Ellis drainhole: The Ellis drainhole is the earliest known somblete mathematical model of traversable wormhole. It is static, scherically symmetric solution of the Einstein vaccum field equations augmented by inclusion of a sodor field p minimally coupled to the geometry of shartime with coupling bolarity of specific to the orthodox bolarity (-ve instead of +ve).

$$\mathcal{R}_{\mu\nu} - \frac{1}{2} \mathcal{R}_{\mu\nu} = -2 \left(\phi_{,\mu} \phi_{,\nu} - \frac{1}{2} \phi^{,\kappa} \phi_{,\kappa} g_{\mu\nu} \right) \longrightarrow \mathbf{D}$$

where
$$\Box \phi = 0 \longrightarrow \emptyset$$

The Einstein's field egns above is similar are equivalent to $R_{\mu\nu} = -2 \phi, \mu \phi, \nu \longrightarrow 3$

$$T_{\mu\nu}$$
: $R = R_{\mu}^{\mu} = -2 \phi^{\prime \mu} \phi_{,\mu} \longrightarrow \Theta$

- Elis dramhole line element: From now on we will were the Ricci tensor with +01 -ve polarilez.

$$c^2 \mathcal{A} \tau^2 = c^2 \mathcal{A} t^2 - \left[\mathcal{A} p - \mathcal{G} f(p) c \mathcal{A} t \right]^2 - r^2(p) \mathcal{A} \Omega^2 \longrightarrow \mathcal{G}$$

$$f(P) = m \left(\frac{1}{1 - e^{-(2m/n)}} \right)$$

$$ro(p) = \sqrt{(p-m)^2 + a^2} e^{(m/n)} \phi \longrightarrow \exists$$

$$\phi = \alpha(p) = \frac{n}{a} \left[\frac{\pi}{2} - \tan^{-1} \left(\frac{p - m}{a} \right) \right] \longrightarrow \emptyset$$

Note: t and p are not temporal and radial Schwarzschill coordinates.

Schwarzschild coordinates are related to Elis coordinates in the following way:

$$T = t + \frac{1}{c} \int \frac{f(\rho)}{1 - f^2(\rho)} d\rho$$

$$0 = 0$$

$$\phi = \phi$$

$$c^{2} dt^{2} = c^{2} dt^{2} - \left[dp + f(p) c dt \right]^{2} - n^{2} d\Omega^{2}$$

$$= c^{2} \left(1 - f^{2}(p) \right) dt^{2} + 2 f(p) c dt dp + dp^{2}$$

$$- n^{2} d\Omega^{2}$$

$$T = \mathcal{L} + \frac{1}{c} \int \frac{f(\rho)}{1 - f^{2}(\rho)} d\rho$$

$$\Rightarrow \partial T = \partial \mathcal{L} + \frac{1}{c} \frac{f(\rho)}{1 - f^{2}(\rho)} d\rho$$

$$\Rightarrow \partial T^{2} = \partial \mathcal{L}^{2} + \frac{2}{c} \frac{f(\rho)}{1 - f^{2}(\rho)} \partial \mathcal{L} \partial \rho + \frac{1}{c^{2}} \frac{f^{2}(\rho)}{(1 - f^{2}(\rho))^{2}} \partial \rho^{2}$$

$$\Rightarrow \partial^{2}(1 - f^{2}(\rho)) \partial T^{2} = c^{2}(1 - f^{2}(\rho)) \partial \mathcal{L}^{2} + 2f(\rho)c \partial \mathcal{L} \partial \rho$$

$$+ \frac{f^{2}(\rho)}{1 - f^{2}(\rho)} \partial \rho^{2}$$

$$+ \frac{f^{2}(\rho)}{1 - f^{2}(\rho)} \partial \rho^{2}$$

$$\Rightarrow \partial^{2}(1 - f^{2}(\rho)) \partial T^{2} = c^{2}(1 - f^{2}(\rho)) \partial \rho^{2}$$

$$\Rightarrow c^{2}(1-f^{2}(\rho))d\ell^{2}+2f(\rho)cd\ell d\rho-d\rho^{2}$$

$$= c^{2}(1-f^{2}(\rho))d\tau^{2}-f^{2}(\rho)-d\rho^{2}-d\rho^{2}$$

$$= c^{2}(1-f^{2}(\rho))d\tau^{2}-f^{2}(\rho)-1+f^{2}(\rho)-1+f^{2}(\rho)-1+f^{2}(\rho)$$

$$= c^{2}(1-f^{2}(\rho))d\tau^{2}+\frac{-f^{2}(\rho)-1+f^{2}(\rho)}{1-f^{2}(\rho)}d\rho^{2}$$

$$= c^{2}(1-f^{2}(\rho))d\tau^{2}-\frac{d\rho^{2}}{1-f^{2}(\rho)}$$

:.
$$c^2 d \tau^2 = c^2 (1 - f^2(\rho)) d \tau^2 - \frac{d \rho^2}{1 - f^2(\rho)} - \rho^2 d \Omega^2 \rightarrow 10$$

Note: For the case of Schwarzschild metric
$$f(\rho) = \int \frac{2MG}{c^2 \rho} \quad \text{and} \quad r = ro(\rho) = \rho.$$
 Therefore ρ and r are equivalent for Schwarzschild metric.

For Ellis drainhole metric,
$$r(\rho) = J(\rho - m)^2 + a^2 e^{(m/n)} \phi$$

Derivation / Ancartz: Eleis's original intention was to replace
the Schwardchild metric will a nonvingular
model for a gravitating elementary particle. The Schwardsthild

motric is found by inserting the most general scherically symmetric and static metric tensor into the vaccum field equations and solving away. On his effort to get hid of the singularity by adding a minimally reached mightine fairly massless scalar field into the history. Excis involved the same ansatz (expressed in Ellis's coordinates) into the field eque given above, to derive the Ellis's drainhole.

On Ellis's orondinates, the most general scherically symmetric and static line element is

 $c^2d\tau^2 = c^2dt^2 - \left[d\rho^2 - f(\rho) \circ dt\right]^2 - t_{\rho}^2 + r_{\rho}^2 d\Omega^2$

Now, because of the spherical symmetry, the soular amonto

$$\phi = \phi \alpha(\rho)$$
 = 03

$$g_{\mu\nu} = \begin{pmatrix} 1 - f^{2}(\rho) & f(\rho) & 0 & 0 \\ f(\rho) & -1 & 0 & 0 \\ 0 & 0 & -h^{2}(\rho) & 0 \\ 0 & 0 & 0 & -h^{2}(\rho) & 0 \end{pmatrix}$$

$$\Gamma_{01}^{1} = \Gamma_{10}^{1} = -f^{2}f'$$
; $\Gamma_{01}^{0} = \Gamma_{10}^{0} = -ff'$

$$\Gamma^{2} = \frac{2}{33} - \sin\theta \cos\theta ; \Gamma^{1}_{33} = (f^{2} - 1) nn' \sin^{2}\theta$$

$$\Gamma_{22}^{0} = fnn'; \Gamma_{13}^{3} = \Gamma_{31}^{2} = \Gamma_{21}^{2} = r'$$

$$\int_{33}^{0} = f n n' \sin^2 \theta ; \int_{22}^{1} = (f^2 - 1) n n'$$

$$R_{\mu\nu} = \partial_{\mu} \Gamma^{\rho}_{\nu\rho} - \partial_{\rho} \Gamma^{\rho}_{\mu\nu} + \Gamma^{\sigma}_{\mu\rho} \Gamma^{\rho}_{\sigma\nu} - \Gamma^{\sigma}_{\mu\nu} \Gamma^{\rho}_{\sigma\rho}$$

$$R_{00} = \partial_{\nu} \Gamma^{\rho}_{\nu\rho} - \partial_{\rho} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\sigma}_{\nu\rho} \Gamma^{\rho}_{\sigma\rho} - \Gamma^{\sigma}_{\nu\rho} \Gamma^{\rho}_{\sigma\rho}$$

$$= -\partial_{\nu} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho}$$

$$= -\partial_{\nu} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho}$$

$$= -\partial_{\nu} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} + \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho}$$

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$$= -\partial_{\nu} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho}$$

$$= -\partial_{\nu} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}_{\nu\rho} - \Gamma^{\rho}_{\nu\rho} \Gamma^{\rho}$$

$$= \pi (f')^{2} (1-f^{2}) + ff'' (1-f^{2}) - ff' \cdot 2ff'$$

$$+ 2(ff')^{2} (1-f^{2}) + 2(f^{2}f')^{2} + 2ff' (1-f^{2}) \frac{p'}{p}$$

$$= (f')^{2} (1-f^{2}) + ff'' (1-f^{2}) - 2ff')^{2} (1-f^{2})$$

$$+ 2(ff')^{2} (1-f^{2}) + 2ff' (1-f^{2}) \frac{n'}{p}$$

$$= (1-f^{2}) \left[(f')^{2} + ff'' - 2(ff')^{2} + 2(ff')^{2} + 2ff' \frac{n'}{p} \right]$$

$$= (1-f^{2}) \left[(f')^{2} + ff'' + 2ff' \frac{n'}{p} \right]$$

$$= (1-f^{2}) \frac{(f'n)^{2} + ff'' + 2ff' \frac{n'}{p}}{n^{2}}$$

$$= (1-f^{2}) \frac{(f'n)^{2} + ff'' + 2ff' \frac{n'}{p}}{n^{2}}$$

$$= (1-f^{2}) \sqrt{\frac{f'(f^{2})^{2}}{n^{2}}}$$

$$= (1-f^{2}) \sqrt{\frac{f'(f^{2})^{2}}{n^{2}}}$$

$$= (1-f^{2}) \sqrt{\frac{f'(f^{2})^{2}}{n^{2}}}$$

$$|R_{00} = (1 - f^2) \nabla^2 \left(\frac{f^2}{2}\right) \qquad \Rightarrow (4)$$

$$\begin{array}{c}
R_{11}^{0} = \partial_{1} \Gamma_{1p}^{\rho} - \partial_{p} \Gamma_{11}^{\rho} + \Gamma_{1p}^{\sigma} \Gamma_{\sigma1}^{\rho} - \Gamma_{11}^{\sigma} \Gamma_{p}^{\rho} \\
= \partial_{1} \Gamma_{10}^{0} + \partial_{1} \Gamma_{11}^{1} + \partial_{1} \Gamma_{12}^{2} + \partial_{1} \Gamma_{13}^{3} \\
- \partial_{1} \Gamma_{11}^{1} + \Gamma_{1p}^{0} \Gamma_{01}^{\rho} + \Gamma_{1p}^{1} \Gamma_{12}^{\rho} \\
+ \Gamma_{1p}^{2} \Gamma_{21}^{\rho} + \Gamma_{1p}^{2} \Gamma_{01}^{\rho} + \Gamma_{1p}^{\sigma} \Gamma_{0p}^{\rho} \\
- \Gamma_{1p}^{1} \Gamma_{1p}^{\rho} \\
= \partial_{1} \Gamma_{10}^{0} + \partial_{1} \Gamma_{12}^{2} + \partial_{1} \Gamma_{3}^{3} + \Gamma_{10}^{0} \Gamma_{01}^{0} \\
+ \Gamma_{11}^{0} \Gamma_{01}^{1} + \Gamma_{10}^{2} \Gamma_{11}^{3} + \Gamma_{11}^{0} \Gamma_{11}^{0} \\
+ \Gamma_{11}^{0} \Gamma_{01}^{1} + \Gamma_{10}^{2} \Gamma_{11}^{0} \Gamma_{11}^{0} + \Gamma_{11}^{1} \Gamma_{11}^{1} \\
+ \Gamma_{11}^{3} \Gamma_{3}^{3} - \Gamma_{11}^{0} \Gamma_{0}^{0} - \Gamma_{11}^{0} \Gamma_{01}^{1} \\
- \Gamma_{11}^{4} \Gamma_{10}^{0} - \Gamma_{11}^{1} \Gamma_{11}^{1} \Gamma_{11}^{1} \\
- \Gamma_{11}^{4} \Gamma_{12}^{3} \Gamma_{12}^{3} + \Gamma_{12}^{2} \Gamma_{11}^{2} \Gamma_{11}^{2} \Gamma_{12}^{2}
\end{array}$$

$$||R_{11}|| = -\nabla^2\left(\frac{1}{2}f^2\right) + 2\frac{p''}{p} \longrightarrow 15$$

$$R_{22} = \frac{1}{2} \left(\left[1 - f^2 \right] (n^2)' \right)' - 1 \longrightarrow \mathbb{Z}$$

$$R_{33} = -\sin^2 \theta \left(\left(\frac{1}{2} \left(\left[1 - f^2 \right] (n^2)' \right)' - 1 \right) \right) \longrightarrow \mathbb{Z}$$

$$R_{01} = R_{10} = f \nabla^{2} \left(\frac{1}{2} f^{2} \right)$$

$$\Box \phi = 0 \Rightarrow g^{\mu\nu} \nabla_{\mu} \nabla_{\nu} \phi = g^{\mu\nu} \nabla_{\mu} \partial_{\nu} \phi = 0$$

$$\Rightarrow g^{0} \nabla_{\nu} g^{\mu\nu} (\partial_{\mu} \partial_{\nu} \phi - \Gamma_{\mu\nu}^{\sigma} \partial_{\sigma} \phi) = 0$$

$$\Rightarrow g^{0} (\partial_{0} \partial_{\nu} \phi - \Gamma_{00}^{\sigma} \partial_{\sigma} \phi) + g^{0} (\partial_{0} \partial_{1} \phi - \Gamma_{01}^{\sigma} \partial_{\sigma} \phi)$$

$$+ g^{10} (\partial_{1} \partial_{0} \phi - \Gamma_{10}^{\sigma} \partial_{\sigma} \phi) + g^{11} (\partial_{1} \partial_{1} \phi - \Gamma_{11}^{\sigma} \partial_{\sigma} \phi)$$

$$+ g^{12} (\partial_{2} \partial_{\nu} \phi - \Gamma_{22}^{\sigma} \partial_{\sigma} \phi) + g^{33} (\partial_{3} \partial_{3} \phi - \Gamma_{33}^{\sigma} \partial_{\sigma} \phi)$$

$$\Rightarrow \Box \alpha = -g^{00} \Gamma^{\frac{1}{00}} \partial_{1} \alpha - g^{01} \Gamma^{\frac{1}{01}} \partial_{1} \alpha \\
-g^{10} \Gamma^{\frac{1}{10}} \partial_{1} \alpha + g^{11} \partial_{1} \partial_{1} \alpha - g^{11} \Gamma^{\frac{1}{11}} \partial_{1} \alpha \\
-g^{22} \Gamma^{\frac{1}{22}} \partial_{1} \alpha - g^{33} \Gamma^{\frac{1}{33}} \partial_{1} \alpha$$

$$\Rightarrow -ff'(f^{2}-1) \alpha' + ff^{2}f' \alpha' + ff^{2}f'\alpha'$$

$$= \begin{bmatrix} 1-f^{2} \end{bmatrix} \alpha'' + \begin{bmatrix} 1-f^{2} \end{bmatrix} ff'\alpha'$$

$$+ \frac{1}{p^{2}} (f^{2}-1) n n' \alpha' + \frac{1}{n^{2} \sin^{2}\theta} (f^{2}-1) n n' \sin^{2}\theta \alpha'$$

$$= \begin{bmatrix} 1-f^{2} \end{bmatrix} \alpha'' + \underbrace{ff'(f^{2}-1)}_{p^{2}} + \underbrace{ff'(f^{2$$

From
$$(9, -\frac{1}{2}, -\frac{1}{2}) = 0 \Rightarrow (f^2 - 1) = 0$$

From $(20, -\frac{1}{2}) = 0 \Rightarrow (f^2 - 1) = 0$

From $(20, -\frac{1}{2}) = 0 \Rightarrow (f^2 - 1) = 0$

$$\Rightarrow n^2 \left(-\frac{f^2}{2} \right)' = 0 \Rightarrow n^2 \left(-\frac{f^2}{2} \right)' = 0$$

$$\Rightarrow n^2 \left(1 - f^2 \right)' = 0 \Rightarrow 0$$

From (21),
$$\frac{ro''}{r} = \infty (\alpha')^2$$

$$\left(\left(1-f^2\right)\left(\frac{n^2}{2}\right)'\right)'=1$$

$$\Rightarrow \left(1-f^2\right)\left(\frac{n^2}{2}\right)' = \rho \overline{\bullet} m \qquad \Rightarrow 6\overline{\bullet}$$

Grant
$$20/2 + 27$$
, — (Integrating)
 $\left(\frac{n^2}{2}\right)(1-f^2)' + (1-f^2)\left(\frac{n^2}{2}\right)' = \rho \overline{a} m \overline{b} m$

$$\Rightarrow \left(\frac{n^2}{2}\left(1-f^2\right)\right)'=\rho$$

$$\Rightarrow \frac{h^2}{2} (1 - f^2) = \frac{\rho^2}{2} + \frac{e}{2} \quad \left(\text{Integrating} \right)$$

$$\Rightarrow n^2(1-f^2) = \rho^2 + \delta C \qquad (28)$$

Ofrom
$$(25)$$
,
$$\alpha' = \frac{n}{(f^2 - 1)^{n^2}} = -\frac{n}{\rho^2 + C} \longrightarrow (25, 2)$$

From
$$(27)$$
,
$$(1-f^2)(n^2)' = 2(\rho - m)$$

$$\Rightarrow (1-f^2) n n' = (\rho - m) \longrightarrow (29)$$

$$\frac{n'}{n} = \frac{\rho - m}{\rho^2 + C}$$
 \Rightarrow 30

Then
$$(\alpha')^2 = \frac{p''}{p} = \frac{-(r')^2 + rn'' + (r')^2}{p^2} = + \left(\frac{p'}{p}\right)^4 + \left(\frac{r'}{p}\right)^2$$

$$= + \left(\frac{p - m}{p^2 + c}\right)' + \left(\frac{p - m}{p^2 + c}\right)^2$$

$$= + \frac{(p^2 + c) - (p - m) 2p + (p - m)^2}{(p^2 + c)^2}$$

$$= + \frac{p^2 + c - 2p^2 + 2mp + p^2 + 2mp}{(p^2 + c)^2}$$

$$\Rightarrow \left(\alpha'\right)^2 = + \frac{C - m^2}{\left(\rho^2 + C\right)^2} \longrightarrow 31$$

$$\alpha' = -\frac{n}{\rho^2 \bar{\bullet} m^2 + n^2} \longrightarrow 33$$

$$\frac{p'}{h} = \frac{\rho - m}{\rho^2 \bar{s} m^2 + n^2}$$

$$\Rightarrow (\log r)' = \frac{\rho - m}{\rho^2 + m^2 - n^2}$$

$$\Rightarrow \log r = \frac{\rho - m}{\rho^2 + m^2 - n^2}$$

$$\frac{p'}{p} - \frac{m}{n} \alpha' = \frac{\rho - m}{\rho^2 + m^2 + n^2} + \frac{m}{\rho^2 + m^2 + n^2}$$

$$= \frac{\rho}{\rho^2 + m^2 + n^2}$$

$$\Rightarrow \left(\ln r - \frac{m}{n} \alpha \right)' = \frac{\rho}{\rho^2 + m^2 + n^2}$$

$$\Rightarrow \ln n - \frac{m}{n} \alpha = \frac{1}{2} \ln \left| p^2 + m^2 + n^2 \right| + constant$$

$$\Rightarrow then ln ro = \frac{1}{2} ln \left| \rho^2 + m^2 + n^2 \right| + \frac{m}{n} \propto$$

Thre we have absorbed the constant in α because as $\phi^{\alpha} = (\alpha(\rho) + \text{const})^{\alpha}$ does not change the situation as in Einstein's equ derivatives of ϕ abbeaus, not any ϕ term

$$\Rightarrow p^{2}(p) = \left| p^{2} \neq m^{2} + n^{2} \right| e^{\frac{2m}{n}} \propto$$

Throw (28), -

$$n^{2}(1-f^{2}) = \rho^{2} \pm m^{2} + n^{2}$$

$$-2 \frac{m}{n} \propto$$

$$\Rightarrow f^{2}(p) = 1 - sgn(\rho^{2} \pm m^{2} + n^{2}) e \longrightarrow 35$$

Note: The learndary condition for the field ϕ is, $\lim_{\rho \to \infty} \phi = \lim_{\rho \to \infty} c(\rho) = 0$ resulting the spacetime to be asymptotically flat.

There are 3 different cases
$$m^2 < m^2$$
 $m^2 = n^2$
 $m^2 > m^2$

For only one rease we will get a drainhole. From the line element we can see that ro(p) gives the radius of the sequence of 2-sheres along the proordinate needed to assemble the live blace hypersurface of the metric. We are looking for a drainhole, which is basically a wormhole with a gravity field that is basically a wormhole with a gravity field that bushes test footicle through the wormhole in one direction. So, to find the rase that contains a drainhole, we must look for a radius function that is rontinuous, never zero, reaches one nouzero minimum and increases monotonically reaches one nouzero minimum and increases monotonically. When moving in either direction away from the minimum. When moving in either direction away from the minimum. Othis corresponds to the double furnel at shape, of this rorresponds to the double furnel at shape, of mormholes. On all three rosses, the integration constant is selected in such a way so that pin a(p) = 0

Case 1:
$$m^2 < m^2$$
; $a = \sqrt{m^2 - m^2}$

$$\alpha' = \frac{-m}{\rho^2 \mp m^2 + m^2}$$

$$\Rightarrow \alpha(\rho) = \int \alpha' d\rho = \int \frac{-n}{\rho^2 - a^2} d\rho = \frac{n}{2a} \ln \frac{(\rho + a)}{(\rho - a)}$$

$$m^2(\rho) = |\rho^2 - a^2| \frac{\rho + a}{\rho - a} = \frac{|\rho + a|}{|\rho - a|} = \frac{m}{|\rho - a|}$$

$$f^{2}(\rho) = 1 - sgn\left(\left|\rho^{2} - a^{2}\right|\right) \left|\frac{\rho}{\rho - a}\right|^{\frac{m}{a}}\right)$$

ro2 (P) relearly possesses both Lero foint and a singularity

Case 2:
$$m^2 = n^2$$

$$\alpha' = \frac{-n}{\rho^2}$$

$$\Rightarrow \alpha = \frac{n}{\rho}$$

$$n^2(\rho) = \rho^2 e^{\rho}$$

$$f^2(\rho) = 1 - sgn \rho^2 e^{\rho} = 1 - e^{\rho}$$

$$n^2(\rho) \text{ has singularity.}$$

$$\frac{\sigma_{ase 3}}{\sigma^{2}}; \quad m^{2} > m^{2}; \quad a = \sqrt{m^{2} \sigma m^{2}}$$

$$\frac{\sigma' = -\frac{n}{\rho^{2} + a^{2}}}{\rho^{2} + a^{2}}$$

$$\Rightarrow \quad \alpha(\rho) = \frac{n}{a} \left[\frac{\pi}{2} - tan^{-1} \left(\frac{\rho}{a}\right)\right]$$

$$\frac{2m}{n^{2}} \left[\frac{\pi}{2} - tan^{-1} \left(\frac{\rho}{a}\right)\right]$$

$$f^{2}(\rho) = (\rho^{2} + a^{2})e$$

$$-\frac{2m}{a} \left[\frac{\pi}{2} - tan^{-1} \left(\frac{\rho}{a}\right)\right]$$

$$f^{2}(\rho) = 1 - e$$

 $p^2(\rho)$ is continuous, never zero, reaches one nonzero minimum at $\rho=m$ and increases monotonically when moving in either direction away from this minimum. It is the case that contains drainhole.

$$T = \pounds + \frac{1}{e} \int \frac{f(\rho)}{1 - f^2(\rho)} d\rho$$

So, T has temporal rure. When m=0 then $f(\rho)=0$ So, m is like gravity here. When m=0, —

$$ds^{2} = e^{2}dT^{2} - dp^{2} - (p^{2} + a^{2}) d\Omega^{2}$$

$$ds^{2} = c^{2} \left(1 - f^{2}(\rho)\right) dT^{2} - \frac{d\rho^{2}}{\left(1 - f^{2}(\rho)\right)} - r^{2}(\rho) d\Omega^{2}$$

$$= c^{2} \left(1 - f^{2}(\rho)\right) \left(dT^{2} - \frac{d\rho^{2}}{\left(1 - f^{2}(\rho)\right)^{2}}\right) - r^{2}(\rho) d\Omega^{2}$$

$$= c^{2}(1-f^{2}(p))(dT^{2}-dp^{2})-p^{2}d\Omega^{2}$$

where
$$dp^{2} = \frac{dp^{2}}{(1 - f^{2}(p))^{2}}$$

$$dp^{2} = e^{\frac{4m}{a} \left[\frac{\pi}{2} - tan^{-1} \left(\frac{p}{a} \right) \right]} dp^{2}$$

$$\Rightarrow dp = e^{\frac{2m}{a} \left(\frac{\pi}{2} - tan^{-1} \left(\frac{p}{a} \right) \right)} dp$$

$$ds^2 = c^2 \left(\frac{\pi}{a} \left(\frac{1}{a}\right)^2 \left(\frac{\rho}{a}\right)^2\right) \left(\rho T^2 d\rho_*^2\right) - r^2 d\Omega^2$$

$$dU dV = dT^{2} - dp_{*}^{2}$$

$$= (dT + dp_{*})(dT - dp_{*})$$

$$\frac{\partial \mathbf{T}}{\partial \rho} = e^{\frac{4m}{\alpha} \left[\frac{\pi}{2} - \tan^{-1} \left(\frac{\rho}{\alpha} \right) \right]} d\rho = -\frac{\partial \mathbf{V}}{\partial \rho}$$

From this we can get UV kenrose diagram.

