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ST

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2.1. $\frac{\Delta P}{P_0}$ where $\Delta P = \frac{M}{\frac{4}{3}\pi R^3}$ and $P_0 = 2.7 \times 10^{-11} \text{ dy/m}^3$

a) $M = 100 \text{ kg}$, $R = 1 \text{ m}$ from a human

$$\frac{\Delta P}{P_0} = \frac{\frac{100}{\frac{4}{3}\pi R^3}}{2.7 \times 10^{-11}} = 8.84 \times 10^{21} \approx 10^{21}$$

b) $M = 2 \times 10^{30} \text{ kg}$, $R = 1.5 \times 10^8 \text{ m} = 1 \text{ AU}$ For the Sun

$$\frac{\Delta P}{P_0} = \frac{\frac{2 \times 10^{30}}{\frac{4}{3}\pi (1.5 \times 10^8)^3}}{2.7 \times 10^{-11}} = 5.24 \times 10^{22} \approx 10^{22}$$

c) $M = 5 \times 10^{11} M_\odot$, $R = 1 \text{ Mpc} = 3.1 \times 10^{22} \text{ m}$ For the Milky Way and Andromeda

$$\frac{\Delta P}{P_0} = \frac{\frac{5 \times 10^{11} \times 2 \times 10^{30}}{\frac{4}{3}\pi (3.1 \times 10^{22})^3}}{2.7 \times 10^{-11}} = 2.47 \approx 10^1$$

d) $M = P_0 \times \frac{\frac{4}{3}\pi (100 \text{ m})^3}{2.7 \times 10^{-11}} = \frac{100 \times 4 \times 3 \times (1 \times 10^{-2})^3}{2.7 \times 10^{-11}} = 3.33 \times 10^{45} \text{ kg}$
 $\frac{\Delta P}{P_0} = \frac{\frac{3.33 \times 10^{45}}{\frac{4}{3}\pi (3.1 \times 10^{24})^3}}{2.7 \times 10^{-11}} \approx 0.75 \approx 10^0$

2.2. If $H_0 = 70 \text{ km/s/Mpc}$,

$$R_0 = \frac{c}{H_0} = \frac{\text{km/s}}{\text{km/s/Mpc}} = \text{Mpc} = \frac{3 \times 10^5}{70} = 4300 \text{ Mpc} = 4.3 \text{ Gpc}$$

$$T_0 = \frac{1}{H_0} = \frac{1}{70 \text{ km/s/Mpc}} = \frac{1}{70 \text{ km/s/Mpc}} \cdot \frac{3.1 \times 10^{19} \text{ km}}{\text{Mpc}} = 4.42 \times 10^{17} \text{ s}$$

$$= 1.4 \times 10^{10} \text{ years}$$

2.2(e) ΔP \wedge $M = 5 M_\odot$ black hole

$$r = r_{sh} = \frac{2 \times 5 M_\odot}{c^2} = \frac{2 \times 6.67 \times 10^{11} \times 5 \times 2 \times 10^{30}}{(3 \times 10^8)^2} = 1.5 \times 10^4 \text{ m}$$

$$\frac{\Delta P}{P_0} = \frac{\frac{5 \times 2 \times 10^{30}}{\frac{4}{3}\pi (1.5 \times 10^4)^3}}{2.7 \times 10^{-11}} = 2.62 \times 10^{44} \approx 10^{44}$$

$$2.3 \text{ a) Wien's Law: } \frac{hc}{\lambda p} \approx 0.99 kT \rightarrow \lambda p = \frac{hc}{0.99 kT}$$

$$= \frac{6.63 \times 10^{-34} \text{ J m}^2 \text{ K}^{-5}}{4.98 \times 1.38 \times 10^{-23} \text{ J m}^2 \text{ K}^{-4} \text{ K}^{-1}} \cdot 3 \times 10^9 \text{ m s}^{-1} \approx \frac{0.0029 \text{ m K}}{T} = \frac{0.29 \text{ cm} \cdot \text{K}}{T} = \lambda$$

$$T_{\odot} = 5778 \text{ K} \rightarrow \lambda_p = \frac{0.29 \text{ cm} \cdot \text{K}}{5778 \text{ K}} \approx 5.02 \times 10^{-5} \text{ cm}$$

This is the wavelength of green visible light, the color of the sun.

$$\text{b) } \bar{E}_{\text{mean}} = \frac{\epsilon \tau}{h \nu} = \frac{\pi^2 k^4}{15 h^3 c^3} T^4 = \frac{\pi^4 k T}{15 \cdot 2 \cdot 4} = \frac{\pi^4 k T}{36} =$$

$$2.7 kT \quad T_0 = 5778 \text{ K} \text{ for He Sun}$$

$$\bar{E}_{\text{mean}} = 2.7 \times 1.38 \times 10^{-23} \text{ J K}^{-1} \cdot 5778 \text{ K} \approx 2.15 \times 10^{-19} \text{ J}$$

$$T = 2.73 \text{ K for CMB}$$

$$\bar{E}_{\text{mean CMB}} = 2.7 \times 1.38 \times 10^{-23} \text{ J K}^{-1} \cdot 2.73 \text{ K} \approx 1.02 \times 10^{-22} \text{ J}$$

Extra credit: If half the temperature of the sun can ionize hydrogen, and both the sun and CMB are subject to the same hydrogen ionization physics, the redshift at which the CMB was generated is very high. The current observed temperature of the CMB is much less than half the sun's temperature, but as it is known that the universe was opaque with electrons ionized from hydrogen or helium interacting with photons, the temperature must have been much higher, also supported by the Planck temperature. To have such a great decrease in temperature then, the CMB must be high. Wien's law states $\lambda = \frac{c}{T}$ so high redshift (redshift) would increase the wavelength, and so lower the temperature observed.

Using the equation $T = T_0(1+z)$, where $T_0 = 273 \text{ K}$ and $T = 300 \text{ K}$, $z = 1090$ for the CMB, as expected +2
high