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Part a) What happens if we do not neglect pressure from the perturbation equation?

First, I am going to solve this equation neglecting pressure so I can compare it later.

Starting w/ d'term by
Plugging in the scalar
Fourier mode (assuming
Perturbations are scalars):

$$\dot{J} = \frac{\partial^2}{\partial t^2} \left[\int_{S}^{2} e^{i(kx-\omega t)} \right] \\
= \frac{\partial}{\partial t} \left[-i\omega \int_{S}^{2} e^{i(kx-\omega t)} \right] \\
= \frac{\partial}{\partial t} \left[-i\omega \int_{S}^{2} e^{i(kx-\omega t)} \right] \\
= \frac{\partial}{\partial t} \left[-i\omega \int_{S}^{2} e^{i(kx-\omega t)} \right] \\
= -\omega^2 \int_{S}^{2} e^{i(kx-\omega t)} dx \\
= -\omega^2 \int_{$$

Now moving onto of

$$d = d [dsei(kx-\omega t)]$$

= -iwdsei(kx-wt)

Now plugging the derivatives into the original equations:

477Gp. (ds e i(kx-ast)) - 0 = 0

Setting pressure

term to 0

- ds e i(kx-ast) = 2Hidge i(kx-ast) = 0

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- de ilkx-ost) 2 - 24ico 447670] = 0

$$\omega = -b \pm \sqrt{b^2 + 4ac} = -(24i) \pm \sqrt{(24i)^2 - 4(1)/444690}$$

Starting with the positive value:

The distre growth term where the distremental is the decay term. Now, I will include pressure in the differential equation:

J+2HJ-47Gg.d- Cs Vot=0

Furier Modes:

P= Ps e (kx-ast)

= is e (kx-ast)

I will use the same derivatives as last time, but I will now derive the $\frac{C_3}{a}$ $\sqrt{3}$ term

 $\nabla^2 \int = \frac{\partial^2}{\partial x^2} \left(\frac{\partial^2}{\partial s^2} e^{i(kx - \omega t)} \right)$

$$= \frac{\partial}{\partial x} \left(\frac{\partial}{\partial s} i k e^{i(kx-\omega t)} \right)$$

$$= -\frac{\partial}{\partial x} \left(\frac{\partial}{\partial s} i k e^{i(kx-\omega t)} \right)$$

$$= -\frac{\partial}{\partial x} \left(\frac{\partial}{\partial s} i k e^{i(kx-\omega t)} \right)$$

Now plugging in all the derivatives into the differential equations:

477Gp. (
$$d_s e^{i(kx-axt)}$$
) + $\frac{C_s}{a}(d_s k e^{i(kx-axt)}) = 0$

$$\omega = \frac{-b \pm \sqrt{b^{2} - 4ac^{2}}}{2a}$$

$$-(-2cH) \pm \sqrt{(-2cH)^{2} - 4(-c)(4\pi Gg_{0} + \frac{k^{2}c_{0}^{2}}{a})^{2}}$$

$$= \frac{2(-1)}{2cH} + \frac{2(-1)^{2} + 4(4\pi Gg_{0} + \frac{k^{2}c_{0}^{2}}{a})^{2}}{-2}$$

$$= \frac{2cH \pm 2c \int_{-2cH}^{2cH} - \frac{k^{2}c_{0}^{2}}{a}}{-2}$$

$$= -(iH \pm i)H^{2} - 4\pi Gg_{0} - \frac{k^{2}c_{0}^{2}}{a}$$

Starting to sub in Are positive term first into the Fourier mode.

$$i(kx - (-iH - iJH^2 - 476p, -\frac{k^2c_3^2}{a^2})t)$$

$$d = de$$

=
$$\sqrt{\frac{i(x+it)t^2-4\pi c_p - \frac{v^2 c_s^2}{ac}}{2}}$$

= $\sqrt{\frac{i(x+it)t^2-4\pi c_p - \frac{v^2 c_s^2}{ac}}{2}}$
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$$\int_{1}^{2} \frac{1}{e^{2}} \frac{1}{e$$

Now plugging in the negative term into the Fourier Mode:

In Aris case, of is the decaying term while of is the growing term.

B) This instability illustrates a competition between which two forces?

This instability illustrates the competition between radiative pressure (The k space) and The gravitational force (The omega space)

C) Do you think a similar equation may also explain one of the gas clumping processes at work in galaxies? If yes, which one?

Yes, there is a similar equation that may explain gas clumping processes in galaxies. It is the Jean's instability. The densest parts of the cold patches in a galaxy will collapse due to gravity since there is a lack of pressure.