

# The fuzzball proposal for black holes

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# Theoretical Problems

- Information paradox
- Microstate problem

Black holes can be assigned an entropy  $S = \frac{A_{hor}}{4G}$

Corresponding microstates:  $e^{\frac{S}{k_b}}$ . Huge number.

Microstates correspond to metrics: due to the no hair theorem there is only one metric!

# The fuzzball conjecture

Both GR and QM are involved, therefore we can try to solve the problem by using String Theory.

→ Fuzzball conjecture (Mathur,2002)

Fuzzballs are solutions that describe black holes in String Theory.

# Structure of the fuzzball

The inner region of black holes is replaced by a ball of strings of radius  $R_S$ , with no event horizon.

- $r > R_S$ : Objects that are far from the surface are affected in the same way as with the classical model of the black holes.
- $r < R_S$ : We begin to observe stringy behaviours. In this region we have different metrics corresponding to different vibration profiles of the strings. These metrics describe different microstates!

Only a restricted set of microstates is explicitly known. The aim of my thesis was to find new metrics.

The AdS/CFT correspondence plays a major role. In the near horizon limit the spacetime contains an  $AdS_3$  factor, therefore we can work with a 2 dimensional CFT instead.

The CFT is weakly coupled, therefore the calculations are way more manageable.

# My contribution

New solutions can be found by applying diffeomorphisms on a known metric, representing  $AdS_3 \times S_3 \times T_4$ .

On the CFT side, this corresponds to applying operators to the state corresponding to  $AdS_3 \times S_3 \times T_4$ .

By using this method, I could find new microstates. The next step consists in generalizing my results in order to obtain microstates that cannot be obtained with this method.

Open problems:

- In what microstates are black holes typically found?
- When do fuzzballs produce observable deviations from a typical black hole?

Thank you for your attention!