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Capítulo 3

1. Descubrimiento rayos X – Electrón

$$\tan \theta = \frac{v_y}{v_x} = \frac{a_y t}{v_0} = \frac{qE}{m} \frac{\ell}{v_0^2}$$

$$\vec{F} = q\vec{E} + q\vec{v} \times \vec{B} = 0$$

$$v_x = \frac{E}{B} = v_0$$

$$\frac{q}{m} = \frac{v_0^2 \tan \theta}{E\ell} = \frac{E \tan \theta}{R^2 \ell}$$

- 2. Carga del electrón
- 3. Línea espectral

$$\lambda = \lambda_{
m limit} rac{n^2}{n^2 - n_0^2} \qquad (n = n_0 + 1, n_0 + 2, n_0 + 3 \dots)$$

$$rac{1}{\lambda} = R_{
m H} \left(rac{1}{n^2} - rac{1}{k^2}
ight) \qquad egin{array}{l} d \sin \theta = n \lambda \\ Figure 1 - M\'axima \\ difracción \end{array}$$

$$(n^2 \quad k^2)$$

Figure 2- Rydberg

- 4. Cuantización
- 5. Radiación de cuerpo negro

$$\lambda_{\max}T = 2.898 \times 10^{-3} \text{ m} \cdot \text{K}$$
 $R(T) = \epsilon \sigma T^4$ $\&(\lambda, T) = \frac{2\pi ckT}{\lambda^4}$ $\&(\lambda, T) = \frac{2\pi c^2 h}{\lambda^5} \frac{1}{e^{\hbar c/\lambda kT} - 1}$ Figure 6 - Wien desplazamiento Figure 5 - Stefan-Boltzmann Figure 4 - Rayleigh-Jeans Figure 3-Radiación de Planck

$$h = 6.6261 \times 10^{-34} \,\text{J} \cdot \text{s}$$
 $E_n = nhf$, $\Delta E = hf$

Efecto fotoeléctrico

Einstein

$$E = hf$$
 $\lambda f = c$ $hf = \phi + \text{K.E. (electron)}$ $hf = \phi + \frac{1}{2} m v_{\text{max}}^2$ $eV_0 = \frac{1}{2} m v_{\text{max}}^2$

Cuántico

$$\frac{1}{2} m v_{\text{max}}^2 = e V_0 = h f - \phi$$
 $e V_0 = \frac{1}{2} m v_{\text{max}}^2 = h f - h f_0 = h (f - f_0)$ $E = h f = \frac{h c}{\lambda}$

7. Producción de X-Ray

$$E_f = E_i - hf$$
 $eV_0 = hf_{ ext{max}} = rac{hc}{\lambda_{ ext{min}}}$ $\lambda_{ ext{min}} = rac{hc}{e} rac{1}{V_0} = rac{1.240 imes 10^{-6} \, ext{V} \cdot ext{m}}{V_0}$

Figure 7-Duane-Hunt

8. Efecto Compton

$$p = \frac{E}{c} = \frac{hf}{c} = \frac{h}{\lambda}$$

$$E_e^2 = (mc^2)^2 + p_e^2 c^2$$

$$Energy \quad hf + mc^2 = hf' + E_e$$

$$p_x \qquad \frac{h}{\lambda} = \frac{h}{\lambda'} \cos \theta + p_e \cos \phi \qquad p_e^2 = \left(\frac{h}{\lambda}\right)^2 + \left(\frac{h}{\lambda'}\right)^2 - 2\left(\frac{h}{\lambda}\right)\left(\frac{h}{\lambda'}\right) \cos \theta$$

$$p_y \qquad \frac{h}{\lambda'} \sin \theta = p_e \sin \phi \qquad \Delta \lambda = \lambda' - \lambda = \frac{h}{mc}(1 - \cos \theta)$$

9. Producción de pares y aniquilación

$$\gamma \rightarrow e^+ + e^ hf = E_+ + E_- + \text{K.E. (nucleus)}$$

Capitulo 4

$$hf > 2m_e c^2 = 1.022 \text{ MeV}$$
 (for pair production)

Energy $2m_e c^2 \approx hf_1 + hf_2$

Momentum $0 = \frac{hf_1}{c} - \frac{hf_2}{c}$ $hf = m_e c^2 = 0.511 \text{ MeV}$

- 1. Estructura del átomo
- 2. Rutherford

$$n = \frac{\rho \left(\frac{\mathrm{g}}{\mathrm{cm}^3}\right) N_{\mathrm{A}} \left(\frac{\mathrm{molecules}}{\mathrm{mol}}\right) N_{\mathrm{M}} \left(\frac{\mathrm{atoms}}{\mathrm{molecule}}\right)}{M_{\mathrm{g}} \left(\frac{\mathrm{g}}{\mathrm{mol}}\right)} = \frac{\rho N_{\mathrm{A}} N_{\mathrm{M}}}{M_{\mathrm{g}}} \operatorname{atoms} \frac{\mathrm{molecule}}{\mathrm{cm}^3}$$

$$n = \frac{\rho\left(\frac{g}{cm^3}\right)N_A\left(\frac{molecules}{mol}\right)N_M\left(\frac{atoms}{molecule}\right)}{M_g\left(\frac{g}{mol}\right)} = \frac{\rho N_A N_M}{M_g} \frac{atoms}{cm^3} \qquad nt = \frac{\rho N_A N_M t}{M_g} \frac{atoms}{cm^2} \qquad N_s = ntA = \frac{\rho N_A N_M t A}{M_g} atoms$$

$$f = \frac{\text{target area exposed by scatterers}}{\text{total target area}} = \frac{ntA\sigma}{A}$$
$$= nt\sigma = nt\pi b^{2}$$
$$f = \pi nt \left(\frac{Z_{1}Z_{2}e^{2}}{8\pi\epsilon \kappa K}\right)^{2} \cot^{2}\frac{\theta}{2}$$

$$N(\theta) = \frac{N_i nt}{16} \left(\frac{e^2}{4\pi\epsilon_0}\right)^2 \frac{Z_1^2 Z_2^2}{r^2 K^2 \sin^4(\theta/2)}$$

3. Modelo atómico clásico

$$\vec{F}_e = \frac{-1}{4\pi\epsilon_0} \frac{e^2}{r^2} \hat{e}_r \qquad a_r = \frac{v^2}{r} \qquad \frac{1}{4\pi\epsilon_0} \frac{e^2}{r^2} = \frac{mv^2}{r} \qquad v = \frac{e}{\sqrt{4\pi\epsilon_0 mr}}$$

4. Modelo de Bohr del átomo de hidrógeno

$$L=\mathit{mvr}=\mathit{n\hbar}$$
 $v=rac{\mathit{n\hbar}}{\mathit{mr}}$ $v^2=rac{e^2}{4\pi\epsilon_0\mathit{mr}}=rac{\mathit{n}^2\hbar^2}{\mathit{m}^2\mathit{r}^2}$

$$r_n = \frac{4\pi\epsilon_0 n^2 \hbar^2}{me^2} \equiv n^2 a_0$$

$$a_0 = \frac{4\pi\epsilon_0 \hbar^2}{me^2}$$

$$= \frac{(1.055 \times 10^{-34} \,\mathrm{J} \cdot \mathrm{s})^2}{\left(8.99 \times 10^9 \,\frac{\mathrm{N} \cdot \mathrm{m}^2}{\mathrm{C}^2}\right) (9.11 \times 10^{-31} \,\mathrm{kg}) (1.6 \times 10^{-19} \,\mathrm{C})^2}$$

$$= 0.53 \times 10^{-10} \,\mathrm{m}$$

$$E_n = -\frac{e^2}{8\pi\epsilon_0 r_n} = -\frac{e^2}{8\pi\epsilon_0 a_0 n^2} \equiv -\frac{E_0}{n^2}$$

$$E_0 = \frac{e^2}{8\pi\epsilon_0 a_0} = \frac{e^2}{(8\pi\epsilon_0)} \frac{me^2}{4\pi\epsilon_0 \hbar^2} = \frac{me^4}{2\hbar^2 (4\pi\epsilon_0)^2} = 13.6 \text{ eV}$$

$$hf = E_u - E_\ell$$

$$\begin{split} \frac{1}{\lambda} &= \frac{f}{c} = \frac{E_u - E_\ell}{hc} \\ &= \frac{-E_0}{hc} \bigg(\frac{1}{n_u^2} - \frac{1}{n_\ell^2} \bigg) = \frac{E_0}{hc} \bigg(\frac{1}{n_\ell^2} - \frac{1}{n_u^2} \bigg) \ \frac{E_0}{hc} = \frac{me^4}{4\pi c \hbar^3 (4\pi \epsilon_0)^2} \equiv R_\infty \\ & v_n = \frac{n\hbar}{mr_n} = \frac{n\hbar}{mn^2 a_0} = \frac{1}{n} \frac{\hbar}{ma_0} \\ & \frac{1}{\lambda} = R_\infty \bigg(\frac{1}{n_\ell^2} - \frac{1}{n_u^2} \bigg) \\ & v_n = \frac{1}{n} \frac{e^2}{4\pi \epsilon_0 \hbar} \qquad \alpha \equiv \frac{v_1}{c} = \frac{\hbar}{ma_0 c} = \frac{e^2}{4\pi \epsilon_0 \hbar c} \approx \frac{1}{137} \end{split}$$

5. El principio de correspondencia

$$f_{\text{classical}} = f_{\text{orb}} = \frac{\omega}{2\pi} = \frac{1}{2\pi} \frac{v}{r} \quad f_{\text{classical}} = \frac{1}{2\pi} \left(\frac{e^2}{4\pi\epsilon_0 m r^3}\right)^{1/2} \quad f_{\text{classical}} = \frac{me^4}{4\epsilon_0^2 h^3} \frac{1}{n^3}$$

$$\text{Orbital radius} \quad r_n = \frac{4\pi\epsilon_0 \hbar^2}{me^2} n^2 = n^2 a_0$$

$$\text{Velocity} \quad v_n = \frac{n\hbar}{mr_n}$$

$$f_{\text{Bohr}} = \frac{me^4}{4\epsilon_0^2 h^3} \frac{1}{n^3} = f_{\text{classical}} \quad E_{\text{nergy}} \quad E_n = -\frac{e^2}{8\pi\epsilon_0 a_0 n^2}$$

6. Errores del modelo de Bohr - Reducción de masa

$$\mu_e = rac{m_e M}{m_e + M} = rac{m_e}{1 + rac{m_e}{M}} \qquad R = rac{\mu_e}{m_e} R_{\infty} = rac{1}{1 + rac{m_e}{M}} R_{\infty} = rac{\mu_e e^4}{4\pi c \hbar^3 (4\pi \epsilon_0)^2}$$
 $rac{1}{\lambda} = Z^2 R \left(rac{1}{n_\ell^2} - rac{1}{n_u^2}
ight)$

7. Espectro Rayos-X y número atómico

$$E (\mathbf{x} \ \mathbf{ray}) = E_u - E_{\ell} \qquad f_{K_{\alpha}} = \frac{3cR}{4} (Z - 1)^2$$

$$\frac{1}{\lambda_{K_{\alpha}}} = R(Z - 1)^2 \left(\frac{1}{1^2} - \frac{1}{2^2}\right) = \frac{3}{4}R(Z - 1)^2 \qquad f_{K_{\alpha}} = \frac{c}{\lambda_{K_{\alpha}}} = \frac{3cR}{4}(Z - 1)^2$$

$$\frac{1}{\lambda_{K}} = R(Z - 1)^2 \left(\frac{1}{1^2} - \frac{1}{n^2}\right) = R(Z - 1)^2 \left(1 - \frac{1}{n^2}\right)$$