

Lec 1 Sep 7th

at the end of this course ask the q: Is the gradient a vector?

This course is useful in 2 aspects

- ① General Relativity needs diff geo because of curved space and time
- ② learn powerful and beautiful tools to describe physics in any geometry

Chapter 1 - on some basic mathematics

We need some maths to be able to define a manifold (Thursday of next week)

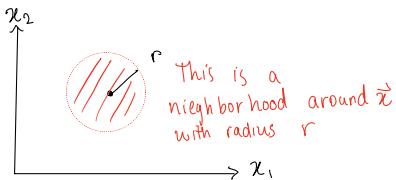
1.1 \mathbb{R}^n and its topology

A point in \mathbb{R}^n is an n-tuple $(x_1, \dots, x_n) \in \mathbb{R}^n$. The idea of continuous is that any 2 points in \mathbb{R}^n have a line connecting them that exists in \mathbb{R}^n .

ex. Integers are not continuous (discrete)

The continuity of a space defines its topology. Here we focus on local vs global topology. We use distance to define the topology. Recall, the distance between $\vec{x}, \vec{y} \in \mathbb{R}^n$ is $d(\vec{x}, \vec{y}) = \sqrt{(x_1 - y_1)^2 + \dots + (x_n - y_n)^2}$

A neighborhood of radius r of $\vec{x} \in \mathbb{R}^n$ is the set of points s.t. $N_r(\vec{x}) = \{ \vec{y} \in \mathbb{R}^n, d(\vec{x}, \vec{y}) < r \}$



A set of points in \mathbb{R}^n are discrete if there exists a neighbourhood about each point that contains no other points.

A set of points $S \subseteq \mathbb{R}^n$ is open $\Leftrightarrow \forall x \in S \exists$ a neighbourhood all in S

Example ① $S = \{x | a < x < b\}$ is open

② $S = \{x | a \leq x \leq b\}$ is not open. because $x=a$ does not have a neighbourhood all within S

Note: Open sets cannot contain boundary points.

\mathbb{R}^n has the Hausdorff property, which means that any 2 points in \mathbb{R}^n have neighbourhoods that do not intersect. $d(\vec{x}, \vec{y})$ induces a topology on \mathbb{R}^n , which says that d determines whether a set is open or not.

Open sets have the following properties:

- ① empty set \emptyset and the whole set S are open
- ② If O_1, O_2 are open sets the $O_1 \cap O_2$ is open
- ③ The union of open sets (finite number) is open

The topology of a set consists of the set and all the open sets in that set.

Any distance function induces the natural topology in \mathbb{R}^n

For example, $d'(\vec{x}, \vec{y}) = [4(x_1 - y_1)^2 + (x_2 - y_2)^2 + \dots + (x_n - y_n)^2]^{1/2}$ has the same induced topology as any other distance function. You can define a topology without distance.

§1.2 Mappings

A map from M to N associates an element $x \in M$ to a unique $y \in N$



S is a subset of M and the image of S under f is $f(S) = T$. The inverse image of T is $f^{-1}(T) = S$

f can be many to one. If all points in $f(S)$ have a unique inverse in S then f is 1-1 and \exists a one-to-one map f^{-1} called the inverse of f

Example: $\sin(x)$ is many to one b/c $\sin(x) = \sin(x+2\pi)$

Notation: $f: M \rightarrow N$ f maps $M \rightarrow N$
 $f: x \mapsto y$ f maps x to y

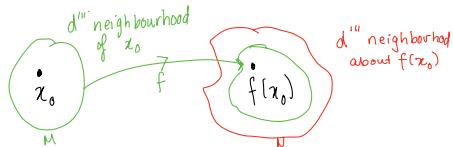
Given $f: M \rightarrow N$ and $g: N \rightarrow P$ then \exists a composition map $g \circ f: M \rightarrow P$ such that $(g \circ f)(x) = g(f(x))$. If $f: M \rightarrow N$ then
f defined \forall points in $M \Rightarrow f$ maps M into N
f defined \forall points in $N \Rightarrow f$ maps M onto N
if f is both 1-1 and onto then f is a bijection. If f have an inverse, then f is 1-1

New [A map $f: M \rightarrow N$ is continuous at $x \in M$, if any open set in N containing $f(x)$ contains the image of an open set in M contains x .

f is continuous on M if it is continuous $\forall x \in M$

old [Look at how this is related to continuous defined in calculus. Recall, f is continuous at x_0 if $\forall \epsilon > 0 \exists \delta > 0$ s.t
 $|f(x) - f(x_0)| < \epsilon$ if $|x - x_0| < \delta$

We define $d'''(x, x_0) = |x - x_0|$ then our definition can be rewritten as follows: f is continuous at x_0 if $\forall d'''$ -neighbourhoods of $f(x_0)$ contains the image of a d''' neighbourhood of x_0



Theorem: $f: M \rightarrow N$ is continuous iff the inverse image of every open set is open in M

Lec 2 Sep 12

Official Assignment will be released tonight. Crowdmark link will be sent out. A MATH 433?

Today's Topics: ① Real analysis ② Group theory ③ Linear Algebra ④ Algebra of Square Matrices

§ 1.3 Real analysis

$f(x)$ is analytic at $x=x_0$ if it has a Taylor expansion about x_0 with a non-zero radius of convergence

Analytic functions (C^∞) which is a subset of C^ω

We will assume functions are analytic, but we'll often say smooth (C^∞)

An operator A of functions is a map that takes a function and yields another function

Examples: $A(f) = g(f)$, g is a function

$$D(f) = \frac{df}{dx} \text{ where } f \in C^1$$

The commutator of 2 operators A, B on f is $[A, B](f) = (AB - BA)f$ or $A(B(f)) - B(A(f))$. If $[A, B] = 0$ functions then A and B commute

Example $A = \frac{d}{dx}$ and $B = x \frac{d}{dx}$

$$\begin{aligned} [A, B](f) &= \frac{d}{dx} \left(x \frac{df}{dx} \right) - x \frac{d}{dx} \left(\frac{df}{dx} \right) \\ &= x \frac{d^2 f}{dx^2} + \frac{df}{dx} - x \frac{d^2 f}{dx^2} = \frac{df}{dx} \end{aligned} \quad \text{hence } A \text{ & } B \text{ do not commute}$$

function spaces don't have
a good intuition on the commutator
but in other contexts there is, such as
Later in the course

§ 1.4 Group theory

A set of elements G with a binary operation \cdot is a group if

$$[G_i] \text{ Associative: } x \cdot (y \cdot z) = (x \cdot y) \cdot z$$

$$[G_{ii}] \text{ Identity: } \exists e \in G \text{ such that } x \cdot e = e \cdot x = x \quad \forall x \in G$$

$$[G_{iii}] \text{ Inverse: } \forall x \in G \exists x^{-1} \in G \text{ s.t. } x \cdot x^{-1} = x^{-1} \cdot x = e$$

And is closed under the operation

A group is abelian (commutative) if

$$[G_{iv}] x \cdot y = y \cdot x$$

Example: ① Set of permutation of n objects

② Rotations of a regular polygon

Aside: the inverse is unique, and the identity is unique

A subgroup is simply a group that is contained within the group

Example: a set of permutation of n objects where the first element is unchanged.

This is (identical) similar to the permutations of $(n-1)$ objects [isomorphism]

§ 1.5 Linear Algebra

A set V is a vector space (over \mathbb{R}) if it has a binary operation $+$ where it is an abelian group and satisfies the following under Multiplication. Let $\vec{x}, \vec{y} \in V$, $a, b \in \mathbb{R}$

$$[V_i] \quad a \cdot (\vec{x} + \vec{y}) = (a \cdot \vec{x}) + (a \cdot \vec{y})$$

$$[V_{ii}] \quad (a+b) \cdot \vec{x} = (a \cdot \vec{x}) + (b \cdot \vec{x})$$

$$[V_{iii}] \quad (ab) \cdot \vec{x} = a \cdot (b \cdot \vec{x})$$

$$[V_{iv}] \quad 1 \cdot \vec{x} = \vec{x}$$

The identity under addition is $\vec{0} = 0$

- Example
- (1) $n \times n$ matrices
 - (2) Continuous real function on $a \leq x \leq b$

* Dual spaces become critical in the next 2 weeks *

Notation; we often drop \cdot & (\cdot) and write $a\vec{x} + b\vec{y}$

A set $\{\vec{x}_1, \vec{x}_2, \dots, \vec{x}_n\}$ is linearly dependent if $\exists \{a_1, \dots, a_n\}$ $a_i \neq 0$ s.t. $a_1\vec{x}_1 + a_2\vec{x}_2 + \dots + a_n\vec{x}_n = 0$

If $\forall a_i = 0$ then the set is linearly independent

A vector space has a basis, which of the dimension of V , and allows us to generate any element of V

If \vec{x}_i , $i=1, \dots, n$ is a basis of V then $\forall \vec{y} \in V$, $\exists a_i$'s such that

$$\vec{y} = \sum_{i=1}^n a_i \vec{x}_i$$

A set of vectors $\{\vec{y}_1, \dots, \vec{y}_m\}$ generates a subspace of V with $a_1\vec{y}_1 + \dots + a_m\vec{y}_m$ $a_i \in \mathbb{R}$ $i=1, \dots, m$

if $m < n \Rightarrow$ proper subspace

A normed Vector Space is one w/ a mapping from V to \mathbb{R} s.t. $a \in \mathbb{R}$, $\vec{x}, \vec{y} \in V$

$$[N_i] \quad n(\vec{x}) \geq 0 \quad \& \quad n(\vec{x}) = 0 \text{ iff } \vec{x} = 0$$

$$[N_{ii}] \quad n(a\vec{x}) = |a| n(\vec{x})$$

$$[N_{iii}] \quad n(\vec{x} + \vec{y}) \leq n(\vec{x}) + n(\vec{y})$$

Examples

If $V = \mathbb{R}^n$ then

$$n(\vec{x}) = d(\vec{x}, 0) = [x_1^2 + \dots + x_n^2]^{\frac{1}{2}}$$

$$n'(\vec{x}) = d'(\vec{x}, 0) = [4x_1^2 + \dots + x_n^2]^{\frac{1}{4}}$$

$$n'''(\vec{x}), d'''(\vec{x}, 0) = \max(|x_1|, \dots, |x_n|)$$

All 3 satisfy N_i, N_{ii}, N_{iii} . In addition, some norms satisfy the parallelogram rule

$$[N_{iv}] \quad [n(\vec{x} + \vec{y})]^2 + [n(\vec{x} - \vec{y})]^2 = 2(n(\vec{x}))^2 + 2(n(\vec{y}))^2$$

n, n' satisfy N_{iv} by n'' does not

If we have all 24 properties then we can define a bilinear symmetric inner product.

$$\vec{x} \cdot \vec{y} = \frac{1}{4} [n(x+y)]^2 - \frac{1}{4} [n(\vec{x}-\vec{y})]^2$$

bilinear : $(a\vec{x} + b\vec{y}) \cdot \vec{z} = a(\vec{x} \cdot \vec{z}) + b(\vec{y} \cdot \vec{z})$
 $\vec{z} \cdot (a\vec{x} + b\vec{y}) = a(\vec{z} \cdot \vec{x}) + b(\vec{z} \cdot \vec{y})$

Symmetry : $\vec{x} \cdot \vec{y} = \vec{y} \cdot \vec{x}$

Positive definite $\vec{x} \cdot \vec{x} \geq 0$ and $\vec{x} \cdot \vec{x} = 0 \iff \vec{x} = \vec{0}$

$n(\vec{x})$ on \mathbb{R}^n is the Euclidean Norm, \mathbb{R}^n with the Euclidean norm is denoted E^n

A Pseudonorm is a norm that violates N_i & N_{iii}. This occurs in special relativity. What is the history of the Pseudo Norm?

§1.6 Algebra of Square Matrices

A linear transformation T on a vectorspace is a map from V onto V which is linear

$$T(a\vec{x} + b\vec{y}) = aT(\vec{x}) + bT(\vec{y})$$

If $\{\vec{e}_i\}_{i=1,\dots,n}$ is a basis for V , then $\vec{x} = \sum_{i=1}^n a_i \vec{e}_i$

and $T(\vec{x}) = T\left(\sum_{i=1}^n a_i \vec{e}_i\right) = \sum_{i=1}^n a_i T(\vec{e}_i)$ The $T(\vec{e}_i)$ can be expressed as $T_{ij}\vec{e}_j$
 $= \sum_{i=1}^n a_i \sum_j T_{ij} \vec{e}_j$

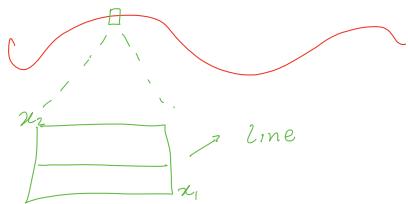
Where T_{ij} are the components of transformed T and are often written in matrix form.

If \vec{A}, \vec{C} are vectors and B is a matrix then $\vec{A}^T B \vec{C} = \sum_{ij} A_i B_{ij} C_j$ Strongly encourage to write in this form and not switch indices

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Newton's law can be expressed without any coordinates. The manifold is locally like \mathbb{R}^n

example:



§ 2.1 Definition of a Manifold

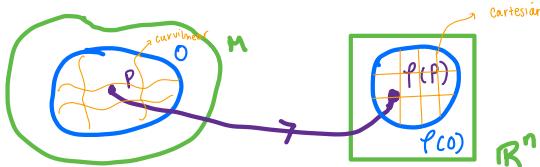
Idea: Any smooth curve/surface/volume [any dimension] looks locally like \mathbb{R}^n

A set of points M is a manifold if

If $x \in M$ has an open neighbourhood that has a continuous map which is 1-1 onto to map (bijective) open set of \mathbb{R}^n . Then M has dimension n .

In this framework, there is no measure of length on M . Distance is a global property and we will discuss this later.

Points in M look like \mathbb{R}^n , not \mathbb{E}^n , unless we have a metric



What is the difference between map, chart, atlas?

$$\varphi: O \rightarrow \mathbb{R}^n [x^1, x^2, \dots, x^n]$$

O is an open set

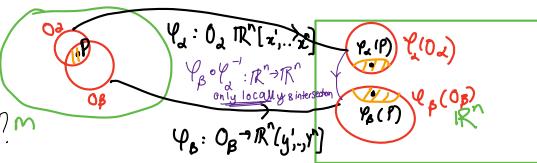
(O, φ) is a chart

(x^1, x^2, \dots, x^n) are the coordinates of \mathbb{R}^n

There can be multiple charts at a given point on a manifold. These charts must overlap

Q: WTF is para-compact

Q: is picture same w/ closed sets?



Q: if they maintain intersection, does that mean area & angle preserved, is possible? \Rightarrow metric

A collection of charts $(O_\alpha, \varphi_\alpha)$ is an atlas. This covers the manifold.

Note: $\varphi_\alpha: O_\alpha \rightarrow \mathbb{R}^n$ & $\varphi_\beta: O_\beta \rightarrow \mathbb{R}^n$ are homeomorphic $\xrightarrow{\text{def continuous and bijective}}$ $\varphi_\beta \circ \varphi_\alpha^{-1}: \mathbb{R}^n \rightarrow \mathbb{R}^n$ is a continuous function and is a coordinate transformation

$$y^i = y^i(x^1, \dots, x^n) \quad i=1, \dots, n$$

We say φ_α & φ_β are C^k related if all the partial derivatives of order k are continuous

If all $p \in M$, for all charts in M , is C^k -related then M is a C^k -manifold.

We assume M is a C^∞ -manifold [differentiable]

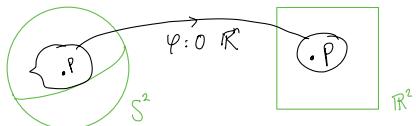
Examples: \mathbb{R}^n is a n -differentiable manifold

- ① \mathbb{R}^n has the natural topology
- ② can use identity map for the charts [any point in \mathbb{R}^n maps to itself]

§ 2.2 The Sphere is a Manifold

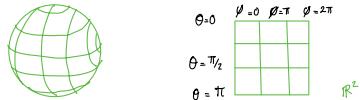
The two-sphere in \mathbb{R}^3 is denoted by S^2 and defined by
 $(x^1)^2 + (x^2)^2 + (x^3)^2 = \text{constant}$

A one-sphere is a circle



We can map small neighbourhoods of P to a disc in \mathbb{R}^2 . This map does not preserve length or angles

Another way to do this mapping is to use Spherical coordinates



$$\begin{aligned}\theta &= x^1 & 0 \leq x^1 \leq \pi & \text{Colatitude} \\ \phi &= x^2 & 0 \leq x^2 \leq 2\pi & \text{Longitude}\end{aligned}$$

The map has problem at:

$\theta=0, \pi$ the line is mapped to a point

$\phi=0, \pi$, this line gets mapped to 2π

A solution restricted to $0 < x^1 < \pi$ $0 < x^2 < 2\pi$ yields a chart for almost the entire sphere.

Another chart could be a similar system but where $\phi=0$ at the equator and then go from $\phi=-\pi/2$ to $\phi=\pi/2$

Assignment 2 will have a Q on stereographic projects.

§ 2.3 Other examples of Manifolds

A set M that can be parameterized continuously is a manifold and its dimension is the number of independent parameters

- ① Set of rotations of a Rigid Object of 3D. Dimension 3 (Euler Angles)
- ② All (pure boosts) Lorentz transformations is a manifold of dimension 3.
The parameters are the components of Velocity
- ③ N particles in 3D, $3N$ dimensions for the position and $3N$ for Velocity. This is a manifold of dim $6N$
- ④ An algebraic or differential equation for a dependent variable y in terms of an indep var x ,
The set (y, x) is a manifold

⑥ A vector space V over \mathbb{R} is a manifold. Suppose V is n -dim with basis $\{\vec{e}_1, \dots, \vec{e}_n\}$. Any $y \in V$ can be written as $\vec{y} = a^1 \vec{e}_1 + \dots + a^n \vec{e}_n$. We have a mapping $\vec{y} \mapsto (a^1, \dots, a^n)$ from V to \mathbb{R}^n . It turns out that V is identical (isomorphic) to \mathbb{R}^n .

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§ 2.5 curves

A map from a C^∞ manifold to another manifold N that is C^∞ and a bijection is a C^∞ diffeomorphism from M to N [How does this compare to a homeomorphism, a homeomorphism is continuous (C^0) and a bijection]

Diffeomorphism's in a specialized case of a homeomorphism [a subset/subspace?]

A differentiable manifold is a set M such that all points in M have an open set that has a map (diffeomorphism) to an open set in \mathbb{R}^n , we say it has dimension N

A curve is a differentiable map, say γ , from an open set of \mathbb{R} into M .

$$\gamma : [a, b] \xrightarrow{\text{gamma}} M \quad \text{or} \quad \gamma \mapsto \gamma(\lambda) \in M$$

We parameterize the curve with lambda λ .

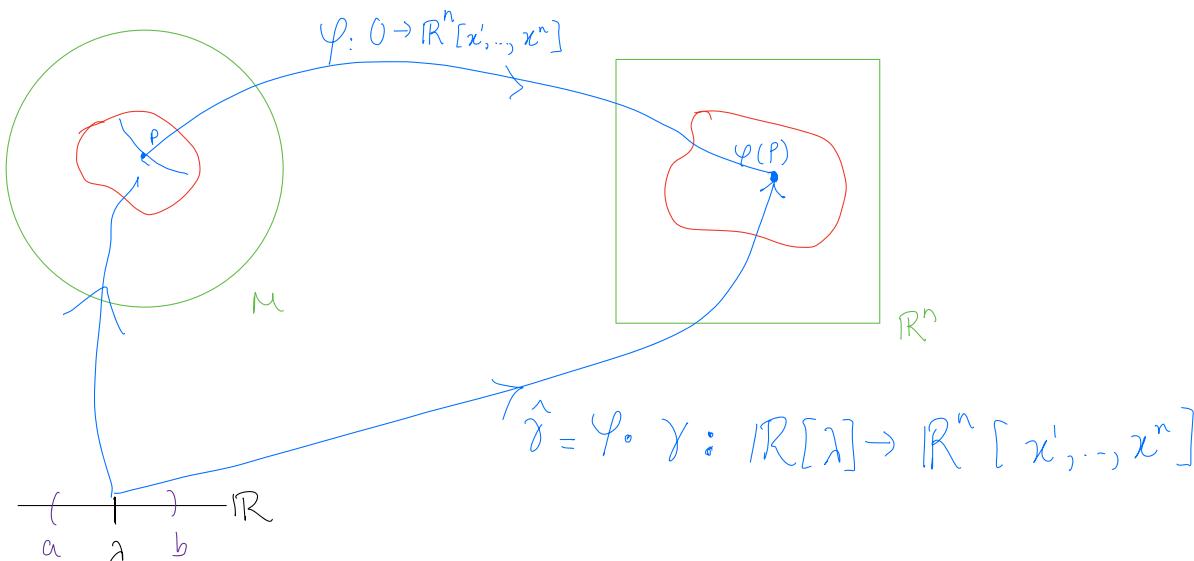
Two curves with the same image but different parameterizations are different.

Suppose the image of the curve is in the open set O with chart ψ :

$$\psi : O \rightarrow \mathbb{R}^n [x^1, \dots, x^n]$$

We obtain a coordinate representation of the curve:

$$\begin{aligned} \hat{\gamma} &= \psi \circ \gamma : \mathbb{R}[\lambda] \rightarrow \mathbb{R}^n [x^1, \dots, x^n] \\ \text{OR} \\ \lambda &\mapsto (x^1(\lambda), \dots, x^n(\lambda)) = [x^1(\gamma(\lambda)), \dots, x^n(\gamma(\lambda))] \end{aligned}$$



Quote:
Thank you for
paying attention
more so than me

γ is differentiable if M is a C^∞ manifold

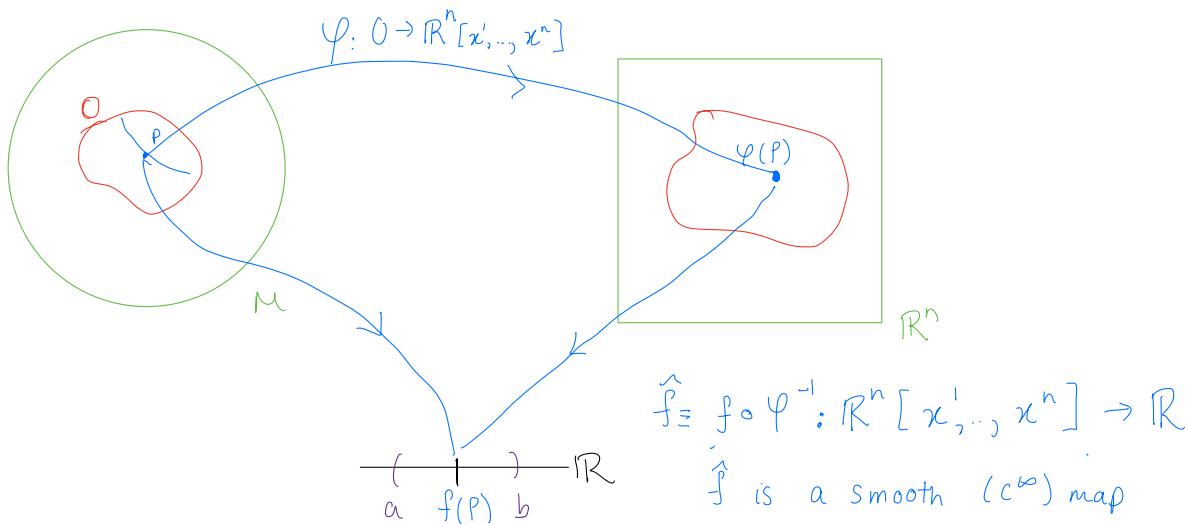
§ 2.6 Functions

A Function, say f , on M is a smooth map from M to \mathbb{R} ,

$$f: M \rightarrow \mathbb{R} \text{ or } x \mapsto f(x) \in \mathbb{R}$$

With chart $\varphi: U \rightarrow \mathbb{R}^n[x^1, \dots, x^n]$ we get a coordinate representation of f :

$$\begin{aligned}\hat{f} &\equiv f \circ \varphi^{-1}: \mathbb{R}^n \rightarrow \mathbb{R} \\ \text{or } (x^1, \dots, x^n) &\mapsto \hat{f}(x^1, \dots, x^n)\end{aligned}$$



On a manifold we always have coordinates but we don't always mention them explicitly

§ 2.7 Vectors and Vector Field

In a manifold, no magnitude for a vector.

Vectors typically have a direction and magnitude. In our definition, we will have a direction but no magnitude because we only have a local description about each point.

Suppose we have a curve, γ , that passes through the point $p \in M$ with coordinates $\{x^i\}_{i=1, \dots, n}$ and also a smooth function for M

$$\begin{aligned}\gamma: [a, b] &\rightarrow M \\ f: M &\rightarrow \mathbb{R}\end{aligned}$$

We can evaluate f on the curve

$$g \equiv f \circ \gamma = f \circ \varphi^{-1} \circ \varphi \circ \gamma = \hat{f} \circ \gamma: [a, b] \rightarrow \mathbb{R}$$

OR

$$\lambda \mapsto f(x^1(\lambda), \dots, x^n(\lambda))$$

OR

$$f(x^i(\lambda)): [a, b] \rightarrow \mathbb{R}$$

Since f and γ are both differentiable, so is g . We can differentiate g w.r.t x using the chain rule

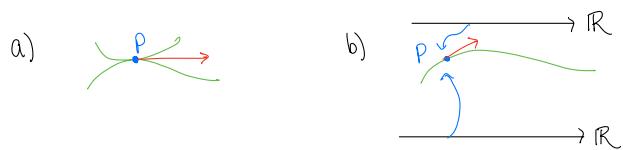
This is a directional derivative. What is a directional derivative?

$$\frac{dg}{d\lambda} = \sum_{i=1}^n \frac{dx^i}{d\lambda} \frac{\partial f}{\partial x^i} \Rightarrow \frac{d}{d\lambda} = \sum_{i=1}^n \frac{dx^i}{d\lambda} \frac{\partial}{\partial x^i}$$

This is an operator

Observe that the above is a directional derivative of f in the direction of $\left\{ \frac{dx^i}{d\lambda} \right\}$. These are the components of the tangent to the curve

Note that each tangent vector $\left\{ \frac{dx^i}{d\lambda} \right\}$ has an infinite number of curves that are tangent to it.



Example:

At the point $P = x^i(0)$, consider the curve

$$\hat{\gamma}_1 = (\psi \circ \gamma_1)(\lambda) = \lambda a^i \quad [a^i \text{ constants}]$$

The Tangent is

$$\left. \frac{dx^i}{d\lambda} \right|_{\lambda=0} = a^i$$

Now consider a different curve at P

$$\hat{\gamma}_2 = (\psi \circ \gamma_2)(\mu) = \mu^2 b^i + \mu a^i = x^i, \quad a^i, b^i \text{ constants}$$

$$\text{The tangent vector is } \left. \frac{dx^i}{d\mu} \right|_{\mu=0} = 2\mu b^i + a^i = a^i$$

The curves γ_1 and γ_2 have the same tangent vector at P

It can be shown that tangents to the curve form a Vector space. We can show all the properties of a vector space are satisfied but we only show closure.

Proof: Suppose $a, b \in \mathbb{R}$ and we have curves $\gamma_1(\lambda)$ and $\gamma_2(\mu)$

$$\text{From } \gamma_1: \frac{d}{d\lambda} = \sum_{i=1}^n \frac{dx^i}{d\lambda} \frac{\partial}{\partial x^i}$$

$$\text{From } \gamma_2: \frac{d}{d\mu} = \sum_{i=1}^n \frac{dx^i}{d\mu} \frac{\partial}{\partial x^i}$$

Consider the linear superposition of the two,

$$a \frac{d}{d\lambda} + b \frac{d}{d\mu} = \sum_{i=1}^n \left(a \frac{dx^i}{d\lambda} + b \frac{dx^i}{d\mu} \right) \frac{\partial}{\partial x^i}$$

we introduce a new parameter ϕ such that

$$\frac{d}{d\phi} = \sum_{i=1}^n \left(a \frac{dx^i}{d\lambda} + b \frac{dx^i}{d\mu} \right) \frac{\partial}{\partial x^i} = \sum_{i=1}^n \frac{\partial x^i}{\partial \phi} \frac{\partial}{\partial x^i}$$

This proves closure, which means the scalar sum of 2 tangent vectors is also a tangent vector ■

If you consider the tangent vectors along the coordinate lines, $\{x^i\}$, we get $\{\frac{\partial}{\partial x^i}\}$. This forms a basis to the vector space.

In our equations for $\frac{d}{d\lambda}$, $\{\frac{dx^i}{d\lambda}\}$ are the components in the vector space.

Lec 5- Sep 21

All future assignments will be due on Thursdays @ 5^{pm}

Previously, we define $g = f \circ \gamma: (a, b) \rightarrow \mathbb{R}$, $\frac{dg}{d\lambda} = \sum_{i=1}^n \frac{dx^i}{d\lambda} \frac{\partial f}{\partial x^i} \Rightarrow \frac{d}{d\lambda} = \sum_{i=1}^n \frac{dx^i}{d\lambda} \frac{\partial}{\partial x^i}$ where $\frac{dx^i}{d\lambda}$ are the components of the tangent to the curve

With Coordinates $\{x^i\}$, we get coordinate lines $\{\frac{\partial}{\partial x^i}\}$ which form a basis.

Since there is a 1 to 1 correspondence between the tangent vectors at P and the space of partial derivatives at P, we use $\{\frac{\partial}{\partial x^i}\}$ to denote the tangent vectors to the curve

§ 2.8 Basis vectors and basis vector fields

Every point P in manifold M has a tangent space denoted by T_p or $T_p M$, which is a vectorspace, $\text{dim}(M) = n$.

We need n linearly independent vectors in $T_p M$ to form a basis. A coordinate system $\{x^i\}$ at P has a coordinate basis of $\{\frac{\partial}{\partial x^i}\}$ of $T_p M$ for all $P \in M$

Suppose $\{\bar{e}_i\}_{i=1,\dots,n}$ is another basis of $T_p M$. Any vector in $T_p M$, say \bar{V} , can be written as

$$\bar{V} = \sum_{i=1}^n V^i \frac{\partial}{\partial x^i} = \sum_{j=1}^n V^j \bar{e}_j$$

$\{V^i\}$ & $\{V^j\}$ are the coordinates w.r.t. $\{\frac{\partial}{\partial x^i}\}$ and $\{\bar{e}_j\}$. These coordinates are functions on M.

Note: In $\{\frac{\partial}{\partial x^i}\}$ i is a superscript but appears in the denominator & is considered a subscript

For Vectors we use subscripts for the basis and superscripts for coordinates

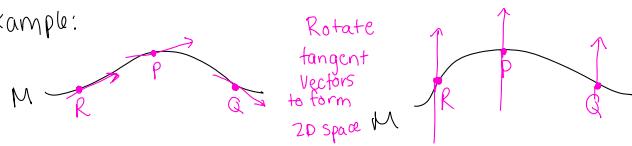
A vector is an object that lives in $T_p M$. A vector field are mappings that defines a vector for all $P \in M$

A vector field is cliffable if its coordinates are differentiable. The basis $\{\frac{\partial}{\partial x^i}\}$ are linearly independent if $\{x^i\}$ are (proper) coordinates.

§2.9 Fiber Bundles

A manifold M with a tangent space $T_p M$ can be combined to form a tangent bundle (TM)

1D example:



The tangent bundle is a 2D space (in general $2n$) is also a manifold. We can define a projection to get the point from the tangent bundle. The tangent bundle is an example of a fiber bundle.

§2.12 Vector Fields and Integral Curves

Any curve on a manifold has a tangent vector at every point. Since this is true \forall points on the curve this defines a vector field. It can be shown that any smooth vector field has a curve associated with it this is an integral curve.

Vector fields correspond to a system of first order ODEs and the integral curve is the solution.

Suppose we have a smooth vector field $\bar{V} \in T_p M$ with components $V^i(p)$ with coordinates $\{x^i\}$. Then if we write $V^i(p) = V^i(x^1, \dots, x^n)$. Then we get a system of DEs

$$\boxed{\frac{dx^i}{d\lambda} = V^i(x^1, \dots, x^n)}$$

If V^i is C^1 $\forall i$ then \exists a soln to the system which is the integral curves.

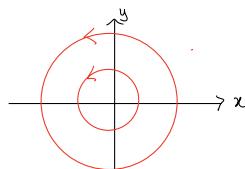
Example:

$$\bar{V} = x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x}$$

$$\frac{dx}{d\lambda} = -y \quad \frac{dy}{d\lambda} = x$$

If you differentiate the first

$$\frac{d^2 x}{d\lambda^2} = -\frac{dy}{d\lambda} = -x \quad \text{Harmonic Oscillator}$$



§2.13 Exponentiation of the Operator $\frac{d}{d\lambda}$

Suppose we have an analytic (smooth) manifold (C^∞) with coordinates $\{x^i(\lambda)\}$ along integral curves

Then $\bar{Y} = \frac{d}{d\lambda}$ are analytic functions of λ and we can \therefore Taylor Expand. We Taylor expand $\{x^i\}$ about λ_0 ,

$$\begin{aligned} x^i(\lambda_0 + \epsilon) &= x^i(\lambda_0) + \epsilon \left(\frac{dx^i}{d\lambda} \right) \Big|_{\lambda_0} + \frac{1}{2} \epsilon^2 \left(\frac{d^2 x^i}{dx^2} \right) \Big|_{\lambda_0} + \dots \\ &= \left(1 + \epsilon \frac{d}{d\lambda} + \frac{\epsilon^2}{2} \frac{d^2}{d\lambda^2} + \dots \right) x^i \Big|_{\lambda_0} \\ &\equiv \exp \left[\epsilon \frac{d}{d\lambda} \right] x^i \Big|_{\lambda_0} \quad [\text{THIS IS NOTATION}] \end{aligned}$$

We have the exponentiation of the operators, which is short hand for the above expression.

Note: $\exp(\epsilon \frac{d}{d\lambda}) = e^{\epsilon \frac{d}{d\lambda}} = e^{\epsilon \bar{Y}}$ This is trying to give us a sense of distance

§ Lie brackets and non coordinate basis

Suppose $\{x^i\}$ is a coordinate system and $\{\frac{\partial}{\partial x^i}\}$ is a basis of vector fields

We know that any n linearly independent vectors form a basis but can a basis form a coordinate system? No

By construction $\frac{\partial}{\partial x^i} \times \frac{\partial}{\partial x^j}$ commute for all i, j

$$\left[\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right] = \frac{\partial}{\partial x^i} \frac{\partial}{\partial x^j} - \frac{\partial}{\partial x^j} \frac{\partial}{\partial x^i} = 0$$

Suppose we have $\bar{V} = \frac{d}{d\lambda}$ and $\bar{W} = \frac{d}{d\mu}$ we can show that they need not always commute

$$\begin{aligned} \left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right] &= \frac{d}{d\lambda} \left(\frac{d}{d\mu} \right) - \frac{d}{d\mu} \left(\frac{d}{d\lambda} \right) \\ &= \left(\sum_{i=1}^n v^i \frac{\partial}{\partial x^i} \right) \left(\sum_{j=1}^n w^j \frac{\partial}{\partial x^j} \right) - \left(\sum_{j=1}^n w^j \frac{\partial}{\partial x^j} \right) \left(\sum_{i=1}^n v^i \frac{\partial}{\partial x^i} \right) \\ &= \sum_{i,j} \left\{ v^i \frac{\partial w^j}{\partial x^i} \frac{\partial}{\partial x^j} + v^i w^j \frac{\partial}{\partial x^i} \frac{\partial}{\partial x^j} \right. \\ &\quad \left. - w^j \frac{\partial v^i}{\partial x^j} \frac{\partial}{\partial x^i} - w^j v^i \frac{\partial}{\partial x^j} \frac{\partial}{\partial x^i} \right\} \end{aligned}$$

$$\left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right] = \sum_{i,j} v^i \frac{\partial w^j}{\partial x^i} \frac{\partial}{\partial x^j} - w^j \frac{\partial v^i}{\partial x^j} \frac{\partial}{\partial x^i}$$

If this is non zero then $\frac{d}{d\lambda}$ & $\frac{d}{d\mu}$ form a non-coordinate basis

The Lie bracket of $\bar{V} = \frac{d}{d\lambda}$ and $\bar{W} = \frac{d}{d\mu}$ is $\left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right]$

Lec 6 - Sep 26

Q1] just show they are inverses

Q4) using C^∞ argue
that commutator is 0

A1 marks very latest next thursday

§ 2.14 Lie brackets & non-coordinate basis

The Lie bracket is defined as the commutator of two vectors: $\left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right]$

Geometric interpretation:

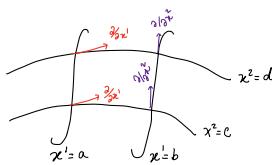
Consider a coordinate basis $\frac{\partial}{\partial x^1}, \frac{\partial}{\partial x^2}$ where $\left[\frac{\partial}{\partial x^1}, \frac{\partial}{\partial x^2} \right] = 0$

The integral curve of $\frac{\partial}{\partial x^1}$ [that are tangent to $\frac{\partial}{\partial x^1}$]

$$\frac{dx^1}{d\lambda} = 1 \quad \text{and} \quad \frac{dx^2}{d\lambda} = 0$$

$$\Rightarrow x^1 = \lambda + c \quad \text{and} \quad x^2 = \text{constant.}$$

Integral Curves of $\frac{\partial}{\partial x^2}$ are $x^1 = \text{constant}, x^2 = \lambda + \text{constant}$



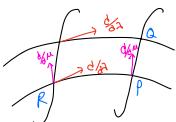
HOW DOES THIS RELATE
TO DYNAMIC SYSTEMS and
Vector fields discussed in Calc 4

Locally things can be curvy but on a curve only 1 parameter changes

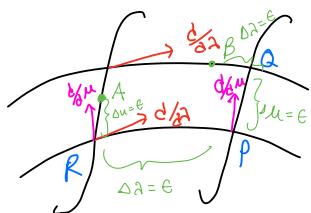
Along each integral curve all the x^i 's are constant except for one that changes

Next consider a non-coordinate basis $\bar{V} = \frac{d}{d\lambda}, \bar{W} = \frac{d}{d\mu}$ with $\left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right] \neq 0$

On the integral curves of $\frac{d}{d\lambda}$, λ increases and μ can also change.



Suppose we start at P and move $\Delta\lambda = \epsilon$ and then $\Delta\mu = \epsilon$ to end up at A



We can also start at P, move $\Delta\mu = \epsilon$ then $\Delta\lambda = \epsilon$ and end up at B

Find the approximate distance between A + B:

1st path, first move to R: $x^i(R) = \exp \left[\epsilon \frac{d}{d\lambda} \right] x^i |_P$

Then to A, $x^i(A) = \exp[\epsilon \frac{d}{du}] \exp[\epsilon \frac{d}{da}] x^i|_p$

Similarly, we move from P to Q to B

$$x^i(B) = \exp[\epsilon \frac{d}{du}] \exp[\epsilon \frac{d}{da}] x^i|_p$$

The difference between the two is

$$\begin{aligned} x^i(B) - x^i(A) &= \left[\exp\left(\epsilon \frac{d}{da}\right), \exp\left(\epsilon \frac{d}{du}\right) \right] x^i|_p \\ &= \dots \quad [\text{Assignment 2.4(b)?}] \\ &= G^2 \left[\frac{d}{da}, \frac{d}{du} \right] + O(\epsilon^3) \end{aligned}$$

\bar{v}, \bar{w} are in a coordinate basis iff $[\bar{v}, \bar{w}] = 0$ \bar{v} and \bar{w} are vector fields

§2.16 One-forms [co vectors]

$T_p M$ is the space of tangent vectors at $P \in M$. A one-form is a linear, real-valued function of vectors $\tilde{\omega}: V \rightarrow a \in \mathbb{R}$

The space of one forms is the dual space to the tangent space $T_p M$

Suppose $\tilde{\omega}$ is a one form and \bar{v} is a vector, both at P, then we have an operation

$$\boxed{\tilde{\omega}(\bar{v}) \in \mathbb{R}}$$

One forms are linear with ($a, b \in \mathbb{R}$, $\tilde{\sigma}$ is a one form)

$$\textcircled{1} \quad \tilde{\omega}(a\bar{v} + b\bar{w}) = a\tilde{\omega}(\bar{v}) + b\tilde{\omega}(\bar{w})$$

$$\textcircled{2} \quad (a\tilde{\omega})(\bar{v}) = a(\tilde{\omega}(\bar{v}))$$

$$\textcircled{3} \quad (\tilde{\omega} + \tilde{\sigma})(\bar{v}) = \tilde{\omega}(\bar{v}) + \tilde{\sigma}(\bar{v})$$

These properties ensure that one forms at P forms a vector space. This is called the dual space of $T_p M$ called $T_p^* M$

Vectors are linear, real-valued functions of one forms and hence $T_p M$ is the dual of $T_p^* M$

$$T_p^{**} M = T_p M \quad [\text{this is always the case}]$$

$$\begin{aligned} \text{Example: } (\alpha \tilde{\omega} + b \tilde{\sigma})(\bar{v}) &= (\alpha \tilde{\omega})(\bar{v}) + (b \tilde{\sigma})(\bar{v}) \\ &= \alpha(\tilde{\omega}(\bar{v})) + b(\tilde{\sigma}(\bar{v})) \end{aligned}$$

Notation: $\tilde{\omega}(\bar{v})$ or $\bar{v}(\tilde{\omega}) = \langle \tilde{\omega}, \bar{v} \rangle$ these are all called contraction

Vectors are sometimes called Contravariant; One-forms (covectors) are covariant

§2.17 Examples of One-forms

example matrix algebra
Column vectors are vectors
Row vectors are one-forms

$$(a, b) : \begin{pmatrix} x \\ y \end{pmatrix} \mapsto (a, b) \begin{pmatrix} x \\ y \end{pmatrix} = ax + by \in \mathbb{R}$$

§2.18 dirac delta function

C^∞ functions are an abelian group under addition and a Vector Space under multiplication.

The dual space of the functions are one forms and called distributions

$$\delta(x) : f(x) \mapsto \int_{-\infty}^{\infty} \delta(x) f(x) dx = f(0)$$

§2.19 The gradient and the pictorial representation of a one-form

A vector field has a unique vector at every point. A field of one-form has a unique one form at every point.

Differentiability of one forms will be determined in terms of diff'ability of vectors and functions

A Tangent Bundle, TM contains $M \times T_p M$. A cotangent Bundle contains T^*M contains $M \times T_p^* M$. Both are fiber bundles

We will show that the gradient of f , denoted $\tilde{d}f$ is a one-form and defined as:

$$\tilde{d}f\left(\frac{d}{dx}\right) = \frac{df}{dx} = \sum_{i=1}^n \frac{dx^i}{dx} \frac{\partial f}{\partial x^i} \in \mathbb{R}$$

gradient "exists" in Dual space and takes in from the Tangent Space

$\tilde{d}f$ is an element of T_p^*M and the contraction with $\frac{dz^i}{d\lambda}$ yields the directional derivative of f along a curve tangent to V

Check $\tilde{d}f$ is a one form:

$$\begin{aligned}\tilde{d}f(a \frac{d}{d\lambda} + b \frac{d}{du}) &= (a \frac{d}{d\lambda} + b \frac{d}{du})f \text{ by above definition} \\ &= a \frac{df}{d\lambda} + b \frac{df}{du} = a \tilde{d}f(\frac{d}{d\lambda}) + b \tilde{d}f(\frac{d}{du})\end{aligned}$$

we see $\tilde{d}f$ is a linear operator on vectors

$\frac{df}{d\lambda}$ at P is computed from $\frac{\partial f}{\partial z^i}$ at P and this forms the components of $\tilde{d}f$

Note $\frac{\partial f}{\partial x^i}$ has a lower index, opposite to the basis of vectors

Lec 7 - Sep 28^m

§ 2.21 Basis 1-forms and components of 1-forms

Any n linearly independent one-forms are a basis of T_p^*M [cotangent space]

Given a basis of $T_p M$ say $\{\bar{e}_i, i=1, \dots, n\}$ this induces a dual basis to T_p^*M $\{\tilde{\omega}^i, i=1 \dots n\}$

If $\bar{v} \in T_p M$ then the dual basis $\tilde{\omega}^i$ is defined by

$$\tilde{\omega}^i(\bar{v}) = v^i$$

$$\bar{v} = \sum_{i=1}^n v^i \bar{e}_i$$

and

$$\tilde{\omega}^i(\bar{e}_j) = \delta_j^i \text{ [Kronecker delta]}$$

We show that $\{\tilde{\omega}^i\}$ are linearly independent and form a basis of T_p^*M .

Consider any one-form \tilde{q}

$$\begin{aligned} \tilde{q}(\bar{v}) &= \tilde{q}\left(\sum_{j=1}^n v^j \bar{e}_j\right) \\ &= \sum_{j=1}^n v^j \tilde{q}(\bar{e}_j) \end{aligned} \quad \text{Linear}$$

Define $q_j = \tilde{q}(\bar{e}_j)$ to be the components of \tilde{q} on the dual basis to $\{\bar{e}_j\}$

$$\text{Also } \tilde{q}(\bar{v}) = \sum_j q_j \tilde{\omega}^j(\bar{v}) \Rightarrow \tilde{q} = \sum_j q_j \tilde{\omega}^j$$

Note: compare with $v = \sum_j v^j \bar{e}_j$

Therefore, $\{\tilde{\omega}^i\}$ are a basis since there are n of them and we can generate any \tilde{q} with this decomposition

It follows,

$$\tilde{q}(\bar{v}) = \sum_j q_j v^j \quad \text{This is a Contraction}$$

If $\{\bar{e}_i\}$ is a basis of $T_p M$ \forall points $U \subset M$ then $\{\tilde{\omega}^i\}$ is a basis of T_p^*M \forall points $U \subset M$

The coordinate basis $\{x^i\}$ on U defines a natural vector field $\{\partial/\partial x^i\}$ (a basis of $T_p M$)

and this defines a natural basis of one forms $\{\tilde{\partial}/\partial x^i\}$ [at each point]

With this notation, $\tilde{\partial}/\partial x^i(\partial/\partial x^j) = \delta_j^i$

§ 2.21 Index Notation

Components of vectors v^i

Components of one-forms ω_j

Vector basis: \bar{e}_i

One form basis $\tilde{\omega}^j$

Coordinate basis; (1-forms) $\tilde{d}x^j$

(vectors) $\frac{\partial}{\partial x^i}$

$$\text{Example: } \tilde{\omega}(\bar{v}) = \sum_j v^j \omega_j = v^j \omega_j \quad [\text{Einstein's Summation Notation}]$$

An index that occurs twice is summed over if one is a subscript and one is a superscript

$$\text{Example: } \tilde{\omega} = \sum_j \omega_j \tilde{d}x^j \rightarrow \omega_j \tilde{d}x^j$$

Examples with no sum: $v^j v^k$, $v^j \omega_j$, $v^j w^j$
 not repeated index
 no repeated index
 no sub

§ 2.22 Tensor and Tensor fields

We build on vectors and 1-forms to get tensors \nearrow operator

At $P \in M$ a tensor of Type (N, N') is a linear map \nearrow that takes N -1 forms and N' vectors and yields a real number.

Example: F is a $(2, 2)$ tensor

We can write this as:

$$F(\tilde{\omega}, \tilde{\sigma}; \bar{v}, \bar{w})$$

Since it is linear in all arguments

$$\begin{aligned} & F(a\tilde{\omega} + b\tilde{\lambda}, \tilde{\sigma}; \bar{v}, \bar{w}) \\ &= a F(\tilde{\omega}, \tilde{\sigma}; \bar{v}, \bar{w}) + b F(\tilde{\lambda}, \tilde{\sigma}; \bar{v}, \bar{w}) \end{aligned}$$

$$F(\tilde{\omega}, \tilde{\sigma}; a\bar{v} + b\bar{u}, \bar{w}) = a F(\tilde{\omega}, \tilde{\sigma}; \bar{v}, \bar{w}) + b F(\tilde{\omega}, \tilde{\sigma}; \bar{u}, \bar{w})$$

§ 2.23 Examples of Tensors

In Linear Algebra, column vectors are vectors are (0^1) tensors, Row "vectors" are One-forms or (1^0) tensors, and Matrix is a (1^1) tensor

§ 2.24 Components of Tensors and the outer product

Consider 2 Vectors \bar{V}, \bar{W} . We can form a (2^0) Tensor with the Outer product.

$$\begin{array}{c} \text{chosen} \\ \nearrow \quad \nwarrow \\ \bar{V} \otimes \bar{W} \end{array} \quad \begin{array}{c} \text{Variable} \\ \nearrow \quad \nwarrow \\ (\tilde{p}, \tilde{q}) \end{array} = a + R \\ \text{Outer product} \\ \text{direct product} \\ \text{tensor product} \end{array} \quad \begin{array}{c} \text{Variable} \\ \nearrow \quad \nwarrow \\ \bar{V}(\tilde{p}) \quad \bar{W}(\tilde{q}) \end{array}$$

If \tilde{p}, \tilde{q} are one forms we can form a (0^2) tensor

$$\begin{array}{c} \tilde{p} \otimes \tilde{q} \\ \downarrow \quad \downarrow \\ \text{chosen} \quad \text{variable} \end{array} (\bar{V}, \bar{W}) = \tilde{p}(\bar{V}) \tilde{q}(\bar{W})$$

\otimes is the outer/direct/tensor product

The outer product of an (N^M) tensor and an $(N'^{M'})$ tensor is a tensor of order $(N+N') \choose M+M'$

The components of a tensor are the values it takes when it has basis vectors and 1-form as arguments

Example: If S is a (3^2) tensor, then on the basis $\{\bar{e}_i\}$ and $\{\tilde{w}_i\}$ has components

$$S_{lm}^{ijk} \equiv S(\tilde{w}^i, \tilde{w}^j, \tilde{w}^k; \bar{e}_l, \bar{e}_m)$$

Lec 8 Oct 3

S 2.25 Contractions

$\bar{V} \otimes \tilde{\omega}$ is a $(1, 1)$ tensor and is written as $V^i \omega_j$.

Consider examples with

S_{jk}^i is a $(1, 2)$ tensor

P^{lm} as a $(2, 0)$ tensor

These can be contracted in various ways

$S_{jk}^i P^{lm}$ is a $(2, 1)$ tensor
 $\underbrace{\quad}_{3} \quad \underbrace{\quad}_{\text{contracted}}$

$S_{jk}^i P_{ls}$ is a $(1, 2)$ tensor

In general, the 2 above differ [unless P is symmetric]

Property: Contractions are independent of the basis

IDEA of Proof:

$$(1) \quad \tilde{\sigma}(\bar{V}) = \dots = \sigma_i V^i \quad \text{see assignment 2 for the details}$$

(2) Consider A a $(2, 0)$ tensor and B a $(0, 2)$ tensor. A contraction of A and B is $A^{ij} B_{jk} = C_k^i$ where C is a $(1, 1)$ tensor

$$\begin{aligned} \text{Consider } C \text{ applied to } \tilde{\sigma} \text{ and } \bar{V}, \quad C(\tilde{\sigma}; \bar{V}) &= C_k^i \sigma_i V^k \\ &= A^{ij} B_{jk} \sigma_i V^k \\ &= \sigma_i A^{ij} B_{jk} V^k \quad \text{components to operators?} \\ &= \sigma_i A(\tilde{\omega}^i; \tilde{\omega}^j) B(\bar{e}_j, \bar{e}_k) V^k \\ &\quad \text{using tensors are linear} \\ &\quad \downarrow = A(\sigma_i \tilde{\omega}^i; \tilde{\omega}^j) B(\bar{e}_j, V^k \bar{e}_k) \\ &= A(\tilde{\sigma}, \tilde{\omega}^j) B(\bar{e}_j, \bar{V}) \\ &= A(\tilde{\sigma}, B(\bar{e}_j, \bar{V}) \tilde{\omega}^j) \end{aligned}$$

$$\begin{aligned} \text{Aside: } B(\bar{e}_j, \bar{V}) \tilde{\omega}^j &= B_{lm} \tilde{\omega}^l \tilde{\omega}^m \bar{e}_j V^n \bar{e}_n \tilde{\omega}^j \\ &= B_{lm} (\tilde{\omega}^l \bar{e}_j) (\tilde{\omega}^m \bar{e}_n) (V^n \tilde{\omega}^j) \\ &\quad \underbrace{\delta_j^l}_{\text{}} \quad \underbrace{\delta_n^m}_{\text{}} \\ \Rightarrow B(\bar{e}_j, \bar{V}) \tilde{\omega}^j &= B_{jn} V^n \tilde{\omega}^j \\ &= B_{jn} \tilde{\omega}^j V^n \\ &= B(\tilde{\sigma}, \bar{V}) \\ \therefore C(\tilde{\sigma}; \bar{V}) &= A(\tilde{\sigma}, B(\tilde{\sigma}, \bar{V})) \quad \rightarrow \text{completely independent from basis.} \\ &\quad \text{Empty space} \end{aligned}$$

$$\begin{aligned}
 \text{Aside: } B(\bar{e}_j, \bar{v}) \tilde{\omega}^j &= (B_m \tilde{\omega}^l \tilde{\omega}^m) \bar{e}_j v^p \bar{e}_p \tilde{\omega}^j \\
 &= B_{lm} (\tilde{\omega}^l \bar{e}_j) (\tilde{\omega}^m \bar{e}_p) v^p \tilde{\omega}^j \\
 &= B_{jp} v^p \tilde{\omega}^j
 \end{aligned}$$

Recall: $\tilde{\omega}^i \bar{e}_j = \delta_j^i$.

This is independent of basis and indices

§ 2.26 Basis Transformations

Recall, a tensor of type (N, N') is a linear function that takes N 1-forms and N vectors as arguments

This definition is modern. Previously tensors were defined as how components change under a change of basis

At $P \in M$ suppose $\{\bar{e}_i\}$, $\{\bar{e}_j\}$ are bases to $T_p M$. There is a linear transformation matrix Λ with $\bar{e}_j' = \Lambda^i_j \bar{e}_i$. Λ^i_j is non-singular

Recall: Oneforms have a basis defined by $\tilde{\omega}^i(\bar{e}_k) = \delta_k^i$

We will determine how $\tilde{\omega}^i$ basis transforms. To do this, we multiply the definition by Λ^k_j

$$\begin{aligned}
 \Lambda^k_j \tilde{\omega}^i(\bar{e}_k) &= \Lambda^k_j \delta_k^i \\
 \tilde{\omega}^i(\Lambda^k_j \bar{e}_k) &= \Lambda^i_j \\
 \tilde{\omega}^i(\bar{e}_j') &\\
 \Rightarrow \tilde{\omega}^i(\bar{e}_j') &= \Lambda^i_j
 \end{aligned}$$

We define the inverse of Λ^i_j to be $\Lambda^{k'}_j$

$$\boxed{
 \begin{aligned}
 \Lambda^{k'}_j \Lambda^i_i &= \delta^{k'}_i \\
 \Lambda^{k'}_j \Lambda^{i'}_{k'} &= \delta^i_j
 \end{aligned}
 }$$

We multiple the previous eqn by Λ^k_i to get $\Lambda^{k'}_i \tilde{\omega}^i(\bar{e}_j') = \Lambda^{k'}_i \Lambda^i_j = \delta^{k'}_j$

$$\Lambda^{k'}_i \tilde{\omega}^i(\bar{e}_j') = \delta^{k'}_j = \tilde{\omega}^{k'}(\bar{e}_j')$$

The functions in front of \bar{e}_j' must be equal.

$$\tilde{\omega}^{k'} = \Lambda^{k'}_i \omega^i$$

$\tilde{\omega}^{k'}$ transforms with $\Lambda^{k'}_i$

Compare with

$$\bar{e}_j' = \Lambda_{j'}^j \bar{e}_i \quad \bar{e}_j' \text{ transforms with } \Lambda_{j'}$$

1-forms transform with the inverse of Λ . we can also transform coordinates

$$v^{i'} = \tilde{\omega}^{i'}(\bar{v}) = \Lambda_{j'}^i \tilde{\omega}^j(\bar{v}) = \Lambda_{j'}^i v^j$$

and similarly

$$q_{k'} = \tilde{q}(\bar{e}_{k'}) = \tilde{q}(\Lambda_{k'}^j \bar{e}_j) = \Lambda_{k'}^j \tilde{q}(\bar{e}_j) = \Lambda_{k'}^j q_j$$

Summary:

v^i and $\tilde{\omega}^i$ Transform with $\Lambda_{j'}^i$

q_i and \bar{e}_i transform with $\Lambda_{k'}^i$

Its because of these difference that

$v^i \bar{e}_i$ & $v^{i'} \bar{e}_{i'}$ are basis independent.

Vector Coordinates (Superscripts) are contravariant since they transforms "opposite" to how there basis transforms

1-form Coordinates (subscripts) are covariant since they transform like \bar{e}_i

§ Coordinate Transformations

Suppose $U \subset M$ has coordinates $\{x^i, i=1,\dots,n\}$. Introduce new coordinates $\{y^i, i=1,\dots,n\}$

$$y^{i'} = f(x^1, \dots, x^n) \quad j' = 1, \dots, n$$

$$y^{i'} = f^{i'}(x^j)$$

Jacobian?

The coordinate transformation is $\frac{\partial y^i}{\partial x^j}$ has a non-zero determinant in U

All $P \in V$ can be described with $\{x^i\}$ or $\{y^i\}$ has coordinate vector basis

$$\left\{ \frac{\partial}{\partial x^i} \right\} \text{ and } \left\{ \frac{\partial}{\partial y^{i'}} \right\}$$

These must be related by

$$\frac{\partial}{\partial y^j} = \frac{\partial x^i}{\partial y^j} \frac{\partial}{\partial x^i}$$

Compare with what we saw previously

$$\bar{e}_{j'} = \Lambda_{j'}^i \bar{e}_i \quad \text{explicit expression for lambda}$$

$$\therefore \Lambda_{j'}^i = \frac{\partial x^i}{\partial y^{j'}} \quad \text{covariant}$$

Similarly

$$\Lambda_j^{k'} = \frac{\partial y^{k'}}{\partial x^j} \quad \text{contravariant}$$

$$\text{Since } \frac{\partial x^i}{\partial y^j} \frac{\partial y^{j'}}{\partial x^k} = \frac{\partial x^i}{\partial x^k} = \delta_k^i$$

Lec 9 Oct 5, 2023

§ 2.27 Tensor Operations on Components

Given a tensor T with components $\{T^i_{j...}\}$ on a basis, the following is basis invariant.

$$aT \text{ or } \{aT^i_{j...}\}$$

This can be denoted as $T \rightarrow aT$

The Outer Product is also basis invariant

$$A, B \rightarrow A \otimes B$$

or

$$\{A^i_{j...}\}, \{B^k_{l...}\} \rightarrow \{A^i_{j...} B^k_{l...}\}$$

A tensor operation is one where operations on components produces components that are the tensor, independent of the basis. This include:

- (1) Addition
- (2) Scalar Multiplication
- (3) Outer Products
- (4) Contractions

§ 2.28 Functions and Scalars

A scalar is a (0) tensor, which is a function on M independent of the basis

Example: $V^i \tilde{\omega}_i$ is a scalar

V^i is not a scalar

§ 2.29 The metric tensor on a Vector Space

An inner product is a rule that associates a number with 2 vectors and it is a (0) tensor

It is also referred to as a metric tensor, g_i^j , ^{bar not one}

$$g_1(\bar{v}, \bar{w}) = g_1(\bar{w}, \bar{v}) \equiv \bar{w} \cdot \bar{v}$$

g_1 is a symmetric tensor with components

$$g_{ij} = g_1(\bar{e}_i, \bar{e}_j)$$

We will require that g_1 has an inverse.

If $g_{ij} = \delta_{ij}$ then it is the Euclidean metric, and the vector space is Euclidean space, \mathbb{E}^n

Given any g_{ij} , we can change to a new basis, say $\{\bar{e}_{i'j'}\}$ such that

$$g_{i'j'} = \Lambda_{i'}^k \Lambda_{j'}^l g_{kl}$$

We can pick a basis where the metric tensor is diagonal and only has +1 or -1 as entries

The convention is to list the -1's first then the +1's:

$$g_{i'j'} = \text{diag}(-1, \dots, -1, 1, \dots, 1)$$

This is written in an orthonormal basis

The sum [trace] of these elements is the signature.

We classify the metric tensor as

positive definite if all +1's

negative definite if all -1's

indefinite if we have both 1's and -1's

Example: Minkowski metric in special relativity

$$(-1, 1, 1, 1) \text{ or } (1, 1, 1, -1)$$

This is indefinite

Euclidean Space, \mathbb{E}^n is called Cartesian, with

$$g_{ij} = \delta_{ij} \text{ or } g_1 = I$$

The Minkowski metric has the Lorentz basis

$$\eta = \text{diag}(-1, 1, 1, 1)$$

A transformation matrix from one Lorentz basis to another can be written as

$$\eta = A^T \eta A$$

The metric tensor maps vectors to 1-forms.

$$\tilde{V} = g_1(\bar{V}, \underline{\quad})$$

In terms of components,

$$\begin{aligned} V_i &= \tilde{V}(\bar{e}_i) \\ &= g_1(\bar{V}, \bar{e}_i) \\ &= g_1(v^j \bar{e}_j, \bar{e}_i) \\ &= v^j g_{ji}(\bar{e}_j, \bar{e}_i) \\ &= v^j g_{ji} \quad \left. \right\} \text{Symmetry in } g_{ij} \\ &= g_{ij} v^j \end{aligned}$$

The inverse matrix g_{ij} is called g^{ij} with

$$g^{ij} g_{ik} = \delta^i_k \quad \text{or} \quad g_{ij} g^{jk} = \delta^k_i$$

This yields,

$$\begin{aligned} &g^{ki} V_i \\ &= g^{ki} g_{ij} V^j \\ &= \delta^k_j V^j \end{aligned}$$

This recovers what we had before

$$\text{Summary: } V_i = g_{ij} V^j$$

and

$$V^k = g^{ki} V_i$$

A $(\frac{?}{2})$ tensor, A , can map to a $(\frac{!}{1})$ tensor

$$A^i_j = g_{jk} A^{ik}$$

This can be mapped to $(\frac{0}{2})$ tensor

$$A_{lj} = g_{lm} A^m_j$$

$$\text{This can be inverted } A^{ik} = g^{il} g^{km} A_{lm}$$

Skip the proof

This is called Index raising and Lowering with a metric tensor, there is much less difference between $\binom{N}{N'}$ and $\binom{N-1}{N'-1}$ and $\binom{N+1}{N'+1}$

Hence Why we often refer to them as tensors of order $N + N'$

In a Euclidean vector space, a cartesian basis is $g_{ij} = g^{ij}$

There is no difference between super scripts and Subscripts. and hence we often only use Subscripts.

§ 2.30 The metric tensor field on a manifold

A metric tensor g_1 on a manifold is a $(\frac{0}{2})$ Symmetric tensor and has an inverse at every point.

For all points in the manifold, g_1 serves as a metric on $T_p M$ and has all the properties mentioned previously. but there's more

Using g_1 , we can define distance and curvature

We can use g_1 to define length on M

Suppose a curve has a tangent vector $\bar{V} = \frac{d\bar{x}}{d\lambda}$

then

$$\begin{aligned} d\ell^2 &= d\bar{x} \cdot d\bar{x} = (\bar{V} d\lambda) \cdot (\bar{V} d\lambda) \\ &= \bar{V} \cdot \bar{V} (d\lambda)^2 \\ &= g_1(\bar{V}, \bar{V}) d\lambda^2 \end{aligned}$$

d is infinitesimal
and not gradient

If g_1 is positive definite then $d\ell^2$ is positive and

$$d\ell = (g_1(\bar{V}, \bar{V}))^{1/2} d\lambda$$

If g_1 is indefinite then curves can have

$d\ell$ positive (space like) or negative (timelike)

$d\ell$ is the proper distance for space like curves and the proper time for time-like curves

§ 2.31 Special Relativity

\mathbb{R}^4 with a metric with signature +2 is a manifold called Minkowski space time from special relativity

We can define coordinates $(\Delta t, \Delta x, \Delta y, \Delta z)$ then

$$\Delta s^2 = -c^2(\Delta t)^2 + (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2$$

c is the speed of light

Define $x^0 = ct$, $x^1 = x$, $x^2 = y$, $x^3 = z$ then

$$\Delta s^2 = -(\Delta x^0)^2 + (\Delta x^1)^2 + (\Delta x^2)^2 + (\Delta x^3)^2$$

$$\text{or } \Delta s^2 = \sum_{\alpha, \beta} \Delta x^\alpha \Delta x^\beta$$

This is a pseudo norm and satisfies

(N_{ii}) factor out a scalar
(N_{iv}) parallelogram rule

These are what we need to define an inner product

$$\bar{V} \cdot \bar{W} = \gamma_{\alpha\beta} V^\alpha W^\beta$$

Midterm: covers lec 1-9, but none of the special Relativity

To study: make sure you understand all the lectures and assignments [soln's on Sunday]

Francis will give a description of questions to expect

Will post sample formula sheet on Monday.

Lec 11 w Oct 19

§ 3.1 Intro: How a vector field maps a manifold to itself

Recall: A vector field \bar{V} induces an integral curve

$$\frac{dx^i}{d\lambda} = v^i(x^j) \quad \text{all possible coordinates}$$

Properties of Integral Curves:

- ① \exists a unique curve through each $P \in M$
- ② These curves fill the manifold M

to cover the entire manifold
curve has 1 parameter,
need $(n-1)$ more to get
 n -manifold

If M is n dimensional then the set of integral curves are $(n-1)$ dimensional

Curves like this that fill the manifold is a congruence.

These integral curves provide a natural mapping from M to M along \bar{V} . If \bar{V} is C^∞ , then the mapping is diffeomorphic

Such a mapping is lie dragging



§ 3.2 Lie dragging a function.

Suppose f is a function on a manifold M .



P and Q are on the same curve

We define $f(P) = f_{\Delta\lambda}^*(Q)$ → Star denotes new

$$f(P) = f_{\Delta\lambda}^*(Q) \rightarrow \text{How to get to } P \text{ from } Q \text{ along an integral curve}$$

$\Delta\lambda$ is very small, but could be big,

If $f(Q) = f_{\Delta\lambda}^*(P)$ then f is invariant under the map.

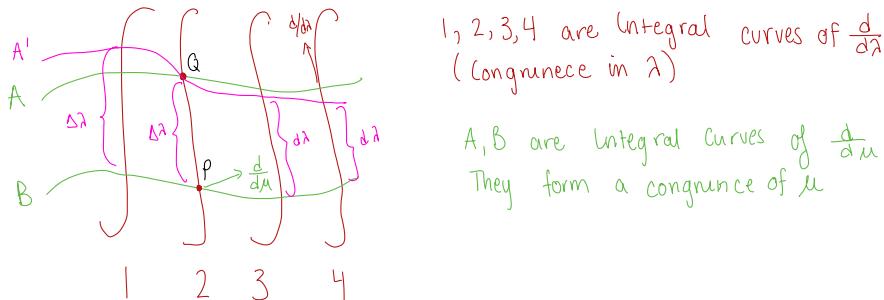
If f is invariant $\forall \Delta\lambda$, then it is said to be lie dragged

If function f that is Lie dragged must satisfy, $\frac{df}{d\lambda} = 0$

§ 3.3 Lie dragging a Vector field

A vector field can be defined by a congruence of curves for which it is tangent.

We now show how to lie drag a vector field



The points along A (μ congruence) are dragged along $\Delta\lambda$ to the curve A' , A' need not be a congruence of μ

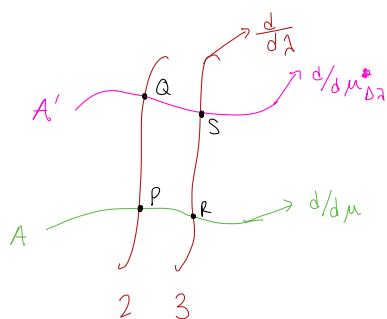
A' defines a new congruence with parameter $\mu_{\Delta\lambda}^*$

This has a tangent vector field $\frac{d}{d\mu_{\Delta\lambda}^*}$, which is the image of $\frac{d}{d\mu}$ under Lie dragging

In general $\mu_{\Delta\lambda}^*$ congruence differs from μ congruence. If they are the same then $\boxed{\frac{d}{d\mu_{\Delta\lambda}^*} = \frac{d}{d\mu}}$ every where

We say a vector field and congruence are invariant under the map

If it is invariant $\forall \Delta\lambda$, then the curves are said to be lie dragged by $d/d\lambda$.



If the distances are infinitesimal and if $\frac{d}{d\mu}$ stretches from P to R on the curve A then:

$\frac{d}{d\mu_{\Delta\lambda}^*}$ stretches from Q to S on A'

If $\frac{d}{d\mu}$ is Lie dragged then B coincides with A and

$$\left(\frac{d}{d\mu} \right)_Q = \left(\frac{d}{d\lambda} \right)_Q$$

This implies $\left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right] = 0$

A vector field is Lie dragged iff $\left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right] = 0$

§ 3.4 Lie derivative.

The derivative of a scalar valued function \mathbb{R} is:

$$\frac{df}{dx} = \lim_{h \rightarrow 0} \frac{f(x+h) - f(x)}{h} \quad \left. \right\} \text{newton's quotient}$$

To compute this we need to compare the function at different points and divide by the distance between them. This has 2 problems:

- ① We don't always have distance but we have the parameter λ along 2 integral curves
- ② Must compare function at different points
we do this by lie dragging

① function:

- METHOD ① evaluate f at $\lambda_0 + \Delta\lambda$, $f(\lambda_0 + \Delta\lambda)$ and drag it back to λ_0
- ② Evaluate f at λ_0
- ③ Find the difference, \div by $\Delta\lambda$ and take the limit as $\Delta\lambda \rightarrow 0$

For ① define f^* such that $\frac{d f^*}{d\lambda} = 0$, hence $f^*(\lambda_0) = f(\lambda_0 + \Delta\lambda)$

Hence we get:

$$\lim_{\Delta\lambda \rightarrow 0} \frac{f^*(\lambda_0) - f(\lambda_0)}{\Delta\lambda}$$

$$= \lim_{\Delta\lambda \rightarrow 0} \frac{f(\lambda_0 + \Delta\lambda) - f(\lambda_0)}{\Delta\lambda} = \left. \frac{df}{d\lambda} \right|_{\lambda_0}$$

The lie derivative of a function is $\mathcal{L}_{\bar{V}} f = \bar{V}(f) = \frac{df}{d\lambda}$

Lec 12 - Oct 24

§ 3.4 Lie derivatives (Continued)

The lie derivative of a function $f: M \rightarrow \mathbb{R}$ along a vector field \bar{V} is computed using lie dragging

Recall, $f(\varphi) = f_{\Delta\lambda}^*(Q)$ and $f(\lambda_0) = f_{\Delta\lambda}^*(\lambda_0 + \Delta\lambda)$

If we $\Delta\lambda \rightarrow -\Delta\lambda$ and $\lambda_0 \rightarrow \lambda_0 + \Delta\lambda$ then

$$f(\lambda_0) = f_{-\Delta\lambda}^*(\lambda_0 - \Delta\lambda)$$

$$f(\lambda_0 + \Delta\lambda) = f_{-\Delta\lambda}^*(\lambda_0)$$

Then the lie derivative of f along \bar{V} is

$$\begin{aligned} (\mathcal{L}_{\bar{V}} f)_{\lambda_0} &= \lim_{\Delta\lambda \rightarrow 0} \frac{f_{-\Delta\lambda}^*(\lambda_0) - f(\lambda_0)}{\Delta\lambda} \\ &= \lim_{\Delta\lambda \rightarrow 0} \frac{f(\lambda_0 + \Delta\lambda) - f(\lambda_0)}{\Delta\lambda} \end{aligned}$$

$$\Rightarrow (\mathcal{L}_{\bar{V}} f)_{\lambda_0} = \left[\frac{df}{d\lambda} \right]_{\lambda_0} = \bar{V}(f)_{\lambda_0}$$

$$\text{In component form } (\mathcal{L}_{\bar{V}} f)_{\lambda_0} = \left(\frac{\partial x^i}{\partial \lambda} \frac{\partial f}{\partial x^i} \right)_{\lambda_0}$$

This generalizes to any differentiable manifold

Most text book's use push back / pull forwards.

→ what is a derivative along a vector?

To compute the lie derivative of a vector field, consider $\bar{U} = \frac{d}{d\mu}$ and $\bar{Y} = \frac{d}{d\lambda}$ and consider an arbitrary function f

At λ_0 and $\lambda_0 + \Delta\lambda$ we know that the lie derivative is

$$(\mathcal{L}_{\bar{U}} f)_{\lambda_0} = \left(\frac{df}{d\mu} \right)_{\lambda_0} = (\bar{U}(f))_{\lambda_0}$$

$$(\mathcal{L}_{\bar{U}} f)_{\lambda_0 + \Delta\lambda} = \left(\frac{df}{d\mu} \right)_{\lambda_0 + \Delta\lambda} = (\bar{U}(f))_{\lambda_0 + \Delta\lambda}$$

Now, we can lie drag $\bar{U}(\lambda_0 + \Delta\lambda)$ to λ_0 with $\bar{U}_{-\Delta\lambda}^*(\lambda_0) = \bar{U}(\lambda_0 + \Delta\lambda) = \frac{d}{d\mu_{-\Delta\lambda}}$

With $[\bar{U}_{-\Delta\lambda}^*, \bar{V}] = 0$ because of what Francis said in lecture [transcribable later]

We taylor expand $\left[\frac{d}{d\mu_{-\Delta\lambda}} f \right]_{\lambda_0}$

$$\left[\frac{d}{d\mu_{-\Delta\lambda}} f \right]_{\lambda_0 + \Delta\lambda} = \left[\frac{d}{d\mu_{-\Delta\lambda}} f \right]_{\lambda_0} + \Delta\lambda \left[\frac{d}{d\lambda} \left(\frac{d}{d\mu_{-\Delta\lambda}} f \right) \right]_{\lambda_0} + O(\Delta\lambda^2)$$

If we solve for the first term on the RHS,

$$\left[\frac{d}{d\mu_{-\Delta\lambda}} f \right]_{\lambda_0} = \left[\frac{d}{d\mu_{-\Delta\lambda}} f \right]_{\lambda_0 + \Delta\lambda} - \Delta\lambda \left[\frac{d}{d\lambda} \left(\frac{d}{d\mu_{-\Delta\lambda}} f \right) \right]_{\lambda_0}$$

BUT $\frac{d}{d\mu_{-\Delta\lambda}} = \frac{1}{d\mu}$

$$\left[\frac{d}{d\mu^*_{-\Delta\lambda}} f \right]_{\lambda_0} = \left[\frac{df}{d\mu} \right]_{\lambda_0 + \Delta\lambda} - \Delta\lambda \left[\frac{d}{d\lambda} \left(\frac{d}{d\mu^*_{-\Delta\lambda}} f \right) \right]_{\lambda_0}$$

↙ Taylor Expand this term
change order because $[\bar{u}_{-\Delta\lambda}, \bar{v}] = 0$

$$\left[\frac{d}{d\mu^*_{-\Delta\lambda}} f \right]_{\lambda_0} = \left[\frac{df}{d\mu} \right]_{\lambda_0} + \Delta\lambda \left[\frac{d}{d\lambda} \left(\frac{df}{d\mu} \right) \right]_{\lambda_0} - \Delta\lambda \left[\frac{d}{d\mu^*_{-\Delta\lambda}} \left(\frac{df}{d\lambda} \right) \right]_{\lambda_0} + O(\Delta\lambda^2)$$

We define the Lie derivative of \bar{u} along \bar{v} is

$$\begin{aligned} (\mathcal{L}_{\bar{v}} \bar{u})(f) &= \lim_{\Delta\lambda \rightarrow 0} \left[\frac{\bar{u}_{-\Delta\lambda(\lambda_0)} - \bar{u}(\lambda_0)}{\Delta\lambda} \right] (f) \\ &= \lim_{\Delta\lambda \rightarrow 0} \left[\left(\frac{d\bar{u}}{d\mu^*_{-\Delta\lambda}} \right)_{\lambda_0} - \left(\frac{d\bar{u}}{d\mu} \right)_{\lambda_0} \right] / \Delta\lambda \end{aligned}$$

Aside *

$$\left[\frac{d}{d\mu^*_{-\Delta\lambda}} f \right]_{\lambda_0} - \left[\frac{df}{d\mu} \right]_{\lambda_0} = \Delta\lambda \left[\frac{d}{d\lambda} \left(\frac{df}{d\mu} \right) \right]_{\lambda_0} - \Delta\lambda \left[\frac{d}{d\mu^*_{-\Delta\lambda}} \left(\frac{df}{d\lambda} \right) \right]_{\lambda_0} + O(\Delta\lambda^2)$$

Substitute *

$$(\mathcal{L}_{\bar{v}} \bar{u})(f) = \lim_{\Delta\lambda \rightarrow 0} \left[\frac{d}{d\lambda} \frac{d}{d\mu} f - \frac{d}{d\mu^*_{-\Delta\lambda}} \frac{d}{d\lambda} f \right] + O(\Delta\lambda)$$

$$(\mathcal{L}_{\bar{v}} \bar{u})(f) = \left[\frac{d}{d\lambda}, \frac{d}{d\mu} \right](f) \quad \text{or} \quad = [\bar{v}, \bar{u}](f)$$

↗ Lie bracket is antisymmetric

→ This is $O(\Delta\lambda)$ b/c \bar{u} divided by $\Delta\lambda$ in the limit

This is equivalent to the directional derivative of \bar{u} in the direction of \bar{v}

§ 3.5 Lie derivative of a one-form

We can determine the Lie derivative of a 1-form in terms of the Lie derivative of a function and Vector field

The lie derivative of a one form can be computed as follows:

$$\mathcal{L}_{\bar{v}} [\tilde{\omega}(\bar{w})] = \frac{d}{d\lambda} [\tilde{\omega}(\bar{w})]$$

Leibniz rule is Product rule

$$= \frac{d\tilde{\omega}}{d\lambda}(\bar{w}) + \tilde{\omega}\left(\frac{d\bar{w}}{d\lambda}\right)$$

OR

$$\mathcal{L}_{\bar{v}} [\tilde{\omega}(\bar{w})] = (\mathcal{L}_{\bar{v}} \tilde{\omega}) \bar{w} + \tilde{\omega}(\mathcal{L}_{\bar{v}} \bar{w})$$

↗ This is the operator, there is an imaginary function at the end of all these terms

This method extends to the outer product of tensors

$$\mathcal{L}_{\bar{v}} (A \otimes B) = \mathcal{L}_{\bar{v}}(A) \otimes B + A \otimes \mathcal{L}_{\bar{v}}(B)$$

Or,

$$\mathcal{L}_{\bar{v}} (T(\tilde{\omega}, \dots; \bar{v}, \dots)) = (\mathcal{L}_{\bar{v}} T)(\tilde{\omega}, \dots; \bar{v}, \dots) + T(\mathcal{L}_{\bar{v}} \tilde{\omega}, \dots; \bar{v}, \dots) + \dots + T(\tilde{\omega}, \dots; \mathcal{L}_{\bar{v}}(\bar{v}), \dots) + \dots$$

for all components of the tensor

This is the Product or Leibniz Rule

S 3.6 Submanifold

The idea is that a sub manifold S of a manifold M is a subset of M which is itself a manifold.

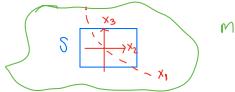
Ex \mathbb{R}^3 is a manifold, a smooth surface is a submanifold.

a smooth curve is a submanifold

A m -dimensional submanifold S of a n -dimensional is a sub set of M with the property that in some neighbourhood of P

$P \in S \subset M$, there exists a coordinate system for M in which the points of S can be written as

$$x^1 = x^2 = \dots = x^{n-m} = 0$$



Sols to the system of DEs $y'_i = f_i(x^1, \dots, x^n)$ with $i=1, \dots, p$ and with coordinates $\{y_1, \dots, y_n, x^1, \dots, x^n\}$. This is a sub manifold

Suppose $P \in S \subset M$ with $\dim S = m$ and $\dim M = n$. A curve in S through P is a curve in both M & S , Through P .

$T_p S$: Tangent space at P in S ($\dim m$)

$T_p M$: Tangent space at P in M ($\dim n$)

$T_p S$ is a vector subspace of $T_p M$ and a submanifold

A tangent vector at P is both in $T_p S$ and $T_p M$

$T_p^* S$: cotangent space at P in S
 $T_p^* M$: cotangent space at P in M

Any $\tilde{\omega} \in T_p^* M$ yields a $\tilde{\omega} \in T_p^* S$ if we restrict the domain to $T_p S$ instead of $T_p M$.

However, $\tilde{\omega} \in T_p^* S$ does not yield a unique $\tilde{\omega}$ in $T_p^* M$

Summary: $\tilde{V} \in T_p S$ is also a vector in $T_p M$ and $\tilde{\omega} \in T_p^* M$ is also a one form in $T_p^* S$

Lec 13 - Oct 26th

Lie derivatives

$$(\mathcal{L}_V U)^i = V^r \frac{\partial}{\partial x^r} U^i - U^r \frac{\partial}{\partial x^r} V^i$$

$$(\mathcal{L}_V \tilde{\omega})_i = V^r \frac{\partial}{\partial x^r} \omega_i + \omega_r \frac{\partial}{\partial x^i} V^r$$

In general for $T^{i...j}_{k...l}$

$$(\mathcal{L}_V T)^{i...j}_{k...l} = V^r \frac{\partial}{\partial x^r} T^{i...j}_{k...l} - T^{r...j}_{k...e} \frac{\partial}{\partial x^r} V^e - \dots - T^{i...r}_{k...e} \frac{\partial}{\partial x^i} V^j + T^{i...j}_{r...e} \frac{\partial}{\partial x^k} V^r + \dots + T^{i...j}_{k...r} \frac{\partial}{\partial x^e} V^r$$

§ 3.7 Frobenius Thm (Vector Field Version)

Suppose a coordinate patch of $S \subset M$ has coordinates y^a $a = \{1, \dots, n\}$ with basis vectors $\{\frac{\partial}{\partial y^a}\}$ for vector fields on S , with $[\frac{\partial}{\partial y^a}, \frac{\partial}{\partial y^b}] = 0 \quad \forall a \neq b$. Since it is a coordinate basis

↑ non coordinate basis

It can be shown that in general for λ , the Lie bracket of any of these two vector fields, yields a vector field tangent to S .

The next theorem says something about the sub manifold if we know a property of the Lie bracket of a Vector Field.

Frobenius' theorem (Vector Field Version)

If a set of m smooth vector fields in $U \subset M$ have Lie brackets which is a linear combination of the m vector field, then the integral curves of the fields mesh to form a family of sub manifolds

Implications

Dim of the submanifold is $\leq m$

Each point in U is on one and only on submanifold. This family of sub manifold is a foliation of U and fills U like the congruence curve do. Each sub manifold is a leaf.

§ 3.9 An Example: the generation of S^2

Consider a ϕ -based vector in spherical coordinates called $\bar{e}_\phi = -y \bar{e}_x + x \bar{e}_y$

Using our notion, this becomes

$$\frac{\partial}{\partial \phi} = -y \frac{\partial}{\partial x} + x \frac{\partial}{\partial y} = \vec{l}_z$$

angular momentum operator
in the Z direction

Similarly

$$\bar{L}_x = -z \frac{\partial}{\partial y} + y \frac{\partial}{\partial z}$$

$$\bar{L}_y = -x \frac{\partial}{\partial z} + z \frac{\partial}{\partial x}$$

It can be shown

$$[\bar{L}_x, \bar{L}_y] = -\bar{L}_z$$

$$[\bar{L}_y, \bar{L}_z] = -\bar{L}_x$$

$$[\bar{L}_z, \bar{L}_x] = -\bar{L}_y$$

Checking $[\bar{L}_x, \bar{L}_y] = (-z \frac{\partial}{\partial y} + y \frac{\partial}{\partial z})(-x \frac{\partial}{\partial z} + z \frac{\partial}{\partial x}) - (-x \frac{\partial}{\partial z} + z \frac{\partial}{\partial x})(-z \frac{\partial}{\partial y} + y \frac{\partial}{\partial z})$

$$= y \frac{\partial^2}{\partial x^2} - x \frac{\partial^2}{\partial y^2} = -\bar{L}_z$$

Since $\{\bar{L}_x, \bar{L}_y, \bar{L}_z\}$ have Lie brackets that are a linear combination of the set, Frobenius theorem yields integral curves that form a Submanifold

Since we have 3 vector fields, we might think the dimension of this set is 3, it turns out the dim is 2.

To see this consider $r = (x^2 + y^2 + z^2)^{1/2}$

We can show that $\tilde{\partial}r$ is the gradient of r

$$\tilde{\partial}r(\bar{L}_x) = \tilde{\partial}r(\bar{L}_y) = \tilde{\partial}r(\bar{L}_z) = 0$$

check: $\tilde{\partial}r(\bar{L}_x) = \bar{L}_x(r)$

$$= (-z \frac{\partial}{\partial y} + y \frac{\partial}{\partial z}) \sqrt{x^2 + y^2 + z^2}$$

B/c of symmetry of the operators, these all equal zero.

$$= -z \left(\frac{y}{r} \right) + y \left(\frac{z}{r} \right) = 0$$

Since the gradients are 0 these exist in the tangent space $\Rightarrow \dim 2$

We can consider $\tilde{\partial}r$ to be a set of surfaces of constant r, S^2 . Since $\bar{L}_x, \bar{L}_y, \bar{L}_z$ are orthogonal to the gradient, they must all lie in the tangent, which is 2 dim $\therefore \{\bar{L}_x, \bar{L}_y, \bar{L}_z\}$ is only dimension 2

§ 3.10 Invariance

Lie derivatives are often used to show that a tensor is invariant in a direction

We say T is invariant under a vector field \bar{V} if

$$\mathcal{L}_{\bar{V}} T = 0$$

T could be ① metric tensor

② a scalar field for PE of a particle

③ Vectors \bar{V} under which T is invariant are important.

§ 3.11 killing Vector fields

Metric tensor can be invariant with respect to a vector field. These vector fields are important.

A killing vector field is a vector field, \bar{v} , such that

$$\mathcal{L}_{\bar{v}} g_{ij} = 0$$

From exercise 3.4, you deduce that $(\mathcal{L}_{\bar{v}} g_{ij})_{ij} = V^k \frac{\partial}{\partial x^k} g_{ij} + g_{kj} \frac{\partial}{\partial x^i} V^k + g_{ik} \frac{\partial}{\partial x^j} V^k = 0$

For a killing vector field

Pick coordinates such that the integral curve are in the x^i direction, then

$$V^i = \delta_i^i$$

Then the above simplifies,

$$(\mathcal{L}_{\bar{v}} g_{ij})_{ij} = \frac{\partial}{\partial x^i} g_{ij} = 0$$

Therefore the metric tensor is invariant with respect to the killing vector

Example: Consider \mathbb{R}^3 in the different coordinates

$$\textcircled{1} \quad \text{Euclidean space} \quad g_{ij} = \delta_{ij}$$

This form is independent of x, y, z and $\therefore \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}$ are all killing vectors

$$\textcircled{2} \quad \text{Spherical coordinates} \quad g_{rr} = \frac{\partial}{\partial r} \cdot \frac{\partial}{\partial r} = 1$$

$$g_{\theta\theta} = \frac{\partial}{\partial \theta} \cdot \frac{\partial}{\partial \theta} = r$$

$$g_{\phi\phi} = \frac{\partial}{\partial \phi} \cdot \frac{\partial}{\partial \phi} = r^2 \sin^2 \theta$$

g_{ij} is independent of ϕ hence \bar{e}_z is a killing vector

$\rightarrow g_{ij}$ is a $(0,2)$ tensor and only diagonal is non zero

It can be shown that \bar{e}_x and \bar{e}_y are also killing vectors

These 6 killing vectors $\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}, \bar{e}_x, \bar{e}_y, \bar{e}_z$ are the only killing vectors possible

§ 3.12 killing Vectors and Conserved quantities in particle dynamics.

In classical mechanics it follows that:

$\textcircled{1}$ if the momentum is conserved P.E function is axially symmetric, then the angular

Q If the PE is independent of say x , then the x component of moment is conserved

these symmetries in the PE energy function give rise to conserved quantities.

However if another symmetry is found in PE function does that mean something else is conserved?

Conserved quantities don't just require the PE is invariant w.r.t to a variable but we also require that is a killing vector

Idea: Newton's 2nd Law

$$m \ddot{V} = -\vec{\nabla} \Phi \quad \text{or} \quad m \dot{V}^i = -\vec{\nabla}^i \Phi \quad \text{or} \quad m \dot{V}^i = -g^{ij} \frac{\partial}{\partial x^j} \Phi$$

Any invariance of this equation that both Φ and g_{ij}^{-1} are invariant w.r.t coordinate

Lec 14 - Oct 13th

Claim: most abstract concepts have been introduced. Chapter 4 is on differential forms

§4 Differential forms

Now, we develop calculus of differential forms or often called exterior calculus or differentiable manifolds

A The algebraic & integral calculus of forms

§4.1 Defn of Volume and the geometric role of differential forms

we now study a class of tensors that enable us to define volume on elements (without an inner product)

A pair of non parallel vectors in Euclidean space defines an infinitesimal area



In our definition of area (volume), we do not need to know the length of the vector or the angle between them.

Consider a 2D manifold and suppose we have two linearly independent infinitesimal vectors. They form a parallelogram

We want to find the area between \bar{a}, \bar{b} . Our definition of area must satisfy the following $\bar{a} \bar{b} \bar{c}$

$$\begin{array}{ccc} \bar{a} & \text{area } (\bar{a}, \bar{b}) + \text{area } (\bar{a}, \bar{c}) = \text{area } (\bar{a}, \bar{b} + \bar{c}) \\ \bar{b} & & \end{array}$$

$$\bar{a} \begin{array}{|c|c|} \hline \bar{b} & \bar{c} \\ \hline \end{array} = \bar{a} \begin{array}{|c|c|} \hline \bar{b} & \bar{b} + \bar{c} \\ \hline \end{array}$$

Since $\text{area}(\cdot, \cdot)$ takes 2 vectors and yields a number it must be a $\binom{0}{2}$ tensor

$$\text{Observe, } \text{area } (\bar{a}, \bar{a}) = 0 \quad \text{for all } \bar{a}$$

Exercise 4.1 If B is a $\binom{0}{2}$ tensor with $B(\bar{u}, \bar{u}) = 0 \quad \forall \bar{u}$ then $B(\bar{u}, \bar{w}) = -B(\bar{w}, \bar{u})$

Proof $B(\bar{u} + \bar{w}, \bar{u} + \bar{w}) = 0$

$$\cancel{B \text{ is linear operator}} \quad B(\bar{u}, \bar{u}) + B(\bar{w}, \bar{u}) + B(\bar{u}, \bar{w}) + B(\bar{w}, \bar{w}) = 0$$

$$B(\bar{u}, \bar{w}) = -B(\bar{w}, \bar{u})$$

- Note (1) that the area function must satisfy this anti symmetry property
(2) Area is not non negative

Recall for Linear algebra, we can find area using

$$\text{area} = \det \begin{vmatrix} V_x & V_y \\ W_x & W_y \end{vmatrix} \quad \text{Anti symmetric}$$

§ 4.2 Notation and definitions for anti symmetric tensors

A $(\frac{0}{2})$ tensor is anti symmetric if

$$\tilde{\omega}(\bar{u}, \bar{v}) = -\tilde{\omega}(\bar{v}, \bar{u})$$

A $(\frac{0}{3})$ tensor is anti symmetric if it changes signs when we exchange any 2 elements

$$\begin{aligned} \tilde{\omega}(\bar{u}, \bar{v}, \bar{w}) &= -\tilde{\omega}(\bar{v}, \bar{u}, \bar{w}) \\ &= -\tilde{\omega}(\bar{w}, \bar{v}, \bar{u}) \\ &= -\tilde{\omega}(\bar{v}, \bar{w}, \bar{u}) \end{aligned}$$

Given any tensor we can build an antisymmetric version of it.

Ex. If $\tilde{\omega}$ a $(\frac{0}{2})$ tensor then

$$\tilde{\omega}_A(\bar{u}, \bar{v}) = \frac{1}{2!} [\tilde{\omega}(\bar{u}, \bar{v}) - \tilde{\omega}(\bar{v}, \bar{u})]$$

This is the anti symmetric part of $\tilde{\omega}$

If \tilde{p} is a $(\frac{0}{3})$ tensor then

$$\begin{aligned} \tilde{p}_A(\bar{u}, \bar{v}, \bar{w}) &= \frac{1}{3!} [\tilde{p}(\bar{u}, \bar{v}, \bar{w}) + \tilde{p}(\bar{v}, \bar{w}, \bar{u}) + \tilde{p}(\bar{w}, \bar{u}, \bar{v}) - \tilde{p}(\bar{w}, \bar{v}, \bar{u}) - \tilde{p}(\bar{v}, \bar{u}, \bar{w})] \\ &\quad \text{Normalizing by number of terms} \end{aligned}$$

\tilde{p}_A is the anti symmetric part of \tilde{p}

Notation:

$$(\tilde{\omega}_A)_{ij} = \frac{1}{2!} (\omega_{ij} - \omega_{ji}) \equiv \omega_{[ij]} \quad \text{this denotes anti symmetric part of } \tilde{\omega}$$

$$(\tilde{p}_A)_{ijk} = \frac{1}{3!} (p_{ijk} + p_{jki} + p_{kij} - p_{kji} - p_{jik} - p_{ikj}) \equiv p_{[ijk]}$$

$[i \dots k]$ denotes a completely anti symmetric set of indices

Notation: we use \sim to denote a completely anti symmetric part of a tensor.

example: T is only a tensor & \tilde{T} is the anti symmetric version of T

Also we say a one form is anti symmetric

Property: For an n -dimensional vector space, a completely antisymmetric $\binom{0}{p}$ tensor ($p \leq n$) has at most
 How many ways can we pick $C_p^n = \frac{n!}{p!(n-p)!} = \binom{n}{p}$ "n choose p" independent components
 p directions in n space?

Why be anti symmetric? will learn later on!

$$\text{ex } \text{In } \mathbb{R}^3 \quad n=3 \quad C_1^3 = 3, \quad C_2^3 = 3, \quad C_3^3 = 1$$

$$\begin{array}{c} x \\ y \\ z \end{array} \quad \begin{array}{c} xy \\ xz \\ yz \end{array} \quad \begin{array}{c} x,y,z \end{array}$$

§ 4.3 Differential forms

A p -form ($p \geq 2$) is a completely anti symmetric tensor of type $\binom{0}{p}$

A one-form is a $\binom{0}{1}$ tensor (by convention anti symmetric)

A zero-form is a $\binom{0}{0}$ tensor (scalar) ???

$\Rightarrow p$ is the degree

Using \otimes (outer product) can take 2 $\binom{0}{1}$ forms to yield a $\binom{0}{2}$ tensor.

A wedge product takes two one forms and yields a 2-form.

i claim

this is
anti symmetric

$$\tilde{p} \wedge \tilde{q} = \tilde{p} \otimes \tilde{q} - \tilde{q} \otimes \tilde{p} = -\tilde{q} \wedge \tilde{p}$$

$$\tilde{q} \wedge \tilde{p} = \tilde{q} \otimes \tilde{p} - \tilde{p} \otimes \tilde{q}$$

Property: If $\{\bar{e}_i : i=1, \dots, n\}$ is a basis of $T_p M$ and $\{\tilde{\omega}^j\}$ is the dual basis of $T_p^* M$
 then $\{\tilde{\omega}^j \wedge \tilde{\omega}^k, j, k=1 \dots n\}$ is a basis for the vector space of two-forms

We can build two-forms in a similar way

$$\tilde{p} \wedge (\tilde{q} \wedge \tilde{r}) = (\tilde{p} \wedge \tilde{q}) \wedge \tilde{r} = \tilde{p} \wedge \tilde{q} \wedge \tilde{r}$$

$$\tilde{p} \wedge \tilde{q} \wedge \tilde{r} = \tilde{p} \otimes \tilde{q} \otimes \tilde{r} + \tilde{q} \otimes \tilde{r} \otimes \tilde{p} + \tilde{r} \otimes \tilde{p} \otimes \tilde{q} - \tilde{r} \otimes \tilde{q} \otimes \tilde{p} - \tilde{q} \otimes \tilde{p} \otimes \tilde{r} - \tilde{p} \otimes \tilde{r} \otimes \tilde{q}$$

We can define the wedge product of a p -form and a q -form

§ 4.4 Manipulating differential forms

Commutation rule of form: $\tilde{p} \wedge \tilde{q} = \tilde{q} \wedge \tilde{p} (-1)^{pq}$

idea: if $\tilde{p} = \tilde{\omega}^i \wedge \dots \wedge \tilde{\omega}^j$ p factors

$\tilde{q} = \tilde{\omega}^k \wedge \dots \wedge \tilde{\omega}^l$ q factors

$$\tilde{p} \wedge \tilde{q} = (\tilde{\omega}^i \wedge \dots \wedge \tilde{\omega}^j) \wedge (\tilde{\omega}^k \wedge \dots \wedge \tilde{\omega}^l)$$

Lecture 15 - Nov 2, 2023

§ 4.4 Manipulation of differential forms

Commutation rule: $\tilde{p} \wedge \tilde{q} = (-1)^{pq} \tilde{q} \wedge \tilde{p}$

Interior Product / Contraction of a vector with a form

If $\tilde{\alpha}$ is a p -form and \bar{v} is a vector, then $\tilde{\alpha}$ requires p vectors

$$\tilde{\alpha}(\bar{v}) \equiv \tilde{\alpha}(\bar{v}, \underbrace{\dots}_{p-1 \text{ arguments left}}) \quad \text{or} \quad \alpha_{ij\dots k} v^i \xrightarrow{\text{text books notation}}$$

$$i_{\bar{v}}(\tilde{\alpha}) = \alpha_{ij\dots k} v^i \xrightarrow{\text{other textbooks notation}}$$

This is an inner product

Example:

$$i_{\bar{v}}(\tilde{\omega}^i \wedge \tilde{\omega}^j \wedge \dots \wedge \tilde{\omega}^k) = V^i \tilde{\omega}^j \wedge \dots \wedge \tilde{\omega}^k - V^j \tilde{\omega}^i \wedge \dots \wedge \tilde{\omega}^k + \dots \xrightarrow{\text{more terms check textbook}}$$

§ 4.5 Restrictions to forms

Suppose W is a subspace of a Vector field V . A p -form, $\tilde{\alpha}$, is a $\binom{p}{p}$ tensor that is (completely) antisymmetric and its arguments could be

$$\underbrace{V \times V \times \dots \times V}_{p \text{-times}}$$

The restriction of $\tilde{\alpha}$ to the subspace W^c is the same p -form but with the domain restricted to W ,

$$\tilde{\alpha}|_W(\bar{x}, \dots, \bar{y}) = \tilde{\alpha}(\bar{x}, \dots, \bar{y}) \quad \text{where } \bar{x}, \dots, \bar{y} \text{ are in } W$$

If $m = \dim W < p$ then $\tilde{\alpha}|_W$ is 0

If $m = p$ then $\tilde{\alpha}|_W$ has one component, $C_p = 1$

Restricted a form is called sectioning

A form is anulled by a Vector space if its restriction to it vanishes

§ 4.6 Fields of forms

A field of p -forms on manifold M gives a p -form Λ points on the manifold M .

A sub manifold $S \subset M$ picks a subspace $T_p S$ for all $p \in S$ and we define the restriction of the

p -form $\tilde{\omega}$ to S by restricting $\tilde{\omega}$ at P to $T_p S$.

§ 4.7 Handedness and Orientability

In a n -dimensional manifold there is a 1-dimensional space of n -forms ($C_n^n = 1$)

Suppose that $\tilde{\omega}$ is an n -form field that we can use to find the Volume. If we have $\{\bar{e}_1, \dots, \bar{e}_n\}$ is a vector basis of $T_p M$ that is linearly independent

It follows that $\tilde{\omega}(\bar{e}_1, \dots, \bar{e}_n) \neq 0$ iff $\tilde{\omega} \neq 0$ at P

Aside: $\tilde{\omega} \Rightarrow \omega_{i_1 \dots i_n} \tilde{\omega}_{i_1} \wedge \tilde{\omega}_{i_2} \wedge \dots \wedge \tilde{\omega}_{i_n}$

Consider

$$\begin{aligned}\tilde{\omega}(\bar{e}_1, \dots, \bar{e}_n) &= \omega_{i_1 \dots i_n} (\tilde{\omega}^1 \wedge \tilde{\omega}^2 \wedge \dots \wedge \tilde{\omega}^n)(\bar{e}_1, \bar{e}_2, \dots, \bar{e}_n) \\ &= \omega_{i_1 \dots i_n} \tilde{\omega}^1(\bar{e}_1) \wedge \tilde{\omega}^2(\bar{e}_2) \wedge \dots \wedge \tilde{\omega}^n(\bar{e}_n) \\ &= \omega_{i_1 \dots i_n}\end{aligned}$$

$\tilde{\omega}$ separates the vector bases into 2 classes

① $\tilde{\omega}(\bar{e}_1, \dots, \bar{e}_n) > 0$ (right hand)

② $\tilde{\omega}(\bar{e}_1, \dots, \bar{e}_n) < 0$ (left hand)

This separation is unique to any n -form.

This manifold is said to be orientable if we define the handedness consistently (all positive or all negative) on the manifold

Example \mathbb{R}^n is orientable

Möbius is not orientable

We only consider orientable manifold

§ 4.8 Volumes and Integration on Oriented Manifolds

A set of n linearly independent vectors (infinitesimal) or an n -dim manifold, can define a non-zero volume. This forms in n -dim a parallelepiped

Integration a function f on M requires multiply f by an infinitesimal Volume then adding this up over all of M

Suppose \tilde{w} is a n -form on an open set U in M with coordinates

$$\{x^1, \dots, x^n\}$$

Since n -forms at $P \in M$ form a 1-D vector space. There exists a function $f(x^1, \dots, x^n)$ such that

$$\tilde{\omega} = f \tilde{d}x^1 \wedge \dots \wedge \tilde{d}x^n$$

We integrate over U by first dividing U into regions (cells) spanned by n -tuples of vectors.

$$\left\{ \Delta x^1 \frac{\partial}{\partial x^1}, \Delta x^2 \frac{\partial}{\partial x^2}, \dots, \Delta x^n \frac{\partial}{\partial x^n} \right\}$$

Where Δx^i 's are infinitesimal.

The Integral of f over a region is f multiplied by the following

$$\Delta x^1 \Delta x^2 \dots \Delta x^n = (\tilde{\Delta}x^1 \wedge \dots \wedge \tilde{\Delta}x^n) (\Delta x^1 \frac{\partial}{\partial x^1}, \dots, \Delta x^n \frac{\partial}{\partial x^n})$$

This looks like d ✓

The integral of f over a cell is written as

$$\int_{\text{Cell}} f(x^1, \dots, x^n) d^n x \approx \tilde{\omega} \quad \begin{matrix} \curvearrowright \\ \text{locally near} \end{matrix} \quad p$$

Add over all the cells and we set the integral

Integrating over region of U

$$\int \tilde{\omega} \stackrel{n\text{-form}}{\equiv} \int f(x^1, \dots, x^n) dx^1 \dots dx^n$$

we will show this is coordinate independent.

Example In 2D with coordinates (λ, u) the above yields

$$\int \tilde{\omega} = \int f(\lambda, u) \hat{d}\lambda \wedge \hat{d}u \rightarrow \text{calculus of manifold}$$

$$= \int f(\lambda, u) d\lambda du \quad \rightarrow \text{Calculus of } \mathbb{R}^2$$

Check transformation of coordinate $(\lambda, \mu) \rightarrow (x, y)$

$$\tilde{d}\lambda = \tilde{d}\lambda(x,y) = \frac{\partial \lambda}{\partial x} \tilde{d}x + \frac{\partial \lambda}{\partial y} \tilde{d}y$$

$$\tilde{d}\mu = \frac{\partial \mu}{\partial x} \tilde{d}x + \frac{\partial \mu}{\partial y} \tilde{d}y$$

We build the form:

$$\tilde{\partial} \lambda \wedge \tilde{\partial} \mu = \left(\frac{\partial \lambda}{\partial x} \tilde{\partial}_x + \frac{\partial \lambda}{\partial y} \tilde{\partial}_y \right) \left(\frac{\partial \mu}{\partial x} \tilde{\partial}_x + \frac{\partial \mu}{\partial y} \tilde{\partial}_y \right)$$

$$= \frac{\partial \lambda}{\partial x} \frac{\partial \mu}{\partial x} \overset{0}{\cancel{\tilde{d}_x \wedge \tilde{d}_x}} + \frac{\partial \lambda}{\partial x} \frac{\partial \mu}{\partial y} \overset{0}{\cancel{\tilde{d}_x \wedge \tilde{d}_y}} + \frac{\partial \lambda}{\partial y} \frac{\partial \mu}{\partial x} \overset{0}{\cancel{\tilde{d}_y \wedge \tilde{d}_x}} + \frac{\partial \lambda}{\partial y} \frac{\partial \mu}{\partial y} \overset{0}{\cancel{\tilde{d}_y \wedge \tilde{d}_y}}$$

$$\begin{aligned}
 &= \left(\frac{\partial \lambda}{\partial x} \frac{\partial \mu}{\partial \lambda} - \frac{\partial \lambda}{\partial y} \frac{\partial \mu}{\partial x} \right) \tilde{d}x \wedge \tilde{d}y \\
 &= \frac{\partial(\lambda, \mu)}{\partial(x, y)} \tilde{d}x \wedge \tilde{d}y \quad \text{Jacobian of the transformation}
 \end{aligned}$$

For a n-dimensional manifold we can integrate an n-form to get a non-zero result

For a submanifold of order p, we can integrate a p-form to get a non-zero result

Lec 1b - online

the lectures are online since francis is out of town. Therefore i will not record sound

S 4.9 N-vectors, duals and the symbol $\epsilon_{ijk\dots k}$

A completely anti symmetric $(^p_0)$ tensor is a p vector. On a n-dimensional manifold this has dim C_p^n
The following spaces all have the same dimension

$$\begin{array}{c} p\text{-forms} \\ (n-p)\text{ forms} \\ p\text{-vectors} \\ (n-p)\text{ vectors} \end{array} \left. \right\} \text{size } (^n_p) = (^n_{n-p})$$

Example

If $n=3$

| | | | |
|-----|-------|-------|-------|
| dim | $p=1$ | $p=2$ | $p=3$ |
| | 3 | 3 | 1 |

We could use the metric tensor, to map a $(^p_0)$ tensor to a (^0_p) tensor and backwards

It can be shown that since the metric tensor is symmetric, this process preserves anti symmetry

Even without a metric, the volume n-form, $\tilde{\omega}$, yields a mapping from p-vectors to $(n-p)$ -forms
this map is the dual map or the hodge-star map [not mentioned in text book]

Suppose T is a 2-vector, with components

$$T^{i\dots k} = T^{[i\dots k]}$$

with $\tilde{\omega}$, we can defin a tensor \tilde{A} such that

$$A_{j\dots l} = \frac{1}{q!} \underbrace{\omega_{i\dots k\dots l}}_{(n-q)\text{ form}} \underbrace{T^{i\dots k}}_{n\text{-form}} \quad \text{or} \quad \tilde{A} = \tilde{\omega}(T) \quad \rightarrow \text{indepent of coordinates}$$

Notation: $\tilde{A} = *T$ and say that \tilde{A} is the dual of T w.r.t. $\tilde{\omega}$

This is invertable and therefore we can bring p-forms to $(n-p)$ vectors

Examples: Consider \mathbb{R}^3 in terms of cartesian coordinates

\bar{U}, \bar{V} are both vectors

\rightarrow end of ch?

In cartesian coordinates, the elements of the associated 1-forms are equal (b/c of δ)

$$\tilde{U} = g_{ij}(\bar{U}) \text{ or } U_i = g_{ij} V^j = \delta_{ij} U^j$$

and $\tilde{V} = g_{ij}(\bar{V}) \text{ or } V_i = g_{ij} V^j = \delta_{ij} V^j$

Typically, we write these as

$$\begin{aligned}\tilde{U} &= U_1 \tilde{dx}^1 + U_2 \tilde{dx}^2 + U_3 \tilde{dx}^3 = U_i \tilde{dx}^i \\ \tilde{V} &= V_1 \tilde{dx}^1 + V_2 \tilde{dx}^2 + V_3 \tilde{dx}^3 = V_i \tilde{dx}^i\end{aligned}$$

basis for 1-forms

With \tilde{U} and \tilde{V} we can use the wedge to find the following 2-form

$$\begin{aligned}\tilde{U} \wedge \tilde{V} &= (U_1 \tilde{dx}^1 + U_2 \tilde{dx}^2 + U_3 \tilde{dx}^3) \wedge (V_1 \tilde{dx}^1 + V_2 \tilde{dx}^2 + V_3 \tilde{dx}^3) \\ &= U_1 V_1 \tilde{dx}^1 \wedge \tilde{dx}^1 + U_1 V_2 \tilde{dx}^1 \wedge \tilde{dx}^2 + U_1 V_3 \tilde{dx}^1 \wedge \tilde{dx}^3 + U_2 V_1 \tilde{dx}^2 \wedge \tilde{dx}^1 + U_2 V_2 \tilde{dx}^2 \wedge \tilde{dx}^2 + U_2 V_3 \tilde{dx}^2 \wedge \tilde{dx}^3 \\ &\quad + U_3 V_1 \tilde{dx}^3 \wedge \tilde{dx}^1 + U_3 V_2 \tilde{dx}^3 \wedge \tilde{dx}^2 + U_3 V_3 \tilde{dx}^3 \wedge \tilde{dx}^3\end{aligned}$$

Switching some bases :

$$\begin{aligned}\tilde{U} \wedge \tilde{V} &= (U_1 V_2 - U_2 V_1) \tilde{dx}^1 \wedge \tilde{dx}^2 \\ &\quad + (U_2 V_3 - U_3 V_2) \tilde{dx}^2 \wedge \tilde{dx}^3 \\ &\quad + (U_3 V_1 - U_1 V_3) \tilde{dx}^3 \wedge \tilde{dx}^1\end{aligned}$$

Coefficients are similar to cross product

Find the dual of this expression.

$$\begin{aligned}\star(\tilde{U} \wedge \tilde{V}) &= \frac{1}{2!} \omega_{ijk} (\tilde{U} \wedge \tilde{V})^{ij} \quad \text{Francis claims in lecture, in piazza, I'm not sure this is true} \\ &= \frac{1}{2!} [\omega_{123} (\tilde{U} \wedge \tilde{V})^{12} + \omega_{132} (\tilde{U} \wedge \tilde{V})^{13} \\ &\quad + \omega_{231} (\tilde{U} \wedge \tilde{V})^{23} + \omega_{213} (\tilde{U} \wedge \tilde{V})^{21} \\ &\quad + \omega_{321} (\tilde{U} \wedge \tilde{V})^{32} + \omega_{312} (\tilde{U} \wedge \tilde{V})^{31}] \\ &= \frac{1}{2} [(U_1 V_2 - U_2 V_1)^* \tilde{dx}^1 \wedge \tilde{dx}^2 + (U_3 V_1 - U_1 V_3)^* \tilde{dx}^3 \wedge \tilde{dx}^1 + (U_2 V_3 - U_3 V_2)^* \tilde{dx}^2 \wedge \tilde{dx}^3 + (U_1 V_2 - U_2 V_1)^* \tilde{dx}^1 \wedge \tilde{dx}^2 \\ &\quad + (U_3 V_1 - U_1 V_3)^* \tilde{dx}^3 \wedge \tilde{dx}^1 + (U_2 V_3 - U_3 V_2)^* \tilde{dx}^2 \wedge \tilde{dx}^3] \\ &= (U_1 V_2 - U_2 V_1)^* (\tilde{dx}^1 \wedge \tilde{dx}^2) + (U_2 V_3 - U_3 V_2)^* (\tilde{dx}^2 \wedge \tilde{dx}^3) + (U_3 V_1 - U_1 V_3)^* (\tilde{dx}^3 \wedge \tilde{dx}^1)\end{aligned}$$

How do we find the dual of the 2-forms?

Note that

$$\begin{aligned}\star(\tilde{dx}^1 \wedge \tilde{dx}^2) &= \frac{\partial}{\partial x^3} \\ \star(\tilde{dx}^2 \wedge \tilde{dx}^3) &= \frac{\partial}{\partial x^1} \\ \star(\tilde{dx}^3 \wedge \tilde{dx}^1) &= \frac{\partial}{\partial x^2}\end{aligned}$$

Dual of a
2-form is a one vector

Check: $\star(\tilde{dx}^1 \wedge \tilde{dx}^2) = \frac{\partial}{\partial x^3}, \star\left(\frac{\partial}{\partial x^1}\right) = \tilde{dx}^2 \wedge \tilde{dx}^3$

Proof $\frac{1}{2!} \omega_{ijk} \star^2 = \omega_{ijk} 1 = A_{jk} \quad (2\text{-form})$

$\Rightarrow A_{23} = 1$ with basis $\tilde{dx}^2 \wedge \tilde{dx}^3$

or $A_{32} = -1$ with basis $\hat{d}x^3 \wedge \hat{d}x^2$

$$*(\tilde{U} \wedge \tilde{V}) = (U_1 V_2 - U_2 V_1) \frac{\partial}{\partial x^3} + (U_2 V_3 - U_3 V_2) \frac{\partial}{\partial x^1} + (U_3 V_1 - U_1 V_3) \frac{\partial}{\partial x^2}$$

Compare with the cross product

$$\bar{U} \times \bar{V} = \begin{vmatrix} \hat{x}_1 & \hat{x}_2 & \hat{x}_3 \\ U_1 & U_2 & U_3 \\ V_1 & V_2 & V_3 \end{vmatrix} = \begin{bmatrix} U_2 V_3 - U_3 V_2 \\ U_3 V_1 - U_1 V_3 \\ U_1 V_2 - U_2 V_1 \end{bmatrix} \quad \text{looks like } \begin{bmatrix} \frac{\partial}{\partial x^1} \\ \frac{\partial}{\partial x^2} \\ \frac{\partial}{\partial x^3} \end{bmatrix}$$

$$\therefore *(\tilde{U} \wedge \tilde{V}) = (\bar{U} \times \bar{V}) \Leftrightarrow (\tilde{U} \wedge \tilde{V}) = *(\bar{U} \times \bar{V})$$

This result is unique to \mathbb{R}^3 .

The map between T and $*T$ is invertible.

Levi-Civita Symbols

$$\epsilon_{ijk} = \epsilon^{ijk} = \begin{cases} +1 & \text{if } ijk \text{ is an even permutation} \\ -1 & \text{if } ijk \text{ is an odd permutation} \\ 0 & \text{otherwise} \end{cases}$$

$$\text{example: } \bar{U} \times \bar{V} = \epsilon_{ijk} U^j V^k$$

B the differential Calculus of form and its applications

Single Variable Calculus States

$$\int_a^b df = f(b) - f(a)$$

We want to derive a derivative operator that reduces to this in the simple case, but is more general

§ 4.14 The exterior derivative

If M is a 1D manifold, $\tilde{d}: \text{0-form} \xrightarrow{\text{takes a function}} 1\text{-form}$. It will agree with the above

If \tilde{d} is a p -form $\tilde{\beta}$, $\tilde{\gamma}$ are q -forms we require

$$(1) \quad \tilde{d}(\tilde{\beta} + \tilde{\gamma}) = (\tilde{d}\tilde{\beta}) + (\tilde{d}\tilde{\gamma}) \quad \text{Distributive}$$

$$(2) \quad \tilde{d}(\tilde{\alpha} \wedge \tilde{\beta}) = (\tilde{d}\tilde{\alpha}) \wedge \tilde{\beta} + (-1)^p \tilde{\alpha} \wedge (\tilde{d}\tilde{\beta}) \quad \text{Anti derivation}$$

$$(3) \quad \tilde{d}(\tilde{\gamma} \tilde{\alpha}) = 0$$

These 3 properties uniquely define \tilde{d}

\tilde{d} is called the exterior derivative

Property (ii) is almost Leibniz but there is an extra $(-1)^p$ to bring \tilde{d} across the p-form

Property (iii) seems odd but is essential

ex) if f is a function then $\tilde{d}f$ is a one form with component $\frac{\partial f}{\partial x^i}$

$\tilde{d}(\tilde{d}f)$ has components of the form $\frac{\partial^2 f}{\partial x^i \partial x^j}$. But this must be a 2-form and since it must be anti-symmetric, it must be 0

§ 4.15 Notation for derivatives

Partial derivatives: $\frac{\partial f}{\partial x^i} = f_{,i}$ 1st derivative

$\frac{\partial^2 f}{\partial x^i \partial x^j} = V_{j,k}^{i,j}$ 1st derivative of a (1,1) tensor

$\frac{\partial^2 f}{\partial x^k \partial x^i} = f_{,ik}$ 2nd Derivative

Recall that a partial derivative is not a tensor operation in general [which assignment went over this?]

example ① $V_{j,k}^{i,j}$ need not be $\underbrace{A_{a'}^{i,j}}_{\text{this part breaks down...}} \underbrace{A_{b'}^{b,c}}_{\text{is coordinate dependent}} A_{b',c'}^{a'}$ What is it?

② $\tilde{d}f = f_{,i}$ (1-form) and is a tensor operation

③ $[\bar{u}, \bar{v}]^i = u^j v_{,j} - v^j u_{,j}$ this is a tensor operator

Each term on the RHS is not a tensor operator but the whole RHS is

Exercise 4.14 on an assignment

a) $\tilde{d}(f \tilde{d}g) = \tilde{d}f \wedge \tilde{d}g$ [with 3rd property means $(-1)^p \tilde{d} \tilde{d}g = 0$]

b) if $\tilde{d} = \frac{1}{p!} d_{i_1 \dots i_p} \tilde{d}x^{i_1} \wedge \dots \wedge \tilde{d}x^{i_p}$ is a p-form

the $\tilde{d}\tilde{\alpha} = \frac{1}{p!} \frac{\partial}{\partial x^k} (d_{i_1 \dots i_p}) \tilde{d}x^k \wedge \tilde{d}x^{i_1} \wedge \dots \wedge \tilde{d}x^{i_p}$

and

$$(\tilde{d}\tilde{\alpha})_{k \dots j} = (p+1) \frac{\partial}{\partial x^k} [\alpha_{i_1 \dots i_p}]$$

or $(\tilde{d}\tilde{\alpha})_{k \dots j} = (p+1) d_{[i_1 \dots i_p, k]}$

Lec 17 - Online

§ 4.1b Familiar Examples of exterior derivatives

We can revisit some old friends with a new perspective

① $\tilde{\delta}$ of a 1-form \tilde{a} in 3D:

$$\begin{aligned} \tilde{\delta} \tilde{a} &= \tilde{\delta}(a_1 \tilde{dx}^1 + a_2 \tilde{dx}^2 + a_3 \tilde{dx}^3) \quad \text{two-form} \\ &= a_{1,j} \tilde{dx}^j \wedge \tilde{dx}^1 + a_{2,j} \tilde{dx}^j \wedge \tilde{dx}^2 + a_{3,j} \tilde{dx}^j \wedge \tilde{dx}^3 \quad \text{all the components that are not zero due to anti symmetry} \\ &= a_{1,2} \tilde{dx}^2 \wedge \tilde{dx}^1 + a_{1,3} \tilde{dx}^3 \wedge \tilde{dx}^1 + a_{2,1} \tilde{dx}^1 \wedge \tilde{dx}^2 + a_{2,3} \tilde{dx}^3 \wedge \tilde{dx}^2 + a_{3,1} \tilde{dx}^1 \wedge \tilde{dx}^3 + a_{3,2} \tilde{dx}^2 \wedge \tilde{dx}^3 \\ &= (a_{3,2} - a_{2,3}) \tilde{dx}^2 \wedge \tilde{dx}^3 + (a_{1,3} - a_{3,1}) \tilde{dx}^3 \wedge \tilde{dx}^1 + (a_{2,1} - a_{1,2}) \tilde{dx}^1 \wedge \tilde{dx}^2 \rightarrow \text{combing like terms by switching the two bases in the wedge} \end{aligned}$$

Consider the dual, or hodge-star

$$*\tilde{\delta}\tilde{a} = (a_{3,2} - a_{2,3})*(\tilde{dx}^2 \wedge \tilde{dx}^3) + (a_{1,3} - a_{3,1})*(\tilde{dx}^3 \wedge \tilde{dx}^1) + (a_{2,1} - a_{1,2})*(\tilde{dx}^1 \wedge \tilde{dx}^2)$$

$$\text{From before } *(\tilde{dx}^2 \wedge \tilde{dx}^3) = \frac{\partial}{\partial x^1} \quad *(\tilde{dx}^3 \wedge \tilde{dx}^1) = \frac{\partial}{\partial x^2} \quad *(\tilde{dx}^1 \wedge \tilde{dx}^2) = \frac{\partial}{\partial x^3}$$

$$\text{Hence, } *\tilde{\delta}\tilde{a} = (a_{3,2} - a_{2,3})\frac{\partial}{\partial x^1} + (a_{1,3} - a_{3,1})\frac{\partial}{\partial x^2} + (a_{2,1} - a_{1,2})\frac{\partial}{\partial x^3} \rightarrow \text{Looks like the curl}$$

or

$$*\tilde{\delta}\tilde{a} = \epsilon_{ijk} \frac{\partial}{\partial x^j} a_k = \epsilon_{ijk} a_{k,j}$$

The RHS is the curl since $\bar{\nabla} \times \bar{a} = \epsilon_{ijk} a_{k,j}$

Summary $*\tilde{\delta}$ = curl in 3D when applied to 1-forms, when exterior derivative is applied to a scalar it gives the gradient, as seen in ③

② First take the dual of a vector then $\tilde{\delta}$

Suppose \bar{a} is a vector field

$$\begin{aligned} *(\bar{a}) &= * (a^1 \frac{\partial}{\partial x^1} + a^2 \frac{\partial}{\partial x^2} + a^3 \frac{\partial}{\partial x^3}) = a^1 *(\frac{\partial}{\partial x^1}) + a^2 *(\frac{\partial}{\partial x^2}) + a^3 *(\frac{\partial}{\partial x^3}) \\ &= a^1 (\tilde{dx}^2 \wedge \tilde{dx}^3) + a^2 (\tilde{dx}^3 \wedge \tilde{dx}^1) + a^3 (\tilde{dx}^1 \wedge \tilde{dx}^2) \end{aligned}$$

$$\begin{aligned} \text{Apply } \tilde{\delta} \quad \tilde{\delta}(*\bar{a}) &= a^1_{,j} (\tilde{dx}^j \wedge \tilde{dx}^2 \wedge \tilde{dx}^3) + a^2_{,j} (\tilde{dx}^1 \wedge \tilde{dx}^3 \wedge \tilde{dx}^j) + a^3_{,j} (\tilde{dx}^1 \wedge \tilde{dx}^2 \wedge \tilde{dx}^j) \\ &= (a^1_{,1} + a^2_{,2} + a^3_{,3}) + \tilde{dx}^1 \wedge \tilde{dx}^2 + \tilde{dx}^3 \end{aligned}$$

Only 1 value for each j that would make the wedges non-zero

$$\text{Note } \tilde{\delta} * \bar{a} = (\bar{\nabla} \cdot \bar{a}) \tilde{w} \quad (\text{divergence})$$

$$③ \tilde{\delta} f = f_{,i} \tilde{dx}^i \quad (\text{gradient})$$

§ 4.17 Integrability conditions for PDEs

Consider the system of 2 PDEs:

$$\frac{\partial f}{\partial x} = g(x, y) \text{ and } \frac{\partial f}{\partial y} = h(x, y)$$

More interested in when a solution exists

Let (x, y) be coordinates of a manifold. Further, define
re write the system compactly as $\alpha_x = g$ and $\alpha_y = h$ and then we can
one form notation $\alpha_i = a_i$ $i=1, 2$

$\alpha_x = g$ and $\alpha_y = h$
vector components

A coordinate-independent version of this is

$$\tilde{\delta}f = \tilde{\alpha} \rightarrow \text{why } \tilde{\alpha} ? \text{ I thought } \alpha_x, \alpha_y \text{ was a vector}$$

This holds because $\frac{\partial f}{\partial x} \tilde{dx} + \frac{\partial f}{\partial y} \tilde{dy} = g \tilde{dx} + h \tilde{dy}$

If f is a soln then it must follow that

$$\tilde{\delta}(\tilde{\delta}f) = \tilde{\delta}\tilde{\alpha} = 0$$

In component form this becomes

$$\begin{aligned} \tilde{\delta}\tilde{\alpha} &= \tilde{\delta}(g \tilde{dx} + h \tilde{dy}) \\ &= \frac{\partial g}{\partial y} \tilde{dy} \wedge \tilde{dx} + \frac{\partial h}{\partial x} \tilde{dx} \wedge \tilde{dy} \quad \rightarrow \frac{\partial g}{\partial x} = 0 \text{ b/c of anti-symmetry} \\ &= \left(\frac{\partial h}{\partial x} - \frac{\partial g}{\partial y} \right) \tilde{dx} \wedge \tilde{dy} = 0 \quad \rightarrow \text{Green's theorem?} \end{aligned}$$

or $\frac{\partial h}{\partial x} = \frac{\partial g}{\partial y} = \frac{\partial}{\partial y} \left(\frac{\partial f}{\partial y} \right) = \frac{\partial}{\partial x} \left(\frac{\partial f}{\partial x} \right)$

$$a_{2,1} - a_{1,2} = 0 \quad \text{or } a_{[i,j]} = 0$$

where does this notation come from?

This condition is necessary and later we will show it is sufficient for a solution to exist

§ 4.18 Exact forms

Observe, if $\tilde{\alpha} = \tilde{\delta}\tilde{\beta}$ then $\tilde{\delta}\tilde{\alpha} = \tilde{\delta}(\tilde{\delta}\tilde{\beta}) = 0$

If $\tilde{\delta}\tilde{\alpha} = 0$ then $\tilde{\alpha}$ is closed

If $\tilde{\alpha} = \tilde{\delta}\tilde{\beta}$ then $\tilde{\alpha}$ is exact

Clearly an exact form is closed. It can be shown that any closed form is exact

§ 4.20 Lie derivatives of forms

If $\tilde{\omega}$ is a p -form then

$$\mathcal{L}_{\bar{v}} \tilde{\omega} = \tilde{\partial} [\tilde{\omega}(\bar{v})] + (\tilde{\partial} \tilde{\omega})(\bar{v})$$

p-form $\xrightarrow{\text{p-1 form}}$ p-form
 p-form $\xrightarrow{\text{p-1 form}}$ p-form

Idea of proof:

case 1 $\tilde{\omega}$ is a 0-form, $\tilde{\omega} = f$

$$\text{LHS } \mathcal{L}_{\bar{v}} \tilde{\omega} = \mathcal{L}_{\bar{v}} f = \bar{V}(f) = \frac{df}{d\lambda}$$

(RHS1) does not make sense, this term is ignored we can't have a function acting on a vector field

$$(\tilde{\partial} \tilde{\omega})(\bar{v}) = (\tilde{\partial} f)(\bar{v}) = \frac{\partial f}{\partial x^i} \tilde{\partial} x^i (\nu \frac{\partial}{\partial x^j})$$

$$(\tilde{\partial} \tilde{\omega})(\bar{v}) = \nu^i f_{,i} \tilde{\partial} x^i \left(\frac{\partial}{\partial x^j} \right)$$

$$= \nu^i f_{,i} = \frac{df}{d\lambda}$$

Case 2: $\tilde{\omega}$ is a 1-form $\tilde{\omega} = \omega_i \tilde{\partial} x^i$ [one form notation]

$$(\text{RHS1}) \quad \tilde{\partial}(\tilde{\omega}(\bar{v})) = \tilde{\partial}(\omega_i v^i) = (\omega_i v^i)_{,j} \tilde{\partial} x^j$$

$$(\text{RHS2}) \quad (\tilde{\partial} \tilde{\omega})(\bar{v}) = (\tilde{\partial}(\omega_i \tilde{\partial} x^i))(\bar{v})$$

$$= (\omega_{i,j} \tilde{\partial} x^j \wedge \tilde{\partial} x^i)(\bar{v}) \xrightarrow{\text{which formula does this use?}}$$

$$= \omega_{i,j} (\tilde{\partial} x^j \otimes \tilde{\partial} x^i - \tilde{\partial} x^i \otimes \tilde{\partial} x^j) (\nu^k \frac{\partial}{\partial x^k}) \xrightarrow{\text{new index for } v} \text{translation of } \Lambda \text{ to outer product}$$

$$= \omega_{i,j} (\tilde{\partial} x^j (\nu^k \frac{\partial}{\partial x^k}) \otimes \tilde{\partial} x^i - \tilde{\partial} x^i (\nu^k \frac{\partial}{\partial x^k} \otimes \tilde{\partial} x^j)) \xrightarrow{\text{distributing } \nu^k \frac{\partial}{\partial x^k}}$$

$$= \omega_{i,j} (\nu^j \tilde{\partial} x^i - \nu^i \tilde{\partial} x^j) \xrightarrow{\text{where did the cross go?}}$$

$$\text{RHS} = (\omega_i v^i)_{,j} \tilde{\partial} x^j + \omega_{i,j} (\nu^j \tilde{\partial} x^i - \nu^i \tilde{\partial} x^j)$$

$$= \omega_{i,j} \nu^i \tilde{\partial} x^j + \omega_i v^i \tilde{\partial} x^j + \omega_{i,j} (\nu^j \tilde{\partial} x^i - \nu^i \tilde{\partial} x^j)$$

$$= \omega_{i,j} \nu^j \tilde{\partial} x^i + \omega_j \nu^j \tilde{\partial} x^i$$

$$= (\omega_{i,j} \nu^j + \omega_j \nu^j) \tilde{\partial} x^i = (\mathcal{L}_{\bar{v}} \tilde{\omega})_i$$

Can fully prove using induction

Lec 18 - Nov 14

§ 4.21 Lie derivatives and Exterior derivative commute

Woo hoo Stokes theorem!!!

Thm: \mathcal{L}_v and $\tilde{\delta}$ commute

We will prove this for a one-form or n-form?

proof: we need a formula from § 4.20

$$\mathcal{L}_v \tilde{\omega} = \tilde{\delta}[\tilde{\omega}(v)] + (\tilde{\delta}\tilde{\omega})(v) \quad \left. \begin{array}{l} \text{Still have} \\ \text{n-form} \end{array} \right\}$$

\downarrow
Gives tensor
of same space?

but replace $\tilde{\omega}$ with $\tilde{\delta}\tilde{\omega}$ → true for n-l exact form

↑ this is not yet able of how $[\tilde{\delta}\tilde{\omega}(v)]$ gets calculated

$$\mathcal{L}_v \tilde{\delta}\tilde{\omega} = \tilde{\delta}[\tilde{\delta}\tilde{\omega}(v)] + (\tilde{\delta}(\tilde{\delta}\tilde{\omega}))(v) \quad \text{B/c of properties of } \tilde{\delta}$$

From the first equation we know

$$\tilde{\delta}\tilde{\omega}(v) = \mathcal{L}_v \tilde{\omega} - \tilde{\delta}[\tilde{\omega}(v)]$$

Substitute into the RHS of the previous equation

$$\begin{aligned} \mathcal{L}_v \tilde{\delta}\tilde{\omega} &= \tilde{\delta}[\mathcal{L}_v \tilde{\omega} - \tilde{\delta}[\tilde{\omega}(v)]] \quad \text{the } [\] \text{ around } \tilde{\omega}(v) \text{ make an n-l form,} \\ &\quad \text{hence } \tilde{\delta} \text{ is acting on an n-l form and} \\ &\quad \text{is zero} \\ &= \tilde{\delta} \mathcal{L}_v \tilde{\omega} \quad \because \text{commutes} \end{aligned}$$

§ 4.22 Stokes thm

We show that the exterior derivative is the inverse of integration, in particular

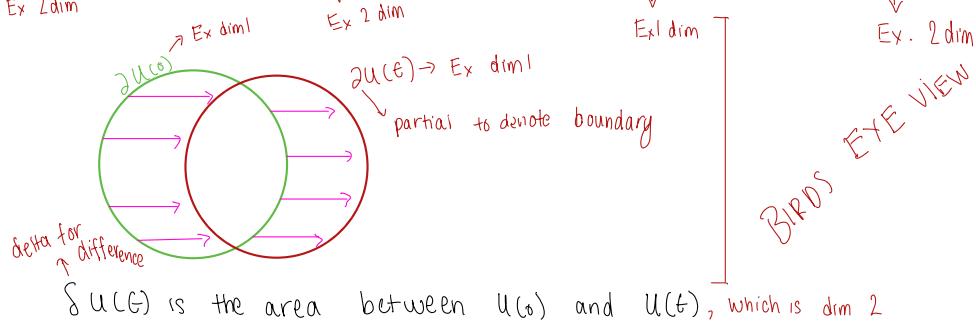
$$\int_U \tilde{\delta}\tilde{\omega} = \int_{\partial U} \tilde{\omega} \quad \text{Different from } \tilde{\omega} \text{ we use}$$

We can integrate n-forms over n-dimension. and we can integrate n-l forms over n-l dimension.

If U is n-dimensional then its boundary is n-l dimensional. The boundary is the exterior of U , and why this is called exterior calculus

Assume, U is a smooth, orientable volume on M that is connected. then ∂U is a submanifold of M . also \vec{z} is a vector field on M

Suppose $U = U(0)$ is a region on M with boundary $\partial U = \partial U(0)$, and $U(t)$ is the Lie dragged region along β



$$\text{or } \delta U(\epsilon) \equiv U(\epsilon) - U(0)$$

To find the change of an integral from $u(0)$ to $u(t)$ we compute the following

$$\int_{U(t)} \tilde{\omega} - \int_{U(0)} \tilde{\omega} = \int_{\tilde{U}(t)} \tilde{\omega}$$

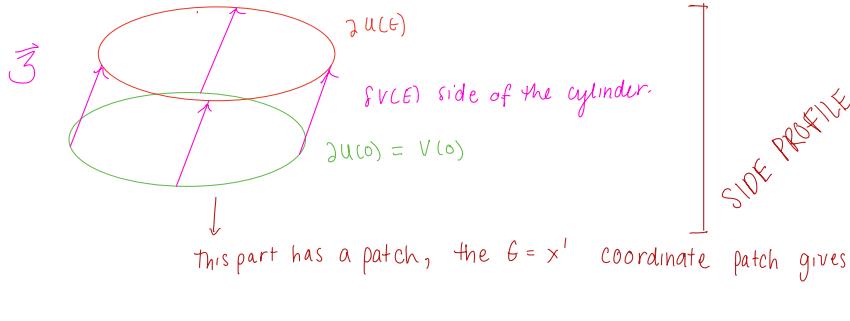
\downarrow

n-form

$\tilde{\omega}$ is an n -form, the n -form that gives us volume. Anti-symmetric and $\binom{0}{n}$ tensor
↗ NO x^i !!

Suppose V is a coordinate patch of $\mathcal{M}(t)$ with coordinates $\{x^1, \dots, x^n\}$

If we lie drag $\mathcal{D}u(\phi)$ along $\vec{\gamma}$ a distance of ϵ , then we have coordinates $\{x^1 = \epsilon, x^2, \dots, x^n\}$



\vec{z} is nowhere tangent to $\partial U(0)$

We investigate the integral of $SV(\epsilon)$ and then extend this to $SU(\epsilon)$

We introduce,

If $\epsilon \ll 1$, then the integral over $\delta V(E)$ is,

This is like
Riemann Sums

$$\tilde{\omega} = \int_{\partial V(E)} v(0) \left[\int_0^t f dx' \right] dx^2 \dots dx^n$$

$$= \epsilon \int_{v(0)}^{\infty} f(0, x^2, \dots, x^n) dx^2 \dots dx^n + O(\epsilon) \xrightarrow{\text{Smaller than } \epsilon} \text{Linearizing since } \epsilon \ll 1$$

$$= \epsilon \int_{\gamma(\sigma)} \tilde{\omega}(\bar{z}) \Big|_{\gamma(\sigma)} + o(\epsilon) \quad \text{This is a change to Calculus on manifolds}$$

Note: $\bar{z} = \frac{\partial}{\partial x^i}$ by design

$$\tilde{\omega}(\frac{\partial}{\partial x^i}) = f(x^1, \dots, x^n) \tilde{\delta}^{x^1} \wedge \dots \wedge \tilde{\delta}^{x^n} (\frac{\partial}{\partial x^i})$$

$$= f(0, x^2, \dots, x^n) \tilde{\delta}^{x^2} \wedge \dots \wedge \tilde{\delta}^{x^n}$$

and $\int_0^\epsilon f dx^i \approx f(0, x^2, \dots, x^n) + o(\epsilon)$

Summary, $\int_{\partial V(\epsilon)} \tilde{\omega} = \epsilon \int_{V(0)} \tilde{\omega}(\bar{z}) \Big|_{\partial U} + o(\epsilon) *$

Now consider,

$$\begin{aligned} \frac{d}{d\epsilon} \int_{U(\epsilon)} \tilde{\omega} &= \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \left[\int_{U(\epsilon)} \tilde{\omega} - \int_{U(0)} \tilde{\omega} \right] \\ &= \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \int_{\partial U(\epsilon)} \tilde{\omega} \quad \text{From first eqn *} \end{aligned}$$

If $\delta U(\epsilon) \approx \delta V(\epsilon)$, which is true if $\epsilon \ll 1$, then

$$\frac{d}{d\epsilon} \int_{U(\epsilon)} \tilde{\omega} = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon} \left(\epsilon \int_{V(0)} \tilde{\omega}(\bar{z}) \Big|_{\partial U} + o(\epsilon) \right)$$

$$\boxed{\frac{d}{d\epsilon} \int_{U(\epsilon)} \tilde{\omega} = \int_{\partial U(0)} \tilde{\omega}(\bar{z}) \Big|_{\partial U(0)}}$$

But using linearization we can approximate the integrand on the LHS using

$$\tilde{\omega} = \epsilon \mathcal{L}_{\bar{z}} \tilde{\omega} + o(\epsilon) \quad ? \text{ Francis doest get where this comes from}$$

on $U(0)$

Sub into LHS

$$\frac{d}{d\epsilon} \int_{U(\epsilon)} \tilde{\omega} = \frac{d}{d\epsilon} \int_{U(0)} \epsilon \mathcal{L}_{\bar{z}} \tilde{\omega} + o(\epsilon)$$

$$\frac{d}{d\epsilon} \int_{U(\epsilon)} \tilde{\omega} = \int_{U(0)} \mathcal{L}_{\bar{z}} \tilde{\omega}$$

But the formula for $\mathcal{L}_{\bar{z}} \tilde{\omega}$ yields

$$\begin{aligned} \int_{U(0)} \mathcal{L}_{\bar{z}} \tilde{\omega} &= \int_{U(0)} \tilde{\delta}[\tilde{\omega}(\bar{z})] + (\tilde{\delta}(\tilde{\omega}))(\bar{z}) \quad \text{this is an nt form} \\ &= \int_{U(0)} \tilde{\delta}[\tilde{\omega}(\bar{z})] \end{aligned}$$

an hence $\rightarrow 0$

If we combine our formulas we set

$$\frac{d}{de} \int_{U(e)} \tilde{\omega} = \int_{U(0)} \tilde{\delta} [\tilde{\omega}(\bar{z})] = \int_{\partial U(0)} \tilde{\omega}(\bar{z})$$

From above

before formula
with L

Define $\tilde{\alpha} = \tilde{\omega}(\bar{z}) = i_{\bar{z}} \tilde{\omega}$ and get

$$\boxed{\int_{U(0)} \tilde{\delta} \tilde{\alpha} = \int_{\partial U(0)} \tilde{\alpha}} \quad \text{Stokes theorem}$$

Lec 19 ~ Nov 16th

PHYSICS NEXT WEEK

§ 4.22 Stokes theorem

Recall $\frac{d}{dt} \int_{U(t)} \tilde{\omega} = \int_{U(0)} \mathcal{L}_{\vec{Z}} \tilde{\omega}$

Justification:

LHS: We integrate $\tilde{\omega}$ (n-form) over the n-volume $U(t)$

To obtain the RHS we Taylor expand $\tilde{\omega}$ at $U(t)$ around $U(0)$

Aside: $f(\vec{x}_0 + t\vec{z}) = f(\vec{x}_0) + t\vec{z} \cdot \vec{\nabla} f(\vec{x}_0) + o(t)$ → little o

Use this idea to appox $\tilde{\omega}$,

Since the Lie derivative is the directional derivative

$$\tilde{\omega}|_{U(t)} = \tilde{\omega}|_{U(0)} + t \mathcal{L}_{\vec{z}} \tilde{\omega}|_{U(0)} + o(t)$$

If you plug this into the LHS,

$$\frac{d}{dt} \int_{U(t)} \tilde{\omega} = \frac{d}{dt} \int_{U(0)} \tilde{\omega} + t \mathcal{L}_{\vec{z}} \tilde{\omega} + o(t)$$

$$= \int_{U(0)} \mathcal{L}_{\vec{z}} \tilde{\omega} + \text{small stuff}$$

Example: In E^2 consider $\tilde{\omega} = \alpha_1 \tilde{\delta}x^1 + \alpha_2 \tilde{\delta}x^2$ can convert to vector

we apply $\tilde{\delta}$ and obtain

$\tilde{\delta}\tilde{\omega} = \tilde{\delta}(\alpha_1 \tilde{\delta}x^1 + \alpha_2 \tilde{\delta}x^2)$ is a product rule going on [Next assignment, comes out Friday Nov 17]

$$= \alpha_{1,2} \tilde{\delta}x^2 \wedge \tilde{\delta}x^1 + \alpha_{2,1} \tilde{\delta}x^1 \wedge \tilde{\delta}x^2$$

$$= \alpha_{2,1} - \alpha_{1,2} \tilde{\delta}x^1 \wedge \tilde{\delta}x^2$$

Stokes' theorem can be written as

$$\int_U \tilde{\delta}\tilde{\omega} = \int_{\partial U} \tilde{\omega}$$

$$\int_U (\alpha_{2,1} - \alpha_{1,2}) \tilde{\delta}x^1 \wedge \tilde{\delta}x^2 = \int_{\partial U} \alpha_1 \tilde{\delta}x^1 + \alpha_2 \tilde{\delta}x^2$$

$$(d_1, d_2) \cdot (\tilde{\delta}x^1, \tilde{\delta}x^2)$$

We can rewrite this in terms of "regular" integrals

$$\int_U \left(\frac{\partial \omega_2}{\partial x} - \frac{\partial \omega_1}{\partial y} \right) dx dy = \oint_{\partial U} \omega \cdot d\bar{x}$$

This is greens theorem,
a special case of stokes thm

§ 4.23 Gauss' theorem and the defn of divergence

Recall Stokes thm can be written as

$$\int_U \tilde{\omega} = \int_{\partial U} \tilde{\omega} \quad \text{or} \quad \boxed{\int_U \tilde{\omega}[\tilde{\omega}(\bar{z})] = \int_{\partial U} \tilde{\omega}(\bar{z}) \Big|_{\partial U}} \quad *$$

not really need

Suppose $\tilde{\omega} = \tilde{\omega}^1 dx^1 \wedge \tilde{\omega}^2 dx^2 \wedge \dots \wedge \tilde{\omega}^n dx^n$ is it because $\frac{\partial}{\partial x^i} \lambda \tilde{\omega}^i = 0$

then $\tilde{\omega}(\bar{z}) = \bar{z}^1 \tilde{\omega}^2 \wedge \dots \wedge \tilde{\omega}^n - \bar{z}^2 \tilde{\omega}^1 \wedge \tilde{\omega}^3 \wedge \dots \wedge \tilde{\omega}^n + \dots$
where does the negative come from?

We compute $\tilde{\omega}$ of the above and get

$$\begin{aligned} \tilde{\omega}[\tilde{\omega}(\bar{z})] &= \cancel{\bar{z}^1 \tilde{\omega}^2 \wedge \dots \wedge \tilde{\omega}^n} + \underline{\bar{z}^2 \tilde{\omega}^1 \wedge \tilde{\omega}^3 \wedge \dots \wedge \tilde{\omega}^n} + \dots + \cancel{\bar{z}^n \tilde{\omega}^1 \wedge \dots \wedge \tilde{\omega}^{n-1}} \\ &\Rightarrow \tilde{\omega}[\tilde{\omega}(\bar{z})] = \bar{z}^i \tilde{\omega}^i \end{aligned}$$

We define the $\tilde{\omega}$ -divergence of \bar{z} as $(\operatorname{div}_{\tilde{\omega}} \bar{z}) \tilde{\omega} \equiv \tilde{\omega}[\tilde{\omega}(\bar{z})]$

If we use components such that ∂U is a surface of constant x^i , then the restriction of $\tilde{\omega}(\bar{z})$ to ∂U is

$$\tilde{\omega}(\bar{z}) \Big|_{\partial U} = \bar{z}^1 \tilde{\omega}^2 \wedge \dots \wedge \tilde{\omega}^n$$

or

$$= \tilde{\omega}^1(\bar{z}) \tilde{\omega}^2 \wedge \dots \wedge \tilde{\omega}^n$$

In general if \tilde{n} is a 1-form normal to the boundary of $U(\partial U)$, which means that $\tilde{n}(\bar{n}) = 0$ $\forall \bar{n}$ tangent to ∂U

and if $\tilde{\omega}$ is an $(n-1)$ -form with

$$\tilde{\omega} = \tilde{n} \wedge \tilde{\omega}$$

then

$$\tilde{\omega}(\bar{z}) \Big|_{\partial U} = \tilde{n}(\bar{z}) \tilde{\omega} \Big|_{\partial U}$$

Therefore the original form of Stokes theorem becomes

$$\int_U (\operatorname{div} \tilde{\omega}) \tilde{\omega} = \int_{\partial U} \tilde{n} \cdot \tilde{\omega}$$

with $\tilde{\omega}$ restricted to ∂U and $\tilde{n} \cdot \tilde{\omega} = \tilde{\omega}$

In component form, this becomes

$$\int_U \tilde{z}_{,i} d^n x = \int_{\partial U} \tilde{z}_{,i} n_i d^{n-1} x$$

Gauss' divergence theorem in \mathbb{R}^n

§ 4.25 Differential forms and Differential Equations

Consider the DE $\frac{dy}{dx} = f(x, y)$ What's the connection between the two?
 we often rewrite it as $dy = f(x, y) dx$

If M is a 2D manifold with coordinates (x, y) , then we consider the following

$$\tilde{d}y - f(x, y) \tilde{d}x = 0 \rightarrow \text{this is inspiration}$$

where f is a function on M .

Suppose \tilde{V} is a vector at $P \in M$ with components $(1, f(P))$

$$\text{Consider } \tilde{d}y(\tilde{V}) = \tilde{d}y(1, f(P)) = f(P)$$

$$\tilde{d}x(\tilde{V}) = \tilde{d}x(1, f(P)) = 1$$

$$\text{This implies, } (\tilde{d}y - f \tilde{d}x)(\tilde{V}) = 0 = \tilde{d}y(\tilde{V}) - f \tilde{d}x(\tilde{V}) = f(P) - f = 0$$

Sols to the DEs define a submanifold of M whose tangent annul the 1-form Send to 0

Sub manifolds that annul the 1-form are Solutions to this can be generalized to n -forms with Frobenius thm

Question Given a DE, what are the equivalent form?

example $\frac{dx^2}{dt^2} + \omega_0^2 x = \omega_0$ is constant. Harmonic oscillator

$$\text{or } \frac{dx}{dt} = \omega_0 y \quad \text{and} \quad \frac{dy}{dt} = -\omega_0 x \quad \text{System of 1st Order Equations}$$

or $\frac{dx}{dt} - \omega_0 y = 0$ and $\frac{dy}{dt} + \omega_0 x = 0$

The 1-forms to consider are:

$$\tilde{\alpha} = \tilde{dx} - \omega_0 y \tilde{dt}$$

$$\tilde{\beta} = \tilde{dy} + \omega_0 x \tilde{dt}$$

Finding submanifolds that annul these forms is equivalent to solving DEs.

The manifold is 3D with coordinates $[x, y, z]$ and the solution is 1D.

§ 4.26 Frobenius' theorem (differential forms version)

The set of forms $\{\tilde{\beta}_i\}$ at $P \in M$ define a subspace of vectors, $T_p S \subset T_p M$, each of which annuls $\tilde{\beta}_i$, i.e. For all $\tilde{v} \in T_p S$, $\tilde{\beta}_i(\tilde{v}) = 0 \quad \forall i = 1, \dots, n$

The set $T_p S$ is called annihilator of $\{\tilde{\beta}_i\}$

The complete ideal consists of all the forms at P whose restriction to $T_p S$ vanishes

Note: if $\tilde{\gamma}$ is a form at P then $\tilde{\gamma} \wedge \tilde{\beta}_i$ is 0 when restricted to $T_p S$ and therefore $\tilde{\gamma} \wedge \tilde{\beta}_i$ is in the complete ideal

A complete ideal has a basis $\{\tilde{\alpha}_j\}$ that generates the ideal ie

the complete ideal of $\{\tilde{\alpha}_j\}$ is the same as the complete ideal of $\{\tilde{\beta}_i\}$

All of this extends from vectors to vector fields

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S 4.26 Frobenius theorem.

$\{\tilde{\beta}_i\}$ defines a subspace, $T_p S \subset T_p M$, each of which annuls $\tilde{\beta}_i$.

$$\forall \tilde{V} \in T_p S \text{ then } \tilde{\beta}_i(\tilde{V}) = 0 \quad \forall i=1,\dots,n$$

$T_p S$ is the annihilator of $\{\tilde{\beta}_i\}$.

The complete ideal consists of all the forms whose restriction to $T_p S$ vanishes.

Note: If $\tilde{\gamma}$ is a form then $\tilde{\gamma} \wedge \tilde{\beta}_i$ is 0 when restricted to $T_p S$ $\therefore \tilde{\gamma} \wedge \tilde{\beta}_i$ is in the complete ideal \star

$\{\tilde{\alpha}_j\}$ is closed if each $\tilde{d}\tilde{\alpha}_j$ is in the complete ideal generated by $\{\tilde{\alpha}_j\}$

Aside: A complete ideal has a basis $\{\tilde{\alpha}_j\}$ that generates ideal

Frobenius Theorem:

Suppose $\{\tilde{\alpha}_i, i=1,\dots,m\}$ is a linearly independent set of 1-form fields in an open set $U \subset M$, where

M is an n -dimensional manifold. The set $\{\tilde{\alpha}_i\}$ is closed iff functions $\{P_{ij}, Q_j | j=1\dots m\}$ such that

$$\tilde{\alpha}_i = \sum_{j=1}^m P_{ij} \tilde{d}Q_j$$

Idea: In general to solve ODEs, we want to find solutions to $\{\tilde{\alpha}_i = 0\}$. The solution to this set of equations by $Q_j = \text{constant}$

This set of Q_j are solution to the equations $\{\tilde{\alpha}_i = 0\}$ and each Q_j defines an m -dimensional Submanifold of M and its tangent Vectors annul $\{\tilde{d}Q_j\}$ and also $\{\tilde{\alpha}_i\}$

example: suppose $\tilde{\alpha} = \tilde{d}f$ this satisfies the above with $P_{ii} = 1$ and $f = Q$. f exists iff $\tilde{d}\tilde{d}f = 0$

exercise 4.30 $\{\tilde{\alpha}_j, j=1\dots m\}$ is a linearly independent set of 1-forms then any form $\tilde{\gamma}$ is in the complete ideal iff $\tilde{\gamma} \wedge \tilde{\alpha}_1 \wedge \tilde{\alpha}_2 \wedge \dots \wedge \tilde{\alpha}_m = 0$

S 5 Applications to Physics

S 5 A Thermodynamics

S 5.1 Simple Systems

Consider a one-component fluid where the conservation of energy dictates that

$$\boxed{\delta Q = P \delta V + dU}$$

1st Law of Thermodynamics

\downarrow
variation (changes?) path dependent?

where U is the internal energy

δQ heat absorbed

$P\delta V$ work done by the fluid

P, V pressure and volume

assuming good coordinates

This law can be written in terms of 1-forms on a 2D manifold with coordinates (V, U)

Then $P(V, U)$ is a function on M that is the equation of state
↑ internal energy
↓ volume

On the RHS, it would make sense to write it as

$$P \tilde{d}V + \tilde{d}U$$

may not be exact

Since $\tilde{d}V$ and $\tilde{d}U$ are one forms, we deduce that the LHS $\tilde{\delta}Q$ is a 1-form as well.

Question is $\tilde{\delta}Q = \tilde{\delta}Q$, is it an exact one-form? If yes, then $\tilde{\delta}\tilde{\delta}Q = 0$ and we deduce

$$0 = \tilde{\delta}(\tilde{\delta}Q) = \tilde{\delta}(P \tilde{d}V + \tilde{d}U)$$

gradient of function again! gradient of P

$$0 = \tilde{\delta}(P \tilde{d}V) = \boxed{\tilde{\delta}P \wedge \tilde{d}V} \quad \text{Assignment 5.}$$

$$\tilde{\delta}P \wedge \tilde{d}V + P \tilde{\delta}\tilde{d}V \rightarrow 0$$

$$\left(\left(\frac{\partial P}{\partial V} \right)_U \tilde{d}V + \left(\frac{\partial P}{\partial U} \right)_V \tilde{d}U \right) \wedge \tilde{d}V = 0$$

$$\left(\frac{\partial P}{\partial U} \right)_V \tilde{d}U \wedge \tilde{d}V = 0$$

But this can only be true if $\left(\frac{\partial P}{\partial U} \right)_V = 0$ This is typically the case.

In general Q does not exist and we can't write $\tilde{\delta}Q$ as $\tilde{\delta}Q$

However, since $\tilde{\delta}Q$ is a 1-form but not exact in 2-space $\tilde{\delta}(\tilde{\delta}Q)$ is a 2-form. This 2-form is in the complete ideal of δQ hence $\tilde{\delta}Q$ is closed.

We can use Frobenius' theorem and deduce that $\exists T(V, U)$ and $S(V, U)$ such that

$$\tilde{\delta}Q = T \tilde{d}S$$

This looks like the 2nd Law of Thermal dynamics.

With this choice the first law becomes

$$T \tilde{d}S = P \tilde{d}V + \tilde{d}U$$

§ 5.2 Maxwell and other mathematical identities

Apply $\tilde{\delta}$ to the above equation

$$\tilde{\delta}(T\tilde{\delta}S) = \tilde{\delta}(P\tilde{\delta}V) + \tilde{\delta}\tilde{\delta}U$$

Assumption 1. presume $T(S, V)$ and $P(S, V)$

partial is proportional to the one form in the gradient calculation.

$$\tilde{\delta}T \wedge \tilde{\delta}S = \tilde{\delta}P \wedge \tilde{\delta}V$$

$$\begin{aligned} \left(\frac{\partial T}{\partial V}\right)_S \tilde{\delta}V \wedge \tilde{\delta}S &= \left(\frac{\partial P}{\partial S}\right)_V \tilde{\delta}S \wedge \tilde{\delta}V \\ &= -\left(\frac{\partial P}{\partial S}\right)_V \tilde{\delta}V \wedge \tilde{\delta}S \end{aligned}$$

$$\Rightarrow \boxed{\left(\frac{\partial T}{\partial V}\right)_S = -\left(\frac{\partial P}{\partial S}\right)_V} \quad \text{One of Maxwell's identities.}$$

Assumption 2 $S(T, V)$ and $P(T, V)$

$$\text{Sub into } \tilde{\delta}(T\tilde{\delta}S) = \tilde{\delta}(P\tilde{\delta}V)$$

$$\tilde{\delta}T \wedge \tilde{\delta}S = \left(\frac{\partial P}{\partial T}\right)_V \tilde{\delta}T \wedge \tilde{\delta}V$$

$$\left(\frac{\partial S}{\partial V}\right)_T \tilde{\delta}T \wedge \tilde{\delta}V = \left(\frac{\partial P}{\partial T}\right)_V \tilde{\delta}T \wedge \tilde{\delta}V$$

$$\Rightarrow \boxed{\left(\frac{\partial S}{\partial V}\right)_T = \left(\frac{\partial P}{\partial T}\right)_V}$$

Assumption : Divide egn by T and apply $\tilde{\delta}$ the you can obtain

$$T \left(\frac{\partial P}{\partial T}\right)_V - P = \left(\frac{\partial U}{\partial V}\right)_T \quad \text{assume } P(T, V) \text{ and } U(T, V)$$

$$T\tilde{\delta}S = P\tilde{\delta}V + \tilde{\delta}U$$

$$\tilde{\delta}P = \left(\frac{\partial P}{\partial V}\right)_T \tilde{\delta}V + \left(\frac{\partial P}{\partial T}\right)_V \tilde{\delta}T$$

Dividing by T

$$\tilde{\delta}S = \frac{P}{T} \tilde{\delta}V + \frac{1}{T} \tilde{\delta}U$$

$$\tilde{\delta}\tilde{\delta}S = \tilde{\delta}\left(\frac{P}{T} \tilde{\delta}V\right) + \tilde{\delta}\left(\frac{1}{T} \tilde{\delta}U\right)$$

$$0 = \left(-\frac{1}{T^2} P \tilde{\delta}T + \frac{1}{T} \tilde{\delta}(P)\right) \wedge \tilde{\delta}V + \left(-\frac{1}{T^2} \tilde{\delta}T \wedge \tilde{\delta}U + \frac{1}{T} \tilde{\delta}\tilde{\delta}U\right) \wedge \tilde{\delta}V \stackrel{?}{=} 0?$$

$$= -\frac{P}{T^2} \tilde{\delta}T \wedge \tilde{\delta}V + \frac{1}{T} \left[\left(\frac{\partial P}{\partial V}\right)_T \tilde{\delta}V + \left(\frac{\partial P}{\partial T}\right)_V \tilde{\delta}T \right] \wedge \tilde{\delta}V - \frac{1}{T^2} \tilde{\delta}T \wedge \tilde{\delta}U$$

$$= -\frac{P}{T^2} \tilde{\delta}T \wedge \tilde{\delta}V + \frac{1}{T} \left(\frac{\partial P}{\partial T}\right)_V \tilde{\delta}T \wedge \tilde{\delta}V - \frac{1}{T^2} \tilde{\delta}T \wedge \tilde{\delta}U + \frac{1}{T^2} \left[\left(\frac{\partial U}{\partial V}\right)_T \tilde{\delta}V + \left(\frac{\partial U}{\partial T}\right)_V \tilde{\delta}T\right] \wedge \tilde{\delta}V$$

$$= -\frac{1}{T^2} \left(\frac{\partial U}{\partial V}\right)_T \tilde{\delta}T \wedge \tilde{\delta}V - \frac{1}{T^2} \left(\frac{\partial U}{\partial T}\right)_V \tilde{\delta}T \wedge \tilde{\delta}V$$

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Phase Space in Mechanics

From section 5.2.3 c of The Geometry of Physics by Frankel

In classical Mechanics we describe a system using generalized coordinates

$$q^1, \dots, q^n \quad \} \text{ compactly called } q$$

These form an n-dimensional manifold M that we call configuration space.

The lagrangian is a function of q_i and \dot{q}_i , where $\dot{q}_i = \frac{dq}{dt}$ which also has n-coordinates

These $2n$ coordinates q_i, \dot{q}_i completely specify the state

The \dot{q}_i are generalized velocities and are in $T_p M$. Therefore, (q, \dot{q}) is in the tangent bundle, TM

The lagrangian, $L(q, \dot{q})$, is a map $L: TM \rightarrow \mathbb{R}$.

For Hamiltonian mechanics, we need the generalized momenta

$$p_i(q_i, \dot{q}_i) \equiv \frac{\partial L}{\partial \dot{q}_i}$$

→ one form, ie in co-tangent space
 ↳ sub script means p is in cotangent space

To build the Hamiltonian, we need the Lagrangian and a transformation

$$(q, \dot{q}) \rightarrow (q, p)$$

This is not simply changing coordinates

To see this suppose we have a change in generalized coordinates

$$q_u \longrightarrow q_v$$

This can be described as

$$q'_u = q_v(q_u) \quad \text{prime denotes new coordinates}$$

$$\dot{q}'_v = \left(\frac{\partial q'_v}{\partial q_u} \right)^j \dot{q}_u^j$$

Compare this with $A_{ij}^{i'} = \frac{\partial y^{i'}}{\partial x^j}$ $v^{i'} = A_{ij}^{i'} v^j$ (contravariant) (vectors do this)

The p 's transform as follows:

$$p_{i'}^v \equiv \frac{\partial L}{\partial \dot{q}_{i'}^v} = \left(\underbrace{\frac{\partial L}{\partial \dot{q}_u^j} \frac{\partial \dot{q}_u^j}{\partial \dot{q}_{i'}^v}}_{\text{covariant}} + \frac{\partial L}{\partial \dot{q}_u^j} \frac{\partial \dot{q}_u^j}{\partial \dot{q}_{i'}^v} \right)$$

$$p_{i'}^v = p_j^u \left(\frac{\partial \dot{q}_u^j}{\partial \dot{q}_{i'}^v} \right) \quad \begin{array}{l} \text{This shows this is covariant and} \\ \text{must live in the cotangent space} \end{array}$$

Compare with $\lambda_{i'}^j = \frac{\partial x^j}{\partial y^{i'}}$

\dot{q} is in the tangent space (vector)

p is in the cotangent space (one-form)

Hence computing p is not only changing variables but is really a map

$$p: TM \rightarrow T^* M \quad \begin{array}{l} \text{right hand side could be } T_p^* M \text{ but} \\ \text{if we add } q, \text{ then } T^* M \end{array}$$

$T^* M$ is the phase space (q, p)

The Hamiltonian is a map H s.t.

$$H: T^* M \rightarrow \mathbb{R} \quad H(q, p)$$

The Lagrangian: $L(q, \dot{q}) = T(q, \dot{q}) - U(q)$

Where the KE is: $T(q, \dot{q}) = \frac{1}{2} g_{jk} \dot{q}^j \dot{q}^k$

Example: Suppose we have 2 masses in 1D

$$M = \mathbb{R}^2 \quad \text{and} \quad TM = \mathbb{R}^4$$

$$T = \frac{1}{2} m_1 (\dot{q}_1)^2 + \frac{1}{2} m_2 (\dot{q}_2)^2$$

need with $g_{ij} = \begin{bmatrix} m_1 & 0 \\ 0 & m_2 \end{bmatrix}$

Example: If we have a mass in 2D

$$T = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2) \quad [\text{cartesian}]$$

with $g_{ij} = \begin{bmatrix} m & 0 \\ 0 & m \end{bmatrix}$

Involved q

$$T = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) \quad g_{ij} = \begin{bmatrix} m & 0 \\ 0 & r^2 \end{bmatrix} \quad [\text{polar coordinates}]$$

In general $p_i \equiv \frac{\partial L}{\partial \dot{q}^i} = \frac{\partial T}{\partial \dot{q}^i} = g_{ij}(q) \dot{q}^i$

∂T can be used to define a Riemannian metric

$$\langle \dot{q}, \dot{q} \rangle = g_{ij}(q) \dot{q}^i \dot{q}^j$$

Kinetic Energy is $\frac{1}{2}$ the length squared of the velocity vector

The generalized momenta p is the covariant version of the generalized velocity.

Example 1: $p_1 = m_1 \dot{q}^1$ and $p_2 = m_2 \dot{q}^2$

In general $p_i = g_{ij} \dot{q}^j$ and $\dot{q}^i = g^{ij} p_j$

§ 5.4 Hamiltonian Vector Fields

Given a Lagrangian, we can obtain the equations of motion from the Euler-Lagrange Equations

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = \frac{\partial L}{\partial q_i}$$

Hamiltonian

$$H(q, p) = p \dot{q} - L$$

Hamilton's eqns

$$\dot{q} \frac{\partial H}{\partial p} + \dot{p} = - \frac{\partial H}{\partial q}$$

Phase Space is the tangent bundle T^*M , which includes M and $T_p M$

On T^*M , which is a manifold, we define a 2-form,

$\tilde{\omega} = \hat{dq} \wedge \hat{dp}$ area in phase space

$\hat{}$ can be called H ?

Take a curve on T^*M of the form

$$\{ q = f(t), p = g(t) \}$$

which is a solution to Hamilton's equations. The tangent vector to the curve is,

$$\bar{U} = \frac{d}{dt} = \underbrace{i}_{\text{basis vectors}} \frac{\partial}{\partial q} + \underbrace{j}_{\text{basis vectors}} \frac{\partial}{\partial p}$$

Theorem: If \bar{u} is a tangent vector to the solution curve then $\mathcal{L}_{\bar{u}} \tilde{\omega} = 0$

Proof: From a formula (4.67)

$$\mathcal{L}_{\bar{u}} \tilde{\omega} = \tilde{\delta}[\tilde{\omega}(\bar{u})] + (\tilde{\delta}\tilde{\omega})(\bar{u})$$

The 2nd term is 0 since $\tilde{\omega}$ is a 2-form on a 2 dim Manifold

$$\begin{aligned} \Rightarrow \mathcal{L}_{\bar{u}} \tilde{\omega} &= \tilde{\delta}[\tilde{\omega}(\bar{u})] \\ &= \tilde{\delta}[\tilde{\delta}_q \wedge \tilde{\delta}_p(\bar{u})] \\ &= \tilde{\delta}[(\tilde{\delta}_q \otimes \tilde{\delta}_p - \tilde{\delta}_p \otimes \tilde{\delta}_q)(\bar{u})] \\ &= \tilde{\delta}[(\tilde{\delta}_q(\bar{u})\tilde{\delta}_p - \tilde{\delta}_p(\bar{u})\tilde{\delta}_q)] \end{aligned}$$

one form
coefficient of 1 form

But $\bar{u} = f \frac{\partial}{\partial q} + g \frac{\partial}{\partial p}$ which yields

$$\tilde{\delta}_q(\bar{u}) = f \quad \text{and} \quad \tilde{\delta}_p(\bar{u}) = g$$

$$\Rightarrow \mathcal{L}_{\bar{u}} \tilde{\omega} = \tilde{\delta}[f \tilde{\delta}_p - g \tilde{\delta}_q]$$

However $f = \frac{\partial H}{\partial p}$ and $g = -\frac{\partial H}{\partial q}$ from hamiltons eqn

$$\begin{aligned} \mathcal{L}_{\bar{u}} \tilde{\omega} &= \tilde{\delta}\left[\frac{\partial H}{\partial q} \tilde{\delta}_q + \frac{\partial H}{\partial p} \tilde{\delta}_p\right] \\ &= \tilde{\delta}[\tilde{\delta}H] = 0 \end{aligned}$$

The area in phase space is conserved along solns to hamiltons equations

A Vector field with $\mathcal{L}_{\bar{u}} \tilde{\omega} = 0$ is a hamiltonian vector field. \bar{u} is tangent to the curves in phase space

The system is conservative (H is constant along solns)

$$\begin{aligned} \mathcal{L}_{\bar{u}} H &= \frac{dH}{dt} = \dot{q} \frac{\partial H}{\partial q} + \dot{p} \frac{\partial H}{\partial p} \\ &= \frac{\partial H}{\partial p} \frac{\partial H}{\partial q} - \frac{\partial H}{\partial q} \frac{\partial H}{\partial p} \\ &= 0 \end{aligned}$$

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§ 5.5 canonical transformation

p and q are not unique. P and Q are canonical if $\tilde{d}q_1 \tilde{d}p = \tilde{d}P_1 \tilde{d}Q$

This requires

$$\frac{\partial Q}{\partial q_1} \frac{\partial p}{\partial P} - \frac{\partial Q}{\partial P} \frac{\partial p}{\partial q_1} = 1$$

Example $Q = p$ and $P = -q$

$$\text{check: } \frac{\partial p}{\partial q_1} \frac{\partial (-q)}{\partial P} - \frac{\partial P}{\partial q_1} \frac{\partial (-q)}{\partial p} = 1$$

§ 5.6 Map between vectors and 1-forms by $\tilde{\omega}$

$\tilde{\omega} = \tilde{d}q \wedge \tilde{d}p$ can be used like the metric tensor to convert vectors to forms and vice versa. Suppose \bar{V} is a

vector field on M . then $\bar{V} = \tilde{\omega}(\bar{V}) = \tilde{d}q_1 \tilde{d}p(\bar{V})$

$$\begin{aligned} &= (\tilde{d}q_1 \otimes \tilde{d}p - \tilde{d}p \otimes \tilde{d}q_1)(\bar{V}) \\ &= \tilde{d}q_1(\bar{V}) \tilde{d}p - \tilde{d}p(\bar{V}) \tilde{d}q_1 \end{aligned}$$

If $\bar{V} = V^1 \tilde{d}q + V^2 \tilde{d}p$ then

$$\begin{aligned} \bar{V} &= V^1 \tilde{d}p - V^2 \tilde{d}q \quad \begin{matrix} \text{when these become 1-form} \\ \text{components doesn't this lower the indices?} \end{matrix} \\ &= -V^2 \tilde{d}q + V^1 \tilde{d}p \end{aligned}$$

We can write $(\bar{V})_i = \omega_{ij} V^j$ and deduce that

$$\omega_{ij} = \begin{bmatrix} 0 & -1 \\ 1 & 0 \end{bmatrix} \quad \text{and} \quad \omega^{ij} = \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix}$$

Using ω^{ij} we can find \bar{V} given \bar{V}

§ 5.7 Poisson Bracket

Say f, g are functions on M and define

$$\bar{X}_f = \overline{df} \quad \text{and} \quad \bar{X}_g = \overline{dg}$$

These are the vector versions of the gradient. From above,

$$\tilde{d}f = \frac{\partial f}{\partial q_1} \tilde{d}q_1 + \frac{\partial f}{\partial p} \tilde{d}p$$

then

$$\bar{X}_f = \bar{\delta f} = \frac{\partial f}{\partial q} \frac{\partial}{\partial q} - \frac{\partial f}{\partial p} \frac{\partial}{\partial p}$$

$$\bar{X}_g = \bar{\delta g} = \frac{\partial g}{\partial p} \frac{\partial}{\partial q} - \frac{\partial g}{\partial q} \frac{\partial}{\partial p}$$

We can then define the Poisson bracket

$$\begin{aligned}\{f, g\} &\equiv \tilde{\omega}(\bar{X}_f, \bar{X}_g) \\ &= \omega_{ij} X_f^i X_g^j \\ &= (X_f)_j X_g^j \\ &= \tilde{\delta} f(\bar{X}_g) = \langle \tilde{\delta} f, \bar{X}_g \rangle\end{aligned}$$

To evaluate this we get,

$$\{f, g\} = \left(\frac{\partial f}{\partial q} \tilde{\delta} q + \frac{\partial f}{\partial p} \tilde{\delta} p \right) \left(\frac{\partial g}{\partial p} \frac{\partial}{\partial q} - \frac{\partial g}{\partial q} \frac{\partial}{\partial p} \right)$$

$$\boxed{\{f, g\} = \frac{\partial f}{\partial q} \frac{\partial g}{\partial p} - \frac{\partial f}{\partial p} \frac{\partial g}{\partial q}} \quad \text{Poisson bracket}$$

Aside $\tilde{\delta} q_i \left(\frac{\partial}{\partial q_j} \right) = 1$

$\tilde{\delta} q_i \left(\frac{\partial}{\partial p_j} \right) = 0$

The above expression is in terms of coordinates. The expression independent of coordinates is

$$\{f, g\} = \tilde{\omega}(\bar{\delta} f, \bar{\delta} g)$$

S 5.8 Many particle systems: symplectic forms

In 3D with no constraints, and N particles we have $6N$ dim'l phase space.

The phase space in general can be said to be $2N$ where N is the number of generalized coordinates
then

$$\boxed{\tilde{\omega} = \sum_{A=1}^n \tilde{\delta} q_A \wedge \tilde{\delta} p_A}$$

Symplectic
Form

The Phase space is a symplectic Manifold.

§ 5.9 Linear Dynamical Systems: the symplectic inner product and conserved quantities

To begin consider the following hamiltonian

$$H = \frac{1}{2} \sum_{A,B=1}^n T^{AB} p_A p_B + V_{AB} q_A^A q_B^B$$

where we assume T^{AB} and V_{AB} are symmetric. If not, we use the fact the product of the asymmetric part and a symmetric function is 0.

For simplicity, assume T^{AB} and V_{AB} are constant.

Hamilton's equations

$$\frac{dp_A}{dt} = -\frac{\partial H}{\partial q_A} = -\sum_B V_{AB} q_B^B$$

$$\frac{dq_A}{dt} = \frac{\partial H}{\partial p_A} = \sum_B T^{AB} p_B$$

$$\begin{aligned} \text{Check: } \frac{\partial H}{\partial q_C} &= \frac{\partial}{\partial q_C} \left(\frac{1}{2} \sum_{A,B=1}^n V_{AB} q_A^A q_B^B \right) \\ &= \frac{1}{2} \sum_{A,B} V_{AB} \delta_C^A q_B^B + \frac{1}{2} V \sum_{A,B} V_{AB} q_A^A \delta_C^B \\ &= \frac{1}{2} \sum_B V_{CB} q_B^B + \frac{1}{2} \sum_A V_{AC} q_A^A \\ &= \sum_B V_{CB} q_B^B \quad \text{VCA, A} \rightarrow \text{B} \end{aligned}$$

If $\bar{Y}_{(1)}$ is a vector with components $\{q_{(1)}^A, p_{(1)A}, A=1\dots n\}$ and \bar{Y}_2 is a vector with components $\{q_{(2)}^A, p_{(2)A}, A=1\dots n\}$ then their symplectic product is

$$\bar{\omega}(\bar{Y}_{(1)}, \bar{Y}_{(2)}) = \sum_A q_{(1)}^A p_{(2)A} - q_{(2)}^A p_{(1)A}$$

If $\bar{Y}_{(1)}$ and $\bar{Y}_{(2)}$ are both solutions, then the symplectic inner product is independent of time

$$\begin{aligned} \frac{d}{dt} \bar{\omega}(\bar{Y}_{(1)}, \bar{Y}_{(2)}) &= \frac{d}{dt} \left[\sum_{A=1}^n q_{(1)}^A p_{(2)A} - q_{(2)}^A p_{(1)A} \right] \\ &= \sum_{A=1}^n \left\{ \frac{dq_{(1)}^A}{dt} p_{(2)A} + q_{(1)}^A \frac{dp_{(2)A}}{dt} - \frac{dq_{(2)}^A}{dt} p_{(1)A} - q_{(2)}^A \frac{dp_{(1)A}}{dt} \right\} \end{aligned}$$

Using Hamilton's eqns next,

$$\frac{d}{dt} \tilde{\omega}(\bar{Y}_{(1)}, \bar{Y}_{(2)}) = \sum_{A,B} \left\{ T^{AB} P_{(1)B} P_{(2)A} - V_{AB} q_{(1)}^A q_{(2)}^B - T^{AB} P_{(2)B} P_{(1)A} + V_{AB} q_{(2)}^B q_{(1)}^A \right\} = 0$$

If T^{AB} and V_{AB} are independent of time then it follows that if $\bar{Y}_{(1)}$ is a soln then so is $\frac{d\bar{Y}_{(1)}}{dt}$

This motivates defining the canonical energy as $E_c(\bar{Y}) = \tilde{\omega}\left(\frac{d\bar{Y}}{dt}, \bar{Y}\right)$

It can be determined that $E_c(\bar{Y}) = H$ evaluated at \bar{Y}

$$E_c(\bar{Y}) = \frac{1}{2} \tilde{\omega}(\dot{\bar{Y}}, \bar{Y}) = \sum_A \left(\dot{q}_{(1)}^A P_{(2)A} - q_{(2)}^A \dot{P}_{(1)A} \right)$$
$$= \frac{1}{2} \sum_{A,B} T^{AB} P_{(1)B} P_{(2)A} + V_{AB} q_{(1)}^B q_{(2)}^A = H$$

So the independence of H w.r.t. yields the conservation of H or the total Mechanical Energy

Other conserved quantities

In general, T^{AB} and V_{AB} can depend on the coordinates \vec{x} .

If $\exists \vec{u}$ such that $\mathcal{L}_{\vec{u}} T^{AB} = 0 = \mathcal{L}_{\vec{u}} V_{AB}$ then there are conserved quantities associated to \vec{u}

This can yield expressions for Linear Momentum or angular momentum conservation

Noether's theorem?

Exam Content

Lec 23 - Nov 20th

§ 5.11 Rewriting Maxwell's equations in differential forms
 § C electromagnetism

We can non-dimensionalize Maxwell's equations in such a way that $C = \mu_0 = \epsilon_0 = 1$ to get

A $\vec{\nabla} \times \vec{B} - \frac{\partial \vec{E}}{\partial t} = 4\pi \vec{J}$ Ampere's Law

B $\vec{\nabla} \times \vec{E} + \frac{\partial \vec{B}}{\partial t} = 0$ Faraday's Law

C $\vec{\nabla} \cdot \vec{B} = 0$ Gaus

D $\vec{\nabla} \cdot \vec{E} = 4\pi \rho$ charge

We will rewrite these using a metric and $\tilde{\delta}$. The relativistic invariant form requires the Faraday 2-form

$$F_{\mu\nu} = \begin{pmatrix} t & x & y & z \\ 0 & -Ex & -Ey & -Ez \\ Ex & 0 & Bz & -By \\ E_y & -Bz & 0 & Bx \\ Ez & By & -Bx & 0 \end{pmatrix}_{\text{rows } \mu, \text{columns } \nu} \quad \mu, \nu = t, x, y, z$$

→ How is this a 2-form?

what does this look like?

Then $\tilde{\delta} F$ is a 3-form on a 4D manifold. Since $\tilde{\delta} F$ is a 3-form on a 4D Manifold, there are $C_3^4 = 4$ different equations,

We can write $F = F_{\mu\nu} \tilde{\delta}^\mu \tilde{\delta}^\nu$, we compute,

$$\tilde{\delta}^P = F_{\mu\nu\rho} \tilde{\delta}^\mu \wedge \tilde{\delta}^\nu \wedge \tilde{\delta}^\rho$$

It is observed that $\tilde{\delta}^P = 0$ iff $F_{[\mu\nu,\rho]} = 0$

① $F_{[xy,z]} = F_{xy,z} + F_{yz,x} + F_{zx,y} = 0$

$$B_{z,z} + B_{x,x} - B_{y,y} = 0 \quad \text{Div of Magnetic field} = 0$$

or $\vec{\nabla} \cdot \vec{B} = 0$ Eqn C

② $F_{[xy,t]} = F_{xy,t} + F_{yt,x} + F_{tx,y} = 0$

$$B_{z,t} + E_{y,x} - E_{x,y} = 0 \quad \text{the } z \text{ equation of eqn B}$$

③ $F_{[yz,t]} = F_{yz,t} + F_{zt,y} + F_{ty,z} = 0$

$$= B_{x,t} + E_{z,y} - E_{y,z} = 0 \quad \text{the } x \text{ eqn in B}$$

$$\textcircled{4} \quad F_{[xz, t]} = F_{x2, t} + F_{zt, x} + F_{+x, z} = 0$$

$$-B_{y,t} + E_{z,x} - E_{x,z} = 0 \quad \text{the } y \text{ eqn of } B$$

For the other equations we need the special relativistic metric

$$(g_{\mu\nu}) = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \rightsquigarrow \text{Lorentzian metric?}$$

This allows us to find the 2-vector $F^{\mu\nu} = g^{\mu\alpha} g^{\nu\beta} F_{\alpha\beta} = g^{\mu\alpha} F_{\alpha\beta} g^{\nu\beta}$

Note: $g^{\mu\nu} = \text{diag}(-1, 1, 1, 1)$

$$F^{\mu\nu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & B_z & -B_y \\ E_y & -B_z & 0 & B_x \\ E_z & B_y & -B_x & 0 \end{pmatrix} \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$F^{\mu\nu} = \begin{pmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ -E_x & 0 & B_z & -B_y \\ -E_y & -B_z & 0 & B_x \\ -E_z & B_y & -B_x & 0 \end{pmatrix}$$

$$F^{\mu\nu} = \begin{pmatrix} 0 & E_x & E_y & E_z \\ -E_x & 0 & B_z & -B_y \\ -E_y & -B_z & 0 & B_x \\ -E_z & B_y & -B_x & 0 \end{pmatrix}$$

The final for eqn's are. $F^{\mu\nu}_{,\nu} = 4\pi J^\mu$ where $J^t = \rho$, $J^i = (J)^i$ $i = x, y, z$

We check 4 different equations

$$\textcircled{1} \quad F^{+z}_{,\nu} = F^{+x}_{,x} + F^{+y}_{,y} + F^{+z}_{,z} = 4\pi J^t$$

$$= E_{x,x} + E_{y,y} + E_{z,z} = 4\pi \rho \quad \text{eqn D}$$

$$\textcircled{2} \quad F^{x\nu}_{,\nu} = F^{x t}_{,t} + F^{x y}_{,y} + F^{x z}_{,z} = 4\pi J^x$$

$$= -E_{x,t} + B_{z,y} - B_{y,z} = 4\pi J^x \quad \text{This is the } x \text{ eqn in A}$$

$$\textcircled{3} \quad F^{y\tau}_{,z\tau} = F^{yt}_{,t\tau} + F^{yx}_{,x\tau} + F^{yz}_{,z\tau} = 4\pi J^y$$

$$= -E_{yt} - B_{z,x} \quad B_{x,z} = 4\pi J^y \quad \text{The } y \text{ eqn in A}$$

$$\textcircled{4} \quad F^{z\tau}_{,z\tau} = F^{zt}_{,t\tau} + F^{zx}_{,x\tau} + F^{zy}_{,y\tau} = 4\pi J^z$$

$$= -E_{zt} + B_{y,x} - B_{x,y} = 4\pi J^z \quad \text{The } z \text{ eqn in A.}$$

Observe that \tilde{F} is coordinate independent, However $F^{u\tau}_{,\tau} = 4\pi J^u$ is coordinate dependent.

Given a basis of the tangent space $(\frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z})$ then, we can define the volume 4-form as

$$\tilde{\omega} = \hat{d}t \wedge \hat{d}x \wedge \hat{d}y \wedge \hat{d}z$$

We define $*\tilde{F} = \frac{1}{2}\tilde{\omega}(F)$ or $(*\tilde{F})_{\mu\nu} = \frac{1}{2}\omega_{\alpha\beta\mu\nu} F^{\alpha\beta}$ dual or Hodge Star

Next we determine the components of this,

$$\textcircled{1} \quad (*F)_{tx} = \frac{1}{2}\omega_{\alpha\beta tx} F^{\alpha\beta}$$

$$= \frac{1}{2}\omega_{yztx} F^{yz} + \underbrace{\frac{1}{2}\omega_{zytx} F^{zy}}_{\text{antisymmetric}}$$

$$yztx \rightarrow -y+zx \rightarrow y+xz \rightarrow -tyxz \rightarrow +xyz \quad \textcircled{+1} \quad ???$$

$$(*F)_{tx} = B_x$$

$$\textcircled{2} \quad (*\tilde{F})_{ty} = \frac{1}{2}\omega_{\alpha\beta ty} F^{\alpha\beta} = \frac{1}{2}(\omega_{zx+ty} F^{zx} + \omega_{x+ty} F^{xz})$$

$$zx+ty \rightarrow -xz+ty \rightarrow x+zy \rightarrow -ty+zy \rightarrow +xyz \quad \textcircled{+1} \text{ coefficient.}$$

$$(*F)_{ty} = B_y$$

$$\textcircled{3} \quad (*\tilde{F})_{+z} = B_z$$

$$\textcircled{4} \quad (*\tilde{F})_{xy} = E_z$$

$$\textcircled{5} \quad (*\tilde{F})_{xz} = -E_y$$

$$\textcircled{6} \quad (*\tilde{F})_{yz} = E_x$$

$$(* \tilde{F})_{\mu\nu} = \begin{pmatrix} 0 & B_x & B_y & B_z \\ -B_x & 0 & E_z & F_y \\ -B_y & -E_z & 0 & E_x \\ -B_z & F_y & E_x & 0 \end{pmatrix}$$

The exterior derivative of this is

$$\tilde{\delta}(*\tilde{F}) = (*\tilde{F})_{\mu\nu\rho} \tilde{\delta}\gamma^\nu \tilde{\delta}u^\rho \tilde{\delta}v^\mu$$

we define $*\tilde{J} = \tilde{\omega}(\tilde{S})$ and apply $\tilde{\omega}$ to the eqn $F^{\mu\nu}{}_\nu = 4\pi J^\mu$

$$\tilde{\omega}(F^{\mu\nu}, \nu) = \tilde{\omega}(4\pi J^\mu)$$

we can write this as

$$\boxed{\tilde{\delta}(*\tilde{F}) = 4\pi * \tilde{J}}$$

and

$$\tilde{\delta} F = 0$$

§ 5.13 Vector Potential

If $\tilde{\delta} F = 0$ then F is closed and since it is a 2-form, \exists a 1-form \tilde{A} such that

$$F = \tilde{\delta} \tilde{A} \quad \text{at least locally}$$

\tilde{A} is the vector potential

Lec 24 Dec 5th

D Dynamics of a perfect fluid

S 5.15 Role of Lie derivatives

A perfect fluid (idealized) is one that conserves certain properties

- (1) Mass
- (2) Entropy
- (3) Vorticity. [will explain] $\nabla \times \text{velocity}$

Today we will express the equations of a fluid using exterior calculus.

S 5.16 the Comoving time derivative

The conservation of mass (continuity eqn) is:

$$\frac{\partial p}{\partial t} + \vec{\nabla} \cdot (p \vec{v}) = 0 \quad \text{things converging, density increases}$$

On an assignment we found

$$(\frac{\partial}{\partial t} + \vec{L}_v) (p \tilde{\omega}) \underset{\text{mass}}{\sim} 0 \quad \left. \begin{array}{l} \text{This is being} \\ \text{written in terms} \\ \text{of differential forms} \end{array} \right\}$$

$$\text{where } \tilde{\omega} = \hat{\delta}x \wedge \hat{\delta}y \wedge \hat{\delta}z$$

The operator $(\frac{\partial}{\partial t} + \vec{L}_v)$ computes the total rate of change following the flow.

Consider the motion of a fluid parcel. If the change happens over a short time,

$$dt \ll 1 \rightarrow \text{makes the following approximation valid}$$

Then the motion is from

$$(x, y, z, t) \rightarrow (x + v^x dt, y + v^y dt, z + v^z dt, t + dt)$$

The difference between the two is

$$(v^x, v^y, v^z, 1) dt = \vec{U} \quad \text{is a 4-vector}$$

In the formulation, the total rate of change following the flow:

$$\mathcal{L}_{\bar{U}} \bar{W} = \left(\frac{\partial}{\partial t} + \mathcal{L}_{\bar{V}} \right) \bar{W} \quad \begin{matrix} \text{Space time} \\ \text{Version.} \end{matrix}$$

Where \bar{W} is a 4-vector and it would need to be decomposed on the RHS

§ 5.17 Eqns of Motion

A perfect fluid conserves entropy. If S is the entropy, then the eqn is

$$1083 \quad \boxed{\left(\frac{\partial}{\partial t} + \mathcal{L}_{\bar{V}} \right) S = 0} \quad \text{Thermodynamics}$$

The conservation of linear momentum (Newton's 2nd law) can be written as:

$$\frac{\partial}{\partial t} V^i + V^j \frac{\partial}{\partial x^j} V^i + \underbrace{\frac{1}{P} \frac{\partial}{\partial x^i} P}_{\substack{\text{total rate of change} \\ F = ma?}} + \frac{\partial}{\partial x^i} \Phi = 0$$

↓ pressure ↓ gravity

P is pressure

Φ is the gravitational potential

V^i is the Velocity

This eqn is a mess as we have both superscripts and subscripts added to each other. Bad!

$\frac{\partial V^i}{\partial x^j}$ is not a $(1,1)$ tensor. \rightarrow Partial derivatives don't transform as vectors

\rightarrow How does a metric relate to a distance fn.

Assume we have a metric

This allows us to convert the vector \bar{V} to the one-form \tilde{V}

$$\tilde{V} = g_{ij} (\bar{V}, *) \quad \text{yields } V_i$$

↳ star means empty in this case

To rewrite the non-linear term, we need the operator

$$\boxed{\left(\frac{\partial}{\partial t} + \mathcal{L}_{\bar{V}} \right) \tilde{V}}$$

To find out what this term looks like, consider
eqn 3.14 in textbook

$$(\mathcal{L}_{\bar{V}} \tilde{V})_i = V^j \frac{\partial}{\partial x^j} V_i + V_j \frac{\partial}{\partial x^i} V^j$$

Can show using $\bar{V} \cdot V = 0$

$$= V^j \frac{\partial}{\partial x^j} V_i + \frac{1}{2} \frac{\partial}{\partial x^i} (V_j V^j)$$

$$\Rightarrow V^j \frac{\partial}{\partial x^j} V_i = (\mathcal{L}_{\bar{V}} \tilde{V})_i - \frac{\partial}{\partial x^i} \left(\frac{1}{2} V^2 \right) \quad \text{if } V^2 = \tilde{V}(\bar{V})$$

Our momentum equation in coordinate independent form becomes

2 of 3

$$\left(\frac{\partial}{\partial t} + \mathcal{L}_{\bar{V}} \right) \tilde{V} + \frac{1}{\rho} \tilde{\delta} p + \tilde{\delta} \left(\tilde{\phi} - \frac{1}{2} V^2 \right) = 0$$

Almost like bernoullis eqn. ??

§ 5.18 Conservation of Vorticity

The vorticity of a fluid with velocity \bar{V} is

$$\vec{\nabla} \times \vec{V} \quad \text{curl of velocity}$$

we have seen that this can be written as

$$\star \tilde{\delta} \tilde{V} \quad \text{curl} \rightarrow 1 \text{-vector}$$

$$\tilde{\delta} \tilde{V} \quad \text{curl} \rightarrow 2\text{-form}$$

To get the vorticity equation we apply $\tilde{\delta}$ to the equation

$$\left(\frac{\partial}{\partial t} + \mathcal{L}_{\bar{V}} \right) \tilde{\delta} \tilde{V} = \frac{1}{\rho} \tilde{\delta} p \wedge \tilde{\delta} p \quad *$$

Case (1) $p = d(p)$ then $\tilde{\delta} p \wedge \tilde{\delta} p = 0$

$$\Rightarrow \left(\frac{\partial}{\partial t} + \mathcal{L}_{\bar{V}} \right) \tilde{\delta} \tilde{V} = 0$$

Vorticity is conserved following the flow

Case 2: $p = p(\rho, s)$ then $\tilde{d}p \wedge \tilde{\delta}p \neq 0$ but

$$\tilde{d}s \wedge \tilde{d}\rho \wedge \tilde{\delta}p = 0 \quad \text{Manifold is 2D } (s, \rho)$$

Apply \tilde{d} to our equation 1 of 3

$$\left(\frac{\partial}{\partial t} + L_{\tilde{V}} \right) \tilde{d}s = 0$$

If we take $\tilde{d}s \wedge$ the vorticity eqn * then

$$\tilde{d}s \wedge \left(\frac{\partial}{\partial t} + L_{\tilde{V}} \right) \tilde{d}\tilde{V} = \frac{1}{\rho^2} \tilde{d}s \wedge \tilde{\delta} \wedge \tilde{\delta}p \xrightarrow{0}$$

or

$$\left(\frac{\partial}{\partial t} + L_{\tilde{V}} \right) \tilde{d}s \wedge \tilde{d}\tilde{V} = 0 \quad \text{Ertels theorem}$$

$\tilde{d}s \wedge \tilde{d}\tilde{V}$ is a 3-form. By mass conservation $\rho \tilde{\omega}$ is conserved these are both 3-forms and on a 3D manifold must be linearly related

$$\tilde{d}s \wedge \tilde{d}\tilde{V} = \alpha \rho \tilde{\omega} \quad \alpha \text{ is a function and must exist!}$$

Since $\tilde{d}s \wedge \tilde{d}\tilde{V}$ and $\rho \tilde{\omega}$ is conserved it follows that α is conserved.

$$\left(\frac{\partial}{\partial t} + L_{\tilde{V}} \right) \alpha = 0$$

claim: $\alpha = \frac{1}{\rho} \vec{\nabla} s \cdot \vec{\nabla} \times \vec{V}$

proof: take the dual of

$$\tilde{d}s \wedge \tilde{d}\tilde{V} = \alpha \rho \tilde{\omega}$$

$$*(\tilde{d}s \wedge \tilde{d}\tilde{V}) = *(\alpha \rho \tilde{\omega})$$

$$\text{Note: } \hat{\delta} \tilde{s}_1 \hat{\delta} V = S_{,i} \hat{\delta} x^i \wedge \epsilon^{ijk} v_{k,j} \hat{\delta} x^j \wedge \hat{\delta} x^k$$

$$= \epsilon^{ijk} S_{,i} v_{k,j} \hat{\delta} x^i \wedge \hat{\delta} x^j \wedge \hat{\delta} x^k$$

Comparing the coefficient of $\tilde{\omega}$ we get.

$$\alpha = \frac{1}{\rho} \vec{\nabla} \tilde{s} \cdot \vec{\nabla} \times \vec{V} \quad \text{Ertel Potential Vorticity.}$$