$

$$\Rightarrow dk_x \hat{=} \frac{2\pi}{L} \underbrace{\Delta n_x}_{=1}$$

$$\Rightarrow \frac{L dk_x}{2\pi} \hat{=} 1$$

$$\text{and } \frac{L^3 dk_x dk_y dk_z}{(2\pi)^3} \hat{=} 1$$

So, we can rewrite our infinite sum as an integral:

$$\langle \vec{r} | e^{-\beta \hat{H}} | \vec{r}' \rangle = \frac{1}{(2\pi)^3} \iiint \exp\left(-\frac{\beta \hbar^2 k^2}{2m} + i\vec{k} \cdot (\vec{r} - \vec{r}')\right) d^3k$$

$$= \frac{1}{(2\pi)^3} \iiint e^{-\frac{\beta \hbar^2 k^2}{2m}} \cos(\vec{k} \cdot (\vec{r} - \vec{r}')) d^3k$$

$$+ \underbrace{\frac{i}{(2\pi)^3} \iiint e^{-\frac{\beta \hbar^2 k^2}{2m}} \sin(\vec{k} \cdot (\vec{r} - \vec{r}')) d^3k}_{=0}$$

$$= \frac{1}{(2\pi)^3} \left(\sqrt{\frac{\pi}{\frac{\beta \hbar^2}{2m}}} \right)^3 \exp\left(-\frac{|\vec{r} - \vec{r}'|}{4 \frac{\beta \hbar^2}{2m}}\right)$$

$$= \left(\frac{m}{2\pi \beta \hbar^2} \right)^{3/2} \exp\left(-\frac{m|\vec{r} - \vec{r}'|}{2\beta \hbar^2}\right)$$

Next, we want to calculate $\text{Tr}(e^{-\beta \hat{H}})$.

$$\text{Tr}(e^{-\beta \hat{x}}) = \iiint \langle \vec{r} | e^{-\beta \hat{x}} | \vec{r} \rangle d^3 r$$

$$\begin{aligned} & \xrightarrow{\text{using } e^{-\beta \hat{x}}} = \iiint \sum_{\vec{k}} \langle \vec{r} | \phi_{\vec{k}} \rangle e^{-\beta \frac{\hbar^2 k^2}{2m}} \langle \phi_{\vec{k}} | \vec{r} \rangle d^3 r \\ & = \sum_{\vec{k}} |\phi_{\vec{k}} \rangle e^{-\beta \frac{\hbar^2 k^2}{2m}} \langle \phi_{\vec{k}} | \end{aligned}$$

$$\begin{aligned} & \xrightarrow{\text{using our previous results (for sum over } \vec{k})} = \underbrace{\iiint \exp\left(-\frac{m|\vec{r}-\vec{r}'|}{2\beta \hbar^2}\right)}_1 \underbrace{\left(\frac{m}{2\pi\beta \hbar^2}\right)^{3/2}}_{\text{indep. of } r} d^3 r \end{aligned}$$

$$\boxed{\text{Tr}(e^{-\beta \hat{x}}) = \frac{V}{L^3} \left(\frac{m}{2\pi\beta \hbar^2}\right)^{3/2}}$$

$$\Rightarrow \langle \vec{r} | \hat{\rho} | \vec{r}' \rangle = \langle \vec{r} | \frac{e^{-\beta \hat{x}}}{\text{Tr}(e^{-\beta \hat{x}})} | \vec{r}' \rangle$$

$$= \frac{\langle \vec{r} | e^{-\beta \hat{x}} | \vec{r}' \rangle}{\text{Tr}(e^{-\beta \hat{x}})}$$

$$= \frac{1}{V} \exp\left(-\frac{m|\vec{r}-\vec{r}'|}{2\beta \hbar^2}\right)$$

Finally:

$$\langle \vec{r} | \hat{\rho} | \vec{r}' \rangle = \frac{1}{V} \exp\left(-\frac{m|\vec{r}-\vec{r}'|^2}{2\beta\hbar^2}\right)$$

(b)

$$\langle \vec{r} | \hat{\rho} | \vec{r} \rangle = \frac{1}{V}$$

What does this mean?

↳ How can we interpret $\langle \vec{r} | \hat{\rho} | \vec{r} \rangle$?

$$\hat{\rho} = \frac{e^{-\beta\hat{x}}}{\text{Tr}(e^{-\beta\hat{x}})}$$

our
earlier expression

Look at $\langle \delta(\vec{r}'' - \vec{r}) \rangle$:

average of finding \vec{r} at position $\vec{r}'' \rightarrow$ it's the operator that tells us where the particle is

$$\langle \delta(\vec{r}'' - \vec{r}) \rangle$$

$$= \text{Tr}(\hat{\rho} \delta(\vec{r}'' - \vec{r})) = \frac{\int \langle \vec{r} | e^{-\beta\hat{x}} \delta(\vec{r}'' - \vec{r}) | \vec{r} \rangle d^3r}{\text{Tr}(e^{-\beta\hat{x}})}$$

$$= \frac{\langle \vec{r}'' | e^{-\beta\hat{x}} | \vec{r}'' \rangle}{\text{Tr}(e^{-\beta\hat{x}})} = \langle \vec{r}'' | \hat{\rho} | \vec{r}'' \rangle$$

So: $\langle \vec{r} | \hat{\rho} | \vec{r} \rangle$ just gives us the density.

The density is independent of \vec{r} !
(this is the case because we have a free particle)

(c) Want to calculate $\langle \hat{\mathcal{H}} \rangle = \text{Tr}(\hat{\rho} \hat{\mathcal{H}})$

$$\text{Tr}(\hat{\rho} \hat{\mathcal{H}}) = \frac{\text{Tr}(e^{-\beta \hat{\mathcal{H}}} \hat{\mathcal{H}})}{\text{Tr}(e^{-\beta \hat{\mathcal{H}}})}$$

$$= - \frac{\partial}{\partial \beta} \log \text{Tr}(e^{-\beta \hat{\mathcal{H}}})$$

$$\nearrow = - \frac{\partial}{\partial \beta} \log \left(V \left(\frac{m}{2\pi\beta\hbar^2} \right)^{3/2} \right)$$

using our
earlier
result

$$= \frac{3}{2} \frac{1}{\beta} = \frac{3}{2} kT$$

$$\text{So: } \boxed{\langle \hat{\mathcal{H}} \rangle = \frac{3}{2} kT}$$