







PhD draft test title

To obtain the academic degree of

Doctor of Science

from the Faculty of Physics of the Karlsruhe Institute for Technology (KIT)

and

from the Institute of Technology "Prof. Jorge A. Sábato" of the Universidad Nacional de General San Martín (UNSAM)

submitted

Dissertation

of

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born in Breisach am Rhein

Day of the oral exam: 01.01.1970

Referent: Prof. Dr. Ralph Engel

Co-referent: Prof. Dr. Missi Ngno

Supervisor(s): Dr. David Schmidt

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Abstract

Resumen

Zusammenfassung

Dies hier ist ein Blindtext zum Testen von Textausgaben. Wer diesen Text liest, ist selbst schuld. Der Text gibt lediglich den Grauwert der Schrift an. Ist das wirklich so? Ist es gleichgültig, ob ich schreibe: "Dies ist ein Blindtext" oder "Huardest gefburn"? Kjift – mitnichten! Ein Blindtext bietet mir wichtige Informationen. An ihm messe ich die Lesbarkeit einer Schrift, ihre Anmutung, wie harmonisch die Figuren zueinander stehen und prüfe, wie breit oder schmal sie läuft. Ein Blindtext sollte möglichst viele verschiedene Buchstaben enthalten und in der Originalsprache gesetzt sein. Er muß keinen Sinn ergeben, sollte aber lesbar sein. Fremdsprachige Texte wie "Lorem ipsum" dienen nicht dem eigentlichen Zweck, da sie eine falsche Anmutung vermitteln.

Contents

english abstract	
resumen español	
deutsche Zusammenfassung	. V
Main Content	
1 Introduction	3
2 The Pierre Auger Observatory	5
2.1 One	
2.1.1 One One	. 6
2.1.2 One Two	6
2.2 Two	. 6
2.2.1 Two One	. 7
2.2.2 Two Two	. 7
2.3 Three	. 7
2.3.1 Three One	. 7
2.3.2 Three Two	. 8
2.3.3 Three Three	. 8
3 Example shit	9
3.1 Section	
3.1.1 Subsection	10
3.2 References and labels	
3.2.1 Labels	
3.2.2 Referencing labels	
3.2.3 Citations	
3.3 Acronoyms	
3.4 Units	
3.5 Figures	
3.6 Tables	
3.7 Mathematical and decay equations	15
3.8 Reminders	
3.9 Miscellaneous	
Supplementary Information	
D Hallo	2 3

Todo list

use 'submitted' for hand in, use 'accepted' for final version	i
fix page vspace formatting once you have the final title	ii
remove todos	1
add pdf hyperref keywords	1
more recent spectrum? proper citation to whom?	16
Figure:	16
remove todos	
add pdf hyperref keywords	





<u>2.1</u>	<u>One</u>		5
	2.1.1	One One	6
		One One One	6
		One One Two	6
	2.1.2	One Two	6
2.2	Two		6
	2.2.1	Two One	
	2.2.2	Two Two	. 7
2.3	Three		7
	2.3.1	Three One	. 7
	2.3.2	Three Two	. 8
	2.3.3	Three Three	8

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

2.1 One

2.1.1 One One

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One One One

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One One Two

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

2.1.2 One Two

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

2.2 Two

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2.3. THREE 7

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2.2.1 Two One

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2.2.2 Two Two

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2.3 Three

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2.3.1 Three One

2.3.2 Three Two

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

2.3.3 Three Three



3.1	Sectio	n	<u> 10</u>
	3.1.1	Subsection	10
		Subsubsection	10
3.2	Refere	ences and labels	10
	3.2.1	Labels	10
	3.2.2	Referencing labels	10
	3.2.3	Citations	11
3.3	Acron	oyms	11
3.4	Units		12
3.5	Figure	es	12
3.6	Tables	5	15
3.7	Mathe	ematical and decay equations	15
3.8	Remin	nders	16
3.9	Misce	llaneous	16

```
\chapter{Chapter: Types of section distinctions}
\blindtext
\section{Section}
\blindtext
\subsection{Subsection}
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\blindtext

\subsubsection{Subsubsection}

\blindtext

3.1 Section

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

3.1.1 Subsection

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

Subsubsection

Hello, here is some text without a meaning. This text should show what a printed text will look like at this place. If you read this text, you will get no information. Really? Is there no information? Is there a difference between this text and some nonsense like "Huardest gefburn"? Kjift – not at all! A blind text like this gives you information about the selected font, how the letters are written and an impression of the look. This text should contain all letters of the alphabet and it should be written in of the original language. There is no need for special content, but the length of words should match the language.

3.2 References and labels

3.2.1 Labels

Make sure to label sections to refer back to throughout your work. Make it an intuitive name.

```
\label{sec:cosmicrays}
\label{subsec:crs_eas}
```

3.2.2 Referencing labels

Use cref instead of ref; it smartly labels if it's a section, chapter, figure, table, etc. Fig. 3.1.

```
\cref{fig:crs_eas_heitler1}
\cref{subsec:fd}
```

3.3. ACRONOYMS

3.2.3 Citations

If you have a ton of reference, just list them with commas. Latex will properly format [1–41]:

```
\cite{Abraham2009b,Abraham2009,Abraham2010a,Abraham2010,Abreu:2011zze,Abreu

→ :2011zzd,Abreu:2011vm,Abreu:2011ki,Abreu:2011fb,Settimo:2012zz,Auger

→ :2012yc,Auger:2012an,Acounis:2012dg,Abreu:2012zz,Abreu:2012zg,Abreu

→ :2012ybu,Abreu:2012pi,Abreu:2012oza,Abreu:2012aniso,Abreu:2011md,

→ Abreu:2013zbq,Abreu:2013qtw,Abreu:2013qfa,Abreu:2013kif,Abreu:2013env,

→ Aab:2014qva,Aab:2014pza,Aab:2014kda,Aab:2014ila,Aab:2014gua,Aab:2014

→ esa,Aab:2014dua,Aab:2014dha,Aab:2014caa,Aab:2014bha,Aab:2014aea,

→ ThePierreAuger:2014nja,PierreAuger:2014yba,Aab2015a,Aab2015,Aab:2015

→ kma}
```

3.3 Acronoyms

There's a neat package called acronyms that will handle them for you. Alex had already set this in the acronym.tex. Just define your acronyms here as he did.

• To use the acronym like **CR!**¹, use:

```
\ac{CR}
```

• To define the acronym in the text—*CMB!* (CMB!), use:

```
\acfi{CMB}
```

If you want this to be the only place in your chapter where the acronym is defined, you need to write:

```
\acfi{CMB}\acused{CMB}
```

as the acronym package does not automatically count this as a definition.

• To makes the acronym plural. CAVEAT is that acronyms ending in an S will add an extra S which is not typically used in English.

```
\acp{CR}
```

 Sometimes acronyms require more complicated definitions, you can define them in the main document and call them throughout. Alex has already defined QGSJET-II.03 and Offline:

```
\qgsjet
\Offline
```

¹CR!

3.4 Units

For defining units, use the SI package, as it will consistently format for you. It sometimes may not recognize something like Mpc.

Examples:

- $10^{20} \, eV$
- 12 km² for multiple units
- 90 % for precentages
- $\approx 5 \times 10^{19} \, \text{eV}$
- 37 g cm⁻² for grammage
- 30 GeV if GeV is not recognized, specify by metric prefix
- $3 \times 10^{15} \, eV$
- 30 % to 60 % a way to consistently format ranges
- km² sr yr

```
\SI{e20}{\eV}
\SI{12}{\square\km}
\SI{90}{\percent}
$\approx \SI{5e19}{\eV}$
\SI{37}{\grammage}
\SI{30}{\giga\eV}
\SI{3e15}{\eV} 3{\times}10^{19}
\SIrange{30}{60}{\percent}
$\square\km \steradian \year}$
```

3.5 Figures

In a PhD thesis you should always use only [t] (top) figure placement. Also note that due to the \graphicspath{{figures/}} command in the preamble, the file paths are relative to the ./figures directory which can thus be dropped from the line. If you also ommit the filename extension (e.g. .pdf or .jpg) your source file will be compilable with both, latex and pdflatex.

```
\begin{figure}[h]
  \centering
  \includegraphics[width=0.8\textwidth]{intro/heitler}
  \caption{Illustration of an \ac{EAS}' particle components.}
  \label{fig:crs_eas_heitler1}
  \end{figure}
```

3.5. FIGURES 13

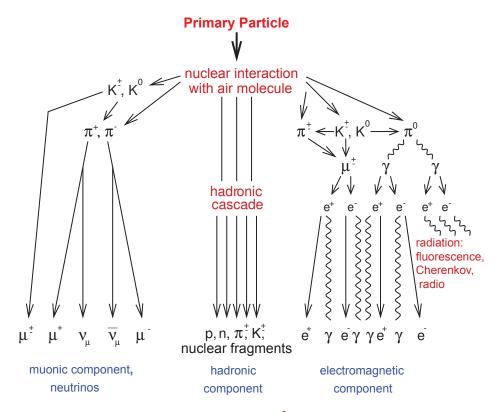


Figure 3.1: Illustration of an **EAS!**²′ particle components.

Use subref to reference elements within a figure for your text or captions.

If you need a footnote in a figure, you have to use footnotemark

```
\begin{figure}[t] \centering
```

³As discussed further in the reconstruction Chapter, quality cuts are performed on reconstructed data from the **SD!**⁴. One of these cuts is known as the 6T5-trigger; it requires that the detector with the highest signal has all of its 6 closest neighbors working at the time of the event. Similarly, a 5T5 only requires 5 of the closest neighbors to be working.

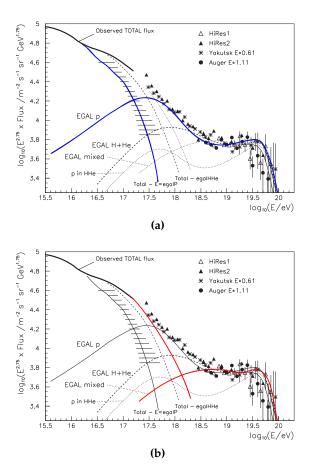


Figure 3.2: Visualization of the (a) pair production dip [42] and (b) mixed composition [43] scenarios that describe the ankle feature.

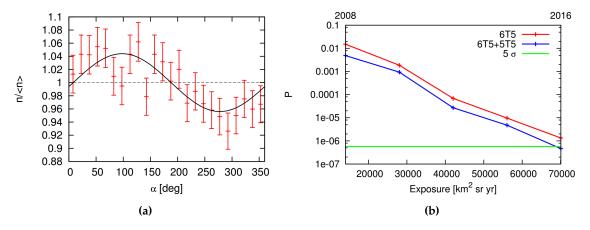


Figure 3.3: (a) (b) Probability for the amplitude of the dipole to arise from an isotropic distribution as a function of the integrated exposure of the Pierre Auger Observatory. Various data sets with different tank triggers are shown ³[44].

```
\subfloat[]{\includegraphics[height=5cm]{intro/auger_dipole}
\label{plot:pao_dipole}}
\subfloat[]{\includegraphics[height=5cm]{intro/auger_dipole_sig}}
\label{plot:pao_dipole_sig}}
```

3.6. TABLES 15

3.6 Tables

Table 3.1: Dipole components and direction in equatorial components [45].

E/EeV	d_{\perp}	d_z	d	α	δ
4-8	-0.024 ± 0.010	0.006 ± 0.006	0.025 ± 0.009	$-75^{\circ} \pm 15^{\circ}$	$82^{\circ}\pm57^{\circ}$
> 8	-0.026 ± 0.015	0.060 ± 0.010	0.065 ± 0.011	$-24^{\circ}\pm12^{\circ}$	$100^{\circ} \pm 10^{\circ}$

3.7 Mathematical and decay equations

For decay equations, use align

$$\gamma_{CMB} + p \rightarrow \Delta^+ \rightarrow p + \pi^0$$
,
$$\gamma_{CMB} + p \rightarrow \Delta^+ \rightarrow n + \pi^+$$
.

For writing 5.5σ , use

```
\sig{5.5}
```

3.8 Reminders

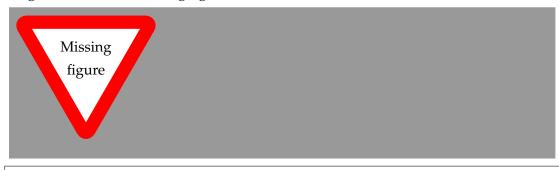
Use to dos so that you don't have to dig through latex code. [inline] makes it so it takes up the line and isn't hanging off the page

• To add a todo inline like this

more recent spectrum? proper citation to whom?

\todo[inline]{more recent spectrum? proper citation to whom?}

• To generate this for missing figures:



\missingfigure{}

To generate a list of all your todos and their page numbers, use

\listoftodos

3.9 Miscellaneous

• 4.6×10^{-7}

 $4.6{\times}10^{-7}$

• For degrees 148.4°

 $\ag\{148.4\}$ or 148.4° circ

• For formatting numbers otherwise in text, -2.0

 $\sum_{-2.0} \text{ or } -2.0$

• Superscripts for text like 20th

20th

• For marking out text —Due to the clean room environment, use:

\deleted{Due to the clean room environment}

This may be useful for editing your thesis later.



- [1] J. Abraham *et al.* (Pierre Auger Collab.), Pierre Auger Collaboration, Limit on the diffuse flux of ultrahigh energy tau neutrinos with the surface detector of the pierre auger observatory, Phys. Rev. D 79 (10) (2009) 102001. doi:10.1103/PhysRevD.79.102001.
- [2] J. Abraham et al. (Pierre Auger Collab.), Upper limit on the cosmic-ray photon fraction at eev energies from the pierre auger observatory, Astroparticle Physics 31 (6) (2009) 399-406. doi:DOI:10.1016/j.astropartphys.2009.04.003.
 URL http://www.sciencedirect.com/science/article/B6TJ1-4W6Y349-1/2/fa58b1b24ace553d6d47e135f06059e1
- [3] J. Abraham et al. (Pierre Auger Collab.), Measurement of the energy spectrum of cosmic rays above 10¹⁸ eV using the Pierre Auger Observatory, Physics Letters B 685 (4-5) (2010) 239 246. doi:DOI:10.1016/j.physletb.2010.02.013.
 URL http://www.sciencedirect.com/science/article/B6TVN-4YC2XC7-3/2/69021ee5b6b08f24e4343776c82fb3ed
- [4] J. Abraham *et al.* (Pierre Auger Collab.), Pierre Auger Collaboration, Measurement of the Depth of Maximum of Extensive Air Showers above 10¹⁸ eV, Phys. Rev. Lett. 104 (9) (2010) 091101. doi:10.1103/PhysRevLett.104.091101.
- [5] P. Abreu, et al., Pierre Auger Collaboration, A Search for Ultra-High Energy Neutrinos in Highly Inclined Events at the Pierre Auger Observatory, Phys.Rev. D84 (2011) 122005. arXiv:1202.1493, doi:10.1103/PhysRevD.85.029902,10.1103/PhysRevD.84.122005.
- [6] P. Abreu, et al., Pierre Auger Collaboration, The Lateral Trigger Probability function for the ultra-high energy cosmic ray showers detected by the Pierre Auger Observatory, Astropart.Phys. 35 (2011) 266–276. arXiv:1111.6645, doi:10.1016/j.astropartphys. 2012.02.005, 10.1016/j.astropartphys.2011.08.001.
- [7] P. Abreu, et al., Pierre Auger Collaboration, Anisotropy and chemical composition of ultra-high energy cosmic rays using arrival directions measured by the Pierre Auger Observatory, JCAP 1106 (2011) 022. arXiv:1106.3048, doi:10.1088/1475-7516/2011/06/022.
- [8] P. Abreu, et al., Pierre Auger Collaboration, The effect of the geomagnetic field on cosmic ray energy estimates and large scale anisotropy searches on data from the

- Pierre Auger Observatory, JCAP 1111 (2011) 022. arXiv:1111.7122, doi:10.1088/1475-7516/2011/11/022.
- [9] P. Abreu, et al., Pierre Auger Collaboration, Advanced functionality for radio analysis in the Offline software framework of the Pierre Auger Observatory, Nucl.Instrum.Meth. A635 (2011) 92–102. arXiv:1101.4473, doi:10.1016/j.nima.2011.01.049.
- [10] P. Abreu, et al., Pierre Auger Collaboration, Measurement of the Cosmic Ray Energy Spectrum Using Hybrid Events of the Pierre Auger Observatory, Eur.Phys.J.Plus 127 (2012) 87. arXiv:1208.6574, doi:10.1140/epjp/i2012-12087-9.
- [11] P. Abreu, et al., Pierre Auger Observatory Collaboration, A Search for Point Sources of EeV Neutrons, Astrophys.J. 760 (2012) 148. arXiv:1211.4901, doi:10.1088/ 0004-637X/760/2/148.
- [12] P. Abreu, et al., Pierre Auger Collaboration, Large scale distribution of arrival directions of cosmic rays detected above 10¹⁸ eV at the Pierre Auger Observatory, Astrophys.J.Suppl. 203 (2012) 34. arXiv:1210.3736, doi:10.1088/0067-0049/203/2/34.
- [13] P. Abreu, et al., Pierre Auger Collaboration, Results of a self-triggered prototype system for radio-detection of extensive air showers at the Pierre Auger Observatory, JINST 7 (2012) P11023. arXiv:1211.0572, doi:10.1088/1748-0221/7/11/P11023.
- [14] P. Abreu, et al., Pierre Auger Collaboration, Search for point-like sources of ultra-high energy neutrinos at the Pierre Auger Observatory and improved limit on the diffuse flux of tau neutrinos, Astrophys.J. 755 (2012) L4. arXiv:1210.3143, doi:10.1088/2041-8205/755/1/L4.
- [15] P. Abreu, et al., Pierre Auger Collaboration, Description of Atmospheric Conditions at the Pierre Auger Observatory using the Global Data Assimilation System (GDAS), Astropart. Phys. 35 (2012) 591–607. arXiv:1201.2276, doi:10.1016/j.astropartphys. 2011.12.002.
- [16] P. Abreu, et al., Pierre Auger Collaboration, Constraints on the origin of cosmic rays above 10¹⁸ eV from large scale anisotropy searches in data of the Pierre Auger Observatory, Astrophys.J. 762 (2012) L13. arXiv:1212.3083, doi:10.1088/2041-8205/762/1/L13.
- [17] P. Abreu, et al., Pierre Auger Collaboration, Antennas for the Detection of Radio Emission Pulses from Cosmic-Ray, JINST 7 (2012) P10011. arXiv:1209.3840, doi: 10.1088/1748-0221/7/10/P10011.
- [18] P. Abreu, et al., Pierre Auger Collaboration, The Rapid Atmospheric Monitoring System of the Pierre Auger Observatory, JINST 7 (2012) P09001. arXiv:1208.1675, doi:10.1088/1748-0221/7/09/P09001.
- [19] P. Abreu, et al., A search for anisotropy in the arrival directions of ultra high energy cosmic rays recorded at the Pierre Auger Observatory, JCAP (2012).

[20] P. Abreu, et al., Pierre Auger Collaboration, Search for signatures of magnetically-induced alignment in the arrival directions measured by the Pierre Auger Observatory, Astropart. Phys. 35 (2012) 354–361. arXiv:1111.2472, doi:10.1016/j.astropartphys. 2011.10.004.

- [21] P. Abreu, et al., Pierre Auger Collaboration, Ultrahigh Energy Neutrinos at the Pierre Auger Observatory, Adv.High Energy Phys. 2013 (2013) 708680. arXiv:1304.1630, doi:10.1155/2013/708680.
- [22] P. Abreu, et al., Pierre Auger Collaboration, Techniques for Measuring Aerosol Attenuation using the Central Laser Facility at the Pierre Auger Observatory, JINST 8 (2013) P04009. arXiv:1303.5576, doi:10.1088/1748-0221/8/04/P04009.
- [23] P. Abreu, et al., Pierre Auger, Identifying Clouds over the Pierre Auger Observatory using Infrared Satellite Data, Astropart.Phys. 50-52 (2013) 92–101. arXiv:1310.1641, doi:10.1016/j.astropartphys.2013.09.004.
- [24] P. Abreu, et al., Pierre Auger Collaboration, Bounds on the density of sources of ultrahigh energy cosmic rays from the Pierre Auger Observatory, JCAP 1305 (05) (2013) 009. arXiv:1305.1576, doi:10.1088/1475-7516/2013/05/009.
- [25] P. Abreu, et al., Pierre Auger Collaboration, Interpretation of the Depths of Maximum of Extensive Air Showers Measured by the Pierre Auger Observatory, JCAP 1302 (2013) 026. arXiv:1301.6637, doi:10.1088/1475-7516/2013/02/026.
- [26] A. Aab, et al., Pierre Auger Collaboration, Origin of atmospheric aerosols at the Pierre Auger Observatory using studies of air mass trajectories in South America, Atmos.Res. (2014)arXiv:1405.7551, doi:10.1016/j.atmosres.2014.05.021.
- [27] A. Aab, et al., Pierre Auger Collaboration, Muons in air showers at the Pierre Auger Observatory: Mean number in highly inclined events, Phys.Rev.D (2014)arXiv:1408.1421.
- [28] A. Aab, et al., Pierre Auger Collaboration, Depth of Maximum of Air-Shower Profiles at the Pierre Auger Observatory: Measurements at Energies above 10^{17.8} eV, Phys.Rev.D (2014)arXiv:1409.4809.
- [29] A. Aab, et al., Telescope Array Collaboration, Pierre Auger Collaboration, Searches for Large-Scale Anisotropy in the Arrival Directions of Cosmic Rays Detected above Energy of 10¹⁹ eV at the Pierre Auger Observatory and the Telescope Array, Astrophys.J. 794 (2) (2014) 172. arXiv:1409.3128, doi:10.1088/0004-637X/794/2/172.
- [30] A. Aab, et al., Pierre Auger Collaboration, Reconstruction of inclined air showers detected with the Pierre Auger Observatory, JCAP 1408 (08) (2014) 019. arXiv:1407.3214, doi:10.1088/1475-7516/2014/08/019.
- [31] A. Aab, et al., Pierre Auger Collaboration, Probing the radio emission from air showers with polarization measurements, Phys.Rev. D89 (2014) 052002. arXiv:1402.3677, doi: 10.1103/PhysRevD.89.052002.

[32] A. Aab, et al., Pierre Auger Collaboration, Muons in air showers at the Pierre Auger Observatory: Measurement of atmospheric production depth, Phys.Rev. D90 (2014) 012012. arXiv:1407.5919, doi:10.1103/PhysRevD.90.012012,10.1103/PhysRevD.90.039904.

- [33] A. Aab, et al., Pierre Auger Collaboration, Search for patterns by combining cosmic-ray energy and arrival directions at the Pierre Auger Observatory, Eur. Phys. J.C (2014)arXiv: 1410.0515.
- [34] A. Aab, et al., Pierre Auger Collaboration, A Targeted Search for Point Sources of EeV Neutrons, Astrophys.J. 789 (2014) L34. arXiv:1406.4038, doi:10.1088/2041-8205/789/2/L34.
- [35] A. Aab, et al., Pierre Auger Collaboration, A search for point sources of EeV photons, Astrophys.J. 789 (2014) 160. arXiv:1406.2912, doi:10.1088/0004-637X/789/2/160.
- [36] A. Aab, et al., Pierre Auger Collaboration, Depth of Maximum of Air-Shower Profiles at the Pierre Auger Observatory: Composition Implications, Phys.Rev.D (2014)arXiv: 1409.5083.
- [37] A. Aab, et al., Pierre Auger, Large Scale Distribution of Ultra High Energy Cosmic Rays Detected at the Pierre Auger Observatory With Zenith Angles up to 80 degrees, Astrophys. J. 802 (2) (2015) 111. arXiv:1411.6953, doi:10.1088/0004-637X/802/2/111.
- [38] A. Aab, et al., Pierre Auger, Searches for Anisotropies in the Arrival Directions of the Highest Energy Cosmic Rays Detected by the Pierre Auger Observatory, Astrophys. J. 804 (1) (2015) 15. arXiv:1411.6111, doi:10.1088/0004-637X/804/1/15.
- [39] A. Aab, et al., Pierre Auger, Muons in air showers at the Pierre Auger Observatory: Mean number in highly inclined events, Phys. Rev. D91 (3) (2015) 032003, [Erratum: Phys. Rev.D91,no.5,059901(2015)]. arXiv:1408.1421, doi:10.1103/PhysRevD.91.059901,10.1103/PhysRevD.91.032003.
- [40] A. Aab, et al., Pierre Auger, Search for patterns by combining cosmic-ray energy and arrival directions at the Pierre Auger Observatory, Eur. Phys. J. C75 (6) (2015) 269. arXiv:1410.0515, doi:10.1140/epjc/s10052-015-3471-0.
- [41] A. Aab, et al., Pierre Auger, Improved limit to the diffuse flux of ultrahigh energy neutrinos from the Pierre Auger Observatory, Phys. Rev. D91 (9) (2015) 092008. arXiv: 1504.05397, doi:10.1103/PhysRevD.91.092008.
- [42] V. Berezinsky, Transition from galactic to extragalactic cosmic rays, arXiv:0710.2750v2 (10 2007).
 URL http://arxiv.org/abs/0710.2750v2
- [43] A. M. Hillas, Cosmic Rays: Recent Progress and some Current Questions, arXiv:astro-ph/0607109 (2006). arXiv:arXiv:astro-ph/0607109. URL http://arxiv.org/abs/astro-ph/0607109v2

- [44] S. Mollerach, E. Roulet, O. Taborda, Evidence for a dipole above 8 EeV (June 2016).
- [45] S. Mollerach, E. Roulet, O. Taborda, Further evidence for a dipole above 8 EeV (Oct. 2016).

