Chapter 1

Canonical Ensemble

1.0.1 Assumptions

- No uncertainty in energy states.

- Each system is a pure system.

- An isolated system is equally likely to be in any available state.

- The average of a property over all available microstates is the macroscopic value.

1.1 MicroCanonical

For a system with constant number, volume, and energy, the number of microstates is
$$\Omega = \Omega(N, V, E).$$

Now take an ensemble or collection of systems like this.Then,

$$E_{tot} = AE$$

$$N_{tot} = AN$$

$$V_{tot} = AV.$$

Since we have no reason to favor one system over the other, we assume each state with the same N, V, E is equally probable. This is known as the principle of equal *a priori* probabilities.

1.2 Canonical

1.2.1 Deriving the Partition Function

Now lets take a ensemble of systems with the same temperature, volume, and number. The ensemble is put into a heat bath until it is fully equilibrated, and then moved into perfect thermal insulation. The entire ensemble then becomes a system in the microcanonical ensemble.

We know that the system must obey,

$$A = \sum_j a_j$$

$$E_{tot} = \sum_j a_j E_j \text{ where } a_j \text{ is the occupancy number of state j.}$$

We also know that for each energy $E_j, \Omega(E_j)$ unique quantum states exist. We do not know the occupancy number of states, however.

We do know (assumption) that for the microcanonical ensemble each state is equally likely. This means that each state for the entire canonical ensemble

is equally likely since the entire canonical ensemble is a system in the micro-canonical ensemble; that is all combinations of occupancy numbers is equally likely.

The number of distributions of occupancy numbers is given by

$$W(\vec{a}) = \frac{A!}{\prod_j a_j!}.$$

This a multinomial coefficient meaning that at large A it is very sharply peaked; in fact, arbitrarily so at large enough A. The probability of being in state j is then,

$$P_j = \frac{\bar{a}_j}{A} = \frac{\sum_a W(\vec{a}) a_j(\vec{a})}{A \sum_a W(\vec{a})}.$$

At infinite number of systems in an ensemble, all contributions besides the most common become completely negligible, thus

$$P_j = \frac{\bar{a}_j}{A} = \frac{a_j^*}{A}.$$

We now have a maximization problem with the following constraints.

$$A = \sum_j a_j$$

$$E_{tot} = \sum_j a_j E_j.$$

Since ln is a monotonic function and, thus, will preserve maxima and minima of inputted functions, $\ln(W(\vec{a}))$ can be maximized instead of $W(\vec{a})$. For further calculations it behooves us to recognize that

$$\ln(W) = \ln(A!) - \sum_j \ln(a_j!).$$

Note the first term is constant and the second is approximately

$$-\sum_j a_j \ln(a_j) + \sum_j a_j.$$

Then,

$$\frac{\partial(\ln W(\vec{a}) - \alpha \sum_j a_j - \beta \sum_j a_j E_j)}{\partial a_k} = 0.$$

For each k then,

$$\ln(a_k^*) + \alpha + 1 + \beta E_k =$$

or

$$a_k^* = e^{-\alpha'} e^{-\beta E_k}.$$

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Summing over all (a_j^*) gives,

$$A = e^{-\alpha'} \sum_j e^{-\beta E_j}.$$

Finally,

$$P_j = \frac{e^{-\alpha'} e^{-\beta E_k}}{e^{-\alpha'} \sum_j e^{-\beta E_j}} = \frac{e^{-\beta E_k}}{\sum_j e^{-\beta E_j}}.$$

We call the denominator Q and label it the partition function. We know show the ensemble energy and some of its derivatives.

1.2.2   Finding β

$$\bar{E} = \frac{\sum_j E_j e^{-\beta E_j}}{Q}.$$

$$p_j = -\left(\frac{\partial E}{\partial V}\right)_N.$$

$$\bar{p} = -\frac{\sum_j \left(\frac{\partial E}{\partial V}\right)_N e^{-\beta E_j}}{Q}.$$

We can then find,

$$\left(\frac{\partial \bar{E}}{\partial V}\right)_{N,\beta} = -\bar{p} + \beta \overline{Ep} - \beta \bar{E}\bar{p}$$

and

$$\left(\frac{\partial \bar{p}}{\partial \beta}\right)_{N,V} = \bar{E}\bar{p} - \overline{Ep}.$$

Comparing the thermodynamic and ensemble averages we see,

$$\left(\frac{\partial \bar{E}}{\partial V}\right)_{N,\beta} + \beta \left(\frac{\partial \bar{p}}{\partial \beta}\right)_{N,V} = -\bar{p},$$

$$\left(\frac{\partial E}{\partial V}\right)_{N,T} - T \left(\frac{\partial p}{\partial T}\right)_{N,V} = -p.$$

One can see from this that $\beta = \frac{k}{T}$. k will be the same for any system as can be seen from a thought experiment where the ensemble consists of two system composites with the same N, V, T. They individual systems will be in equilibrium and from prior calculations the partition function and probability functions are completely separable. This requires they have the same constant k.

1.2.3 Entropy

Let

$$f = \ln Q.$$

Then,

$$df = \left(\frac{\partial f}{\partial \beta}\right)_{E_j} d\beta + \sum_k \left(\frac{\partial f}{\partial E_k}\right)_{\beta, E_{j \neq k}} dE_k.$$

This derivative is equivalent of slightly changing the volumes of all systems which changes the energy levels (which is nothing but PV work) and slightly change the temperature before isolated the ensemble again.

This becomes with some manipulation,

$$df = -\bar{E}d\beta - \beta \sum_k P_k dE_k$$

$$d(f + \beta\bar{E}) = \beta \left(d\bar{E} - \sum_k P_k dE_k\right).$$

The first term is the average energy increase in a system, while the second is the average work done on a system meaning the argument in parenthesis is nothing but the reversible heat. This requires that β is an integrating factor of f.

$$d(f + \beta\bar{E}) = \beta q_{rev}$$

$$d\left(f + \frac{\bar{E}}{kT}\right) = \frac{q_{rev}}{kT}.$$

Thermodynamically the left side must be dS/k, so finally,

$$dS = \frac{d\bar{E}}{T} + kd(\ln Q).$$

1.2.4 Spontaneous Processes

One can think of spontaneous processes as processes that occur because an increase in available states.

We take the known relation,

$$A(N, V, T) = -kT \ln Q(N, V, T).$$

In addition, one can also recognize that a summation over energy levels rather than states can be done.

$$Q = \sum_l \Omega(N, V, E_l)e^{-\beta E_l}.$$

For a spontaneous process we require that,

$$\Omega_2 \geq \Omega_1.$$

Then,
$$Q_2 - Q_1 = \sum_l (\Omega_2 - \Omega_1)e^{-\beta E_l} \geq 0.$$

Finally,
$$\Delta A = A_2 - A_1 = -kT \ln \frac{Q_2}{Q_1} < 0.$$

Chapter 2

Other Ensembles and Fluctuations

2.1 Grand Canonical Ensemble

The grand canonical ensemble is a collection of systems with the same volume, temperature, and chemical potential. The ensemble can be thought of as a collection of systems which have been equilibriated w.r.t chemical potential and temperature by placing the ensemble in a heat bath and particle bath. The ensemble once equilibriated is then isolated. Each system within the ensemble is free to exchange particles or energy though.

With the concept of the GC ensemble established, these equations are easily acertained.

$$\sum_N \sum_j a_{Nj} = A$$

$$\sum_N \sum_j a_{Nj} E_{Nj} = E_{tot}$$

$$\sum_N \sum_j a_{Nj} N = N_{tot}$$

Note that a_{Nj} is the occupancy number of systems with N particles and the j-th energy state. Like the canonical ensemble the GC ensemble is a system in the microcanonical ensembe so we can use the principle of equal *a priori* probabilities to the distributions of occupancy numbers. The number of states given a distribution $a_{\bar{N}j}$ is

$$W(a_{\bar{N}j}) = \frac{A!}{\prod_N \prod_j a_{Nj}!}.$$

The solution to maximizing W with the constrants given is nearly identical to the C ensemble method and simply becomes

$$a_{Nj}^* = e^{-\alpha} e^{-\beta E_{Nj}} e^{-\gamma N} \text{ thus,}$$

$$P_{Nj} = \frac{a_{Nj}^*}{A} = \frac{e^{-\alpha} e^{-\beta E_{Nj}} e^{-\gamma N}}{\sum_N \sum_j e^{-\alpha} e^{-\beta E_{Nj}} e^{-\gamma N}}$$

We note that the denominator is a sum of all states with a two biasing factors based on βE_{Nj} and γN. Like the C ensemble this will be called the partition function of the GC ensemble. We label it Ξ.

2.1.1 Finding β and γ

Lets think of the GC ensemble as collections of canonical ensembles with varying numbers. Each ensemble is in thermal equilibrium since the entire GC is. If we

suddenly prevent molecular transport. Then Ξ breaks down to

$$\Xi = e^{\gamma N} \sum_j e^{-\beta E_j}, \ \forall N.$$

This is the canonical partition function with an extra constant which can be ignored since this constant in probabilities would be applied to both the numerator and denominator. Thus, β must be the same as the canonical ensemble.

γ is a little more involved to derive. If we take the full derivative of $\ln \Xi$, we find

$$df = \left(\frac{\partial f}{\partial \beta}\right)_{\gamma, E_{Nj}} d\beta + \left(\frac{\partial f}{\partial \gamma}\right)_{\beta, E_{Nj}} d\gamma + \left(\frac{\partial f}{\partial \gamma}\right)_{\beta, \gamma, E_{N_i k \neq Nj}} dE_{Nj}.$$

This ends up being

$$df = -\bar{E}d\beta - \bar{N}\gamma + \beta\bar{p}dV$$
$$d(f + \beta\bar{E} + \gamma\bar{N}) = \beta d\bar{E} + \beta\bar{p}dV + \gamma d\bar{N}.$$

For multicomponent systems this is the enemble equivalent to the differential form of entropy. Then γ must be

$$\gamma = -\frac{\mu}{kT}.$$

An important note is that γ is essential another energy scale (chemical potential) compared to thermal energy.

2.1.2 Other Points

Ξ can be written as,

$$\sum_N Q(N)e^{\beta\mu N}$$

Often times we see $e^{\beta\mu}$ as λ which is called the fugacity and is the absolute activity of a state. This can be seen since

$$\mu = kT \ln \lambda$$
$$\Delta\mu = kT \ln(\lambda_2/\lambda_1)$$

The function pV is the natural thermodynamic function for the GC ensemble. The proof is simple,

$$S = \frac{\bar{E}}{T} - \frac{\bar{N}\mu}{T} + k \ln \Xi$$
$$G = \mu N = E + pV - TS$$
$$pV = \mu N - E + TS = \mu N - E + T(\frac{\bar{E}}{T} - \frac{\bar{N}\mu}{T} + k \ln \Xi)$$

One final note, the summation over N goes to infinity since in the limit of infinite systems there is a nonzero chance of at least one system having infinite particles. However, we will see that practically speaking the summation could be cut off for reasons of fluctuations.

2.2   Gibb's Ensemble

Gibb's ensemble is a collection of systems that have been thermally equilibriated and have impermeable but flexible walls. Thus pressure is also equal each system. Number is also fixed by impermeability. Δ is the standard denotation of the partition function which is,

$$\sum_E \sum_V \Omega(V, E) e^{-\beta E} e^{\beta p V}$$

True to its name the natural thermodynamic function for the Gibb's ensemble is Gibb's energy.

2.3   MicroCanonical Ensembles

2.3.1   Universality of Microcanonical Ensemble

For each ensemble, when summing over some combination of energy levels, volumes, or number, we find the degeneracy by the microcanonical partition function Ω. Thus all other ensembles can be thought of as deviations from the microcanonical. The exponential terms are biasings based on intensive properties and the sums represent summations over the extensive properties.

2.3.2   Boltzman Equation

If we start from the definition of entropy derived from the GC partition function, we find the famous Boltzman equation which is actually on Boltzman's tomb, and gives intuition as to the nature of entropy.

$$S = k \ln \Xi + k \left(\sum_{N,j} \beta E_{Nj} \frac{e^{-\beta E_{Nj}} e^{-\gamma N}}{\Xi} + \sum_{N,j} \gamma \frac{N e^{-\beta E_{Nj}} e^{-\gamma N}}{\Xi} \right)$$

$$S = k \ln \Xi + k \sum_{N,j} (\beta E_{Nj} + \gamma N) \frac{e^{-\beta E_{Nj}} e^{-\gamma N}}{\Xi}$$

Recall that,

$$\frac{a^*_{Nj}}{A} = \frac{e^{-\beta E_{Nj}} e^{-\gamma N}}{\Xi}$$

and

$$\ln \frac{a^*_{Nj}}{A} = \ln a^*_{Nj} - \ln A = \ln \left(e^{-\beta E_{Nj}} e^{-\gamma N} \right) - \ln \Xi$$

$$= -(\beta E_{Nj} + \gamma N + \ln \Xi)$$

Thus,

$$S = k \ln \Xi - k \sum_{Nj} \left(\ln a_{Nj}^* - \ln \Xi + \ln A \right) \frac{a_{Nj}^*}{A}$$

$$= k \ln A - \frac{k}{A} \sum_{N,j} a_{Nj}^* \ln a_{Nj}^*$$

This is for each system in the GC ensemble. The entire entropy is then,

$$S = kA \ln A - k \sum_{N,j} a_{Nj}^* \ln a_{Nj}^*$$

$$= k \ln \frac{A}{\sum_{N,j} a_{Nj}^* \ln a_{Nj}^*}$$

$$= k \ln W(a_{Nj}^{\vec{}})$$

$$= k \ln \Omega.$$

2.4 Fluctuations

For the assumption that large systems only assume states in their ensemble average, fluctuations around the mean must be small. If they were not, then statistical mechanics would predict that we should observe macroscopic fluctuations in state. Two canonical ways of showing that fluctuations are indeed small (as compared to the mean) the energy of the canonical and number of the grand canonical ensemble will be shown. The most important equation when dealing with fluctuations or second moments around the mean is,

$$\sigma_M^2 = \overline{\left(M - \bar{M} \right)^2} = \overline{M^2} - \bar{M}^2$$

2.4.1 Energy Fluctuations

For the canonical ensemble $\overline{E^2}$ is

$$\sum_j E_j^2 P_j = \frac{1}{Q} \sum_j E_j^2 e^{-\beta E_j}$$

$$= \frac{-1}{Q} \frac{\partial}{\partial \beta} \sum_j E_j e^{-\beta E_j}$$

The last sum is just the average energy without the partition function's normalization, so

$$\overline{E^2} = \frac{-1}{Q} \frac{\partial}{\partial \beta} (\bar{E} Q).$$

Via product rule we get

$$\overline{E^2} = -\frac{\partial \bar{E}}{\partial \beta} - \bar{E}\frac{\partial \ln Q}{\partial \beta}$$

$$= kT^2\frac{\partial \bar{E}}{\partial \beta} + \bar{E}^2.$$

Finally, via the definition of the varience,

$$\sigma_E^2 = kT^2\left(\frac{\partial \bar{E}}{\partial T}\right)_{N,V} \qquad\qquad = kT^2 C_v$$

Compared to the average energy then the average fluctuations are,

$$\frac{\sigma_E}{\bar{E}} = \frac{(kT^2 C_v)^{\frac{1}{2}}}{\bar{E}}$$

For an ideal gas $\bar{E}$ scales as $O(NkT)$ while C_v scales as $O(Nk)$. Thus the fluctuations are proportional to $O(N^{\frac{1}{2}})$ and in a large system are very small. It is worth noting the same result can be achieved from a Taylor expansion of the probability distribution around $\bar{E}$.

2.4.2   Density Fluctuations

For the GC ensemble we shall look at number fluctuations. The fluctuations are

$$\sum_{N,j} N^2 P_{Nj} = \frac{1}{\Xi}\sum_{N,j} N^2 e^{-\beta E_{Nj}}e^{-\gamma N} = \frac{-1}{\Xi}\frac{\partial}{\partial \gamma}\sum_{Nj} N e^{-\beta E_{Nj}}e^{-\gamma N}$$

$$= \frac{-1}{\Xi}\frac{\partial}{\partial \gamma}(\bar{N}\Xi)$$

$$= \frac{\partial \bar{N}}{\partial \gamma} - \bar{N}\frac{\partial \ln \Xi}{\partial \gamma}$$

$$= kT\left(\frac{\partial \bar{N}}{\partial \mu}\right)_{V,T} + \bar{N}^2$$

This looks very similar to the canonical example and should because it is analogous. All we need now is the thermodynamic relationship,

$$\left(\frac{\partial \mu}{\partial N}\right)_{V,T} = -\frac{V^2}{N^2}\left(\frac{\partial p}{\partial \mu}\right)_{N,T}.$$

The fluctuations are,

$$\sigma_N = \left(\frac{\bar{N}^2 kT\kappa}{V}\right)^{\frac{1}{2}}$$

$$\frac{\sigma_N}{\bar{N}} = \left(\frac{kT\kappa}{V}\right)^{\frac{1}{2}}.$$

Here κ is the isothermal compressibility which is $\frac{1}{p}$ for the ideal gas. Then the fluctuations for the ideal gas is once again the property to the $\frac{-1}{2}$. Furthermore, like before the Taylor expansion can be performed to get similar results.

For number fluctuation, this has measurable effects on the Raleigh scattering. The functions make it so that the scattering is a function of wavelength to the -4th. Therefore, in our atmosphere, blue light gets scattered much more and the sky is blue and sunsets are red.

Chapter 3

Boltzman, Fermi-Dirac, and Bose-Einstein Statistics

When we wish to decouple a system into non-interacting parts, different statistics apply based on the underlying physics of the system. Three major types of statistics exist with respect to molecular systems: Boltzmann, Fermi-Dirac, and Bose-Einstein. To be exact Boltzmann statistics is the limit of the other two at sufficiently high T.

- Fermi-Dirac

  - Occurs when the exchange of basis particles is antisymmetric with respect to the wave function.
  - Example fermions are electrons which the extrapolation of the Fermi-Dirac statistics is the Pauli exclusion principle.

- Bose-Einstein

  - For systems where the exchange of basis particles is symmetric with respect to the wave function, Bose-Einstein statistics apply.
  - Examples are protons and neutrons.

- Boltzmann

  - The high temperature limit of the other two.
  - Ends up being the classical approximation of the quantum mechanical equations.

3.1   Boltzmann Statistics

3.1.1   Hamiltonian Decoupling

If a multiparticle system is dilute enough the interactions between particles is negligible, and the system can be treated as a N independent body problem. Even in liquids and solids convenient "basis particles" can be chosen so the system is modeled by a N independent body system. In these cases where interactions can be ignored, the Hamiltonian can be decoupled.

$$\mathcal{H} = \sum_{i=1}^{N} \mathcal{H}_i$$

In addition, for individual molecules if the individual modes of molecular energy are decoupled,

$$\mathcal{H}_i = \sum_{m \in M} \mathcal{H}_i^m,$$

where M is the set of all decoupled energy modes.

3.1.2 Partition Function Decoupling

We now consider a N distinguishable, independent, body system. Then,

$$\mathcal{Q}(N,V,T) = \sum_j e^{-\beta E_j} = \sum_{j,k,\dots} e^{-\beta \cdot (\epsilon_j^a + \epsilon_k^b + \epsilon_l^c + \dots)}$$
$$= q_a q_b q_c \dots$$

Here we see the canonical partition function becomes N individual particle partition functions. Since the particles are identical, we have

$$q(V,T) = \sum_j e^{-\beta \epsilon_j}$$
$$\mathcal{Q} = q^N.$$

This equation is only true in a distinguishable system though which in molecular systems in rarely the case. Including distinguishably leads to some problems. Consider, two equivalent energy states,

$$E_{tot} = \epsilon_i^a + \epsilon_j^b + \epsilon_j^c + \dots$$
$$E_{tot} = \epsilon_j^a + \epsilon_i^b + \epsilon_j^c + \dots \quad .$$

In fact, there are N such equivalent arrangements. However, in the case that all individual energies are different, there are $N!$ combinations. The problem, thus, is that different energy combinations have differing numbers of equivalent arrangements. This would imply that a simple divisor cannot be applied to reduce the number of configurations of a distinguishable system to a indistinguishable one.

Despite this fact, we can still often make the approximation of dividing the distinguishable partition function by $N!$. At room temperature ($m = 10^{-22}g$, $a = 10$cm), the number of energy states available to a particle in a box is $O(10^{30})$. This number signifies that in all but extremely dense systems at around room temperature or even colder is much greater than the number of particles, so it is unlikely that any two particles adopt the same energy, so

$$\mathcal{Q} = \frac{q^N}{N!},$$

is an excellent approximation for all but 'the lightest molecules at very low temperatures.'

3.1.3 Getting Molecular Properties

Since we now are dealing with individual molecular partition functions, we can get averages of molecular properties. The probability of being in state j is

$$\pi_j = \frac{e^{-\beta \epsilon_j}}{q}.$$

Any mechanical properties is then,

$$\bar{M} = \sum_j M_j \pi_j.$$

The parallelism to the ensembles already mentioned is not an accident as Boltzmann's original examination of statistical mechanics used a system of molecules in thermal equilibrium as his start. In fact, we could from *a priori* principles derive the previous two equations.

3.2 Fermi-Dirac and Bose-Einstein Statistics

3.2.1 Common Derivation

Derivation of the exact distribution of states in an independent system obeying either Fermi-Dirac or Bose-Einstein statistics is most readily done in the grand canonical ensemble. We let $E_j(N, V)$ be the energy available to a N body system, ϵ_k the molecular quantum states, and n_k the occupancy number of a molecular quantum state. Thus it follows,

$$E_j = \sum_k \epsilon_k n_k$$

$$N = \sum_k n_k.$$

From here, we take the Ξ and find its form on a molecular basis.

$$\Xi = \sum_N \lambda^N \sum_j e^{-\beta E_j} = \sum_N \lambda^N \sum_{n_k \in S} e^{-\beta \cdot \sum_i \epsilon_i n_i},$$

where S is the set of all occupancies numbers that obeys $\sum_k n_k = N$. Some more manipulation leads to,

$$\Xi = \sum_{N=0}^{\infty} \sum_{n_k \in S} \lambda^{\sum_i n_i} e^{-\beta \cdot \sum_i \epsilon_i n_i}$$

$$= \sum_{N=0}^{\infty} \sum_{n_k \in S} \prod_k (\lambda e^{-\beta \epsilon_k})^{n_k}.$$

We have moved λ into the second summation which just distributes λ^N to each term. We then isolated n_k by using a series product. From here, we must notice that since $N \to \infty$, every distribution of $n_k \in \mathbb{N}$ compatible with the occupancy number requirements of the given statistics will be summed over. Thus each n_k will vary over all its possible values in combination with every other occupancy

number. The knowledge allows us to write the previous equation as,

$$\Xi = \sum_{n_1}^{\max(n_1)} \sum_{n_2}^{\max(n_2)} \sum_{n_3}^{\max(n_3)} \cdots \prod_k (\lambda e^{-\beta \epsilon_k})^{n_k}$$

or

$$\Xi = \sum_{n_1}^{\max(n_1)} (\lambda e^{-\beta \epsilon_1})^{n_1} \sum_{n_2}^{\max(n_2)} (\lambda e^{-\beta \epsilon_2})^{n_2} \cdots \quad .$$

This is simple the product of N terms so we have,

$$\Xi = \prod_k \sum_{n_k}^{\max(n_k)} (\lambda e^{-\beta \epsilon_k})^{n_k}$$

3.2.2   Fermi-Dirac

For Fermi-Dirac statistics each state is either uniquely occupied or not occupied at all. Therefore, the summation above expands to two terms for occupied and unoccupied states.

$$\Xi = \prod_k (1 + \lambda e^{-\beta \epsilon_k}).$$

3.2.3   Bose-Einstein

In the case of Bose-Einstein statistics each state can be occupied any number of times without restriction. This becomes,

$$\Xi = \prod_k \sum_{n_k}^{\infty} (\lambda e^{-\beta \epsilon_k})^{n_k}$$

$$= \prod_k \sum_{i_k}^{\infty} x^{i_k}$$

$$= \prod_k (1 + x)^{-1} \quad \text{for } x < 1$$

$$= \prod_k (1 + \lambda e^{-\beta \epsilon_k})^{-1}$$

The last two lines uses the result from the geometric series. One way to write this result is,

$$\Xi_{BE}^{FD} = \prod_k (1 + \lambda e^{-\beta \epsilon_k})^{\pm 1}.$$

From here we note that,

$$\ln \Xi = \sum_k \ln (1 \pm \lambda e^{-\beta \epsilon_k}).$$

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Thus $\bar{E}$ can be calculated by first noting,

$$\bar{n}_k = \frac{\lambda e^{-\beta \epsilon_k}}{1 \pm \lambda e^{-\beta \epsilon_k}}$$

then,

$$\bar{E} = \sum_k \bar{n}_k \epsilon_k = \sum_k \frac{\epsilon_k \lambda e^{-\beta \epsilon_k}}{1 \pm \lambda e^{-\beta \epsilon_k}}.$$

By allowing $T \to \infty$ or $\frac{N}{V} \to 0$, $\lambda \to 0$ which means $\bar{n}_k \to 0$. This can be shown to make both Fermi-Dirac and Bose-Einstein statistics go to the Boltzmann limit.

Chapter 4

The Monotonic Ideal Gas

4.1 The Monotonic Ideal Gas

This chapter will concern gases that are sufficiently dilute as to have negligible interactions with other particles. In addition, these molecules will be monotonic i.e. only consist of one atom. Therefore, no rotational or vibrational energy will be taken into consideration. Furthermore, we will assume that the partition function of a molecule can be decoupled into its individual energy modes.

Assumptions

- Gas is dilute enough to neglect individual particle interactions.

- Individual energy modes in the atom are decoupled.

4.2 The Individual Molecular Partition Functions

First we must assess the different modes of energy that a molecule can carry. Since there are no interactions, there is no potential energy in the system unless a body force were to be acting on the entire system. The only available modes for storing energy into the system are then,

1. Translational

2. Electronic

3. Nuclear.

The partition functions are then,

$$q_{mol} = q_{trans}q_{elec}q_{nucl}$$

$$\mathcal{Q} = \frac{q_{mol}^N}{N!}$$

$$\mathcal{Q} = \frac{(q_{trans}q_{elec}q_{nucl})^N}{N!}.$$

4.2.1 Translational Partition Functions

We begin with the exploration of the translational partition function. From one of the first problems solved in quantum mechanics, the particle in an infinite potential box, we know,

$$\epsilon_{n_x,n_y,n_z} = \frac{h^2}{8ma^2}(n_x^2 + n_y^2 + n_z^2).$$

From here we can directly plug in ϵ into q_{trans}.

$$q_{trans} = \sum_{j,k,l} e^{-\beta(\epsilon_{n_x} + \epsilon_{n_y} + \epsilon_{n_z})}$$

$$= \sum_i e^{-\beta\epsilon_{n_i}} \sum_j e^{-\beta\epsilon_{n_j}} \sum_l e^{-\beta\epsilon_{n_l}}$$

$$= \left(\sum_{i=0}^{\infty} e^{-\beta\epsilon_{n_l}} \right)^3$$

$$= \left(\sum_{i=0}^{\infty} \exp\left(-\frac{\beta h^2 n^2}{8ma^2} \right) \right)^3$$

This summation cannot be expressed exactly analytically. However, the density of available states is so large, that the available energy level can be treated as continuous which means the summation can be approximated by an integral. The difference between two energy levels is given by,

$$\Delta = \frac{\beta h^2 (n_x + 1)^2}{8ma^2} - \frac{\beta h^2 n_x^2}{8ma^2} = \frac{\beta h^2 (2n_x + 1)}{8ma^2}.$$

At $T = 300K$, $m = 10^{-22}$, and $a = 10cm$, Δ is on the order of $(2n_x + 1) \cdot 10^{-20}$. For these conditions $n_x \approx 10^{10}$. As one can see then the approximation is a very good one.

Thus, q_{trans} becomes,

$$\left(\int_0^{\infty} e^{-\beta h^2 n^2 / 8ma^2} \mathrm{d}n \right)^3 = \left(\frac{2\pi mkT}{h^2} \right)^{3/2} V.$$

V represents the volume or a^3. This result follows that the integrand and lower limit of the integral forms the error function. Another way to come to this result is to use the density of states approximation where the density of transitional energy levels is,

$$\omega(\epsilon)\mathrm{d}\epsilon = \frac{\pi}{4} \left(\frac{8ma^2}{h^2} \right)^{3/2} \epsilon^{1/2}\mathrm{d}\epsilon.$$

Some observations are worth noting. First that the average momentum p is $(mkT)^{1/2}$. This comes from equating the translational energy stated above using an ensemble average and the momentum form of translational energy. From here, we can see that q_{trans} is,

$$\frac{V}{(h/p)^3} = \frac{V}{\Lambda^3}.$$

Λ is simple the DeBroglie wavelength.

4.2.2 Electronic Partition Functions

For the electronic energy we use energy levels so that,

$$q_{elec} = \sum_i \omega_i e^{-\beta \epsilon_i}.$$

We then fix the ground state energy to zero which gives us,

$$q_{elec} = \omega_1 + \sum_{i=2} \omega_i e^{-\beta \Delta \epsilon_{1i}}.$$

The argument in the exponential is usually large since the energy difference is on the order of electron volts. Specific cases like the halogen series require more than just the first term, however. Regardless, the sum converges quickly and energy levels of available states for many atoms and materials have been tabulated. The following is usually enough,

$$q_{elec} = \omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}}.$$

4.2.3 Nuclear Partition Functions

While there are excited nuclear states to access them requires temperatures of $10^{10} K$. Thus only the ground state degeneracy matters and

$$q_{nucl} = \omega_1.$$

For most cases the nuclear partition function is just omitted.

4.3 Thermodynamic Relations with the Ideal Gas

Taking the definition of the Helmholtz energy in terms of the canonical ensemble we have for the monotonic ideal gas,

$$A(N, V, T) = -kT \ln \mathcal{Q}$$

$$= -kT \ln \left[\left(\frac{V}{\Lambda^3} \right)^N (\omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}})^N \frac{1}{N!} \right]$$

$$= -NkT \ln \frac{V}{\Lambda^3} - NkT \ln (\omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}}) + NkT \ln N - NkT$$

$$= NkT \left(-\ln \frac{V}{\Lambda^3} - \ln (\omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}}) + \ln N - 1 \right).$$

The electronic term is small so,

$$A \approx -NkT \left(\ln \frac{V}{\Lambda^3} - \ln N + 1 \right)$$

or

$$A \approx -NkT \left(\ln \frac{V}{N\Lambda^3} + 1 \right).$$

At this point it is helpful to note that,

$$\ln \mathcal{Q} = N \left(\ln \frac{V}{N\Lambda^3}(\omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}}) + 1 \right).$$

Pressure is given by,

$$p = kT \left(\frac{\partial \ln \mathcal{Q}}{\partial T} \right)_{N,V} = \frac{NkT}{V}.$$

Energy is,

$$E = kT^2 \left(\frac{\partial \ln \mathcal{Q}}{\partial T} \right)_{N,V} = \frac{3}{2}NkT + \frac{N\omega_2 \Delta \epsilon_{12} e^{-\beta \Delta \epsilon_{12}}}{q_{elec}}$$

Entropy is,

$$S = \frac{E}{T} - \frac{A}{T}$$

$$S = \frac{3}{2}Nk + \frac{N\omega_2 \Delta \epsilon_{12} e^{-\beta \Delta \epsilon_{12}}}{q_{elec}T} - NkT \left(-\ln \frac{V}{\Lambda^3} - \ln(\omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}}) + \ln N - 1 \right)$$

$$= Nk \left(\frac{5}{2} + \ln \frac{V}{N\Lambda^3} \right) + Nk \ln(\omega_1 + \omega_2 e^{-\beta \Delta \epsilon_{12}}) + \frac{N\omega_2 \Delta \epsilon_{12} e^{-\beta \Delta \epsilon_{12}}}{q_{elec}}$$

$$= Nk \left(\frac{5}{2} + \ln \frac{V}{N\Lambda^3} \right) + S_{elec}.$$

Using,

$$\mu(T,p) = -kT \left(\frac{\partial \ln \mathcal{Q}}{\partial N} \right)_{V,T},$$

one can similarly find the chemical potential. The procedure is the same for all thermodynamic properties.

Chapter 5

Ideal Diatomic Gas

5.1   Introduction

An ideal diatomic gas has two more means of storing energy compared to the monotonic case: vibrational and rotational. Strictly speaking vibrational and rotational modes are coupled and rotational-electronic and are coupled. However, there are cases where this can be ignored and others where reality requires a more nuanced quantum mechanic approach. One important assumption we will be making is the **Born-Oppenheimer approximation**. This simply states that electronic movement occurs on a timescale where nuclei are effectively stationary. Therefore, we can solve the Schrdinger equation for the electrons with stationary nuclei and for nuclei in an electron cloud. The interatomic interactions determined by the electron cloud often is difficult to solve so empirical and semi-empirical potentials like **Moore** or **Leonard-Jones** are often used.

5.2   Rigid Rotor Harmonic Oscillator

For a diatomic atom, we first notice we can define translational movement as that movement with respect to the center of mass, thus the Hamiltonian becomes

$$\mathcal{H} = \mathcal{H}_{trans} + \mathcal{H}_{int},$$

where the second term is the energy from the relative motion of the two masses. This relative motion is clearly either rotational or vibrational. Technically the rotational energy is effected by the current interatomic distance; however, in most cases the changing of the effective radius is not consequential.

Proof If we let $u(r)$ represent the interatomic potential, then a Taylor series around the current distance, R, is,

$$u(r) = u(R) + (r - R) \left(\frac{\partial u}{\partial r}\right)_{r=R} + \frac{1}{2}(r - R)^2 \left(\frac{\partial^2 u}{\partial r^2}\right)_{r=R}$$

$$= u(R) + \frac{1}{2}k(r - R)^2.$$

The second steps substitutes the second derivative for k. The reason we can drop the linear term stems from the fact that in most cases the interatomic distance is very close to the potential minimum. At the minimum, the first derivative goes to zero. Therefore, if k or the second derivative is large then the bond is stiff and distance will not fluctuate much. *Note:* that k is similar to Hooke's constant.

We can then state,

$$\mathcal{H}_{int} = \mathcal{H}_{rot} + \mathcal{H}_{vib}.$$

Of course, the above is equivalent to,

$$q_{int} = q_{rot}q_{vib}.$$

This approximation is known as the rigid rotor harmonic oscillator approximation. The approximation allows vibrational-rotational decoupling.

5.3 Vibrational Partition Function

5.3.1 Energy Levels and Spectra

For a quantum harmonic oscillator, the energies are given by

$$\epsilon_{vib} = h\nu(n + \frac{1}{2}).$$

Note: that the degeneracy of each level is one, so each state is its own level. ν is given by,

$$\nu = \frac{1}{2\pi}\left(\frac{k}{\mu}\right)^{1/2}.$$

The selection rules for vibrational transitions is that transitions can only occur between energy adjacent states or $\Delta n = \pm 1$. Taking the above equation for ϵ_{vib}, and finding the energy difference between adjacent levels gives

$$\Delta\epsilon = \epsilon_{n+1} - \epsilon_n$$
$$= \nu h((n + 3/2) - (n + 1/2)) = \nu h(1)$$

ν then is,

$$\nu = \frac{\Delta\epsilon}{h} = \frac{1}{2\pi}\left(\frac{k}{\mu}\right)^{1/2}.$$

This gives the spectra for vibrational transitions. For there to be radiation, a changing dipole in the molecule must exist. Using this equation, k can be found for different diatomic molecules. *Note,* the vibrational spectra consists of one line. One other quick fact is that for a given electronic state there is an vibrational energy D_o that corresponds to a disassociating molecule.

5.3.2 Deriving the Partition Function

In order to derive the partition function, a zero state energy must be assigned. Two possibilities exist 0 or $h\nu/2$. The book uses the latter and so shall I. The derivation of the vibrational partition function goes as follows,

$$q_{vib} = \sum_n e^{-\beta\epsilon_n}$$
$$= \sum_n e^{-\beta h\nu(n+1/2)} = \sum_n e^{-\beta h\nu n}e^{-\beta h\nu/2}$$
$$= e^{-\beta h\nu/2}\sum_{n=0}^{\infty} e^{-\beta h\nu n}$$
$$= \frac{e^{-\beta h\nu/2}}{1 - e^{-\beta h\nu}} \qquad e^{-\beta h\nu} < 1.$$

The first step follows by definition. The second is just the expansion and simplification of ϵ_{vib}. The third is just taking the constant out of the sum. The

fourth step involves seeing that the sum is a geometric series if $e^{-\beta h\nu} < 1$. The partition function here is exact given the constraint, but we can still take a high temperature limit using an integral. This then is,

$$q_{vib} = e^{-\beta h\nu/2} \int_{n=0}^{\infty} e^{-\beta h\nu n} \mathrm{d}n = \frac{kT}{h\nu}.$$

5.3.3 Properties of Vibrational Mode

We will look at level occupancy and average vibrational energy. The average energy given by the exact partition function is,

$$E_v = NkT^2 \frac{\mathrm{d}\ln q_v}{\mathrm{d}T} = NkT^2 \frac{\mathrm{d}}{\mathrm{d}T}\left(\frac{h\nu}{2kT} - \ln\left(1 - e^{-h\nu/kT}\right)\right)$$

$$= \frac{Nk\Theta}{2} + Nk\Theta\frac{e^{-\Theta/T}}{1 - e^{-\Theta/T}} = Nk\Theta\left(\frac{1}{2} + \frac{1}{e^{\Theta/T} - 1}\right).$$

Where, $\Theta = h\nu/k$. For the high temperature limit, we have,

$$E_v = NkT^2 \frac{\mathrm{d}\ln q_v}{\mathrm{d}T} = NkT^2 \frac{\mathrm{d}}{\mathrm{d}T}\left(\ln\frac{T}{\Theta}\right) = NkT.$$

The specific heat in this case is simply Nk since C_v is the derivative of energy with respect to temperature.

The fraction of molecules in a given energy state are,

$$f_n = \frac{e^{-\beta h\nu(n+1/2)}}{q_{vib}}.$$

Therefore the fraction of molecules in an excited state are,

$$f_{n>0} = \frac{\sum_{n=1}^{\infty} e^{-\beta h\nu(n+1/2)}}{q_{vib}} = 1 - f_0 = \frac{e^{-\beta h\nu/2}}{q_{vib}} = 1 - (1 - e^{-\beta h\nu}) = e^{-\beta h\nu}.$$

The solution comes from the fact that the fraction in a excited state is the inverse of the fraction in the ground state.

5.4 Rotational Partition Function

5.4.1 Energy States and Spectra

For a rigid rotor the energy states and degeneracy are given by,

$$\epsilon_j = \frac{\hbar^2 J(J+1)}{2I}$$
$$\omega_j = 2J + 1.$$

Here, I is the moment of inertia, which is μr_e^2 for the diatomic case. μ is the reduced mass of the molecule.

The transition rules for rotational energy are that the transition must be adjacent like the vibrational modes and the molecule must have a permanent dipole moment. The electromagnetic spectra that comes from the rotational energy is then given by taking the different in adjacent energy levels like before.

$$\Delta\epsilon = \epsilon_{j+1} - \epsilon_j = \frac{\hbar^2}{2I}((J+1)(J+2) - J(J+1)) = \frac{\hbar^2}{I}(J+1)$$

$$\nu = \frac{\Delta\epsilon}{h} = \frac{h}{4\pi^2 I}(J+1)$$

To get the final form, $\hbar$ was expanded to $h/2\pi$. What this says is unlike vibrational energy, rotational energy transitions emit multiple wavelengths. *Note: we will set the rotational ground energy to zero.*

5.4.2 Deriving the Rotational Partition Function

5.4.3 Heteronuclear Derivation

For reasons of wavefunction symmetry under the exchange of particles we must separate the heteronuclear and homonuclear cases. We begin by examining the heteronuclear.

By definition,

$$q_{rot} = \sum_{J=0}^{\infty} (2J+1)e^{-\beta\bar{B}J(J+1)}.$$

The variable $\bar{B}$ is $h/8\pi^2 Ic$ where c is the speed of light. Then, $\epsilon_J = \bar{B}J(J+1)$. We can further define Θ_r to be $\bar{B}/k$ which is a rotational temperature. In order to write the summation in terms of a integral we require that the difference between adjacent values be small. This is fulfilled if $\Theta_r/k(J+1)$ is small which is true at small J. By the time large J are reached the contributions become negligible though, so we can just take the integral.

$$q_{rot} = \int_0^{\infty} (2J+1)e^{-\Theta_r J(J+1)/T}dJ$$

$$= \int_0^{\infty} e^{-\Theta_r J(J+1)/T}d[J(J+1)] = \left[\frac{T}{\Theta_r}e^{-\Theta_r J(J+1)/T}\right]_0^{\infty} = \frac{T}{\Theta_r}$$

$$= \frac{8\pi^2 IkT}{h^2}.$$

The second line recognizes that a change of variable can drastically simplify the integral. For cases where the above is a bad approximation say $\Theta_r > 0.7T$, the summation can be expanded to 4 terms with good accuracy. Of course higher order approximations can be taken.

The two approximation, however, leave an intermediate gap. The solution is to use a Euler-MacLaurin summation formula approximation. This is given

by

$$\sum_{n=a}^{b} f(n) = \int_{a}^{b} f(n)\mathrm{d}n + \frac{1}{2}[f(b)+f(a)] + \sum_{j=1}^{\infty}(-1)^j \frac{B_j}{(2j)!}[f^{(2j-1)}(a) - f^{(2j-1)}(b)],$$

where B_j is the j-th Bernoulli number and the last bracketed term uses the $(2j-1)$-th derivative. This give for the rotational partition function,

$$q_{rot} = \frac{T}{\Theta_r}\left(1 + \frac{1}{3}\left(\frac{\Theta_r}{T}\right) + \frac{1}{15}\left(\frac{\Theta_r}{T}\right)^2 + \frac{1}{315}\left(\frac{\Theta_r}{T}\right)^3 + \cdots\right).$$

5.4.4 Heteronuclear Thermodynamics

The ensemble rotational energy in the high temperature limit is,

$$E_r = NkT^2\left(\frac{\partial \ln q_{rot}}{\partial T}\right) = NkT^2\left(\frac{\partial}{\partial T}\right)(\ln T - \ln \Theta_r) = NkT.$$

This means that C_v like in the vibrational case is Nk. The fraction of molecules in a particular vibrational state can be taken just like the rotational state, but since it is an identical procedure it is not shown here. However, an important feature of rotational energy levels is that the most occupied level at room temperature is an excited state. The most occupied level can be derived by taking the derivative of the occupancy fraction,

$$\frac{\mathrm{d}f_J}{\mathrm{d}J} = \frac{\mathrm{d}}{\mathrm{d}J}\left(\frac{N_J}{N}\right) = \frac{\mathrm{d}}{\mathrm{d}J}\left(\frac{(2J+1)e^{-\Theta_r J(J+1)/T}}{q_{rot}}\right),$$

and setting it equal to zero.

5.4.5 Homonuclear Derivation

Note: The homonuclear case can be approximated by

$$q_{rot,homo} = \frac{q_{rot,heter}}{2}.$$

This deals with the fact that in the homonuclear case a two-fold symmetry exists perpendicular to the interatomic axis. This prevents the over counting of states. This factor of 2 can be generalized to a factor σ which is called the symmetry number and represents the number of indistinguishable rotational states or the number or rotational symmetries.

However, this will not work at low temperatures since homonuclear diatomic molecules must obey certain wavefunction symmetries when an exchange of nuclei is performed. This symmetry depends on the spins of the nuclei. If the spins are integers the wavefunction is symmetric; otherwise, the wavefunction is anti-symmetric with respect to exchange.

Digression on Symmetry Requirements

To determine the individual symmetry contributions of the electronic, translational, vibrational, rotational, and nuclear energies to the wavefunction, we imagine that we take a diatomic molecule and exchange both nuclei along with their electrons. We then flip the electrons again. For molecules where only the ground electronic state is useful and ground state is Σ_g^+, the electronic wavefunction is symmetric to such operations. The translational and vibration wavefunctions are obviously unaffected by such transformations. This only leaves the nuclear and rotational.

By looking at the rotational wavefunctions for hydrogen, one can see that even J are symmetric to the swap, but odd J are anti-symmetric to it. In addition, for hydrogen, the nuclei spins are $\pm 1/2$. This leads to four spin combinations: three are symmetric and one is anti-symmetric. Since the total wavefunction must be anti-symmetric we must pair symmetric rotational wavefunctions with anti-symmetric nuclear wavefunctions and vice versa. The opposite holds true for integer spin molecules.

For a generic system the number of spin states for a nucleus is $2I + 1$. For a diatomic molecule there are $(2I+1)^2$ states then. The number of antisymmetric states can be calculated using $(2I + 1)(2I)/2$. The number of symmetric is just the total minus the number of antisymmetric wavefunctions. This leads to

- Integral spin

  - Odd J - I(2I+1)
  - Even J - (I+1)(2I+1)

- Half-integral spin

  - Odd J - (I+1)(2I+1)
  - Even J - I(2I+1).

One can see that the two are just mirrors of each other. The partition function is then,

$$q_{rot,nucl} = w_{k,even} \sum_{J\ even} (2J + 1)e^{-\Theta_r J(J+1)/T} + w_{k,odd} \sum_{J\ odd} (2J + 1)e^{-\Theta_r J(J+1)/T},$$

where $w_{k,even/odd}$ represents the appropriate spin weighting. The rotation and nuclear partition functions are then coupled. However, at high temperatures,

$$\sum_{J\ odd} \approx \sum_{J\ even} \approx \frac{1}{2}\sum_{J}.$$

This is just one half of the heteronuclear case. Then all the previous equations apply with the adding factor of $1/2$. In this limit, we can separate the partition functions to get,

$$q_{rot} = \frac{T}{2\Theta_r} = \frac{T}{\sigma\Theta_r}$$

$$q_{nucl} = (2I + 1)^2.$$

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This requires that $\Theta_r < 0.2T$. σ is just the symmetry number mentioned previously. When the temperature is not high enough however the coupled partition function must be used. In general, this is only necessary with molecules like hydrogen where Θ_r is appreciable at gas temperatures. The splitting of H_2 into two different types para (parallel spins) and ortho (anti-parallel spins) comes from this weighting. At room temperature H_2 is 25% para and 75% ortho. At lower temperatures though this ratio changes as the summations over even and odd J's are not approximately equivalent.

All thermodynamic quantities can be calculated normally and will either not vary in the high temperature limit (as is the case for energy and specific heat) or will vary by factors of σ. This is also true in the Euler-MacLaurin approximation.

5.5   Total Partition Function

The total partition function then becomes at the high temperature limit of the harmonic-oscillator rigid rotor approximation,

$$q = q_{trans}\ q_{rot}\ q_{vib}\ q_{nucl}\ q_{elec}$$

$$= \left(\frac{2\pi mkT}{h^2}\right)^{3/2} V \frac{8\pi^2 IkT}{\sigma h^2} \frac{e^{-\beta h\nu/2}}{1 - e^{-\beta h\nu}} \omega_{e1} e^{D_e/kT}.$$

This equation ignores the nuclear partition function.

Note: In cases where the ground electronic state is not a Σ state, the electronic and nuclear angular momentum must be coupled. At high temperatures compared to Θ_r, they can still be decoupled, however. This means that for all but low temperatures for molecules like NO the previous approach would work.

Chapter 6

Classical Mechanics

6.1 Formulation of the Classical Partition Function

Quantum mechanics approaches classical mechanics when quantum numbers become sufficiently high. This occurs at relatively high T for a given quantum quantity. This means for many purposes we can begin our derivation from a classical mechanics standpoint and above a certain threshold quantity we achieve accurate results. For instance the translation quantum number at room temperature, $n \approx 10^8$, thus a classical mechanical approximation is very accurate. However, at the same conditions the vibrational quantum number is too low to achieve accurate results using a classical approach.

6.1.1 Molecular Partition Function

We know for a quantum mechanical derivation of molecular partition functions we arrive at

$$q_p = \sum_{j \in J} e^{-\beta \mathcal{H}_p},$$

for a particular separable mode of energy p. Hence, we assume that for a classical system we take the argument of the summation and integrate over accessible energy levels,

$$q_{p,class} = \int_{\mathcal{H}_p} e^{-\beta \mathcal{H}_p} \mathrm{d}\mathcal{H}_p.$$

The whole molecular partition function is then,

$$q_{class} = \int_{\mathcal{H}} e^{-\beta \mathcal{H}} \mathrm{d}\mathcal{H}$$
$$= \int \cdots \int e^{-\beta \mathcal{H}(\vec{p},\vec{q})} \mathrm{d}\vec{p}\mathrm{d}\vec{q},$$

where the second line takes the Hamiltonian to be a function of coordinates $\vec{q}$ and conjugate momenta $\vec{p}$. The dimensionality of the vectors q and p is simply the dimensionality of the molecule or the number of degrees of freedom that must be satisfied to fully describe the molecule.

Translational Partition Function

We know classically the energy of translation is

$$\mathcal{H} = \frac{1}{2m}(p_x^2 + p_y^2 + p_z^2).$$

Then, the translational partition function is,

$$q_{trans} = \int \cdots \int \exp\left(-\frac{\beta(p_x^2 + p_y^2 + p_z^2)}{2m}\right) dp_x dp_y dp_z dx dy dz$$

$$= V \iiint \exp\left(-\frac{\beta(p_x^2 + p_y^2 + p_z^2)}{2m}\right) dp_x dp_y dp_z$$

$$= V \left(\int_{-\infty}^{\infty} e^{-\frac{\beta p^2}{2m}} dp\right)^3$$

$$= V \cdot \left(\sqrt{\frac{2m\pi}{\beta}}\right)^3$$

$$= V(2\pi mkT)^{3/2}.$$

V comes from the lack of spacial dependence so the volume integral evaluates to the volume. The last two lines come from the error function. When we compare this to the classical approximation made in chapter 5,

$$V\left(\frac{2\pi mkT}{h^2}\right)^{3/2},$$

we find that a factor of h^{-3} is missing.

Rotational Partition Function

For a classical rigid rotor the Hamiltonian is,

$$\mathcal{H} = \frac{1}{2I}\left(p_\theta^2 + \frac{p_\phi^2}{\sin^2\theta}\right).$$

The partition function is then,

$$q_{rot} = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \int_0^{2\pi} \int_0^{\pi} e^{-\beta\mathcal{H}} dp_\theta dp_\phi d\phi d\theta$$

$$= 2\pi \int_{-\infty}^{\infty} e^{-\beta p_\theta^2/2I} dp_\theta \int_{-\infty}^{\infty} \int_0^{\pi} e^{-\beta p_\phi/\sin^2\theta} dp_\phi d\theta.$$

Note that the first argument of the integral is another error function so we can further simplify to,

$$q_{rot} = 2\pi \cdot \sqrt{2\pi IkT} \int_{-\infty}^{\infty} \int_0^{\pi} e^{-\beta p_\phi/\sin^2\theta} dp_\phi d\theta.$$

From here the last two integrals must be carried out. I am not currently sure how to do them though. The answer becomes,

$$q_{rot} = 8\pi^2 IkT,$$

however. Again notice that a factor of h^{-2} is missing from the partition function.

Vibrational Partition Function

For a harmonic oscillator the Hamiltonian is,

$$\mathcal{H} = \frac{p^2}{2\mu} + \frac{k}{2}x^2,$$

where the first term captures the kinetic energy and the second potential. Then, q_{vib} is,

$$
\begin{aligned}
q_{vib} &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \exp\left(-\beta\frac{p^2}{2\mu} + \frac{k}{2}x^2\right) \mathrm{d}x\mathrm{d}p \\
&= \int_{-\infty}^{\infty} e^{-\beta p^2/2\mu}\mathrm{d}p \int_{-\infty}^{\infty} e^{-\beta k x^2/2}\mathrm{d}x \\
&= (2\mu\pi k_b T)^{1/2} \left(\frac{2\pi k_b T}{k}\right)^{1/2} \\
&= \frac{k_b T}{\nu}, \text{ where } \nu = \left(\frac{k}{4\pi^2\mu}\right)^{1/2}.
\end{aligned}
$$

Note once again this integral was done using the error function and that this is off by a factor of h.

Adding Planck's Constant

For each of the molecular partition functions we have examined thus far, the final result is off by a factor of h^s where s is the dimension of $\vec{p}$ or $\vec{q}$. From this we hypothesis,

$$q = \sum_{j \in J} e^{-\beta\mathcal{H}} \rightarrow \frac{1}{h^s} \int \cdots \int e^{-\beta\mathcal{H}} \prod_{j=1}^{s} \mathrm{d}p_j \mathrm{d}q_j.$$

6.1.2   Ensemble Partition Function

Taking the high temperature limit of the ensemble partition function for noniteracting molecules,

$$\mathcal{Q} = \frac{q^N}{N!},$$

we apply the classical molecular partition function to arrive at,

$$
\begin{aligned}
\mathcal{Q} &= \frac{1}{N!} \prod_{j=1}^{N} \left(\frac{1}{h^s} \int \cdots \int e^{-\beta\mathcal{H}_j} \prod_{i=1}^{s} \mathrm{d}p_{ji}\mathrm{d}q_{ji}\right). \\
&= \frac{1}{h^{sN} N!} \prod_{j=1}^{N} \left(\int \cdots \int e^{-\beta\mathcal{H}_j} \prod_{i=1}^{s} \mathrm{d}p_{ji}\mathrm{d}q_{ji}\right).
\end{aligned}
$$

To produce the standard formulation of the classical partition function, we re-label the indices $ji \to k$ so that $k = (j-1)s + i$. That is $k \in [0, s]$ represents the former indices $1i$ or the position and momenta of the first particle. Doing this allows us to simplify the equation as,

$$Q = \frac{1}{h^{sN} N!} \int \cdots \int e^{-\beta \sum_j \mathcal{H}_j} \prod_{k=1}^{sN} \mathrm{d}p_k \mathrm{d}q_k$$

$$= \frac{1}{h^{sN} N!} \int \cdots \int e^{-\beta \mathcal{H}} \prod_{k=1}^{sN} \mathrm{d}p_k \mathrm{d}q_k.$$

The Hamiltonian in the last line is the many body Hamiltonian not to be confused with a molecular or sub-molecular Hamiltonian. This is the correct form of the classical limit even though we assumed noninteracting molecules and merely postulated the form from an observation of the quantum derivation.

Examining a monatomic gas with a Hamiltonian of,

$$\mathcal{H} = \frac{1}{2m} \sum_{j=1}^{N} (p_x^2 + p_y^2 + p_z^2) + U(\vec{r}_1, \vec{r}_2, \ldots, \vec{r}_N),$$

where $\vec{r}_i$ is the position vector of particle i. Then taking the momentum integrals (using the error function), we get

$$Q_{class} = \frac{1}{N!} \left(\frac{2\pi m k T}{h^2} \right)^{3N/2} Z_N,$$

where Z_N is

$$Z_N = \int_V e^{-\beta U(\vec{r}_1, \vec{r}_2, \ldots, \vec{r}_N)} \mathrm{d}\vec{r}_1 \mathrm{d}\vec{r}_2 \cdots \mathrm{d}\vec{r}_N.$$

Z_N has a special name, *the classical configuration integral*. This integral encompasses all the intermolecular interactions present in a system.

6.1.3 Hybrid Partition Functions

When some degrees of freedom can be treated classically and others must be treated quantumly, if the Hamiltonian is separable in those degrees of freedom, so is the partition function into a classical and quantum part.

$$\mathcal{H} = \mathcal{H}_{class} + \mathcal{H}_{quant}$$
$$Q = Q_{class} Q_{quant}$$
$$= \frac{Q_{quant}}{h^{sN} N!} \int e^{-\beta \mathcal{H}_{class}} \mathrm{d}\vec{p}_{class} \mathrm{d}\vec{q}_{class}.$$

Similarly, if the molecular Hamiltonian is separable we have,

$$\mathcal{H} = \mathcal{H}_{class} + \mathcal{H}_{quant}$$
$$q = q_{class} q_{quant}$$
$$= \frac{q_{quant}}{h^s} \int e^{-\beta \mathcal{H}_{class}} \mathrm{d}\vec{p}_{class} \mathrm{d}\vec{q}_{class}.$$

6.2 Phase Space and the Liouville Equation

6.2.1 Defining of Phase Space and Density of States

We take a system of N particles, where each particle has s degrees of freedom, that is s coordinates can uniquely define the particle. To complete the description of the classical system we also need to define s conjugate momenta that correspond to the s coordinates. The system is then uniquely defined by $l = sN$ coordinates and conjugate momenta. The total system then consists of $2l$ values.

Having an uniquely defined classical system, we now have motivation to define a phase space that projects the classical system into a new dynamical space which we will call the phase space. We project the classical system to a subset of $\mathbb{R}^{2l}$ which works to a prism of $2l$ dimensional Euclidean space where the extent of each dimension is the physical limits of the system. Each system at a specified time is then a point in this space. This space is the *phase space* and those points are *phase points*.

For a given phase point, its past and future is uniquely defined by the laws of motion and the current position in phase space. We postulate that trajectories are continuous which for a Hamiltonian system is true since

$$\dot{q}_j = \frac{\partial \mathcal{H}}{\partial p_j} \text{ and } \dot{p}_j = -\frac{\partial \mathcal{H}}{\partial q_j} \text{ for } j \in 2l.$$

Thus, trajectories in this frame work are uniquely defined just like phase points. This means that the evolution of q and p can be defined as some function of t.

We now show the preservation of density on a particular subset of the phase space. Using the microcanonical ensemble conservation of N, V, E, we define a mapping $M : P \to \Omega$, where P is the phase space and Ω is the subset defined by $\{(x_i)_1^{2l} : (x_i)_1^{2l} \sim N, V, E\}$ that is the set of points that correspond to a particular N, V, E. This is an energy "surface" of a specified number and volume. We let all $X \in \Omega$ be occupied. That is all states are equally likely, this is the classical meaning of the *postulate of equal a priori probabilities*. This also suggests that each subdomain of Ω is equally dense since if this was not the case then particular regions of the microcanonical ensemble would be favored. However, it is important to note, that this **does not imply uniform density in** P.

6.2.2 Deriving the Louiville Equation

To begin with deriving the Louivile equation, we first create a function f that gives the infinitesimal density around a phase point. We then have

$$\rho = f(p(t), q(t)) \mathrm{d}p \mathrm{d}q.$$

If we integrate f over the entire domain we retrieve $\mathcal{A}$,

$$\mathcal{A} = \int \cdots \int f(p(t), q(t)) \mathrm{d}p \mathrm{d}q.$$

Any property averaged over all systems is

$$\overline{\phi} = \frac{\int \cdots \int \phi(p(t), q(t)) f(p(t), q(t)) \mathrm{d}p \mathrm{d}q}{\int \cdots \int f(p(t), q(t)) \mathrm{d}p \mathrm{d}q}$$

$$\overline{\phi} = \frac{\int \cdots \int \phi(p(t), q(t)) f(p(t), q(t)) \mathrm{d}p \mathrm{d}q}{\mathcal{A}}.$$

Gibb's postulate was originally formulated by saying this average corresponds to thermodynamical observed variables.

We now turn our attention to the flux of phase points into and out of an arbitrary volume about a point X. We will prefix quantities in this volume with δ. The instantaneous number of phase points in δV is given by,

$$\delta N = f(p(t), q(t)) \delta p_1 \delta q_1 \delta p_2 \cdots \delta q_l \delta p_l.$$

This is just taking f at X and smearing it over the volume δV. To find the total flux we have to consider $2l$ dimensions. However, each dimension is equivalent in phase space, i.e. momenta and position are not treated differently. Therefore, we look at the dimension q_1 and generalize the results. The net flux in the q_1 dimension is given by the difference of the individual fluxes through the orthogonal hyper-surfaces defined by q_1 and $q_1 + \delta q_1$. The flux through first face is given by,

$$f(q_1, \ldots, q_l, p_1, \ldots, p_l) \dot{q}_1 \delta p_1 \cdots \delta p_l \delta q_1 \cdots \delta q_l.$$

Mathematically, this is taking the number of particles per δq_1 or a linear density as can be seen by,

$$\frac{\delta N}{\delta q_1} = f(q_1, \ldots, q_l, p_1, \ldots, p_l) \delta p_1 \cdots \delta p_l \delta q_2 \cdots \delta q_l,$$

and multiplying it by the "velocity" in the normal direction to the face $\partial q_1 / \partial t$. This is analogous to a flux taken in real space. Likewise, the flux at the face defined by $q_1 + \delta q_1$ is given by,

$$f(q_1 + \delta q_1, \ldots, q_l, p_1, \ldots, p_l) \dot{q}_1 (q_1 + \delta q_1, q_2, \ldots, q_l, p_1, \ldots, p_l) \delta p_1 \cdots \delta p_l \delta q_2 \cdots \delta q_l.$$

Taking a Taylor expansion around q_1 we find the flux at δq_1 is approximately,

$$\left(f_{q_1} + \left(\frac{\partial f}{\partial q_1} \right)_{q_1} \delta q_1 \right) \left((\dot{q}_1)_{q_1} + \left(\frac{\partial \dot{q}_1}{\partial q_1} \right)_{q_1} \right) \delta p_1 \cdots \delta p_l \delta q_2 \cdots \delta q_l$$

where we have discarded all higher order terms. If we expand out the terms in the parenthesis we have,

$$f_{q_1} (\dot{q}_1)_{q_1} + \left(\frac{\partial f}{\partial q_1} \right)_{q_1} (\dot{q}_1)_{q_1} + f_{q_1} \left(\frac{\partial \dot{q}_1}{\partial q_1} \right)_{q_1} \delta q_1 + \left(\frac{\partial f}{\partial q_1} \right)_{q_1} \left(\frac{\partial \dot{q}_1}{\partial q_1} \right)_{q_1} \delta q_1^2.$$

Noting again that we will ignore all high order terms, δq_1^2 is sufficiently small to ignore it. Then if we take the difference between the fluxes at the two faces we have,

$$-\left(f_{q_1} \left(\frac{\partial \dot{q}_1}{\partial q_1} \right)_{q_1} + (\dot{q}_1)_{q_1} \left(\frac{\partial f}{\partial q_1} \right)_{q_1} \right) \delta p_1 \cdots \delta p_l \delta q_1 \cdots \delta q_l.$$

Likewise, for p_1

$$-\left(f_{p_1} \left(\frac{\partial \dot{p}_1}{\partial p_1} \right)_{p_1} + (\dot{p}_1)_{p_1} \left(\frac{\partial f}{\partial p_1} \right)_{p_1} \right) \delta p_1 \cdots \delta p_l \delta p_1 \cdots \delta p_l.$$

Here, we recall that there are a total of $2l$ dimensions to analyze the flux over l for momentum and l for position. The total change in number in a instance of time is,

$$-\sum_{i=1}^{l} \left(f_{p_1} \left(\frac{\partial \dot{p}_1}{\partial p_1} \right)_{p_1} + (\dot{p}_1)_{p_1} \left(\frac{\partial f}{\partial p_1} \right)_{p_1} + f_{q_1} \left(\frac{\partial \dot{q}_1}{\partial q_1} \right)_{q_1} + (\dot{q}_1)_{q_1} \left(\frac{\partial f}{\partial q_1} \right)_{q_1} \right) \delta p_1 \cdots \delta p_l \delta p_1 \cdots \delta p_l.$$

The expression is equivalent to,

$$\frac{\partial \delta N}{\partial t}.$$

If we divide by the volume element we get the density f previously defined.

6.2.3 The Conservation of Extension and Incompressibility

6.3 Equipartition of Energy