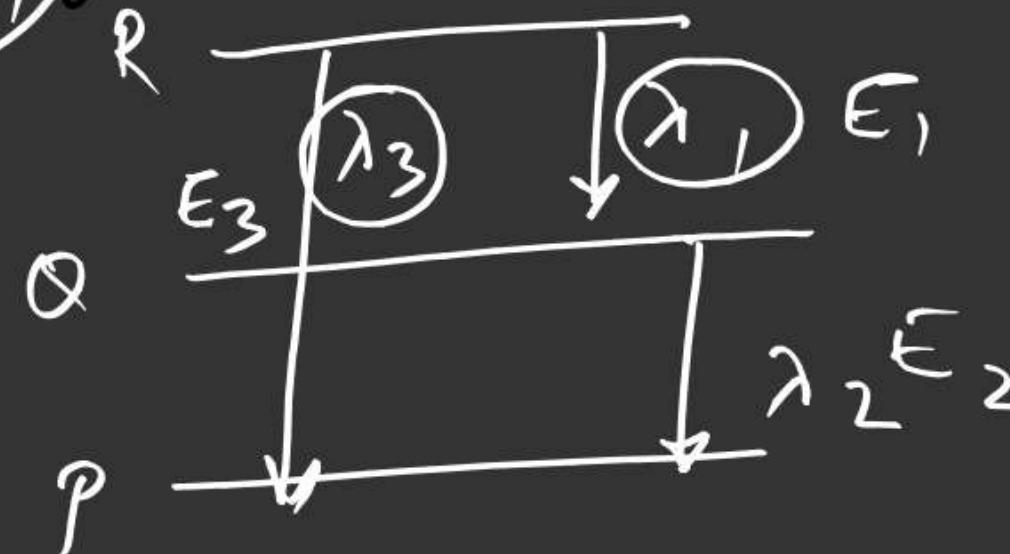


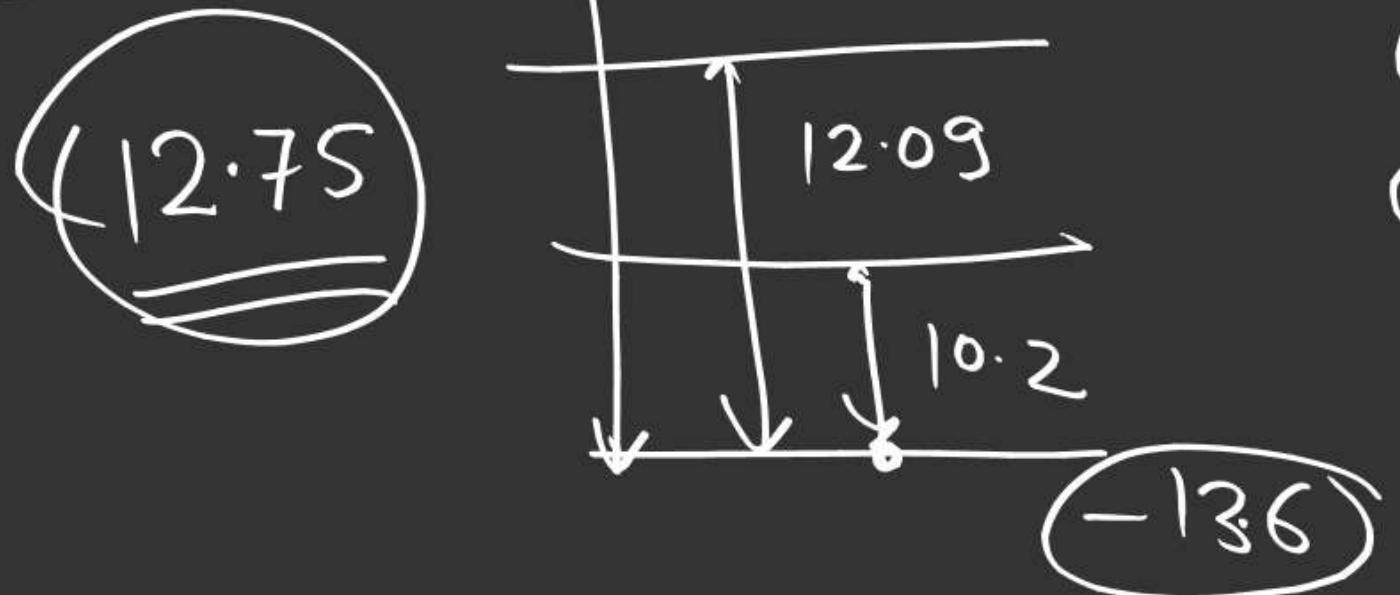
(14)



$$E_3 = E_1 + E_2$$

$$\frac{hc}{\lambda_3} = \frac{hc}{\lambda_1} + \frac{hc}{\lambda_2}$$

(17)



$$-0.85 \quad \left. \begin{array}{l} 12.75 = 13.6 \left[\frac{1}{f} - \frac{1}{n^2} \right] \\ \end{array} \right\}$$

$$mv\tau = J$$

$$KE = \frac{1}{2} \frac{mv^2 \times m\tau^2}{m\tau^2}$$

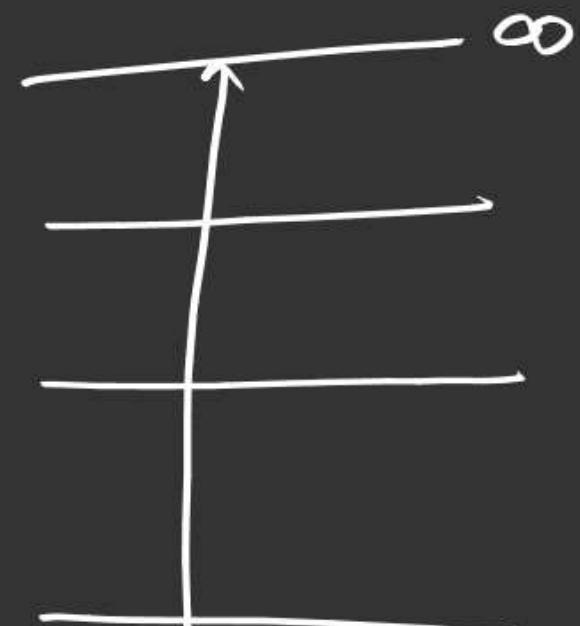
$$= \frac{1}{2} \frac{J^2}{m\tau^2}$$

(22)



$$IE_1 + IE_2 + IE_3 = 19800$$

$$520 + \underline{IE_2} + \underline{IE_3} = 19800$$



$$T.E = -13.6 \frac{z^2}{n^2}$$

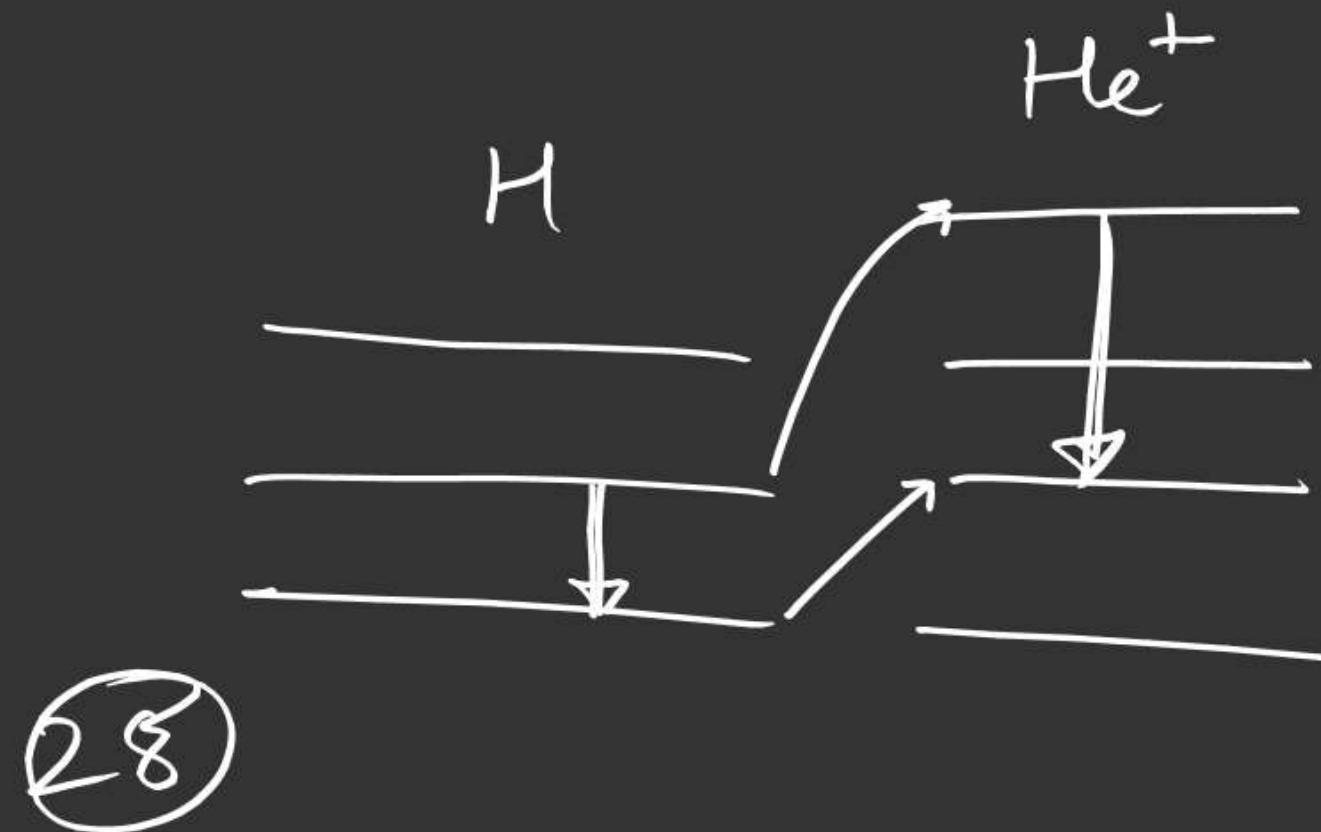
-13.6

13.6

$$T.E = -13.6 \times \frac{9}{1}$$

$$= -13.6 \times 9 \times 1.6 \times 10^{-19} \times N_A$$

↗



28

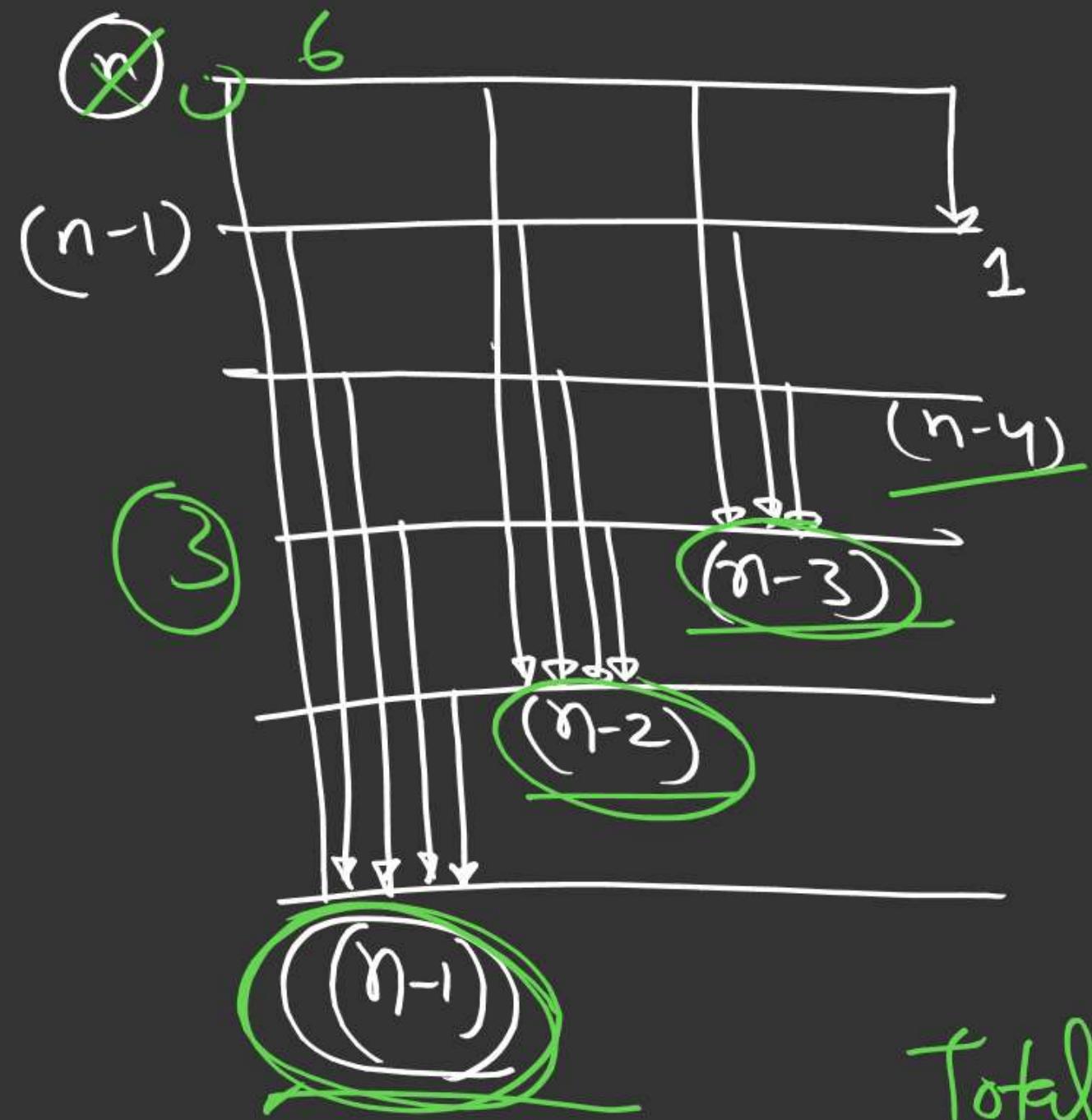
$$n_1 + n_2 = 4$$

$$n_2^2 - n_1^2 = 8$$

$$n_2 - n_1 = 2$$

$$\boxed{n_2 = 3}$$
$$n_1 = 1$$

No. of spectral lines



Total no. of spectral line

$$= 1 + 2 + \dots + n-1$$

$$= \frac{(n-1)n}{2} = nC_2$$

n = no. of energy levels involved

If e's jumps $n \rightarrow 1$

no. of spectral lin

$3 \rightarrow 1$	$4 \rightarrow 1$	$5 \rightarrow 1$	$6 \rightarrow 1$
6	10	15	

Total $\rightarrow 3$

e^- s in 'H' atoms sample jump from 6th level
to ground level. find $= \frac{n(n-1)}{2} = 15$

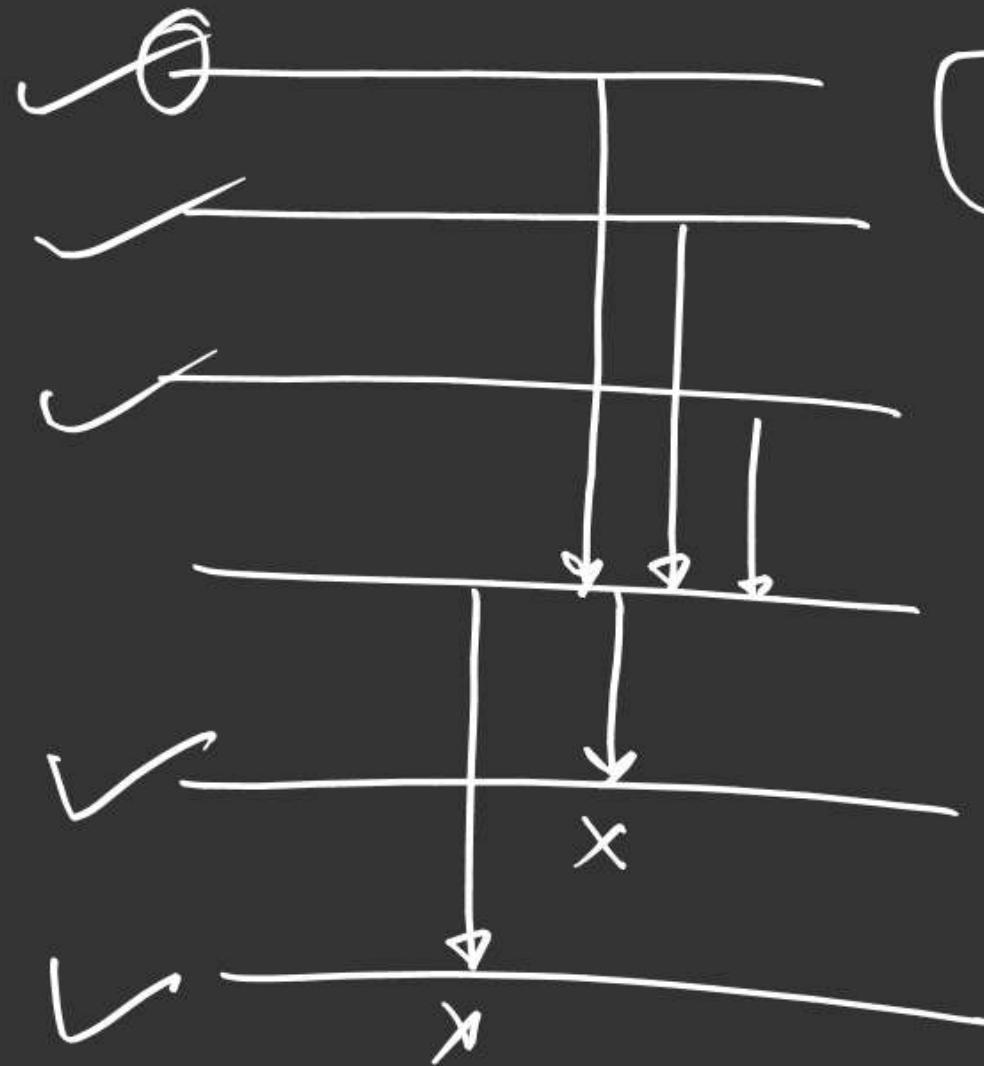
① Total no. of spectral line = 6-1 = 5

② no. of Lyman series radiation = 6-1 = 5

③ no. of Paschen series radiations = 6-3 = 3

Q. e^- in 'H' atoms sample jumps from 6th level to
ground level without emitted any Paschen Series lines

find total no. of spectral lines. Ans 10



$$15 - 5 = 10$$

$$\frac{(5-1)(5)}{2} = 10$$

Ionisation energy: Amt of energy required to remove an e^- from an atom

$$\underline{E_\infty} = -13.6 \frac{z^2}{\infty} = 0$$

$$\text{Ionisation} = E_\infty - E_n$$

$$\text{Energy} = 0 + 13.6 \frac{z^2}{n^2}$$

$$I.E = 13.6 \frac{z^2}{n^2} = -\Delta E$$

Binding energy = $I.E$ # extra energy above $I.E$ will be converted into $K.E$ of released e^-

$$I.E \text{ of H} = 13.6 \text{ eV}$$

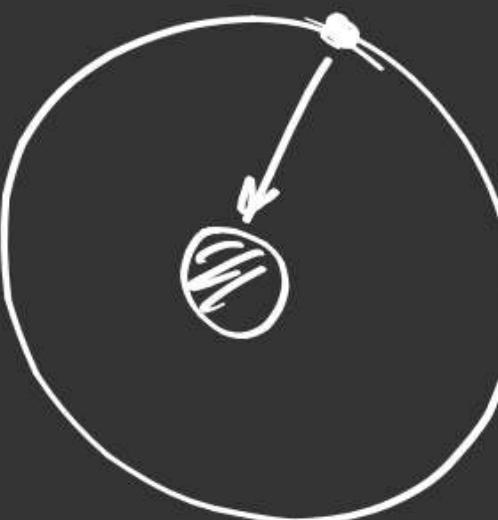
$$\text{II } I.E \text{ of He}^+ = 54.4$$

Q. Assuming Bohr quantization of angular momentum is applicable on our solar system. Considering only one planet in solar system i.e. earth and using gravitational force find expression for r , v & T_E of earth.

$$\text{Force} = \frac{GMm}{r^2}$$

$$\begin{aligned} \text{Potential Energy} &= -\frac{GMm}{r} \end{aligned}$$

$$\left. \begin{aligned} F &= -\frac{dU}{dr} \\ &= +\frac{d}{dr} \left(+\frac{GMm}{r} \right) \\ &= \Theta \frac{GMm}{r^2} \end{aligned} \right\}$$



$$\frac{mv^2}{r} = \frac{GMm}{r^2} - \textcircled{1}$$

$$mv^2 r = \frac{nh}{2\pi} - \textcircled{2}$$

~~$$\frac{mv^2}{r} \cdot \frac{n^2 h^2}{4\pi^2 m^2 r^2} = \frac{GMm}{r^2}$$~~

$$r = \frac{n^2 h^2}{4\pi^2 GMm^2}$$

$$mv \frac{n^2 h^2}{4\pi^2 GMm^2} = \frac{nh}{2\pi}$$

$$v = \frac{2\pi GMm}{nh}$$

$$\overbrace{T.E = \frac{1}{2}mv^2 - \frac{GMm}{r}}$$

$$= -\frac{GMm}{2r}$$

$$T.E = -\frac{2\pi^2 G^2 M^2 m^3}{n^2 h^2}$$

Q. Using Bohr quantization of angular momentum

find expression of r, v & $T.E$ for a single e^-

hypothetical system for which $PE = k \ln r$

$$F = -\frac{dU}{dr} = -\frac{k}{r}$$

$$\frac{mv^2}{r} = \frac{k}{r} \quad \textcircled{1} \Rightarrow v = \sqrt{\frac{k}{m}}$$

$$mv r = \frac{nh}{2\pi} \quad \textcircled{2}$$

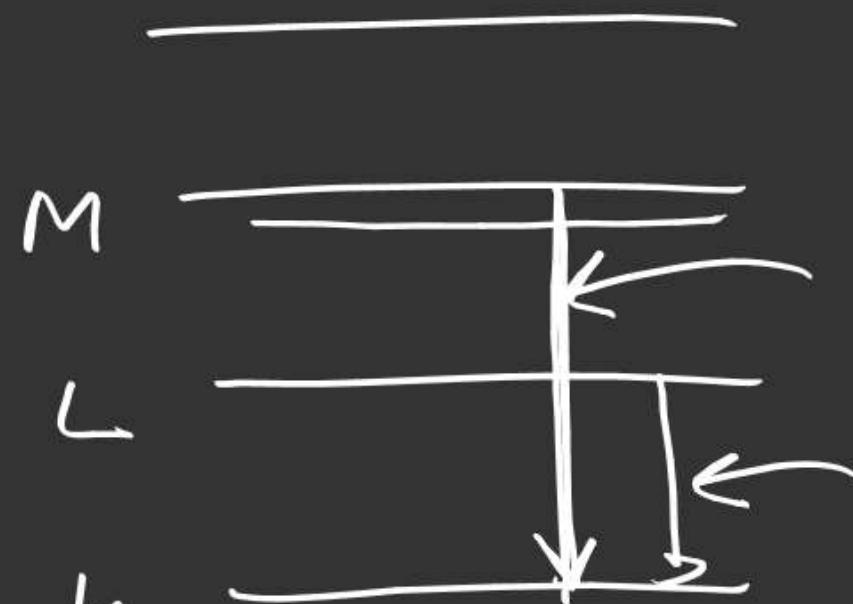
$$jmr \sqrt{\frac{k}{m}} = \frac{nh}{2\pi}$$

$$r = \frac{nh}{2\pi \sqrt{km}}$$

$$\begin{aligned} T.E &= \frac{1}{2} mv^2 + k \ln r \\ &= \frac{1}{2} m \frac{k}{r} + k \ln r \\ &= k \left(\frac{1}{2} + \ln r \right) \end{aligned}$$

Drawbacks

- 1) Bohr Model failed to explain the fine spectrum of hydrogen atom.
- 2) Bohr model failed to explain the splitting of spectral lines in the presence of external electric field called (Stark effect) and magnetic field called Zeeman effect



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