



AST3003S Observational Project

Measuring the Rotational Velocity of the Galactic Plane Using HI Gas 1.4 GHz Emission

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Overview

- Telescope specifications
- Planning your observations
- Telescope operation
- Extracting the spectra
- Brightness temperature
- LSR velocity
- Calculating rotational velocity & distance
- Writing your report



The Milky Way curves over the MeerKAT telescope in South Africa. (SARAO)



Telescope Specifications



TFTO Radio Telescope Specifications

Diameter	3 m
Configuration	Prime Focus
Mount	Alt-azimuth
Wind stow	50 km/h
Central frequency	1428.75 MHz
Bandwidth	62.5 MHz
Spectrometer	1024 channels
Resolution	61 kHz
A/D converter	14 bit
Max slew speed	2°/s



Planning Observations

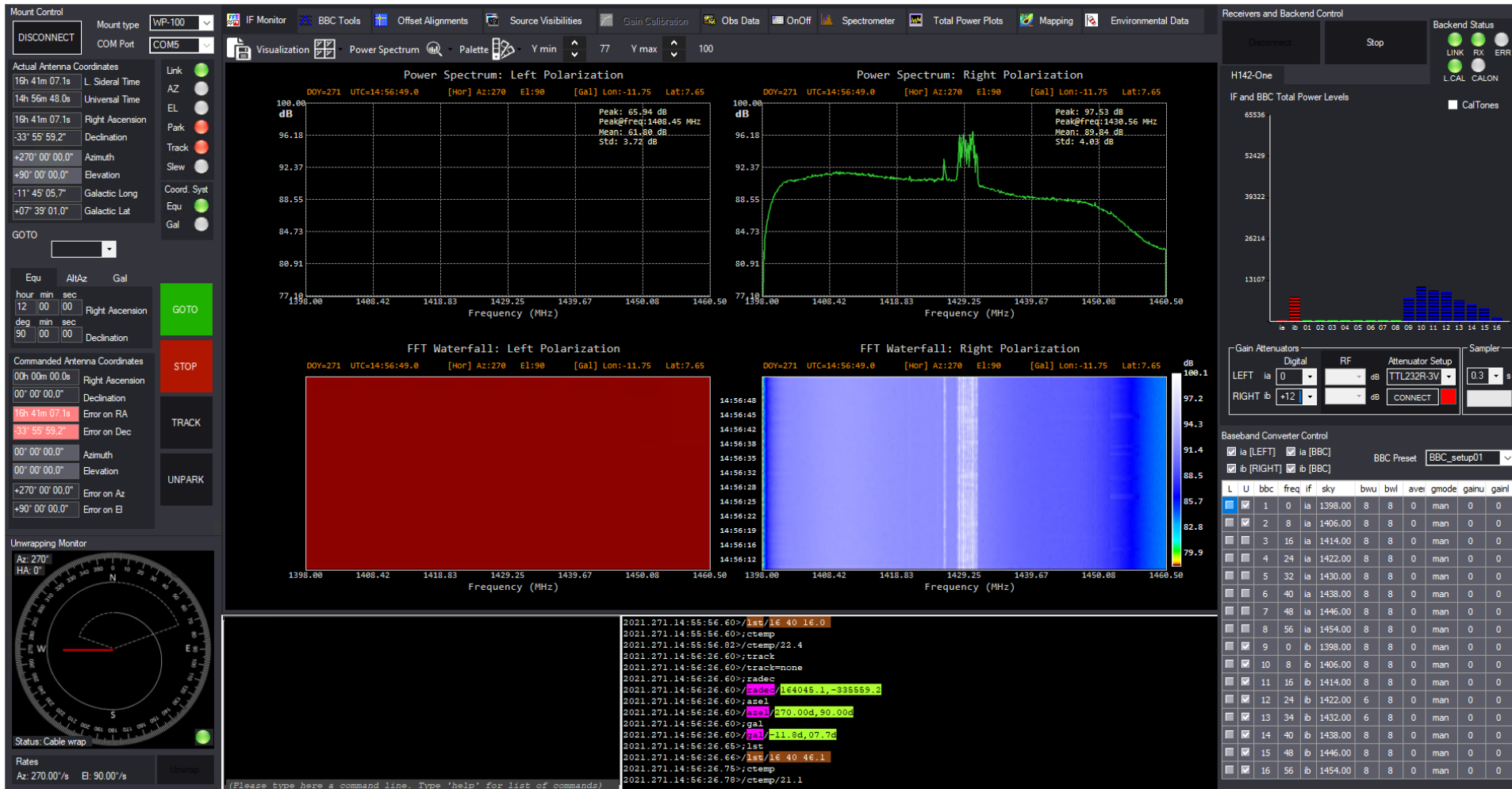


- Each group will observe 3 points along the galactic plane between 270° and 90° longitude
- Check to see what times your coordinates are visible
- Schedule your observing session
- Check weather conditions on day
- Observations done remotely from RW James

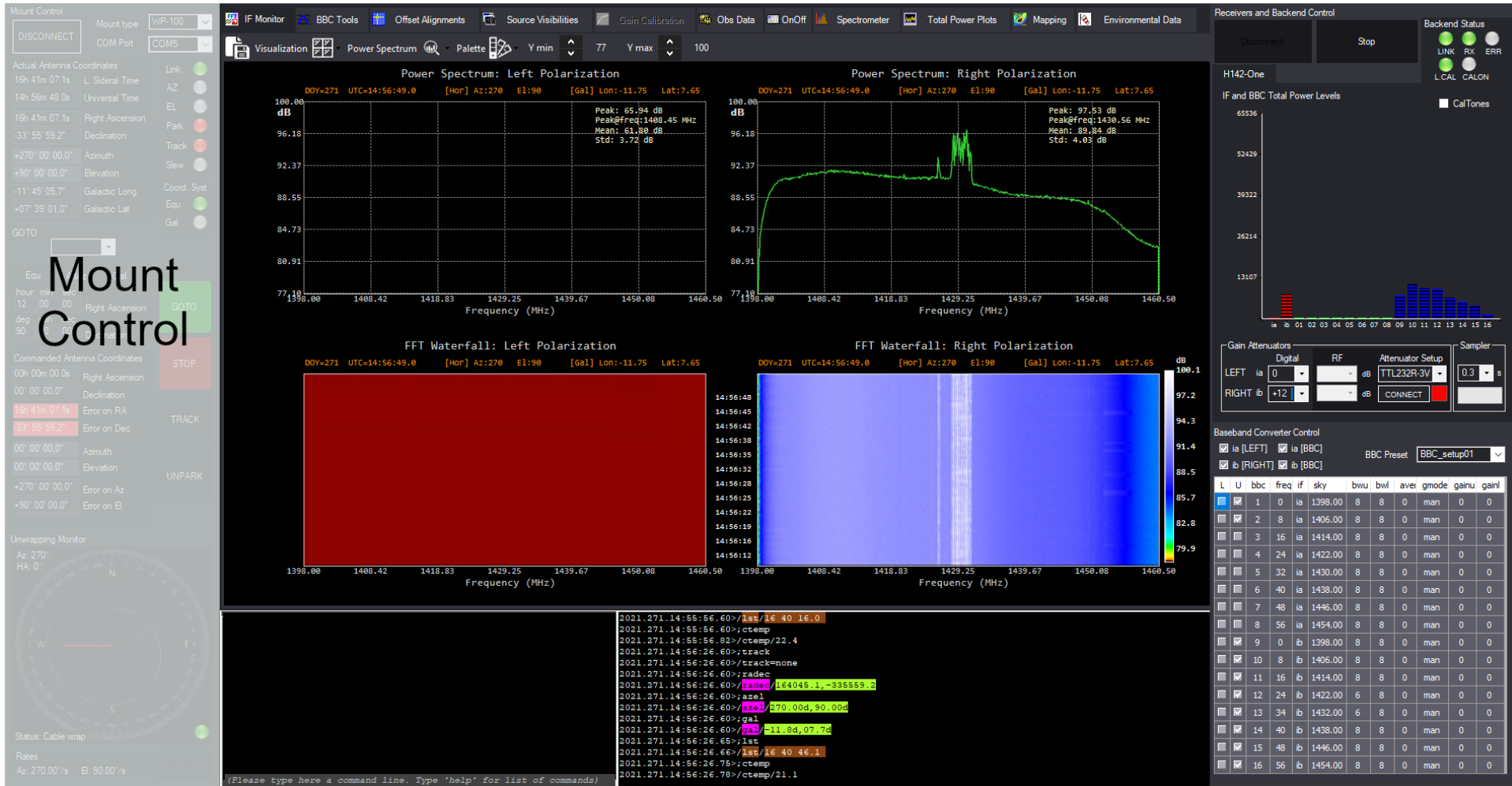


Telescope Operation

- Software UI

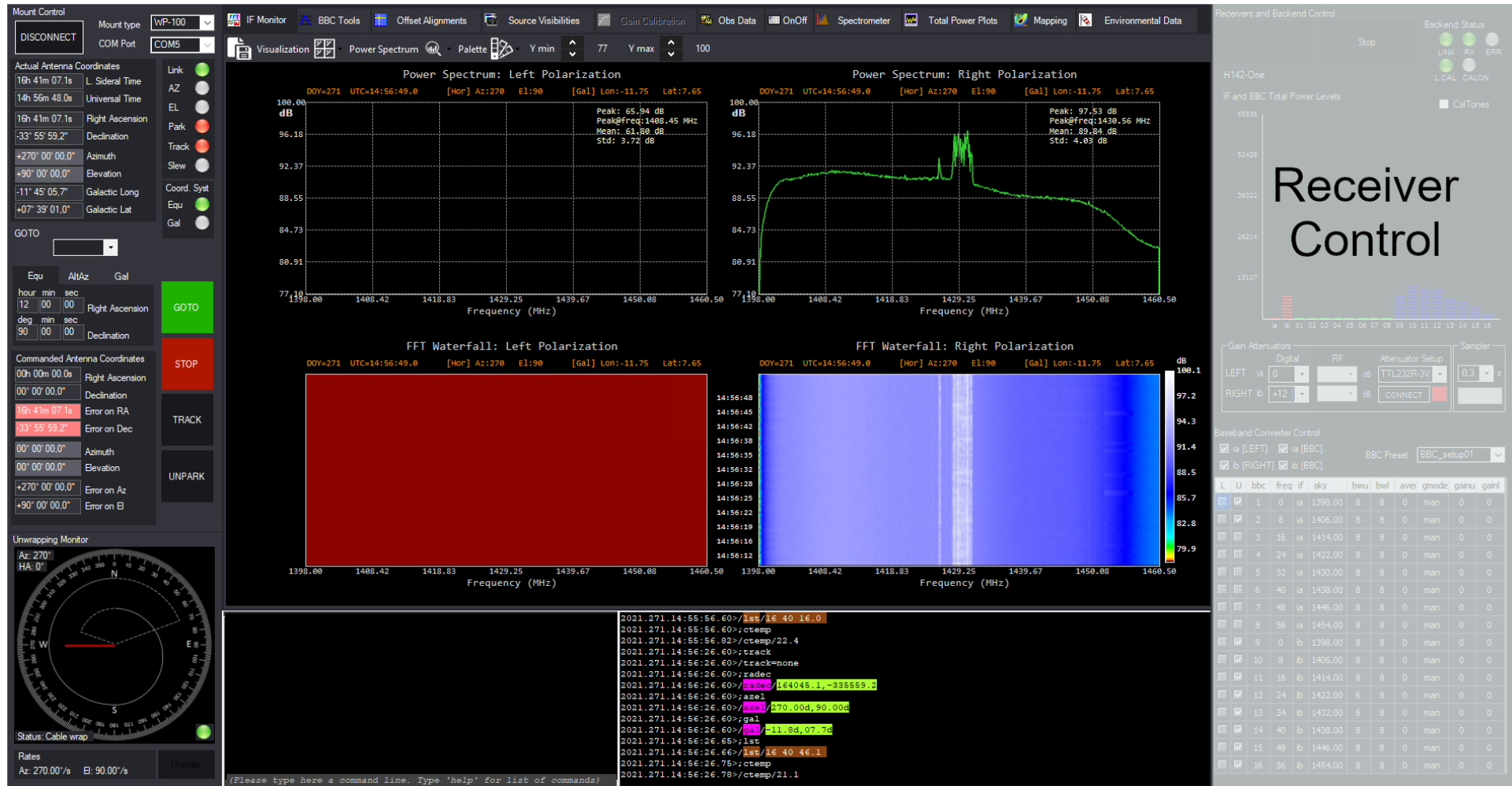


Telescope Operation



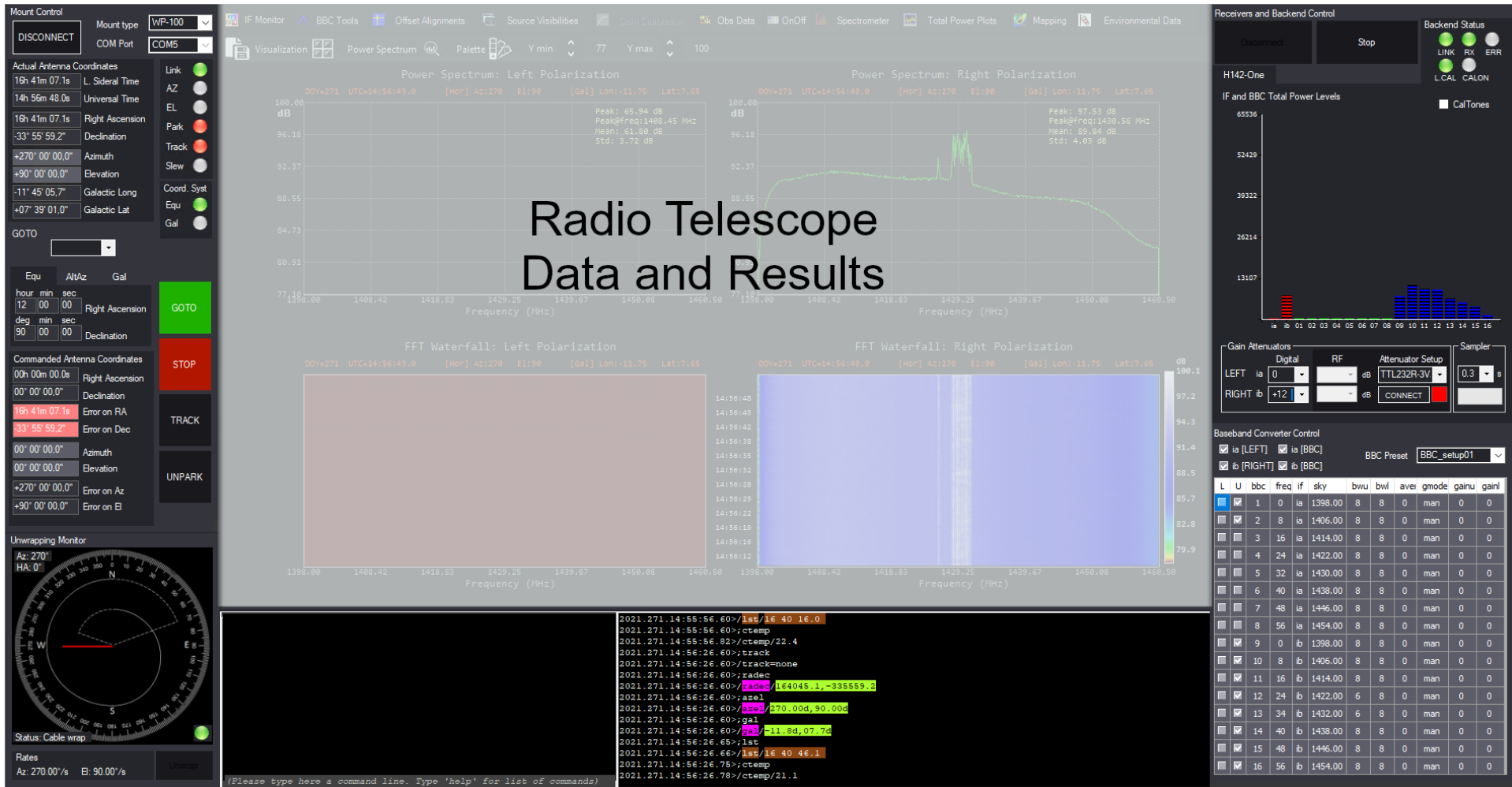
- Software UI
- Displays mount status

Telescope Operation



- Software UI
- Displays mount status
- Monitors power levels

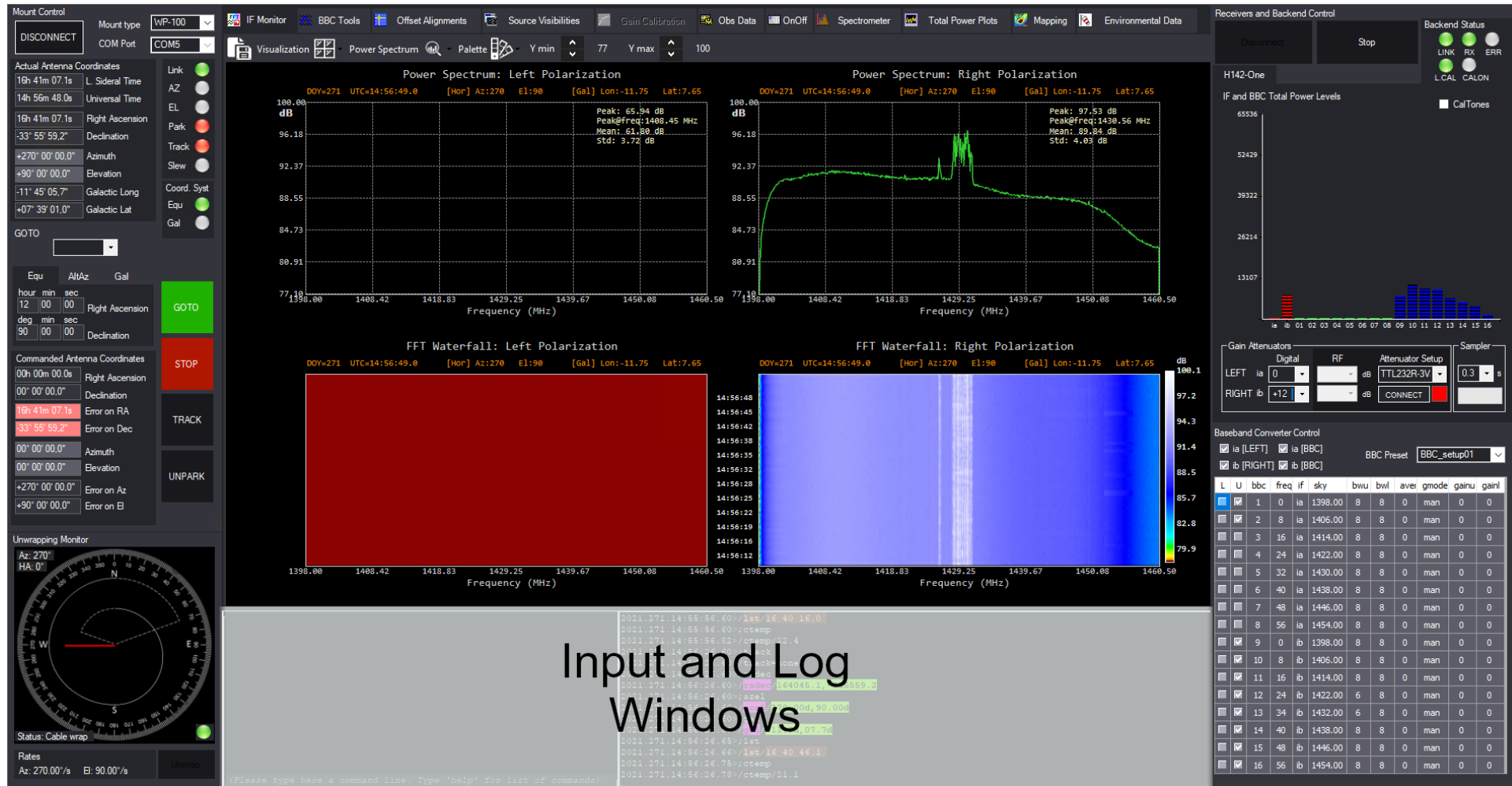
Telescope Operation



- Software UI
- Displays mount status
- Monitors power levels
- Tabs with various functions



Telescope Operation



- Software UI
- Displays mount status
- Monitors power levels
- Tabs with various functions
- Command entry and log

Telescope Operation

Nodding Data Control

Antenna offsets

Commanded	Actual
Az: 0.00°	Az: -0.01°
El: 0.00°	El: -0.01°

OnOff parameters

Repetitions: 10

Averages [s]: 10

Elev cutoff [°]: 60

Distance [beam]: 0 ☒ galoff

Spectrometer setup

☐ Acquire OnOff Spectra

☒ Save FITS Spectra

FoV: 42.00°

HPBW: 4.03°

PHASE

OVERALL

Measure Phases

ON OFF CAL WAIT ACQ BUSY READY

Confirm setup

Stop OnOff

OnOff IF and BBC outputs

rep	dev	sb	on_tpi	off_tpi	cal_tpi	signal	tsys	src_cal
1	bbc09	u	6791,64	6785,12	10670,39	00,02	17,46	00,65
1	bbc10	u	9794,64	9798,00	16154,09	-00,01	15,42	-00,34
1	bbc11	u	8961,27	8961,30	16131,12	00,00	12,50	00,00
1	bbc12	u	8725,30	8730,52	17077,18	-00,01	10,46	-00,52
1	bbc13	u	6194,70	6194,30	14024,97	00,00	07,91	00,04
1	bbc14	u	5055,24	5061,70	11501,39	-00,01	07,86	-00,65
1	bbc15	u	3733,45	3740,73	7571,61	-00,02	09,76	-00,73
1	bbc16	u	1210,18	1215,48	2115,52	-00,06	13,50	-00,53
2	bbc09	u	6784,15	6801,67	10693,03	-00,05	17,48	-01,75
2	bbc10	u	9800,55	9807,36	16160,58	-00,01	15,44	-00,68
2	bbc11	u	8959,12	8970,36	16147,52	-00,02	12,50	-01,12
2	bbc12	u	8696,82	8649,70	17060,30	00,06	10,28	04,71
2	bbc13	u	6190,94	6192,06	14035,48	00,00	07,89	-00,11
2	bbc14	u	5052,06	5053,36	11519,91	00,00	07,81	-00,13
2	bbc15	u	3735,82	3735,73	7581,00	00,00	09,72	00,01
2	bbc16	u	1210,82	1209,97	2114,85	00,01	13,37	00,08
3	bbc09	u	6813,27	6814,76	10687,42	00,00	17,60	-00,15
3	bbc10	u	9833,88	9795,42	16154,91	00,06	15,40	03,85
3	bbc11	u	8999,18	8957,15	16139,45	00,06	12,47	04,20
3	bbc12	u	8746,76	8684,73	17053,88	00,07	10,38	06,20
3	bbc13	u	6212,18	6191,55	14037,88	00,03	07,89	02,06
3	bbc14	u	5066,70	5051,85	11518,97	00,02	07,81	01,48
3	bbc15	u	3745,64	3731,76	7572,15	00,04	09,72	01,39
3	bbc16	u	1214,48	1210,24	2114,97	00,05	13,38	00,42
4	bbc09	u	6764,67	6780,76	10664,61	-00,04	17,46	-01,61
4	bbc10	u	9782,48	9779,45	16143,91	00,00	15,37	00,30
4	bbc11	u	8946,09	8942,27	16133,09	00,01	12,44	00,38
4	bbc12	u	8644,39	8640,94	17064,00	00,00	10,26	00,35
4	bbc13	u	6181,48	6176,48	14034,47	00,01	07,86	00,50

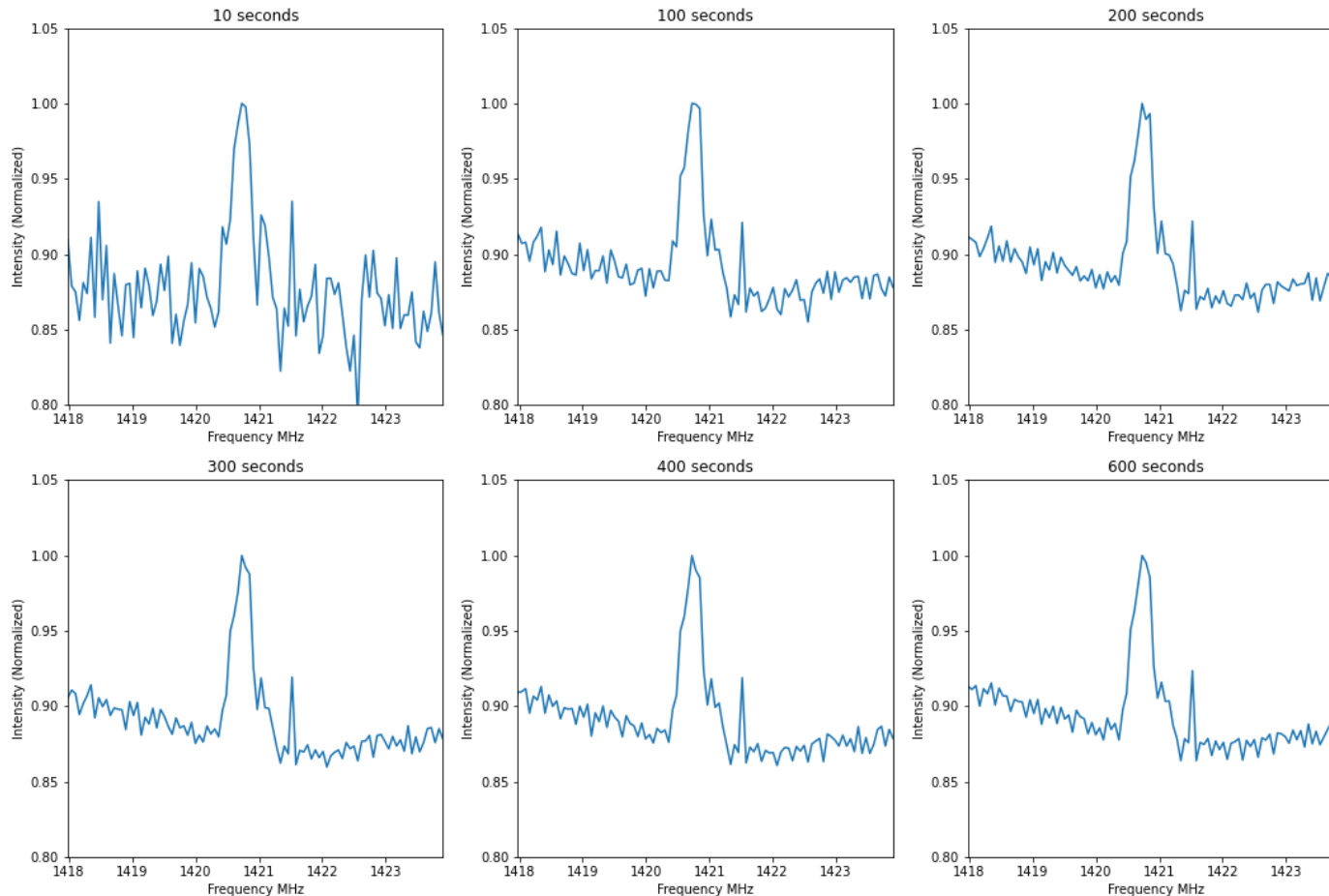
- On/Off used to collect spectral data
- Parameters can be set on top left
- Zero distance will allow on and off to be used in spectra
- Monitor progress bars after starting process



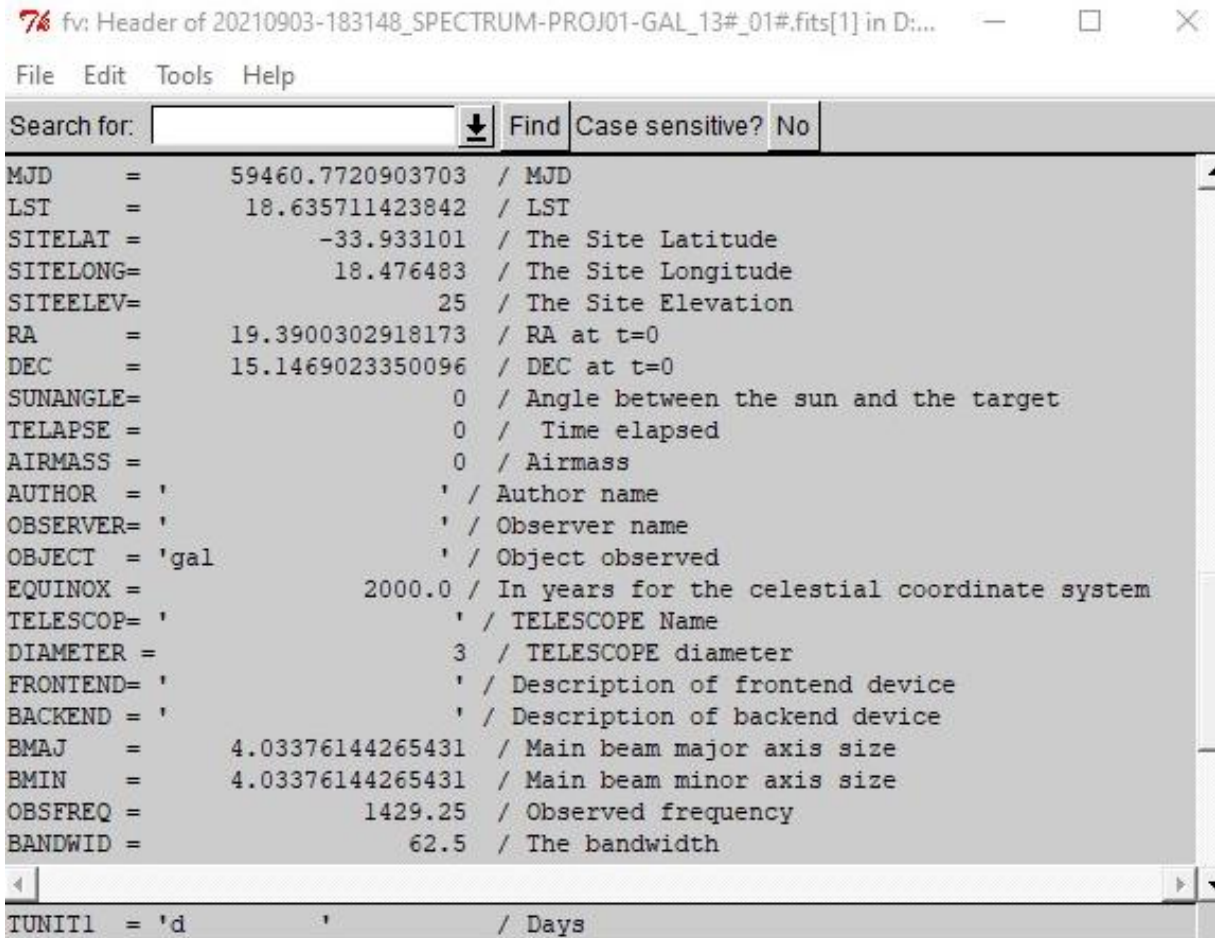
Extracting the Spectra

Galactic coordinates $l = 40^\circ, b = 0^\circ$

- Take multiple acquisitions and stack to increase SNR



Extracting the Spectra



The screenshot shows a text editor window titled "74 fv: Header of 20210903-183148_SPECTRUM-PROJ01-GAL_13#_01#.fits[1] in D:...". The window contains the header information for a FITS file, listing various astronomical parameters and their values. The parameters include MJD, LST, site coordinates (SITELAT, SITESLONG, SITEELEV), celestial coordinates (RA, DEC), observation parameters (SUNANGLE, TELAPSE, AIRMASS), observer information (AUTHOR, OBSERVER, OBJECT), coordinate system (EQUINOX), telescope details (TELESCOP, DIAMETER, FRONTEND, BACKEND), beam sizes (BMAJ, BMIN), observation frequency (OBSFREQ), bandwidth (BANDWID), and time unit (TUNIT1).

```
MJD      =      59460.7720903703 / MJD
LST      =      18.635711423842 / LST
SITELAT  =      -33.933101 / The Site Latitude
SITESLONG=      18.476483 / The Site Longitude
SITEELEV=      25 / The Site Elevation
RA       =      19.3900302918173 / RA at t=0
DEC      =      15.1469023350096 / DEC at t=0
SUNANGLE=      0 / Angle between the sun and the target
TELAPSE  =      0 / Time elapsed
AIRMASS  =      0 / Airmass
AUTHOR   =      ' / Author name
OBSERVER =      ' / Observer name
OBJECT   =      'gal / Object observed
EQUINOX  =      2000.0 / In years for the celestial coordinate system
TELESCOP =      ' / TELESCOPE Name
DIAMETER =      3 / TELESCOPE diameter
FRONTEND =      ' / Description of frontend device
BACKEND  =      ' / Description of backend device
BMAJ     =      4.03376144265431 / Main beam major axis size
BMIN     =      4.03376144265431 / Main beam minor axis size
OBSFREQ  =      1429.25 / Observed frequency
BANDWID  =      62.5 / The bandwidth
TUNIT1   =      'd / Days
```

- Take multiple acquisitions and stack to increase SNR
- Get all important information needed from FITS header

Extracting the Spectra

76 fvt: Binary Table of 20210912-060233_SPECTRUM-PROJ01-GAL_01#_01#.fits[1] in D:/OneDrive/UCT/Work/Telescopes/Radio Data/20210912/FITS/l_297_b_-15/

File Edit Tools Help

Select	Gal_Long	Gal_Lat	LEFT_POL	RIGHT_POL	STATUS	MARKER
	1D	1D	1024K	1024K	4A	16A
<input type="checkbox"/> All	deg	deg	count	count		
Invert	Modify	Modify	Modify	Modify	Modify	Modify
1	2.072986601615E+002	-1.441479236178E+00:	Plot	Plot	on	SPC 000
2	2.073076724812E+002	-1.441633710287E+00:	Plot	Plot	off	SPC 001
3	2.072962723984E+002	-1.441531794248E+00:	Plot	Plot	cal	SPC 002
4	2.072964274120E+002	-1.441763803679E+00:	Plot	Plot	on	SPC 003
5	2.073056305814E+002	-1.441932470252E+00:	Plot	Plot	off	SPC 004
6	2.073046538389E+002	-1.441978510431E+00:	Plot	Plot	cal	SPC 005
7	2.073057542580E+002	-1.441661321302E+00:	Plot	Plot	on	SPC 006
8	2.073048971718E+002	-1.441719105075E+00:	Plot	Plot	off	SPC 007
9	2.073041000439E+002	-1.441782748170E+00:	Plot	Plot	cal	SPC 008
10	2.073053803509E+002	-1.441483137816E+00:	Plot	Plot	on	SPC 009
11	2.073039959442E+002	-1.442126936024E+00:	Plot	Plot	off	SPC 010
12	2.073040867571E+002	-1.441639648127E+00:	Plot	Plot	cal	SPC 011
13	2.073048544026E+002	-1.441926742247E+00:	Plot	Plot	on	SPC 012
14	2.073050518680E+002	-1.441450360458E+00:	Plot	Plot	off	SPC 013
15	2.073039371880E+002	-1.442118923379E+00:	Plot	Plot	cal	SPC 014
16	2.073055868614E+002	-1.441854614762E+00:	Plot	Plot	on	SPC 015
17	2.073052798487E+002	-1.441965203799E+00:	Plot	Plot	off	SPC 016
18	2.073056916614E+002	-1.441510545716E+00:	Plot	Plot	cal	SPC 017
19	2.072966027257E+002	-1.441725007283E+00:	Plot	Plot	on	SPC 018
20	2.072964769194E+002	-1.441854173797E+00:	Plot	Plot	off	SPC 019

Go to: Edit cell:

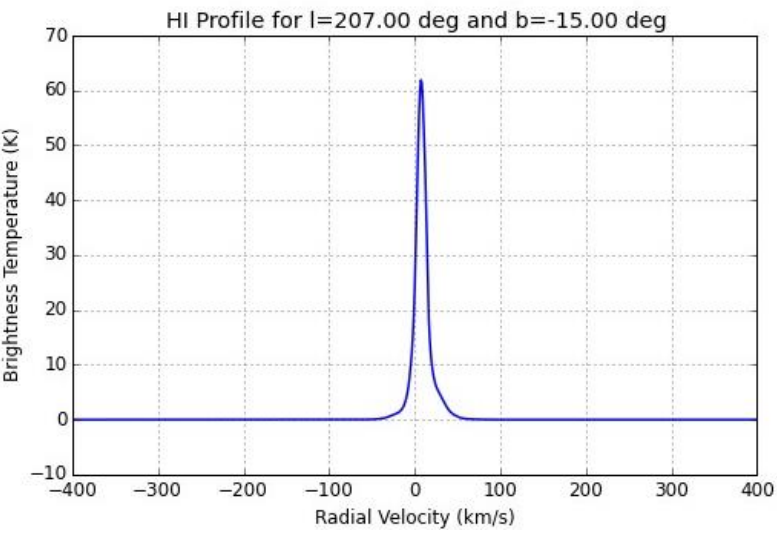
- Take multiple acquisitions and stack to increase SNR
- Get all important information needed from FITS header
- Pay attention to column headers and rows
- Make use of Astropy from your JupyterHub
- Helpful commands in prac manual



Brightness Temperature

Select Position		
Coordinate system	Galactic coordinates (l, b) ▼	
Position	RA [h m s]/ l [°]	207
	DEC [±° ' "]/ b [°]	-15
Effective beamsize FWHM [°]		4.03
Velocity Window for display	Minimum [km/s]	-400.00
	Maximum [km/s]	400.00
<input type="button" value="Search data"/>		

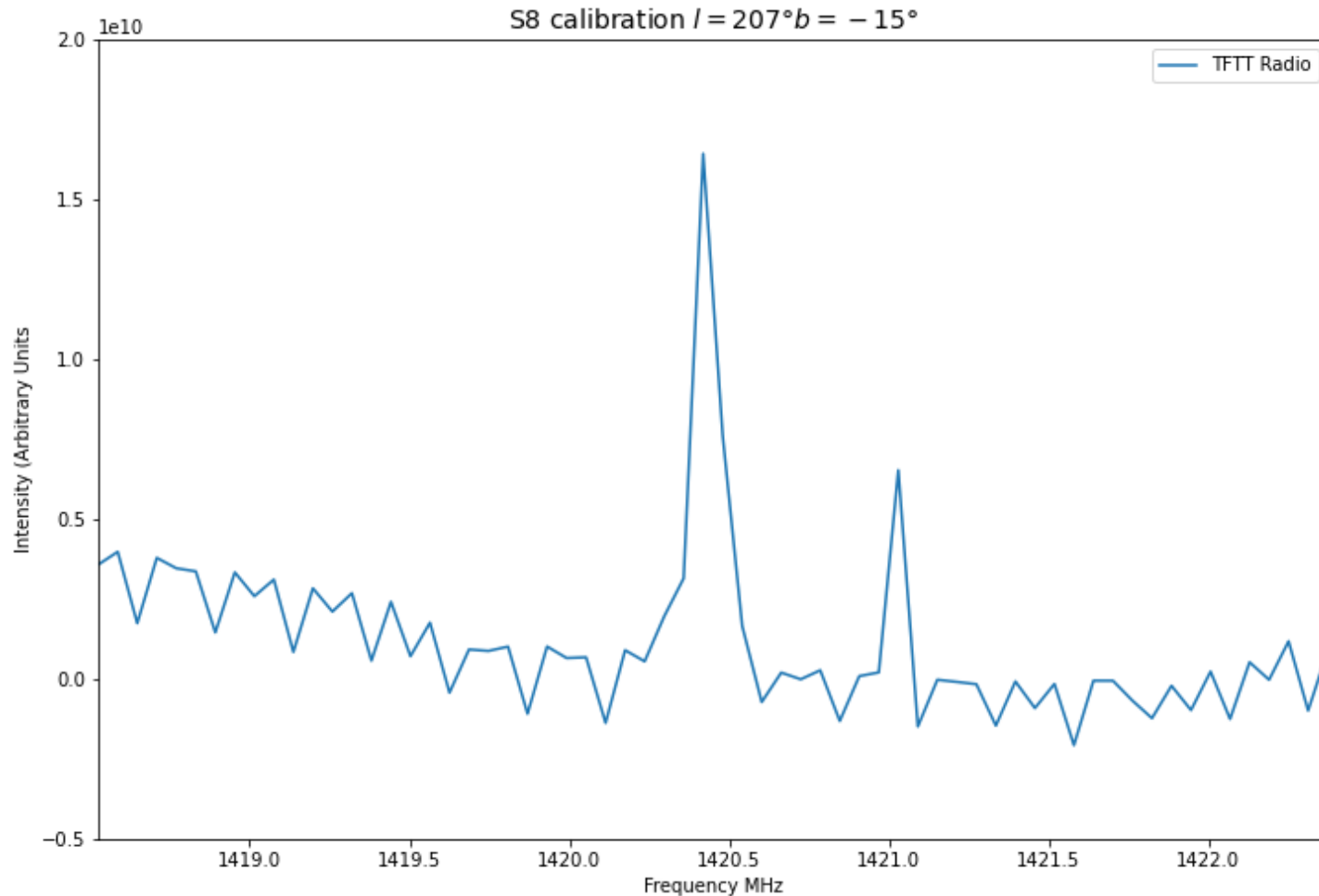
Result



- Use IAU S8 data as source of known brightness
- Enter location into University of Bonn HI Profile Search
- Download HI profile



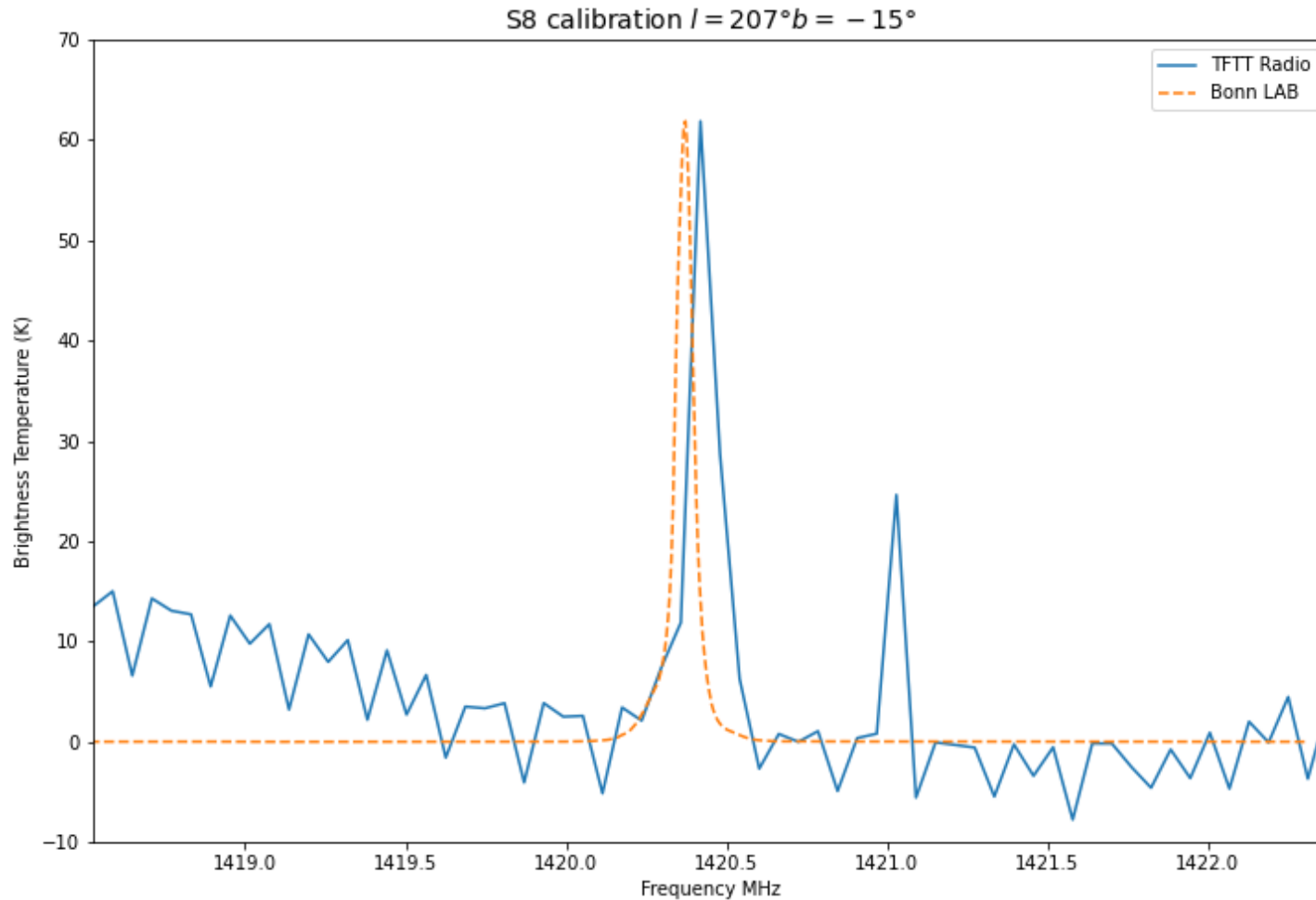
Brightness Temperature



- Use IAU S8 data as source of known brightness
- Enter location into University of Bonn HI Profile Search
- Download HI profile
- Use the mean of your data around HI line to set floor



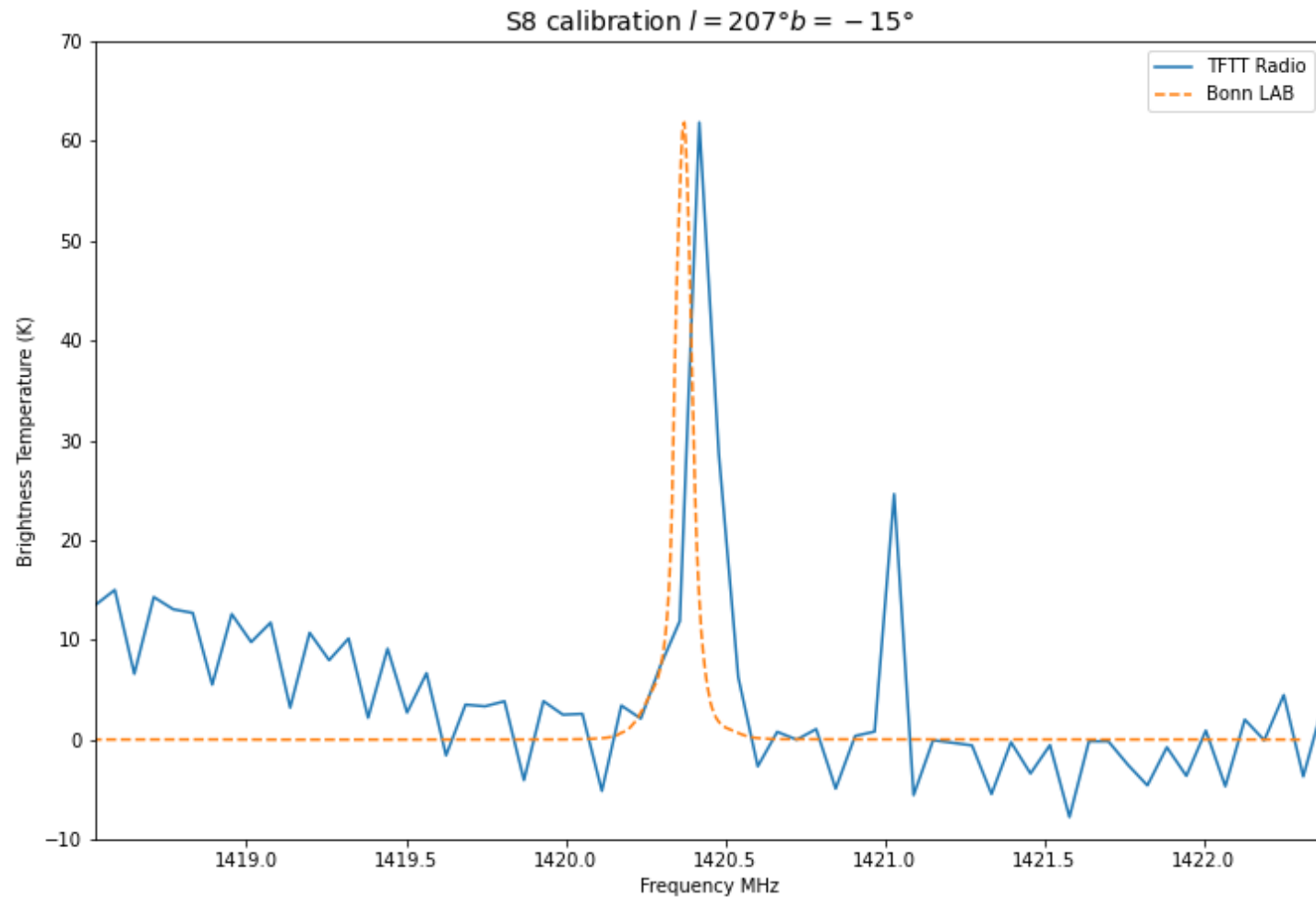
Brightness Temperature



- Use IAU S8 data as source of known brightness
- Enter location into University of Bonn HI Profile Search
- Download HI profile
- Use the mean of your data around HI line to set floor
- Use Bonn profile to determine scaling factor
- Apply to all spectra

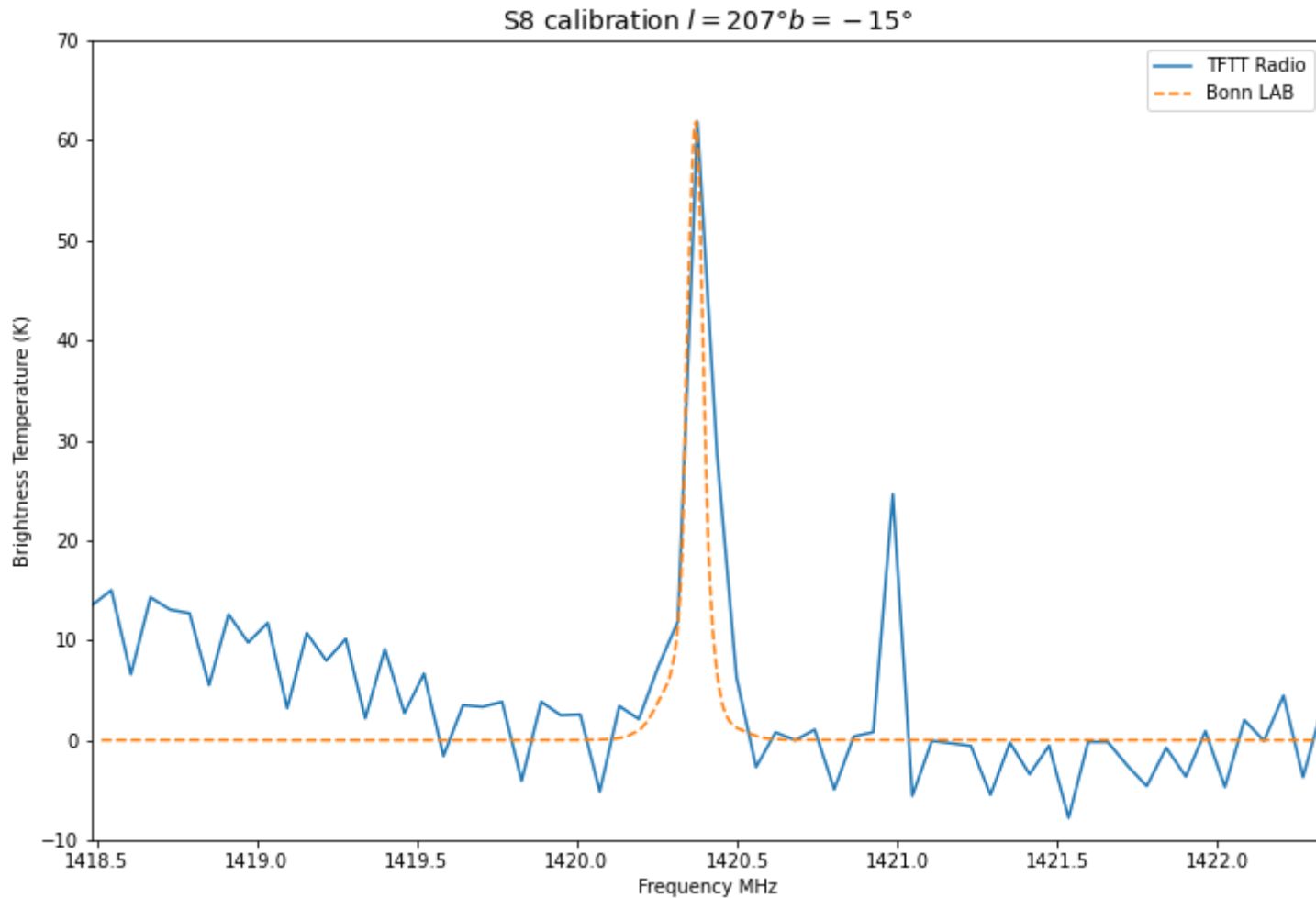


LSR Velocity



- Correct for Earth's motion in direction of coordinates
- Enter required information in Python given script

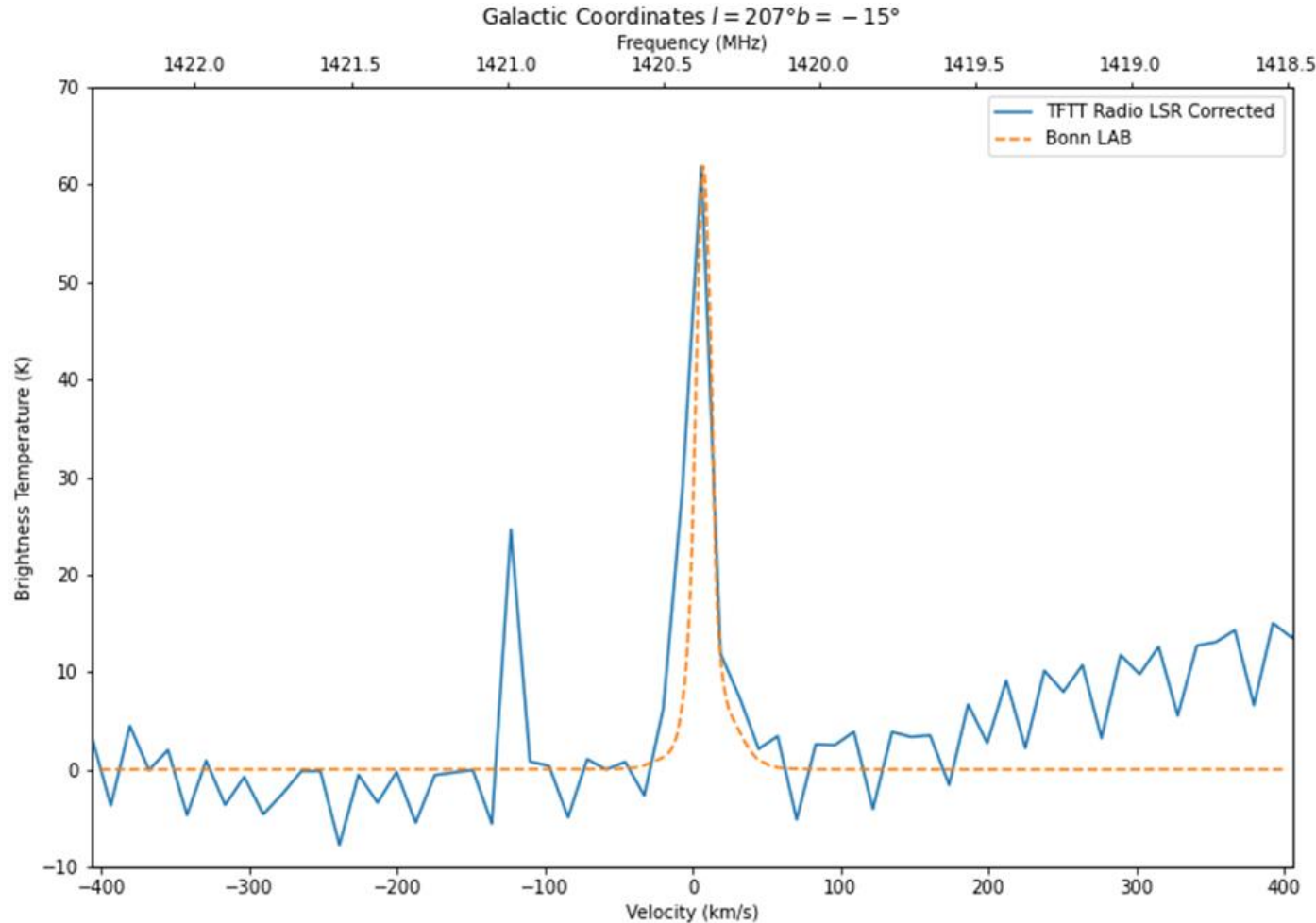
LSR Velocity



- Correct for Earth's motion in direction of coordinates
- Enter required information in Python given script
- Use observed and apparent rest frequency to correct LSRV



LSR Velocity

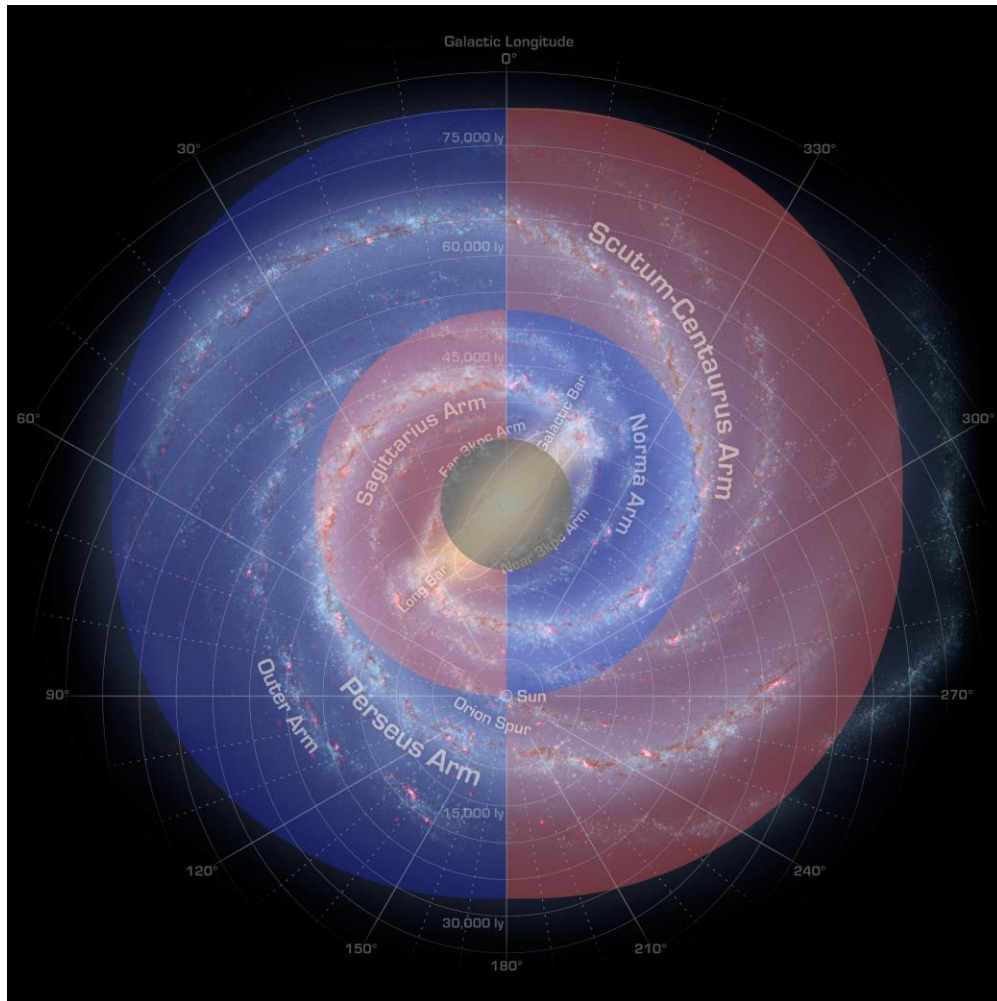


- Correct for Earth's motion in direction of coordinates
- Enter required information in Python given script
- Use observed and apparent rest frequency to correct LSRV
- Covert frequency to radial velocity scale using

$$v = c \frac{\Delta f}{f_0}$$

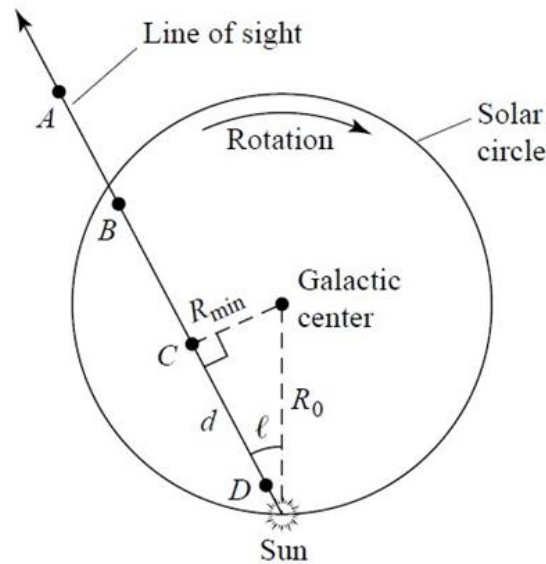
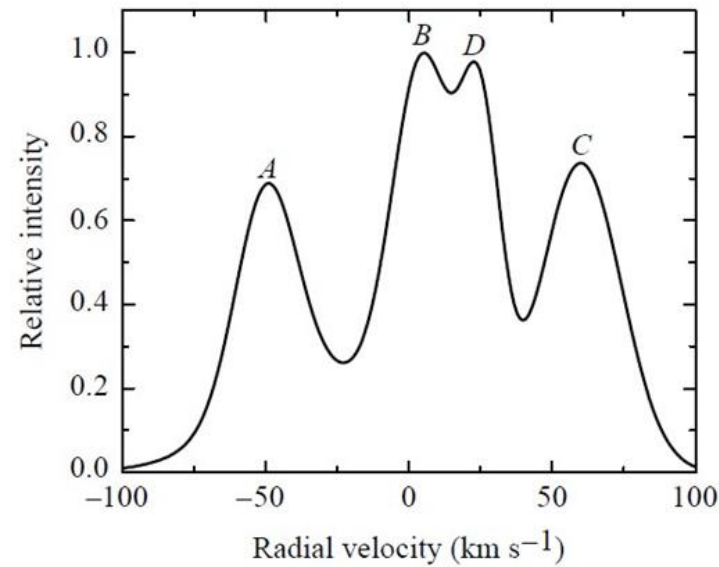


Rotational Velocity & Distance



- Rotational velocity decreases with increased distance from center
- Longitudes between 0° and 90° are redshifted within Sun's orbital radius and blueshifted between 270° and 0°

Rotational Velocity & Distance



- Rotational velocity decreases with increased distance from center
- Longitudes between 0° and 90° are redshifted within Sun's orbital radius and blueshifted between 270° and 0°
- The maximum velocity in this region point tangent to orbital circle
- Distance to point is

$$R_{\min} = R_0 \sin(\ell)$$

Report

- Method
 - Explain your observations and every step of your practical
 - Use paragraphs and not lists
 - Avoid explaining unimportant details (*e.g. “then I clicked open”*)
- Results
 - Report all final results obtained and any calculations done
- Discussion
 - Interpret and explain your results
 - Compare your results to other findings if appropriate
- Conclusion



Report

- References
 - Reference all information from sources which are not your own (including images)
 - Use a known referencing style (e.g. Harvard)
 - Make sure references are cited within your report
- Additional comments
 - Observations are done in groups but each person must submit their own report
 - All figures and tables should be numbered and captioned
 - Plagiarism will be heavily penalized
 - Remember significant figures and errors
 - Submit in PDF form
 - If you have any questions email me (bret.yotti@uct.ac.za)

