Nuclear Saturation with QCD sum rules

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Abstract

The scalar and vectorial self energies obtained through QCD sum rules are introduced in the Quantum Hadrodynamics (QHD) equations. This QHD and QCD mixing show us that the effect of the density on the coupling constants is very small.

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I. INTRODUCTION

The research for the exact equilibrium properties of the nuclear matter [1,2] is an old problem of nuclear physics and it continues to have a great interest nowadays. Quantum Hadrodynamics (QHD) [3] is one of the models with larger success to describe the properties of the infinite nuclear matter as well as those of finite nuclei. QHD describes the nucleon-nucleon interaction through the mesons exchange $(\pi, \sigma, \omega, \rho, \text{etc.})$. Several calculations of nuclear structure using QHD and its extensions were made with success in the explanation of experimental data [4]. For the purposes of this work only the simplest QHD model will be considered.

The basic QHD-model that explains the nuclear matter includes the nucleon (ψ) coupled with sigma (σ) and omega (ω) mesons. In spite of the pion be the principal component of the nucleon-nucleon interaction, it is not included because the nuclear matter is an isotropic system with parity conservation. Analyzing the model we see that the real part of the scalar term (σ -meson) is typically of the order of several hundred MeV attractive while the real part of the time component of the vector term (ω meson) is typically of the order of several hundred MeV repulsive. However the energies involved in problems of nuclear structure are only of a few tens of MeV. That order of energy is obtained on the QHD model due to a large cancellation among the scalar and vector pieces and this process of cancelation can be controlled through an appropriate choice of the coupling constants. However, these so called QHD constants are quite different from those ones given in Bonn potential [5] or empirical data [3]. So it is interesting to discuss the validity of these coupling constants.

The success of the Walecka model, which is based on Dirac's equation, is due to the mutual cancellation between the large scalar and vector potentials. However in some works [6,7] it is discussed that the composed nature of the nucleon can suppress the scalar optical potential. However, as it is argued in [8], it can be shown [9] that due

to the nucleon being immersed in the nuclear media, such suppression does not exist. This result makes possible the use of the Dirac phenomenology for composed particles. Besides, recently the effective field theory (EFT) [10–15] has verified that the QHD models are consistent with the symmetries of the Quantum Chromodynamics (QCD), the correct theory of the strong interaction. This last fact motivates the mixing between QHD and QCD accomplished in this work.

It is known that QCD is the correct theory of the strong interaction. However, there is not a perturbative treatment for QCD at energies involved in nuclear matter problems, so it is interesting to use the nonperturbative method given by QCD sum rules that was introduced by Shifman, Vainshtein, and Zakharov in the late 1970's [16]. The method consists in describing a correlation function in terms of hadron degree of freedom as well as quarks degree of freedom. This last one can be written in an operator product expansion (OPE) where the perturbative part (Wilson coefficients) is separate from the nonperturbative (condensates). The power of this technique is in the fact that the nonperturbative part is the same for all problems.

The objective of this work is to get the results obtained from QCD sum rules in the media [8] and use them on QHD equations to analyze the necessary constants to obtain the saturation point of the infinite nuclear matter.

II. THE QHD MODEL

The model used to describe the nuclear matter is from Serot and Walecka [3] and it includes nucleons (ψ) interacting with σ and ω mesons. So that the Lagrangian density is given by

$$\mathcal{L} = \bar{\psi}(\gamma^{\mu}\partial_{\mu} + g_{s}\sigma - g_{v}\gamma^{\mu}\omega_{\mu} - M)\psi + \frac{1}{2}(\partial_{\mu}\sigma\partial^{\mu}\sigma - m_{s}^{2}\sigma^{2}) - \frac{1}{2}\left(G_{\mu\nu}G^{\mu\nu} - m_{v}^{2}\omega^{\mu}\omega_{\mu}\right) , \quad (1)$$

where g_s and g_v are the mesons coupling constants, $G_{\mu\nu} = \partial_{\mu}\omega_{\nu} - \partial_{\nu}\omega_{\mu}$ is the strength tensor for the vector field and M, m_s and m_v are the nucleon, σ -meson and ω -meson masses, respectively. The Green function for the nucleon is given by

$$G(k) = \left(\gamma_{\mu}\tilde{k}^{\mu} - M^{*}\right) \left\{ \frac{1}{\tilde{k}^{2} - M^{*2} + i\epsilon} + \frac{i\pi}{E^{*}(k)} \delta\left(k^{0} - E(k)\right) \theta(k_{F} - |\mathbf{k}|) \right\} , \qquad (2)$$

where it was defined

$$\tilde{k}^{\mu} \equiv k^{\mu} + \Sigma^{\mu}_{(v)} , \qquad (3)$$

$$M^* \equiv M + \Sigma_{(s)} , \qquad (4)$$

$$E^*(k) \equiv \sqrt{\left|\tilde{\mathbf{k}}\right|^2 + M^{*2}} \,. \tag{5}$$

Here $\Sigma_{(s)}$ and $\Sigma_{(v)}^{\mu}$ represent the scalar and vector self-energies, respectively.

In the more simple model only the contributions of the baryons from the Fermi sea will be considered. This is equivalent to disregard the contributions coming from antibarions (Dirac sea). So in agreement with the definition of energy-momentum tensor

$$\varepsilon = \langle \Psi | \hat{T}_{00} | \Psi \rangle - VEV$$

$$= \frac{1}{2} \frac{m_s^2}{g_s^2} \Sigma_{(s)}^2 + \frac{1}{2} \frac{m_v^2}{g_v^2} \Sigma_{(v)}^2 + \frac{\gamma}{(2\pi)^3} \int_0^{k_F} d^3k \sqrt{|\mathbf{k}|^2 + M^{*2}} ,$$
(6)

where VEV is the vacuum expectation value of the T_{00} and γ is the spin-isospin degeneracy. The self-energies can be obtained through "tadpole" Feynman diagrams and are given by the following relationships

$$\Sigma_{(s)} = M^* - M$$

$$= -\frac{g_s^2}{m_s^2} \frac{\gamma M^*}{4\pi^2} \left[k_F E_F^* - M^{*2} \ln \left(\frac{k_F + E_F^*}{M^*} \right) \right]$$
(7)

and

$$\Sigma_{(v)}^{\mu} = -\delta^{\mu 0} \frac{g_v^2}{m_v^2} \frac{\gamma}{6\pi^2} k_F^3 , \qquad (8)$$

The expression (7) should be solved in a self-consistent way. Therefore all the ingredients are available to calculate the energy density. This result is also known as Mean Field Theory (MFT), because it can be obtained with a mean field approach for the meson fields.

A simple analysis shows us that the self-energies $\Sigma_{(s)}$ and $\Sigma_{(v)}$ are proportional to the scalar and vectorial densities, respectively. Then the form presented by Eq. (6) represents an expansion on powers of these densities. On the other hand, in agreement with the EFT concepts, the energy density is a functional and has an expansion in powers of the densities (scalar, vectorial, etc.) organized

through Georgi's naive dimensional analysis (NDA). So, the energy densities have the same form that the expansions obtained by EFT [13–15]. Therefore, since QHD has a foundation on the symmetries of QCD, the next step is to calculate the self-energies through the QCD methods. However, as the energy levels treated in the nuclear problems are in a nonperturbative regime for QCD, the sum rules is the appropriate method to be used.

III. QCD SUM RULES

The QCD sum rules method [16] begins with the following time-ordered correlation function, defined by

$$\Pi_{\alpha\beta}(q) \equiv i \int d^4x \ e^{iq \cdot x} \langle 0|T \left[\eta_{\alpha}(x)\bar{\eta}_{\beta}(0)\right]|0\rangle , \qquad (9)$$

where $|0\rangle$ is the physical nonperturbative vacuum state and $\eta_{\alpha}(x)$ is an interpolating field with the quantum numbers of a nucleon. As proposed by Ioffe [17], the proton field is given by

$$\eta(x) = \epsilon_{abc} \left(u_a^T C \gamma_\mu u_b \right) \gamma_5 \gamma^\mu d_c , \qquad (10)$$

where u and d are the quark fields, a, b, c are the color indexes and C is the charge-conjugation matrix. Now the correlation function can be written as an operator product expansion (OPE) whose nonperturbative part (condensates) can be separated from the perturbative one (Wilson coefficients). The OPE can be generated starting from the following expansion for the quark propagator [18,19]

$$S_{ab}(x) = \langle 0|T \left[q_a(x)\bar{q}_b(0)\right]|0\rangle$$

$$= i\frac{\delta_{ab}}{2\pi^2}\frac{\cancel{x}}{x^4} - \frac{\delta_{ab}}{12}\langle\bar{q}q\rangle_{vac} + \cdots, \qquad (11)$$

where $\langle \bar{q}q \rangle_{vac}$ is the quark condensate which can be determined from the Gell-Mann-Oakes-Renner relation,

$$2m_q \langle \bar{q}q \rangle_{vac} = -m_\pi^2 f_\pi^2 \left(1 + \mathcal{O}(m_\pi^2) \right) .$$

In this relation, $m_{\pi}=138$ MeV is the pion mass, $f_{\pi}=93$ MeV is the pion decay constant and $m_q=(m_u+m_d)/2\simeq 7\pm 2$ MeV is the average of the up and down quark masses.

The same correlation function, Eq (9), can be written in a phenomenological way, which mean that the correlator has a hadronic description through the nucleon Green function, Eq (2). The match of the theoretical side (OPE) and the phenomenological one is the essence of the QCD sum rules. Using the fundamental state of the nuclear matter as vacuum and the interacting propagator in the phenomenological side, it is possible to obtain the following sum rules [8]

$$M_N^* = -\frac{8\pi^2}{M^2} \langle \bar{q}q \rangle_{\rho_B} , \qquad (12)$$

$$\Sigma_{(v-QCD)} = \frac{64\pi^2}{3M^2} \langle q^{\dagger} q \rangle_{\rho_B} . \tag{13}$$

Here M^2 represent the borel mass with value near $M^2 = 1 \text{GeV}^2$ [17], and

$$\langle q^{\dagger}q\rangle_{\rho_B} = \frac{3}{2}\rho_B \;, \tag{14}$$

$$\langle \bar{q}q \rangle_{\rho_B} = \left(1 - \frac{\sigma_B \rho_B}{m_\pi^2 f_\pi^2} + \cdots \right) \langle \bar{q}q \rangle_{vac} .$$
 (15)

The expression (12) is a generalization of the Ioffe's formula [17] to finite density. The sigma term is estimate in Ref [20] as $\sigma_B \simeq 45 \pm 10$ MeV. Taking the ratios of the expressions (12) and (13) in relation to the mass M_N (Ioffe's formula), they are obtained the following results:

$$\Sigma_{(s-QCD)} = -\frac{\sigma_B M_N}{m_\pi^2 f_\pi^2} \rho_B , \qquad (16)$$

$$\Sigma_{(v-QCD)} = \frac{8m_q M_N}{m_{\pi}^2 f_{\pi}^2} \rho_B \ . \tag{17}$$

Although a self-consistent relation is not obtained as in Eq (7), these expressions have a more fundamental nature based on quark and gluon degree of freedom. At this point, the idea is to use the expressions (16) and (17) to calculate the energy density of the nuclear matter.

IV. NUMERICAL RESULTS AND CONCLUSION

Firstly it should be noted that only the ratios between the coupling constants and the mesons mass appear in the expressions where the energy density of the nuclear matter is calculated. Some values for these ratios are presented in the table I. In the first line of table I the coupling constants were adjusted so that the MFT might reproduce the saturation properties of the nuclear matter. With the use of these coupling constants, the result obtained with MFT are the following (dotted curve on Fig. 1): saturation point on $k_F = 1.12 \text{ fm}^{-1}$ with a energy density by nucleon given by -15.75 MeV. This result reproduces the exact equilibrium properties of the nuclear matter given by Refs [1,2], which is the expected result, once the coupling constants were chosen for such adjustment to occur. However, these coupling constants are very different of the empirical values [3].

	g_s^2/m_s^2	g_v^2/m_v^2
MFT	$3.029 \ 10^{-4}$	$2.222 \ 10^{-4}$
Bonn Potential	$3.233 \ 10^{-4}$	$4.104 \ 10^{-4}$
Empirical Values	$3.402 \ 10^{-4}$	$3.516 \ 10^{-4}$

TABLE I. Coupling constants used in MFT and in Bonn potential and the empirical values.

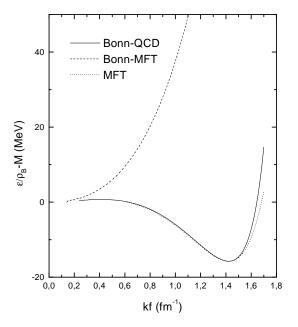


FIG. 1. Binding energy by nucleon in nuclear matter as function of Fermi momentum. In the continuous line is adopted the Bonn coupling constants and the self-energies obtained with QCD sum rules. In the dotted and dashed lines, the original MFT equations are used together with the coupling constants from the MFT and Bonn potential, respectively.

The most accepted values for these constants in free space are found in the Bonn potential [5], for which the nucleon-nucleon interaction is adjusted to describe the scattering data. The Bonn coupling constants for the sigma and omega mesons are given on the second line of table I while the empirical values are presented on the third line of that table. It must be noted that, in MFT the value of the coupling constant for the scalar meson is higher than the respective value for the vector meson. But in the Bonn potential and in the empirical data there is an inversion in the magnitude of these constants. Using the Bonn's values in Eq. (6), the result presented by the dashed curve on Fig. 1 is obtained. That curve shows us that the energy density for MFT does not present a good behavior. There is no saturation point, in other words, there is no formation of nuclear matter. The same result is obtained with the empirical constants. The problem is in the fact that the use of the Bonn or empirical coupling constants implicates in a small increase of the attraction but with a much larger increase in the repulsion. It can

be argued that the difference between the QHD coupling constants and the Bonn values is due to the fact that the latter are obtained in free space while the QHD values are obtained in nuclear media, but this argument is not valid for empirical constants. However, using the constants given by Bonn potential, and the self-energies found through QCD sum rules, Eqs. (16) and (17), with $\sigma_B=49$ MeV and $m_q=6.8$ MeV, the result represented by the continuous line on Fig. 1 is obtained. That curve has a saturation point $k_F=1.42$ fm⁻¹ with -15.75 MeV for the energy density by nucleon, which is the saturation point of the nuclear matter. Therefore the QCD sum rules allows us to reproduce the properties of nuclear matter with the simplest QHD model using more realistic values for the coupling constants.

Since the empiric constants are not so different from the Bonn values when they are compared to the QHD values, it can be that the effects of the density do not alter the values of these constants significantly. On the other hand, the exact determination of the parameters σ_B and m_q should allow a better evaluation of the coupling constants in the media for the various QHD models and consequently to determine with more precision their real variation in relation to the values of the constants in the vacuum. Furthermore this result is another indicative that the quark degree of freedom takes an important place in nuclear problems.

Finally, it is known that the QCD sum rules work with more accuracy when there is high transferred momentum. So I hope this mix between QCD sum rules and QHD model can be applied, with success, in systems in a regime of high density and high temperature as neutron stars. Besides, in these calculations the contributions of the gluon condensate and four quark condensate were not included. But it is known that the contribution of these terms for the sum rule is very small. However, if more precise answers are required, these terms have to be included as well as the most sophisticated versions of QHD. These and other questions are left for future works.

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