

# SOLUTIONS FOR EQUILIBRIUM OF MODEL GALAXIES

Prescribe  
temperature  $T(R, Z)$   
(Eq. (8)) and  
potential  $\phi(R, Z)$   
(Eqs. (9) & (10))

Solve for  $v_\phi(R, Z)$   
(Eq. (6))  
assuming  
that  $v_\phi \rightarrow 0$  at  $\infty$

- ✓ Obtain  $v_\phi(R)|_{Z=0}$  from last step
- ✓ Create a 1D interpolation table of  $v_\phi(R)|_{Z=0}$

- ✓ Numerically integrate the integrands in the exponential terms in Eq. (5) using the analytical forms of  $\frac{\partial \phi}{\partial R}$  and  $\frac{\partial \phi}{\partial Z}$  and  $v_\phi$  from interpolation table
- ✓ Obtain the density field  $\rho(R, Z)$  as in Eq. (5), with the choice of value of  $\rho(0, 0)$ .

Set the pressure  $P(R, Z) = \rho(R, Z) \times c_s^2(R, Z)$   
where,  $c_s^2(R, Z) = \frac{k_B T}{\mu m_p}$