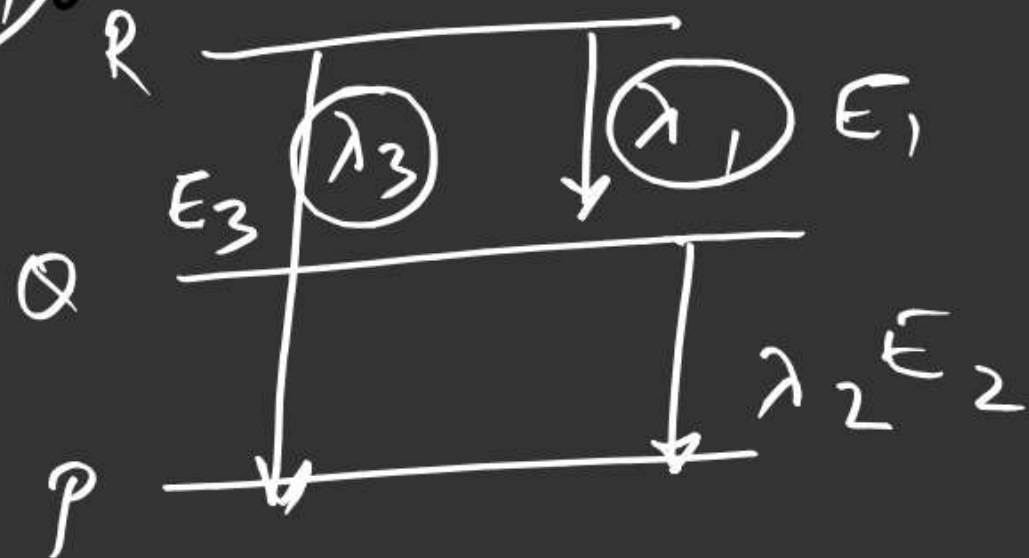


(14)

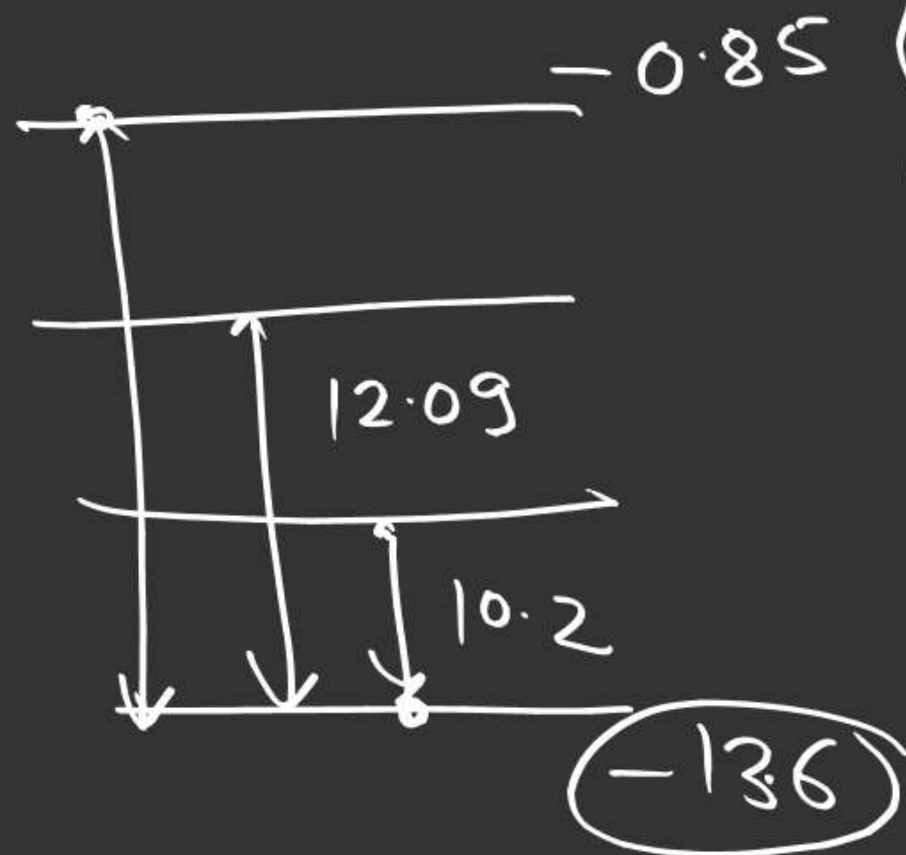


$$E_3 = E_1 + E_2$$

$$\frac{hc}{\lambda_3} = \frac{hc}{\lambda_1} + \frac{hc}{\lambda_2}$$

(17)

(12.75)

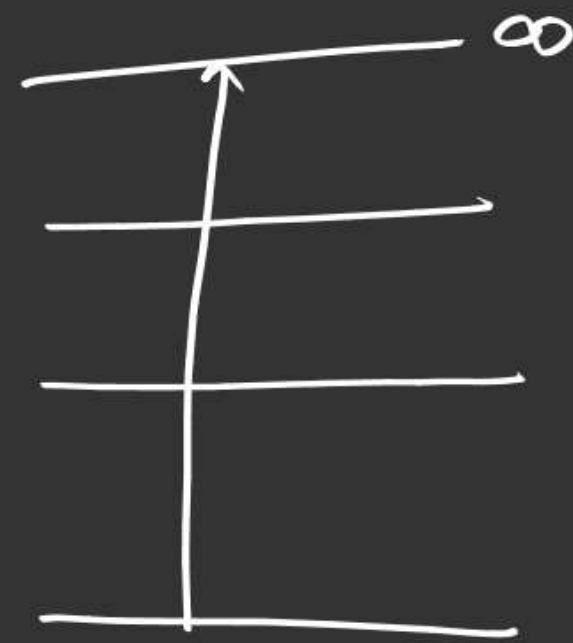


$$12.75 = 13.6 \left[\frac{1}{1} - \frac{1}{n_2^2} \right]$$

$$mvr = J$$

$$KE = \frac{1}{2} \frac{mv^2 \times mr^2}{mr^2}$$

$$= \frac{1}{2} \frac{J^2}{mr^2}$$



$$TE = -13.6 \frac{Z^2}{n^2}$$

-13.6

13.6

(22)



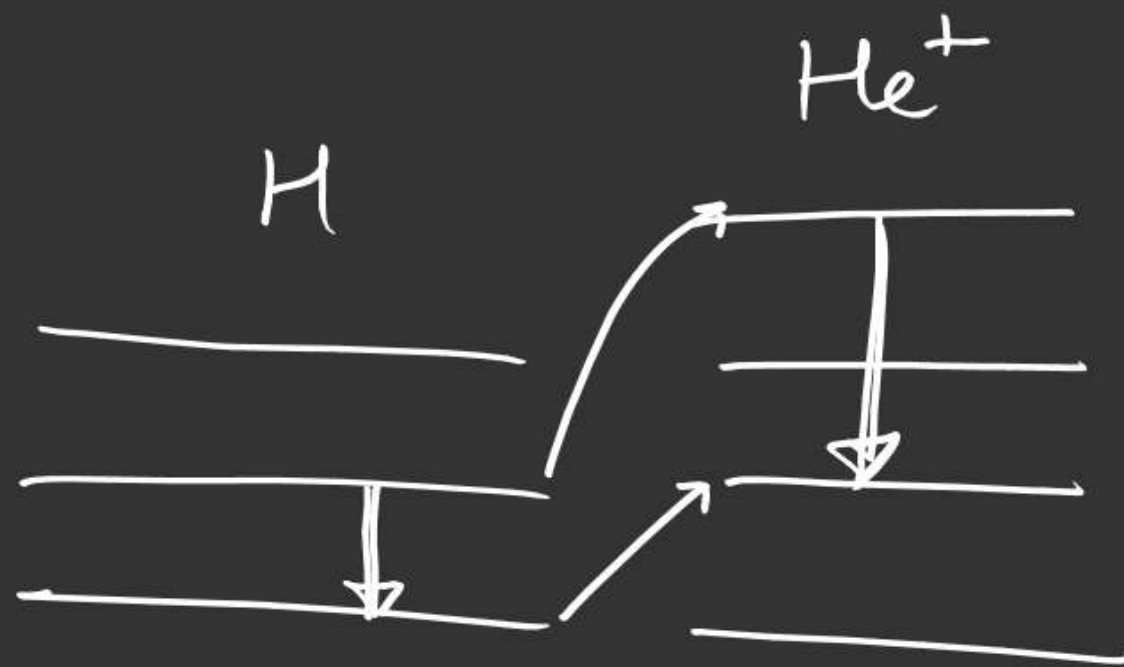
$$IE_1 + IE_2 + IE_3 = 19800$$

$$\underline{520} + IE_2 + IE_3 = 19800$$

$$T.E = -13.6 \times \frac{9}{1}$$

$$= -13.6 \times 9 \times 1.6 \times 10^{19} \times N_A$$

(28)



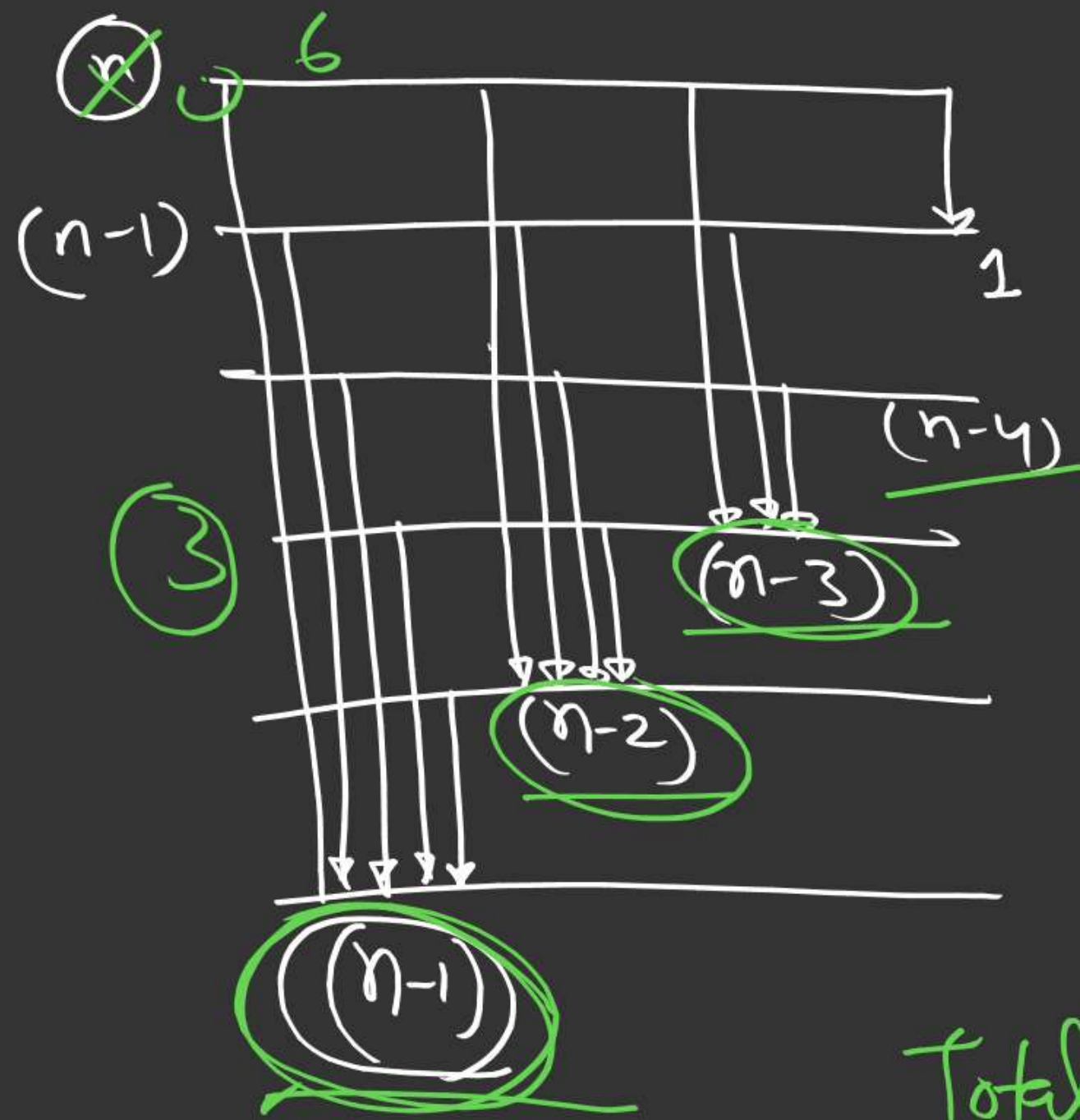
$$n_1 + n_2 = 4$$

$$n_2^2 - n_1^2 = 8$$

$$n_2 - n_1 = 2$$

$$\begin{aligned} n_2 &= 3 \\ n_1 &= 1 \end{aligned}$$

no. of spectral lines



Total no spectral line

$$= 1 + 2 + \dots + n-1$$

$$= \frac{(n-1)n}{2} = nC_2$$

n = no. of energy levels involved

if e^- jumps $n \rightarrow 1$

no. of spectral line

3 \rightarrow 1	4 \rightarrow 1	5 \rightarrow 1	6 \rightarrow 1
Total \rightarrow 3	6	10	15

e^- s in 'H' atoms sample jump from 6th level to ground level. find

$$\text{Total no. of spectral line} = \frac{n(n-1)}{2} = 15$$

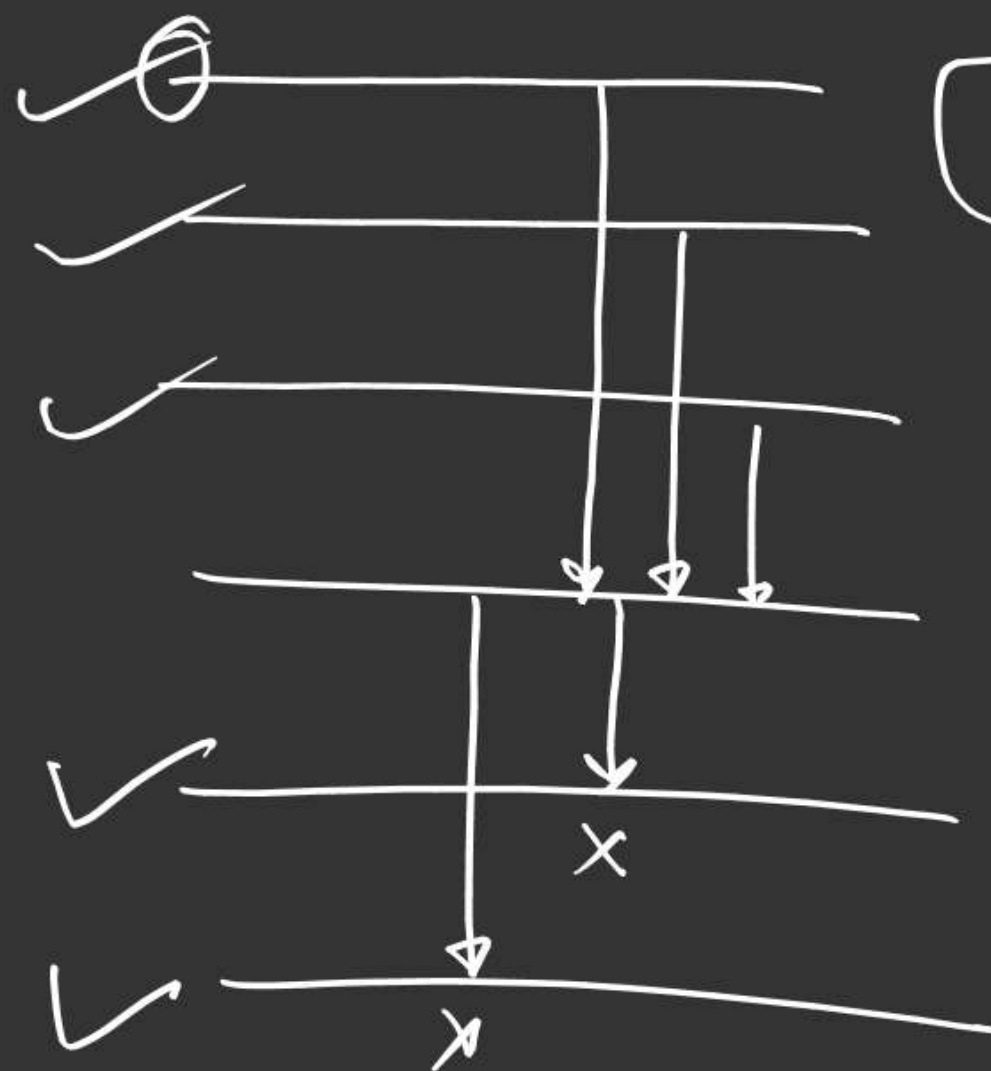
$$\text{① Total no. of spectral line} = 6-1 = 5$$

$$\text{② no. of Lyman series radiation} = 6-1 = 5$$

$$\text{③ no. of Paschen series radiations} = 6-3 = 3$$

Q. e^- in 'H' atoms sample jumps from 6th level to ground level without emitted any Paschen series lines
find total no. of spectral lines.

Ans 10



$$15 - 5 = 10$$

$$\frac{(5-1)(5)}{2} = 10$$

Ionisation energy: Amt of energy required to remove an e^- from an atom

$$\underline{E_{\infty}} = -13.6 \frac{Z^2}{\infty} = 0$$

$$\text{Ionisation} = E_{\infty} - E_n$$

$$\text{Energy} = 0 + 13.6 \frac{Z^2}{n^2}$$

$$\text{I.E} = 13.6 \frac{Z^2}{n^2} = -TE$$

$$\text{I.E of H} = 13.6 \text{ eV}$$

$$\text{I.E of He}^+ = 54.4$$

Binding energy = I.E

extra energy above I.E will be converted into KE of released e^-

Q. Assuming Bohr quantization of angular momentum is applicable on our solar system. Considering only one planet in solar system i.e. earth and using gravitational force find expression for r , v & TE of earth.

$$\text{Force} = \frac{GMm}{r^2}$$

$$\text{Potential Energy} = -\frac{GMm}{r}$$

$$F = -\frac{dU}{dr}$$

$$= +\frac{d}{dr} \left(+\frac{GMm}{r} \right)$$

$$= -\frac{GMm}{r^2}$$



$$\frac{mv^2}{\cancel{r}} = \frac{GMm}{r^2} \quad \text{--- (1)}$$

$$mv r = \frac{nh}{2\pi} \quad \text{--- (2)}$$

$$\frac{\cancel{m}}{r} \cdot \frac{n^2 h^2}{4\pi^2 m^2 \cancel{r^2}} = \frac{GM\cancel{m}}{\cancel{r^2}}$$

$$r = \frac{n^2 h^2}{4\pi^2 GM m^2}$$

$$mv \frac{n^2 h^2}{4\pi^2 GM m^2} = \frac{nh}{2\pi}$$

$$v = \frac{2\pi GMm}{nh}$$

$$\begin{aligned} T.E &= \frac{1}{2}mv^2 - \frac{GMm}{r} \\ &= -\frac{GMm}{2r} \end{aligned}$$

$$T.E = -\frac{2\pi^2 G^2 M^2 m^3}{n^2 h^2}$$

Q. Using Bohr quantization of angular momentum find expression of r , v & T.E for a single e^- hypothetical system for which $PE = k \ln r$

$$F = -\frac{dU}{dr} = -\frac{k}{r}$$

$$\frac{mv^2}{r} = \frac{k}{r} \quad \text{--- (1)}$$

$$mv^2 = \frac{nh}{2\pi} \quad \text{--- (2)}$$

$$\Rightarrow v = \sqrt{\frac{k}{m}}$$

$$\sqrt{m} r \sqrt{\frac{k}{m}} = \frac{nh}{2\pi}$$

$$r = \frac{nh}{2\pi \sqrt{km}}$$

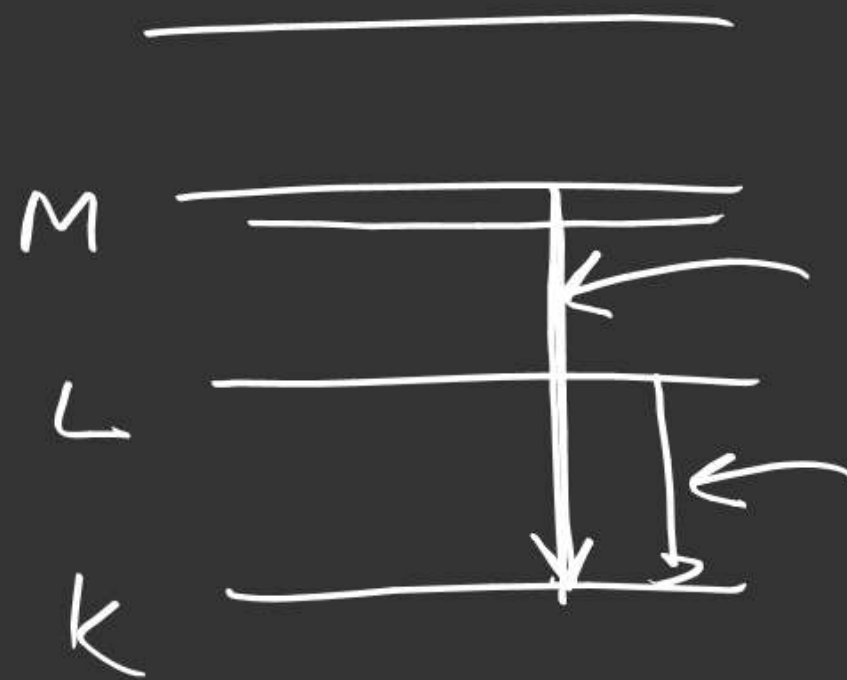
$$T.E = \frac{1}{2}mv^2 + k \ln r$$

$$= \frac{1}{2} \cancel{m} \frac{k}{\cancel{m}} + k \ln r$$

$$= k \left(\frac{1}{2} + \ln r \right)$$

Drawbacks

- 1) Bohr Model failed to explain the fine spectrum of hydrogen atom.
- ② Bohr model failed to explain the splitting of spectral lines in the presence of external electric field called (Stark effect) and magnetic field call Zeeman effect



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