8 18 Ch -5

- Reconstruct of selected as inage as whele -) Restoration - modelly the degradition Applying mucise process - recover - Kerbiation techaques i) Spatial domain ii) frequency domain

Adelitive Noise Image bluer i Degradation

Model of mage Degradation / Resto eation:

Model of mage Degradation / Resto eation: f(n,y)) Degradition) Restoration)

(n,y)

Noir

(n,y)

Restoration)

Restoration $\int_{1n}^{n} f(n,y) = h(n,y) + f(n,y) + h(n,y) + h(n,y)$ Let H & linear & position is variant process

Important Noise probability density for (PDF) Noise - Image alquistion / transmission - Spatial populies of noise of spatial characteristics of noise + fig nopeibles of noise of freg contents of noise i) Gaussian noise [Normal noise model] $1(2) = \frac{1}{\sqrt{2}} \left(\frac{3}{\sqrt{2}}\right)$ $\frac{1}{\sqrt{2}} \left(\frac{3}{\sqrt{2}}\right)$ 2 - geay level

A - mean ang of 2

- Stol deviation Je-1 le 140 2 # due to electronic Cinuit rose, sensor noise, T) variance 20/. J valus ((u-t), (u+t)) 15/. Jalues [le-20, ll+20] high temp. $l(3) = \frac{32}{6}(3-a)e^{-(3-a)^{2/6}}$ $= \frac{32}{6}(3-a)e^{-(3-a)^{2/6}}$ ii) Rayleigh Noik $\frac{1}{220}$ $\frac{220}{0.601}$ M= a+ /1/4 A used in eage image's digital convers

111) Edang Gama Noise $P(2) = 5a^{b} 2^{b-1}$ (b-1)!the laser imaging

e - a2

e

j 270 k = a(b-1)(b-1) k = a(b-1)(b-1)(b-1)U = 5/a $\int_{0}^{\infty} = b/a^{2}$ It used is laser imaging iv) Exportabal roise if billing = exporential le = 1/a

--= 1/a

--= 1/a

---= 1/a

---= 1/a y Vijam noik a \le 2 \le b

Show ia 1/a - - - 2 1(2) 5 to $H: a \xrightarrow{+b} 2$ $= (b+a)^{2}$ $= (12)^{2}$ # landon no. generation

vi) Salt le pepper noise [Impulstroise] \langle \langl - light dot I glay level b glay level a - deste dot -> unipla Pa=0 or 16=0 Pa & Pp) salt & pippe garales shot & spike noise -ve inpules - black (sepper) point eve inpules - while (salt) point 8 6,18 a = 0 (s buch) 6 = 258 (white)

Kuidic noise -> Electrical en electronechanical interprener ~ Image a equisition - Spahally dependent noise # Estimation: - Inspection of fourier spectrum

- Automated analysis - knowledge - gregal

becation of frequency components

treation of frequency components

therefore systemare available - study

characteristics of cystem noise
capture a set of images - "Plat" (unifoundy illuminable)

capture a set of images - "Plat" (unifoundy illuminable)

capture a set of images ships

calculation - mean an variance of gray

level.

skip (subinab) - C - Stip (subinege) - S $\mathcal{U} = \underbrace{2i}_{2i} \underbrace{2i}_{5} \underbrace{(2i)}_{2} - \underbrace{(3)}_{2i}$ LT = \(\lambda \left(2i) - \left(2i) - \left(2) \)

2i = \quad \text{gay levels alves} \(\rho \) \\
\(\rho \) \(\rho

-) Mistogram shape - don't filt match - if the slope is Gaussian mean l'vaiance Helprasion of inage in presence of roise only Spatial filling g(x,y) = f(my) + h(x,y)Legladation - only due to noise (additive) f(x,y) = f(x,y) + h(x,y) $g(n,y) = f(n,y) + \mathcal{N}(n,y)$ $f(u,v) + \mathcal{N}(u,v)$ I. Mean films i) Auhanatic Mean files -) Simplest mean files 2001 f(n,y)= I & g(s,t) - (1)

f(n,y)= In (s,t) & Sny

Sny -1 set of wordinates is a rectangular

substrage window of size man

contract of the size of the - Aug of corrupted image g(n,y) in the area I defined by Sny I smoothers local variation - noise is reduced due to blureing

ii) Geometric mean flue:

f(n,y) = TI

g(s,t) (s,t) (s,t) (mn)

- product of pixels in the arbitrage

whom window - revised to power /mn

- smoothing similar to authoretic mean

film

- loce less inage details

in) Contra harmonic mean film: f(2,y) = {5, \in 1 \in 5, \in 1 \in 5, \in 1} (S,t) E Sny

(S,t) Q-ve : eliminates salt noise -1 cannot do both 1 inultaneously A Reduces to animalic mean files

Order Shahishics filler Rushorahion in presence of noise somly 1. Order statistics filter: Spahal filhrig. - Response ordering (renting) the pixels

I depend by the ranking results

Meddar Ither: f(n,y) = median $\frac{2}{3}g(s,t)$ -(i)

replaces the value of pixel by

redian of geny levels

sexuelled noise reduction with less

blevering i) Meddar fitter: reflective is preserve inputer noise.

i) Min & Max files f(h,y) = max & g(s,t) - ii $(s,t) \in S_{hy}$ (pepper noise)

f(n,y)=min 2g(s,t)) q(ii) Salt some (s,t)ESmy mas - finding heightest point min - finding darkest point iii) Mid point film: - computes the orid pint b/w man & -> combines order statistic Daveraging - works best - landonly distributed is) Alpha -teinmed mean fleu $\hat{f}(xy) = \frac{1}{mn-d} \left(\frac{1}{s,t} \right) \in \mathcal{S}_{ny}$ When \$=0 -> Reduced to Achmetic fifther Remove highest d/2 pixels & lowest d/2 pixels -> Remaining pixels (mn-d) d = mn-1 -s median filher Ther calus - us eful in much tiple Types of noin reduction

and the tony # Adaphire Filkes i) Adaptive, Local roise reduction mean à variante neasure of ang.

measure of ang.

contrast

ang gray level files - based on 4 quantities a) g(r, y) realise of roisy inage (r, y) b) of source of noise country fin, y) to form g(x; y)

y m, should mean of the pixels in Sny

of 2 should variance of the pixels in Sny $f(n,y) = g(n,y) - I_n$ $f(g(n,y) - m_L)$ i) Adaphine median f(y)median flu works well mly when
the inpulse arise is not large i'e la l's
is less man 0.2 Aderphive median - can hardle injulse noise

of layer value

- pleuve détaits sonoohig non impulse roise Imm = min value of glay level in Sny
Zman = man
zmed = median I Let Zmin Smar = mar allowed size of Sny e Algorithm works in 2 levels Level A: A1 = 2 med - Zmin A2 = 2 med - 2 mes if A, >0 & A2 20, go b level B else is wease rindow size if window \(\xi\) Somes upeat level A

else O/A Zmy Level B: B1 = 2my - 2min
B2 = 2my - 2min if \$170 &B2 <0 joll 2my Advantages - remove salt paper noise

- provide smoothing - realiste distribution

Perioder noise reduction virg peg. domain i) Band reject flows;

Afterwate a band of frequencies above the origin of the F. T Ideal

SI; $D(u,v) < D_0 - \omega/2$ H(u,v) = h > 0; $D_0 - \omega/2$ $\leq D(u,v) \leq D_0 + \omega$ which I; $D(u,v) > D_0 + \omega/2$ which I; $D(u,v) > D_0 + \omega/2$ Notice for the certical function of the certical function of the certical $\omega \rightarrow 0$. If $\omega \rightarrow 0$. Windh of the band

Do - rapid al wrter/wt of freg.

Sinilarly Ruthersonth: $H(u,v) = \frac{1}{17} \int D(u,v) \omega \int_{0}^{2\pi} du \int F^{\mu\nu}$ Gaussian $IP(u,v) = 1 - 2 \int D(u,v) \omega \int_{0}^{2\pi} du \int F^{\mu\nu}$ L Gaussian

ii) Bandpass filker -> opposite operation of a bandrefect Hbp (u,v) = 1- Hbr (u,v) radies Do Lenhes at (40, 40) with Symmetry at (-40, Vo) is $H(u,v) = 50 \text{ j.} D_{1}(u,v) \leq D_{2}(u,v) \leq D_{3}$ or $D_{2}(u,v) \leq D_{3}$ where $D_{1}(u,v) = \begin{cases} u-M/2 - u_{0} \\ 1 - v_{0} \end{cases}$ $\Delta D_{2}(W,v) = (u-M/2+40)^{2} + (v-N/2+40)^{2}/2$ Butterwork $H(u,v) = \frac{1}{1+\left(\frac{D_0^{-1}}{D_1(u,v)}\right)^n} - \frac{1}{2}$

Gamsian $H(u,v) = 1-e^{-V_{\perp} \int D_{\ell}(u,v)} \frac{D_{z}(u,v)}{D_{r}^{2}} \frac{D_{z}(u,v)}{D_{z}^{2}} \frac{D_{z}^{2}}{D_{z}^{2}} \frac{D_{z}^{2}}{D_{z}^$