

Science Olympiad
UT Invitational

October 25, 2025

Astronomy C Answer Key



ANSWER KEY ANSWER KEY
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Section A [23 points]

1. [1 pt] D
 2. [2 pts] White dwarf : VII, Giant : III, Main sequence : V, Subgiant : IV, Supergiant : I (**1 pt for 2/5**)
 3. [1 pt] White dwarf
 4. [1 pt] Neutron star
 5. [2 pts] Microwave, Infrared, Visible, Far ultraviolet, Gamma (**1 pt for 4 in right order**)
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6. [1 pt] B
7. [1 pt] F
8. [1 pt] I
9. [1 pt] A
10. [1 pt] Main sequence
11. [2 pts] Yes (**1 pt**), far bottom right (or off the diagram) (**1 pt**) (Do NOT accept the region around A)
12. [2 pts] x : color/spectral classification, $B - V$ index (**1 pt**); y : abs./app. magnitude, flux, intensity (**1 pt**)

13. [1 pt] Blackbody
14. [2 pts] Absorption line (**1 pt**), photons at specific energies are absorbed by an intervening, cool medium (**1 pt**)
15. [2 pts] Object A (**1 pt**). Explanation mentions qualitative difference in spectra (e.g., the wavelength at peak intensity of A is shorter than B; OR, the total luminosity of A is greater than B) (**0.5 pts**) and connects it to a ‘law’ (e.g., Wien’s; OR, Stefan–Boltzmann [resp.]) (**0.5 pts**)
16. [2 pts] Blackbody spectra should not intersect (**1 pt**). As a blackbody becomes warmer, it radiates more light at all wavelengths (**1 pt**). (Do NOT accept any mention of the two spectra having ‘different shapes’)

Section B [25 points]

17. [1 pt] Orion Molecular Complex
18. [1 pt] Lambda Orionis Ring
19. [1 pt] Barnard's Loop
20. [1 pt] Expansion of ejecta (**0.5 pts**) from a supernova (**0.5 pts**)
21. [2 pts] Infrared (**1 pt**), stellar formation is often obscured by gas and dust in the visible spectrum, but infrared light can shine through (**1 pt**) (Only half credit for explanation if missing mention of 'dust')
22. [2 pts] 10–30 million years old (**1 pt**). The stars and gas will disperse into the field (**1 pt**) (Half credit for explanation mentioning there will be less gas)

23. [1 pt] Molecular clouds/stellar nurseries (Do NOT accept cloud complex)
24. [2 pts] Kinetic (or thermal) energy/heat (Also accept rotational energy)
25. [2 pts] Does not have a coherent “core” velocity structure; OR, the stars move at a rapid pace away from each other.
26. [1 pt] CO
27. [1 pt] Radio/millimeter/microwave (Do NOT accept infrared)

28. [3 pts] 0.766 (Accept 76.6 %)
29. [3 pts] Cluster B is older (**1 pt**). Observe that cluster B has a lower ratio of high mass stars (**1 pt**). Since high-mass stars live shorter lives than low-mass stars, the older cluster would have a lower ratio of high mass stars (**1 pts**).

30. [1 pt] Image 15
31. [1 pt] T Tauri. Half credit for young stellar object (YSO) or pre-main sequence (PMS) star
32. [2 pts] Star/protostar, accretion/protoplanetary disk, bipolar flow/jet, starspot

Section C [41 points]

33. [1 pt] PSR B1509-58
34. [1 pt] Pulsar, pulsar wind nebula, neutron star. Half credit for nebula.
35. [2 pts] 15:13:[30, 80], -59:[06:xx, 09:xx]. Half credit for coordinates slightly outside of the accepted range
36. [2 pts] [1400, 1800] eV
37. [2 pts] [10000, 13500] counts
38. [2 pts] Additional emission lines on the continuous peak
39. [1 pt] Dimmer
40. [2 pts] Material from the outflowing jets (**1 pt**) is colliding with the interstellar medium (**1 pt**)

41. [1 pt] Helix Nebula
42. [1 pt] 200
43. [2 pts] Giant stars shed their outer layer (**1 pt**) as radiation pressure overcomes the weakened gravity. This stellar wind glows because it is eventually ionized (**1 pt**) by the exposed core of the star.
44. [2 pts] White dwarf (**1 pt**); carbon-oxygen (or oxygen-neon) (**1 pt**)

Section C continued...

45. [1 pt] B
46. [1 pt] Crab Nebula/M1/Crab Pulsar
47. [2 pts] Chandra, Hubble, and Spitzer (**1 pt for 1/3**)
48. [1 pt] It was created by a supernova in 1054 (which was observed by astronomers of that time!), which makes it 971 years old today.
49. [2 pts] It has a rotational period of 33.4 milliseconds, meaning that it completes 29.94 revolutions per second.
50. [3 pts] These pulsars are called millisecond pulsars (or recycled pulsars) (**1 pt**), and they spin much faster because they are accreting material from a companion star (**1 pt**) in a close binary system (point given for anything mentioning the two stars orbiting closely, e.g. Roche lobe overflow) (**1 pt**)
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51. [1 pt] Image 14
52. [1 pt] Type IIb
53. [1 pt] Observation of a scattered light echo
54. [1 pt] Hydrogen absorption lines
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55. [5 pts] (**0.5 pts for each valid/invalid**)
A: Invalid, the stars must move in an elliptical orbit with the center of mass at one focus (**1 pt**)
B: Valid, 1:1 (**1.5 pts**)
C: Invalid, the stars are in the wrong phase or center of mass at wrong location (**1 pt**)
56. [1 pt] Kepler's third law
57. [1 pt] WD J181058.67+311940.94
58. [1 pt] Gravitational wave radiation

Section D [16 points]

59. [2 pts] 5800 [5250, 6450]. Use Wien's law
60. [3 pts] 6.96×10^5 [5.62, 8.70]. Use magnitude to luminosity (notice it is equal to the Sun) and Stefan–Boltzmann
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61. [2 pts] 5 (Exact). Use $d = 1/p$. (Half credit for not converting mas to arcsec)
62. [3 pts] Use dist. modulus (**1 pt**) (half credit for attempt). Correct M (**1.5 pts**) and $\log_{10} P$ (**1.5 pts**)

Star	M	$\log_{10} P$
Jeff	-4.1	0.699
Britta	-2.9	1.30
Abed	-2.1	1.74
Troy	-1.2	2.28

63. [2 pts] Plotted correctly (**1 pt**), axes labels (**0.5 pts**), tick marks (**0.5 pts**)
64. [3 pts] Draw a best fit line (**1 pt**), $\alpha = 1.84$ [1.65, 2.03] (**1.5 pts**), $\beta = -5.34$ [-5.87, -4.81] (**1.5 pts**)
65. [1 pt] Type Ia supernova