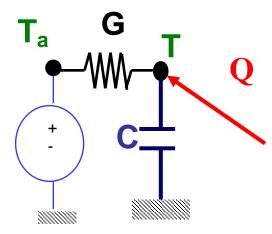
Partie 1-Solution

1)Biot number Bi = $h*D/\lambda$

$$Bi = 7.81 \cdot 10^{-5}$$

- 2) When the Biot number is very small in front of 1, the object can be considered as isothermal
- 3) In a nodal representation, the fuse can be represented by a single node T.



 $C = \rho$ c Volume = ρ c s L with s = $3.14*(D/2)^2$ C= $7.07\ 10^{-3}$ J/K

G = h S with S=3.14*D*L $G = 3.14 \cdot 10^{-4} \text{W/K}$

 $Q=R_{\rm e}~I^2$ with $R_{\rm e}=\rho_{\rm e}~L/s$ Q=1.22 W

4) Nodal equation

$$C\frac{dT}{dt} = G(T_a - T) + Q$$

5)
$$\frac{dT}{dt} = 0$$
, then $0 = G(T_a - T_{eq}) + Q$, and $T_{eq} = T_a + \frac{Q}{G}$
Teq = 3910 °C

6)
$$\tau = C/G$$

 $\tau = 22.5 \text{ s}$

7) Solution

$$T = T_a + \frac{Q}{G} \left(1 - \exp\left(-\frac{t}{\tau}\right) \right)$$

8)
$$\frac{t_f}{\tau} = Ln \left(\frac{1}{1 - \frac{G}{Q} \left(T_f - T_a \right)} \right)$$

 $t_f = 1.26 \ s$

The time to get the fusion is too long to protect efficiently!

- Radiation enhances the transfer, which means that the cooling of the fuse is more intense
- 10) As a consequence, the time to reach the steady state is shorter. In fact the fusion occurs at a time slightly shorter than without radiation

Solution de la Partie rayonnement

First exercice

1) Solar absorptivity

$$\varepsilon_1 = \alpha_1 = 0.9$$
 $\varepsilon_2 = \alpha_2 = 0.1$

$$\lambda_1 = 3~\mu~T_S = 5800~K \qquad \qquad \lambda_1~T_S~= 17400~\mu K \qquad F_1 = 0.995$$

$$\alpha_S = \alpha_1 F_1 + \alpha_2 (1-F_1)$$
 $\alpha_S = 0.896$

IR emisivity

$$\lambda_1 = 3 \mu$$
 $T = 500 K$ $\lambda_1 T = 1500 \mu K$ $F_1 = 0.0128$

$$\epsilon_{IR} = \epsilon_1 \; F_1 + \epsilon_2 \; (1 \text{-} F_1) \qquad \qquad \epsilon_{IR} = 0,11 \label{eq:epsilon}$$

$$\lambda_1 = 3 \mu$$
 $T = 600 K$ $\lambda_1 T = 1800 \mu K$ $F_1 = 0.0393$

$$\varepsilon_{IR} = \varepsilon_1 F_1 + \varepsilon_2 (1-F_1)$$
 $\varepsilon_{IR} = 0.13$

2) Radiative balance equation

$$S \alpha_S \phi_S = 2 \epsilon_{IR} S \sigma (T^4 - T_e^4),$$

where T_e is the temperature of the environment. $T_e = 300 \text{ K}$ (usual ambient temperature in on earth conditions, which is the value given at the end of the text).

The equilibrium temperature is given by:

$$T = \left(T_e^4 + \frac{\alpha_S \varphi_S}{2\varepsilon_{IR} \sigma}\right)^{0.25}$$

$$T = 573.8 \text{ K}$$

Comment:

The choice, for the calculation of $~\epsilon_{IR}$, of the equilibrium temperatures, T=500~K or T=600~K (question 1) has a small influence. For instance, T=573~K for $~\epsilon_{IR}=0,11$ or T=551~K for $\epsilon_{IR}=0,13$. The exact value of T should be obtained with an iterative procedure.

Second exercice

1) View factor matrix

The single value which is given is $F_{23} = 0.059$

$$F_{22} = F_{33} = 0$$
 (Plane surfaces)

$$\begin{array}{lll} F_{21} = 1 - F_{22} - F_{23} \ (flux \ conservation) & F_{21} = 0,941 \\ F_{12} = S_2 * F_{21} / S_1 & F_{12} = 0,118 \\ F_{13} = F_{12} \ (Symetry) & F_{13} = 0.118 \\ F_{11} = 1 - F_{12} - F_{13} \ (flux \ conservation) & F_{11} = 0,765 \\ F_{32} = F_{23} \ (Symetry) & F_{32} = 0.059 \\ F_{31} = 1 - F_{32} - F_{33} \ (flux \ conservation) & F_{31} = 0,941 \\ \end{array}$$

2) The Gebhart factor B_{ij} represents the fraction of flux emitted by surface S_i and absorbed by surface S_i , when taking into account all the possible paths from S_i to S_j .

It has a first contribution corresponding to the direct transfer, that is the flux emitted by S_i which is absorbed by S_j , in this direct transfer: $\alpha_j \ F_{ij}$.

The second contribution represents the flux emitted by S_i , reflected by all the S_k , which is finally absorbed by S_j , after all the possible reflections. For one such surface S_k it is given by :

$$\rho_k F_{ik} B_{ki}$$

Then for the complete cavities, we have:

$$B_{ij} = \alpha_j F_{ij} + \sum_k \rho_k F_{ik} B_{kj}$$

The flux conservation is written as:

$$\sum_{i} B_{ij} = 1$$

The reciprocity formula, between Gebhart factors can be written as:

$$\varepsilon_i S_i B_{ii} = \varepsilon_i S_i B_{ii}$$

3) To obtain B_{11} , B_{21} , B_{31} , the following system should be solved:

$$\begin{split} B_{11} &= \alpha_1 F_{11} + \rho_1 F_{11} B_{11} + \rho_2 F_{12} B_{21} + \rho_3 F_{13} B_{31} \\ B_{21} &= \alpha_1 F_{21} + \rho_1 F_{21} B_{11} + \rho_2 F_{22} B_{21} + \rho_3 F_{23} B_{31} \\ B_{31} &= \alpha_1 F_{31} + \rho_1 F_{31} B_{11} + \rho_2 F_{32} B_{21} + \rho_3 F_{33} B_{31} \end{split}$$

4) Matrix of Gebhart factors

The datas are the following coefficients:

$$B_{11}$$
=0,803 B_{21} =0,818 B_{31} =0,818 B_{22} =0,092. From the reciprocity formula we get:

$$B_{12} = \varepsilon_2 S_2 B_{21} / \varepsilon_1 S_1$$
 $B_{12} = 0.103$

The flux conservation gives then:
$$B_{13} = 1 - B_{11} - B_{12}$$
 $B_{13} = 0,094$ The flux conservation gives also : $B_{23} = 1 - B_{21} - B_{22}$ $B_{23} = 0,09$ $B_{32} = B_{23}$ (Symetry) $B_{32} = 0,09$

5)

$$g_{12} = \varepsilon_1 S_1 B_{12}$$

$$g_{13} = \varepsilon_1 S_1 B_{13}$$

$$g_{23} = \varepsilon_2 S_2 B_{23}$$

$$g_{2e} = \varepsilon_{2e} S_2$$

$$g_{3e} = \varepsilon_{IR} S_3$$

$$Q_3 = \alpha_S \varphi_S S_3$$

6)
$$S_1 = 3,53 \cdot 10^{-2} \qquad S_2 = 4,42 \cdot 10^{-3} \qquad S_3 = 4,42 \cdot 10^{-3}$$

$$g_{12} = 3,62 \cdot 10^{-4} \qquad g_{13} = 3,62 \cdot 10^{-4} \quad g_{23} = 4,31 \cdot 10^{-5} \quad g_{2e} = 3,97 \cdot 10^{-3}$$

$$Q_3 = 5.54 \text{ W}$$

7)
$$\varepsilon_{1}S_{1}B_{13}\sigma(T_{3}^{4}-T_{1}^{4})+\varepsilon_{1}S_{1}B_{12}\sigma(T_{2}^{4}-T_{1}^{4})=0$$

$$\varepsilon_{1}S_{1}B_{12}\sigma(T_{1}^{4}-T_{2}^{4})+\varepsilon_{2}S_{2}B_{23}\sigma(T_{3}^{4}-T_{2}^{4})+\varepsilon_{2e}S_{2}\sigma(T_{e}^{4}-T_{2}^{4})=0$$

$$\varepsilon_{1}S_{1}B_{13}\sigma(T_{1}^{4}-T_{3}^{4})+\varepsilon_{2}S_{2}B_{23}\sigma(T_{2}^{4}-T_{3}^{4})+\varepsilon_{IR}S_{3}\sigma(T_{e}^{4}-T_{3}^{4})+Q_{3}=0$$

8) T_3 is the highest, T_2 the lowest and T_1 is in between.

$$T_1 = 534 \text{ K}$$

$$T_2=353\ K$$

$$T_3 = 620 \text{ K}$$