

## Binet

Determinar fuerza, potencial, aceleración radial y tangencial

$$r = a(1 + \cos\phi)$$

$$(a) \quad \bar{r} \quad ; \quad u = \frac{1}{a(1 + \cos\phi)} \Rightarrow u' = \frac{\sin\phi}{a(1 + \cos\phi)^2} = a \sin\phi u^2$$

$$u'' = a \cos\phi u^2 + 2a \sin\phi u u' = a \cos\phi u^2 + 2a^2 \sin^2\phi u^3$$

$$* \quad u = \frac{1}{a(1 + \cos\phi)} \Rightarrow 1 + \cos\phi = \frac{1}{ua} \Rightarrow \cos\phi = \frac{1}{ua} - 1$$
$$\sin^2\phi = 1 - \cos^2\phi$$

$$\therefore u'' = au^2 \left( \frac{1}{ua} - 1 \right) + 2a^2 u^3 \left( 1 - \frac{1}{a^2 u^2} + \frac{2}{au} - 1 \right)$$

$$= u - au^2 + 2a^2 u^3 - 2u + 4au^2 - 2a^2 u^3 = 3au^2 - u$$

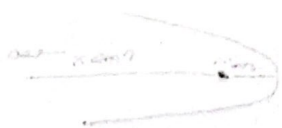
$$\therefore u^2 u'' = 3au^4 - u^3 \Rightarrow u^2 u'' + u^3 = 3au^4$$

$$\therefore \bar{F}(u) = -\frac{l^2}{m} (3au^4) \Rightarrow \bar{F}(r) = -3a \frac{l^2}{m} \frac{1}{r^4}$$

$$b) \quad \frac{dV}{dr} = 3a \frac{l^2}{m} \frac{1}{r^4} \Rightarrow V(r) = 3a \frac{l^2}{m} \int_{r_0}^r r^{-4} dr = 3a \frac{l^2}{m} \left( -\frac{1}{3r^3} \right)_{r_0}^r$$

$$\Rightarrow V(r) = -a \frac{l^2}{m} \frac{1}{r^3} + C \quad ; \quad C, \text{cte integración}$$

$$c) \quad a_c = \frac{\bar{F}}{m} = -\frac{3a l^2}{m^2} \frac{1}{r^4} \quad ; \quad a_t = 0$$



$$(a) \quad u = \frac{1}{a(1 + \cos\phi)}$$

# Potencial Gravitacional

Gráficar  $V_{eff}(r)$  para el potencial gravitacional

$$V(r) = - \frac{GMm}{r}$$

$$* V_{eff}(r) = V(r) + \frac{l^2}{2mr^2} = - \frac{GMm}{r} + \frac{l^2}{2mr^2}$$

$$* E - V_{eff}(r) = \frac{1}{2} m \dot{r}^2 \geq 0$$

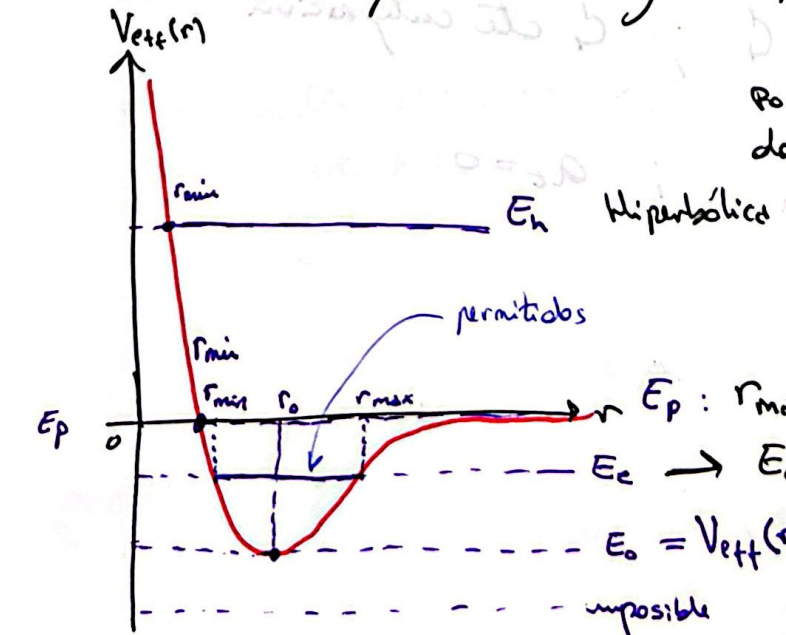
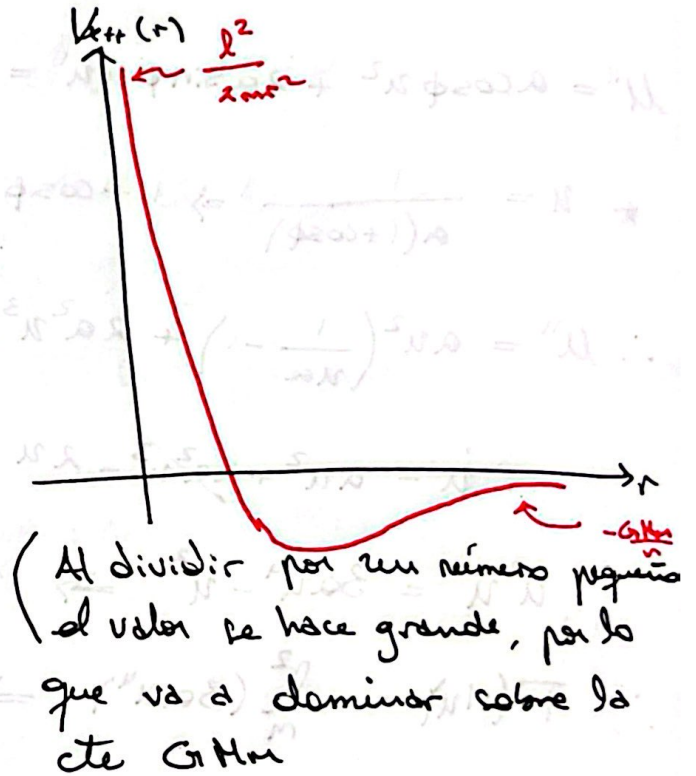
$$\therefore V_{eff}(r) = \frac{1}{r} \left( \frac{l^2}{2mr} - GMm \right)$$

Para graficar, se toman los casos límites donde  $r \rightarrow 0$  y  $r \rightarrow \infty$

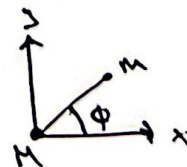
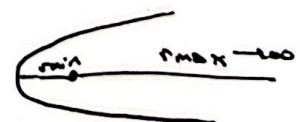
$$\therefore V_{eff} \rightarrow 0 \rightarrow V_{eff} \sim \frac{l^2}{2mr^2} \rightarrow \infty$$

$$V_{eff} \rightarrow \infty \rightarrow - \frac{GMm}{r} \rightarrow 0$$

(La cte  $GMm$  será mayor que algo que se acerca a 0, además se acerca a 0 por el lado negativo)



$\therefore E - V_{eff}(r) \geq 0$   
Posiciones radiales permitidas son aquellas donde la línea  $E$  está por encima o tocando la curva  $V_{eff}(r)$





Condición órbita circular, r. punto de equilibrio si la fuerza radial neta es cero;

$$\left. \frac{dV_{\text{eff}}}{dr} \right|_{r=r_0} = 0 \quad ; \quad V_{\text{eff}} = -\frac{GMm}{r} + \frac{l^2}{2mr^2}$$

$$\Rightarrow \frac{dV_{\text{eff}}}{dr} = \frac{GMm}{r^2} - \frac{l^2}{mr^3} \xrightarrow{\text{Cond. eq.}} \frac{GMm}{r_0^2} - \frac{l^2}{mr_0^3} = 0$$

$$\Rightarrow GMm = \frac{l^2}{mr_0} \Rightarrow \boxed{r_0 = \frac{l^2}{GMm^2}} \Rightarrow \boxed{l^2 = GMm^2 r_0}$$

\* Fuerza :  $F_{\text{eff}}(r) = -\frac{dV_{\text{eff}}}{dr} \Rightarrow F_{\text{eff}}(r) = -\frac{GMm}{r^2} + \frac{l^2}{mr^3}$

\* Periodo :  $V'_{\text{eff}}(r_0) > 0 \quad ; \quad V'_{\text{eff}}(r_0) = k, \quad T = 2\pi\sqrt{\frac{m}{k}}$

$$\text{Si: } V'_{\text{eff}}(r) = \frac{GMm}{r^2} - \frac{l^2}{mr^3} \Rightarrow V''_{\text{eff}}(r) = -\frac{2GMm}{r^3} + \frac{3l^2}{mr^4}$$

$$\Rightarrow V''_{\text{eff}}(r_0) = -\frac{2GMm}{r_0^3} + \frac{3}{mr_0^4} GMm \cancel{r_0} = -\frac{2GMm}{r_0^3} + \frac{3GMm}{r_0^3}$$

$$V''_{\text{eff}}(r_0) = \frac{GMm}{r_0^3} \Rightarrow \boxed{T = 2\pi\sqrt{\frac{r_0^3}{GM}}}$$