

SpaghettLens

Gravitational Lens Modelling

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Motivation

- Gravitational Lenses (GL) hard to find
- Let volunteers help find them: SpaceWarps
- But post processing? SpaghettiLens

Outline

1 Theory

2 Results

3 Outlook

Fermat's Principle

Fermat's Principle¹

Rays of light traverse the path of stationary optical length with respect to variations of the path.

Fermat's Principle

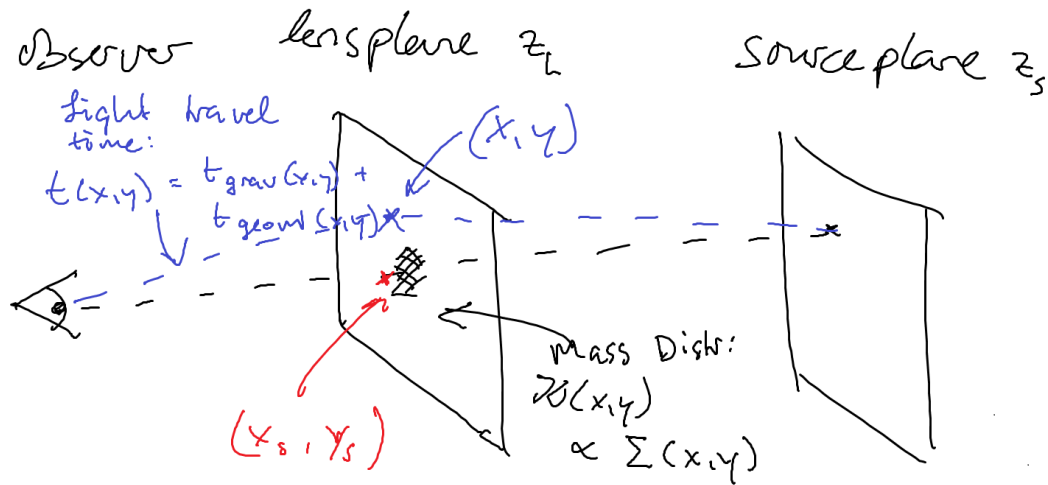
Time t for path X :

$$t[X] = \frac{1}{c} \int_{t_1}^{t_2} n(\vec{x}(t)) \sqrt{1 + \left(\frac{d\vec{x}(t')}{dt'} \right)^2} dt'$$

Path X where t stationary.

¹Ghatak, Ajoy (2009), Optics

Setup



light travel time

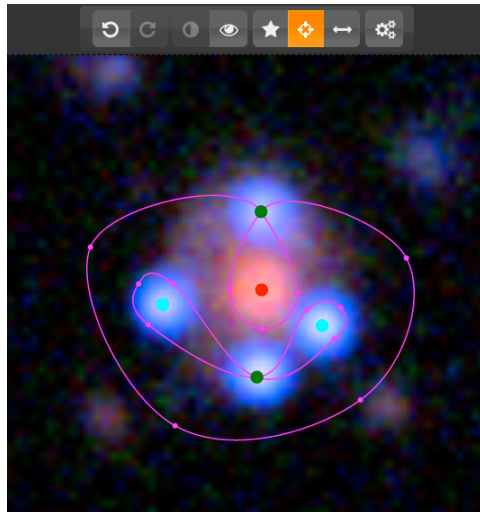
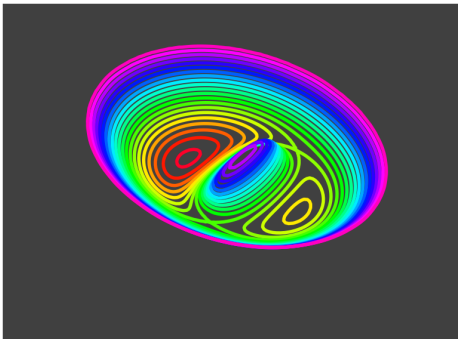
Fermat's Principle

$$t(x, y) = t_{\text{geom}} + t_{\text{grav}} \quad (1)$$

$$t_{\text{geom}} \propto (x - x_s)^2 + (y - y_s)^2 \quad (2)$$

$$t_{\text{grav}} = \langle t_{\text{grav}}(x_o, y_o) \rangle + (1 + z_L) \frac{2G}{c^3} M(x_{\bullet}, y_{\bullet}) \quad (3)$$

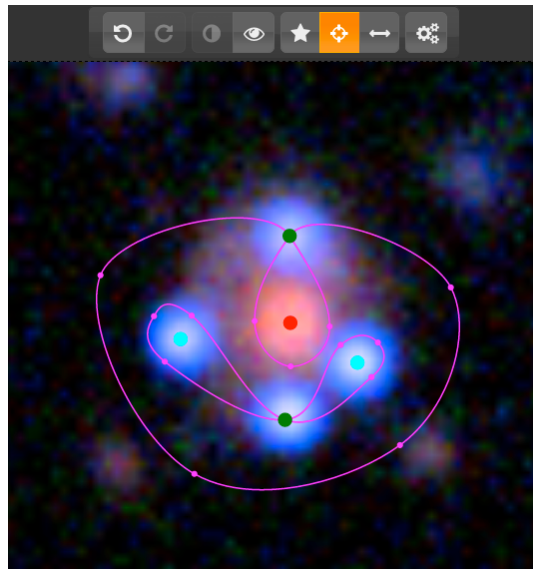
Arrival Time Surface



SpaghettiLens

- Extremal Points (Images)
- Self Intersecting Contour Lines

<http://labs.spacewarps.org/spaghetti/>



SpaceWarps Setup & Results

- CFHT Legacy Survey
- about 11 million classifications
- 29 promising (59 total) new lens candidates

SpaceWarps: II New Gravitational Lens Candidates...²

SPACE WARPS: II. New Gravitational Lens Candidates from the CFHTLS Discovered through Citizen Science

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²arxiv:1504.05587

SpaghettiLens Results: Tests of Performance

- use simulated lenses
- let volunteers model them
- recover Einstein Radii
- Volunteers perform well!

Gravitational Lens Modelling in a Citizen Science Context³

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(MN L^AT_EX style file v2.2)

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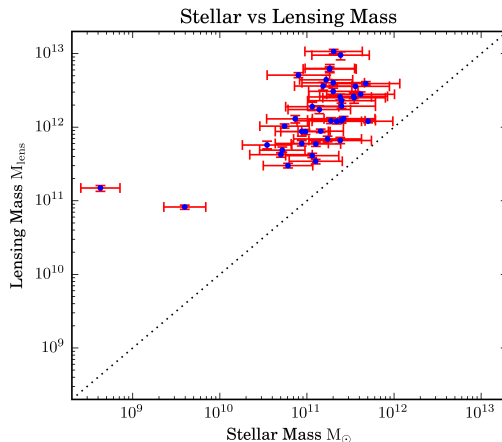
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SpaghettiLens Results: Stellar vs Lensing Mass

- lensing mass against the stellar mass of the candidate lens galaxies
- stellar mass fraction of order 20 percent
- with decreasing trend for the most massive galaxies
- expected for early type galaxies
- outliers? Maybe non-lenses (not yet spectroscopically confirmed)



Outlook

We are currently working on:

- fit parametrized models to the free-form mass distributions (Lucy Oswald)
- determination of photometric red shifts
- estimate stellar populations (using galfit, SExtractor)

Questions?

Questions?

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Appendix

$$A_t = A_{\text{geom}} + A_{\text{grav}} \quad (4)$$

$$A_{\text{geom}} = \frac{1}{2} (x^2 + y^2) \quad (5)$$

$$\nabla^2 A_{\text{grav}}(x, y) = -2\kappa(x, y) \quad (6)$$

$$A = \frac{cD_L}{(1+z_L)^2} \frac{D_{LS}}{D_S} \times t \quad (7)$$

$$\kappa(x, y) = \frac{4\pi G}{c^2} \frac{D_L}{1+z_L} \frac{D_{LS}}{D_S} \times \Sigma(x, y) \quad (8)$$