

Overview of activities

Day 1

- Introduction, overview
- Obtaining and exploring LAT data
- Inspecting the data: count map



Day 2

- Basics of modeling
- Likelihood analysis of a blazar
- Create a SED
- Produce a light curve

Structure of this talk

see jupiter notebook

- Basic theory of likelihood analysis
- Likelihood fit → Characterize spectra of a source
- obtain spectral energy distribution (SED)
- o build a light curve

Fundamental questions of science statistics

Model fitting: given this model, what parameters best fit my data?

Model selection: given two potential models, which better describes my data?

Model fitting

given this model, what parameters best fit my data?

Maximum likelihood approach

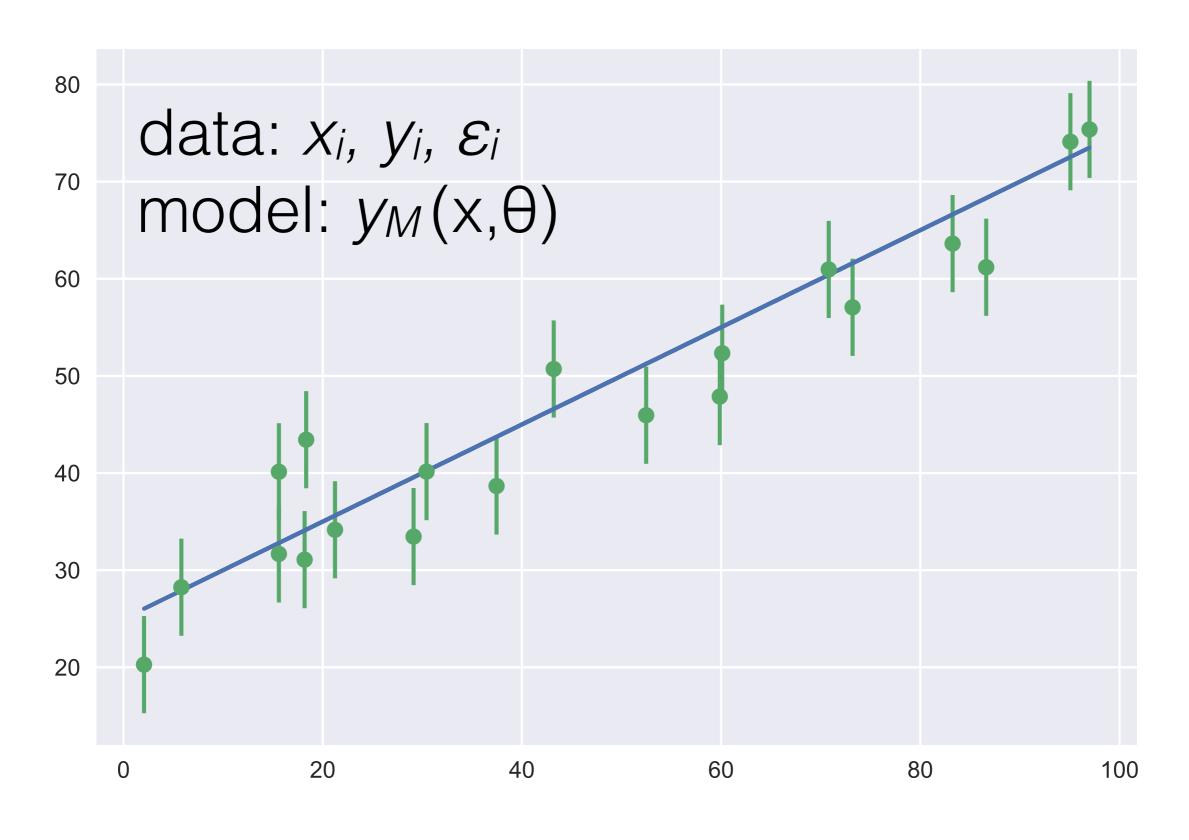
Frequentist

P(data | scientific model)

$$P(D \mid \theta)$$

likelihood

How to write the likelihood function for the problem below?



Probability for a single measurement

$$P_i = \frac{1}{\sqrt{2\pi\epsilon_i^2}} e^{-\frac{(y_i - y_M)^2}{2\epsilon_i^2}}$$

Likelihood:
$$\mathcal{L} = \sum_i P_i = \sum_i \frac{1}{\sqrt{2\pi\epsilon_i^2}} e^{-\frac{(y_i - y_M)^2}{2\epsilon_i^2}}$$

Log-likelihood:

$$\log \mathcal{L} = -\frac{1}{2} \sum_{i=1}^{N} \left\{ \log(2\pi\varepsilon_i^2) + \frac{\left[y_i - y_M(x_i; \theta)\right]^2}{\varepsilon_i^2} \right\}$$

What about Fermi LAT data?

Likelihood for Poisson distribution

$$\mathcal{L} = \prod_{i} P_i = \prod_{i} \frac{m_i^{n_i} e^{-m_i}}{n_i!}$$

Binned likelihood

$$\mathcal{L} = \prod_{i} e^{-m_i} \prod_{i} \frac{m_i^{n_i}}{n_i!} = e^{-N_{\text{pred}}} \prod_{i} \frac{m_i^{n_i}}{n_i!}$$

For bin sizes infinitesimally small, $n_i=1$

$$\mathcal{L} = e^{-N_{\text{pred}}} \prod_{i} m_{i}$$

Log-likelihood

$$\log \mathcal{L} = \sum_{i} \log m_i - N_{\text{pred}}$$

We need to maximize this likelihood function numerically

$$\log \mathcal{L} = \sum_{i} \log m_i - N_{ ext{pred}}$$
 total predicted number of counts $dE \, d\mathbf{p} \, dt$ function $dE \, d\mathbf{p} \, dt$ source model $dE \, d\mathbf{p} \, dt$ sources + other sources + stuff

Model selection

given two potential models, which better describes my data?

Is there a source in a given position of the sky?

maximum likelihood: model

without an additional source

model with additional

source

 $TS = -2\log\left(rac{\mathcal{L}_{ ext{max},0}}{\mathcal{L}_{ ext{max},1}}
ight)$ (the "null hypothesis")

 $TS = 100 \Rightarrow \approx 10\sigma$

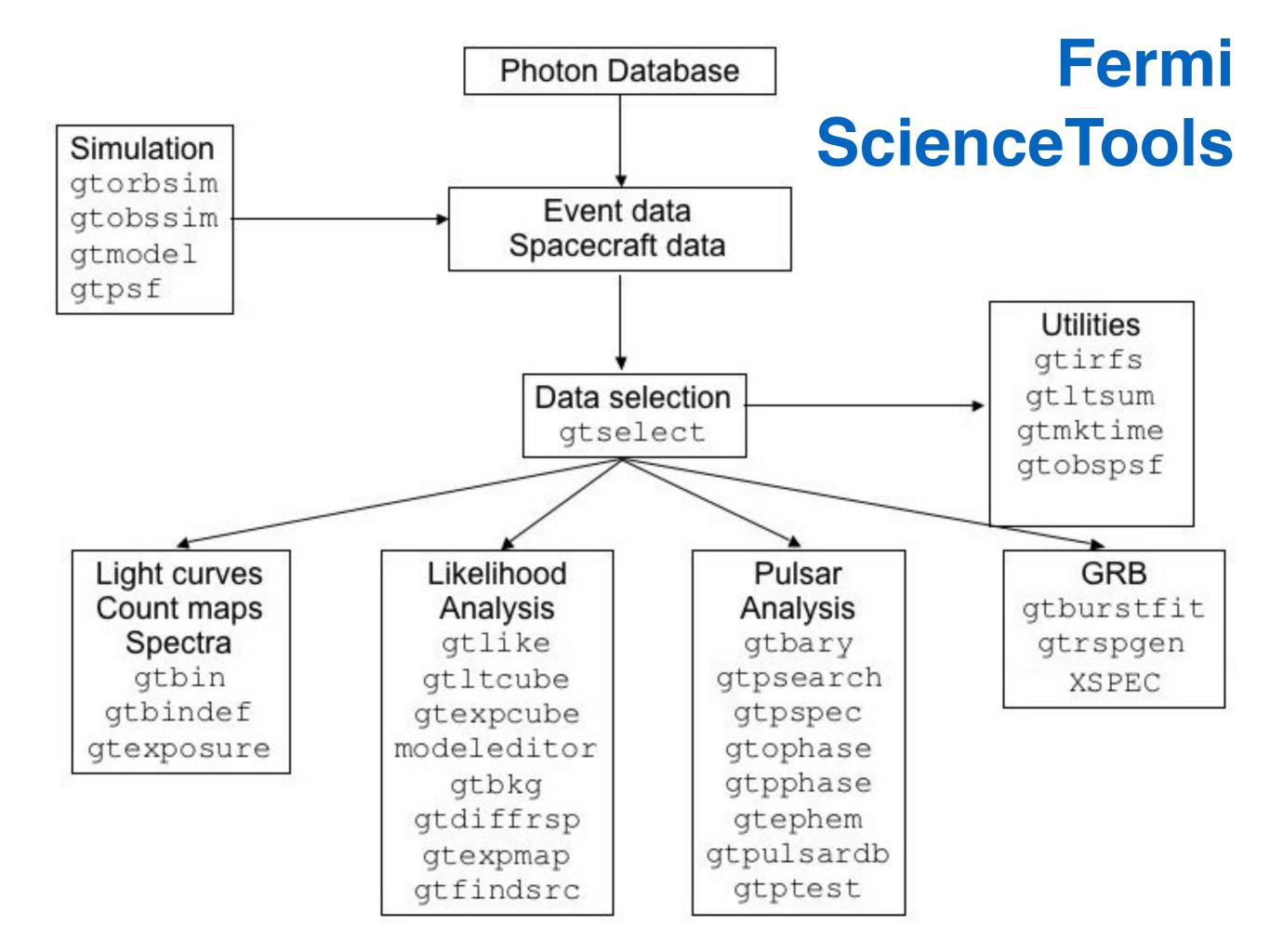
Rule of thumb:

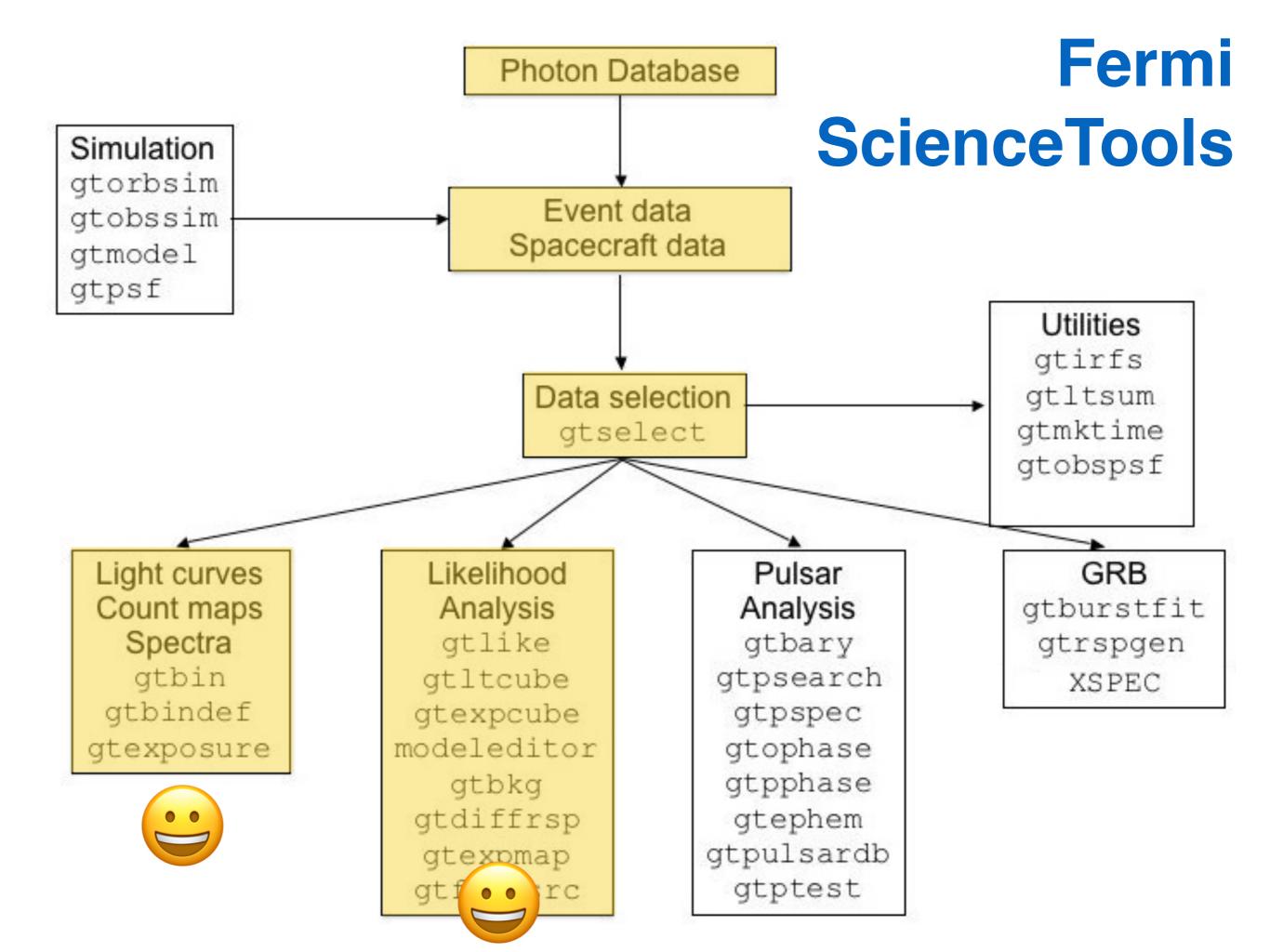
 $TS = 25 \Rightarrow \approx 5\sigma$

Hands-on activity

Steps for modeling of Fermi LAT data

- 1 Define your source model
- 2 Select data + ROI cuts gtselect
- 3 Select good time intervals (GTI) gtmktime
- 4 Bin data, create a count map gtbin
- 5 Compute useful quantities: live time cube, binned exposure cube gtltcube, gtexpcube2
- 6 Maximize likelihood numerically, get initial estimate of parameters gtlike
- 7 Optimize fit: improve parameter estimate





Python wrappers make analysis of Fermi data much easier: Enrico and fermipy

We will use Enrico:

http://enrico.readthedocs.io/en/latest/



https://github.com/rsnemmen/Fermi-LAT-tutorial



Please do not download large files during the tutorial or the WIFI network will *overload*

We will distribute the software and data you need via USB sticks





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Director's cut

(cf separate file)