

4.5. Quantum Field Theory in Rindler Space

Let's consider 2-dim Minkowski Space with metric
(3.8) & (3.10) i.e.,

$$ds^2 = d\bar{u} d\bar{v} = dt^2 - dx^2 \quad (4.66)$$

Coordinate transform²

$$\begin{cases} t = a^{-1} e^{a\xi} \sinh a\eta & (4.67) \\ x = a^{-1} e^{a\xi} \cosh a\eta & (4.68) \end{cases}$$

$a = \text{constant} > 0$ and $-\infty < \eta, \xi < \infty$ or equivalently

$$\bar{u} = -a^{-1} e^{-a\eta} \quad (4.69)$$

$$\bar{v} = a^{-1} e^{a\eta} \quad (4.70)$$

where $u = \eta - \xi$ and $v = \eta + \xi$, then (4.66)

becomes

$$ds^2 = e^{2a\xi} d\bar{u} d\bar{v} = e^{2a\xi} (d\eta^2 - d\xi^2)$$

The coordinates (η, ξ) cover only a quadrant of

Rindler Space, namely the wedge $x > |t|$ shown in fig. 15. Lines of constant η are straight ($x \propto t$) while lines of constant ξ are hyperbolae.

$$x^2 - t^2 = a^{-2} e^{2a\xi} = \text{constant} \quad (4.72)$$

They therefore represent the world lines of uniformly accelerated observers treated in § 3.3. Comparison of (3.62) with (4.72) shows that

$$a e^{-a\xi} = a^{-1} = \text{proper acceleration} \quad (4.73)$$

Thus, lines of large positive ξ (far from $x=t=0$) represent weakly accelerated observers, while the hyperbolae that closely approach $x=t=0$ have large $-ve \xi$ and hence a higher proper acceleration. All the

hyperbolae are asymptotically to the null rays $\bar{u} = 0, \bar{v} = 0$
 (or $\bar{u} = \infty; \bar{v} = -\infty$), which means that the accelerated
 observers approach the speed of light as $\eta \rightarrow \pm \infty$. These
 observer's proper time τ is related to ξ and η by

$$\tau = e^{a\xi} \eta \quad (4.74)$$

Let's consider the quantization of massless scalar field ϕ in
 2-d Minkowski spacetime. The wave eqn.

$$\square \phi = \left(\frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} \right) \phi \equiv \frac{\partial^2 \phi}{\partial \bar{u} \partial \bar{v}} = 0 \quad (4.75)$$

posses standard orthonormal modes soln.

$$\bar{u}_k = (4\pi\omega)^{-1/2} e^{i(\vec{k} \cdot \vec{x} - \omega t)} \quad (4.76)$$

where $\omega = |\vec{k}| > 0$ and $\vec{k} \in (-\infty, \infty)$.

$$\mathcal{L}_{\partial_t} \bar{u}_k = -i\omega \bar{u}_k \quad (4.77)$$

The modes with $k > 0$ consist of right moving wave

$$(4\pi\omega)^{-1/2} e^{-i\omega \bar{u}} \quad (4.78)$$

left moving

$$\bar{u}_k = (4\pi\omega)^{-1/2} e^{-i\omega\bar{v}} \quad (4.79)$$

space under C.T. $g_{\mu\nu} \rightarrow e^{-2a\xi} g_{\mu\nu}$
 The metric (4.71) is conformal to the whole of Minkowski,
 reduces to $d\eta^2 - d\xi^2$ with $-\infty < \eta, \xi < \infty$.

$$e^{2a\xi} \square \phi = \left(\frac{\partial^2}{\partial \eta^2} - \frac{\partial^2}{\partial \xi^2} \right) \phi = \frac{\partial^2 \phi}{\partial u \partial v} = 0 \quad (4.80)$$

the mode solns, —

$$u_k = (4\pi\omega)^{-1/2} e^{i(k\xi \pm \omega\eta)} \quad (4.81)$$

$$\omega = |k| > 0, \quad -\infty < k < \infty$$

$$\mathcal{L}_{\partial_\eta} u_k = -i\omega u_k \quad (4.82)$$

$$\phi = \sum_{k=-\infty}^{\infty} (a_k \bar{u}_k + a_k^+ \bar{u}_k^*) \quad (4.85)$$

$$\phi = \sum_{k=-\infty}^{\infty} (b_k^{(1)} L_{u_k} + b_k^{(1)+} L_{u_k}^* + b_k^{(2)} u_k + b_k^{(2)+} u_k^*) \quad (4.86)$$

$$a_k |0_M\rangle = 0 \quad (4.87)$$

$$b_k^{(1)} |0_R\rangle = b_k^{(2)} |0_R\rangle = 0 \quad (4.88)$$