



基于多模态算法的 发射线星系测光红移估计

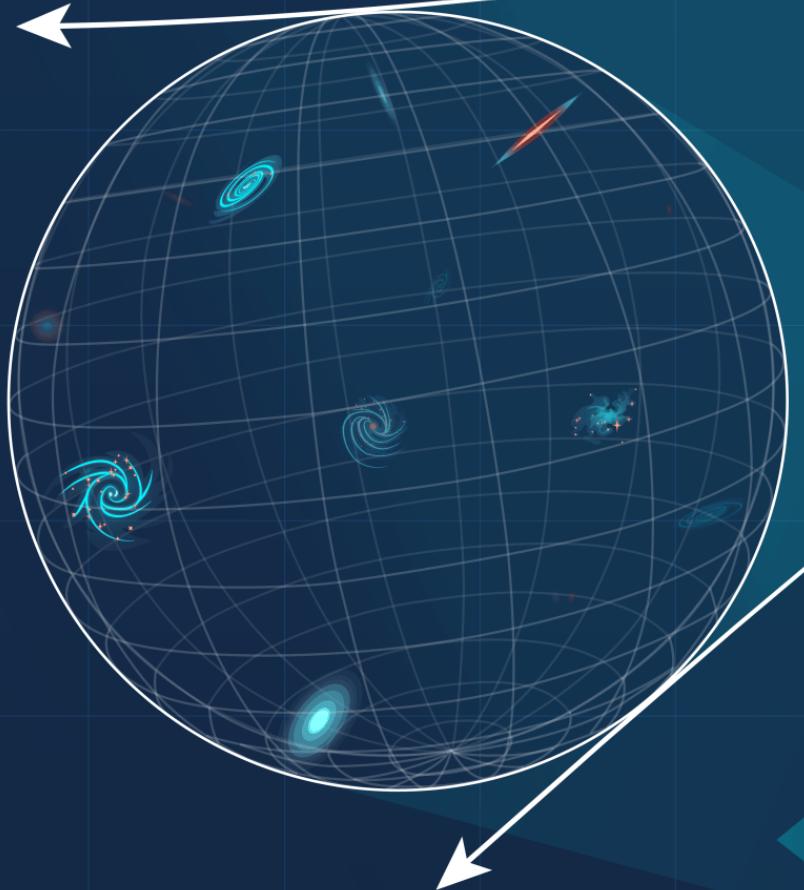
韦诗睿，李长华，张彦霞，崔辰州

weisr@bao.ac.cn

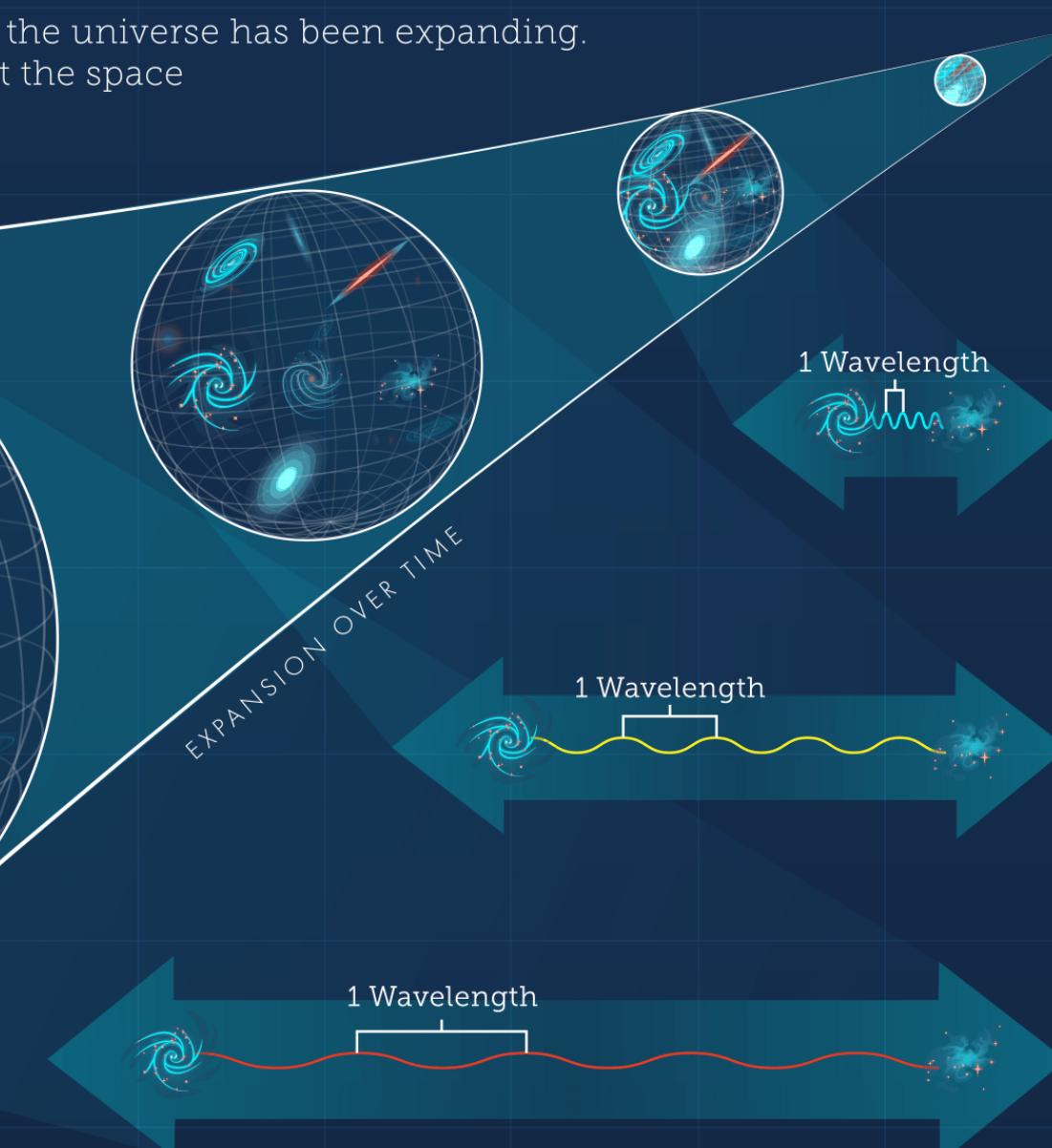
中国科学院国家天文台
国家天文科学数据中心

WHEN SPACE EXPANDS, LIGHT STRETCHES

Since the big bang, the physical space of the universe has been expanding. Stars and galaxies maintain their size, but the space **between** them grows.



EXPANSION OVER TIME



Credit: NASA, ESA, Leah Hustak (STScI)

2025-11-02

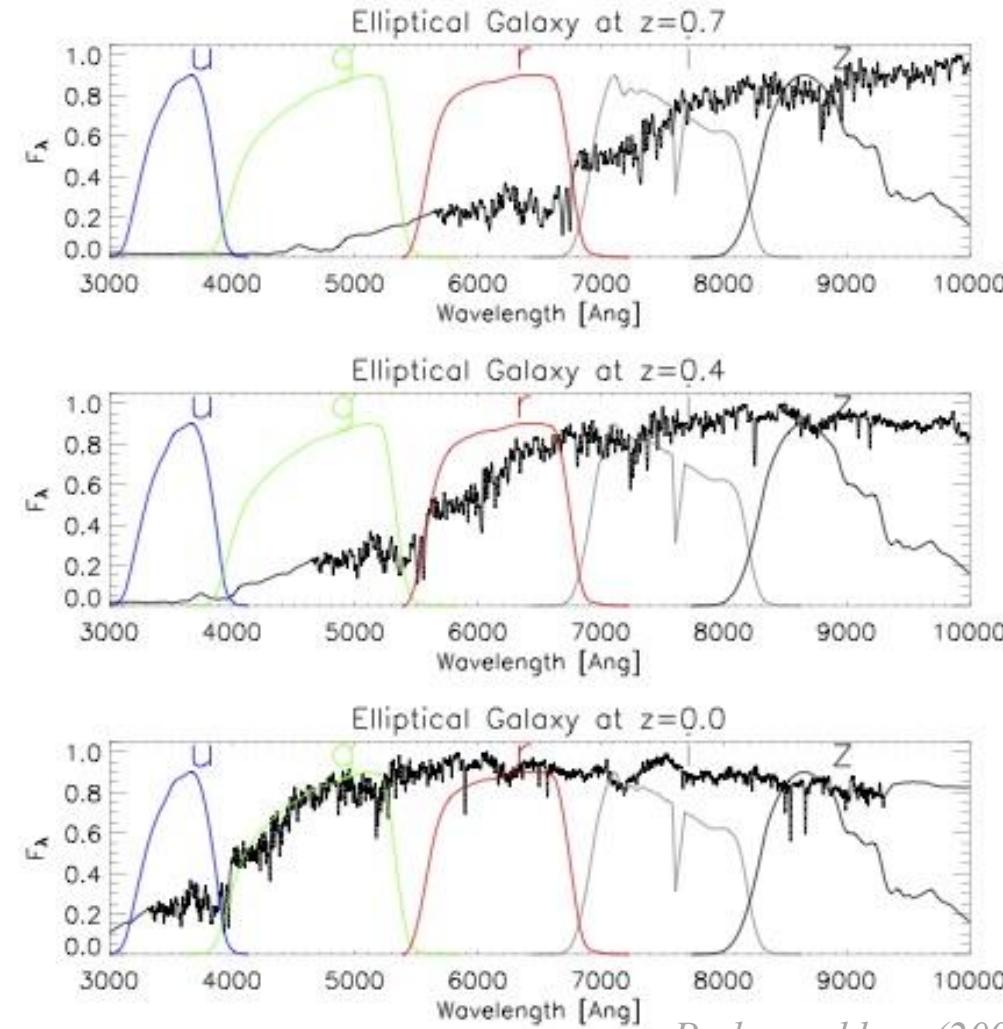
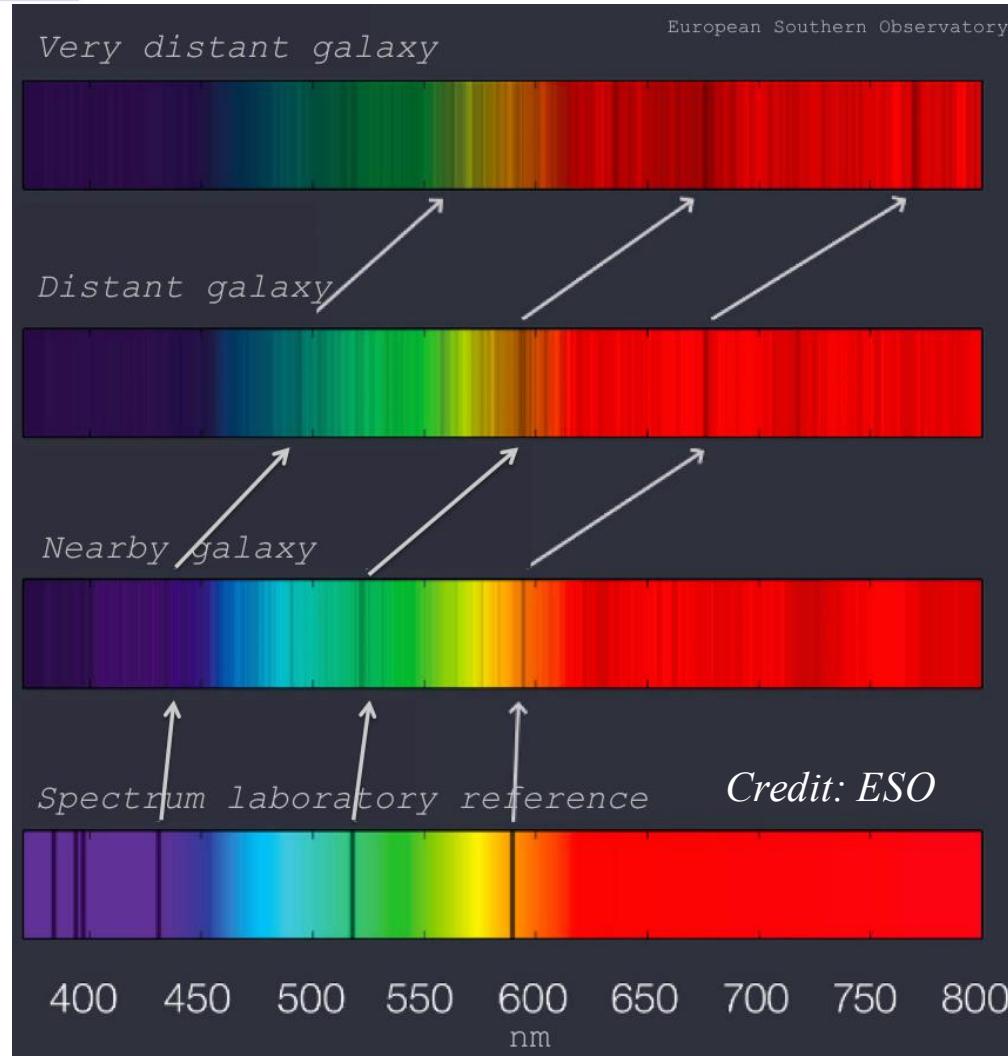
As light travels through expanding space, it is stretched to longer wavelengths.
中国天文学会2025学术年会

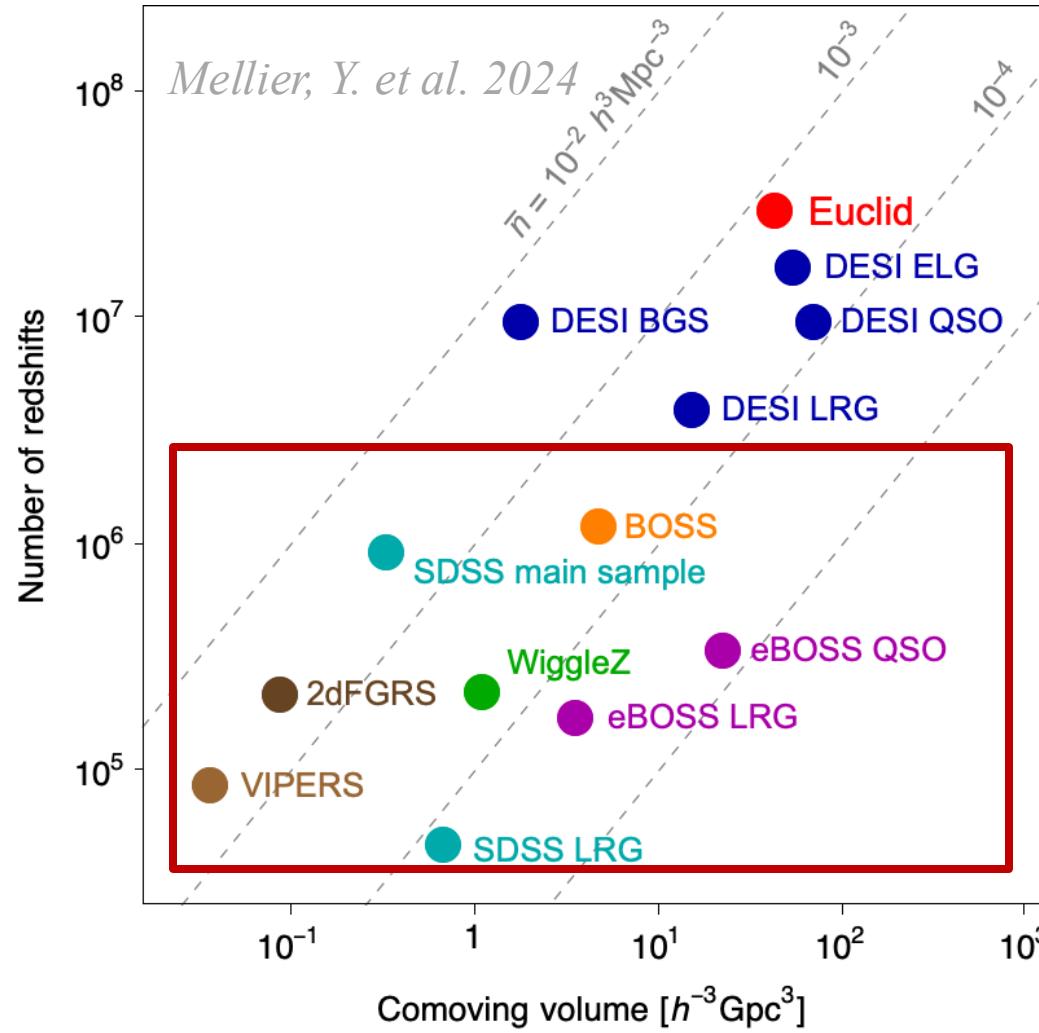
红移

- 天体距离
- 星系演化
- 暗物质与暗能量
- 宇宙大尺度结构

...

光谱红移与测光红移



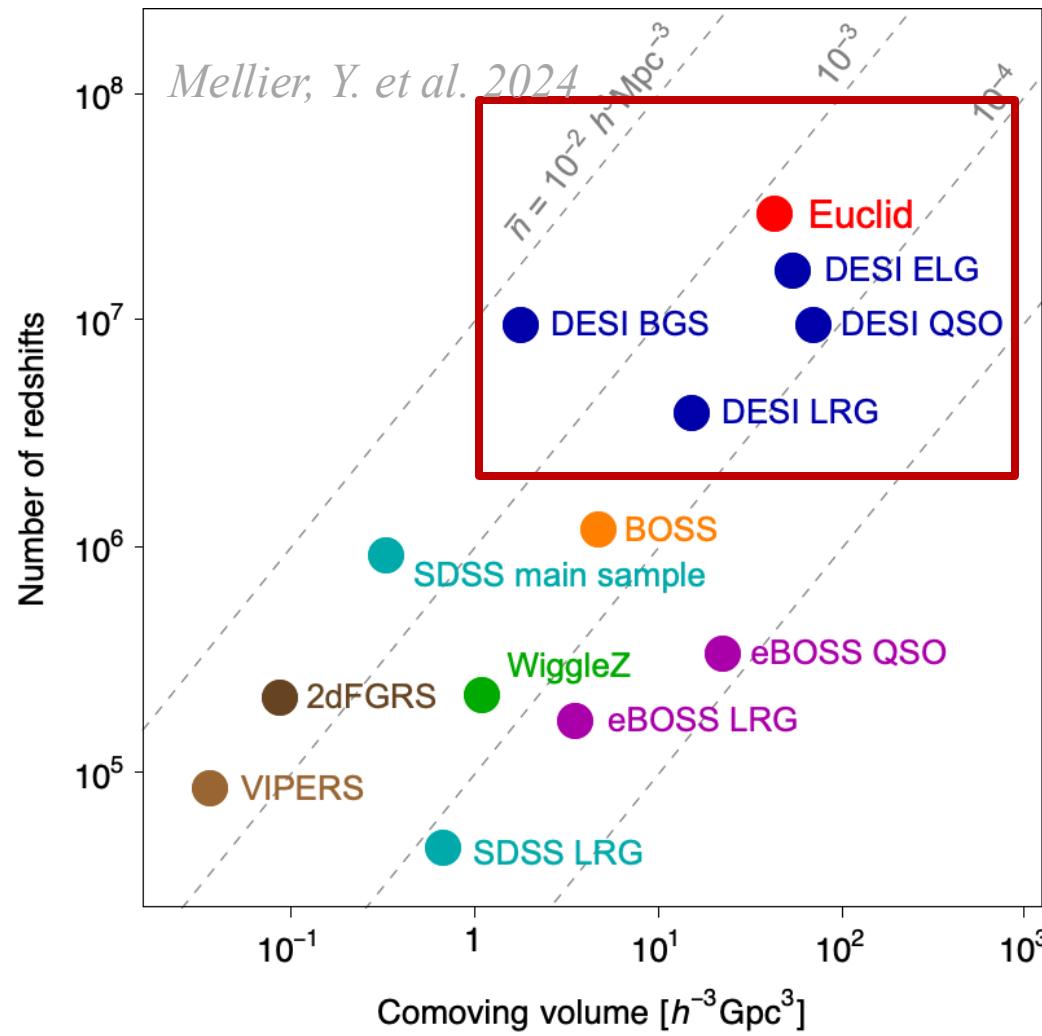


1997-2003
 $z < 0.2$



2000-2021(DR17)
 $z < 1$
Galaxy & QSO

光谱红移巡天



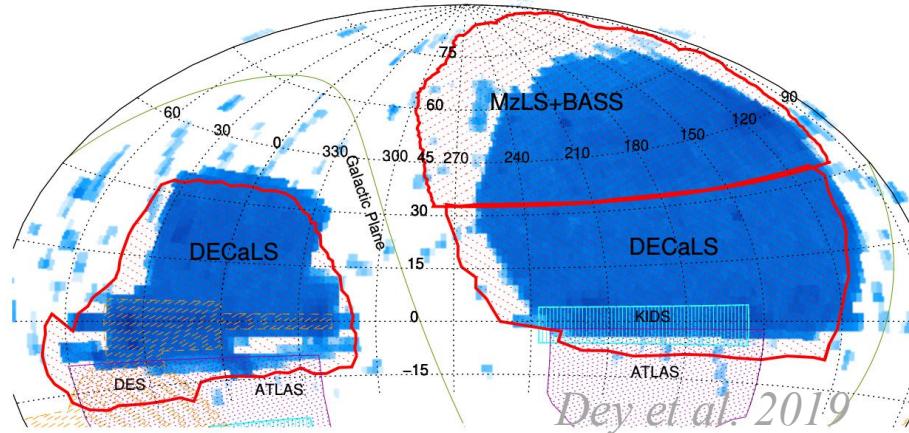
2021 -
 $z \sim 0 - 4$
4千万



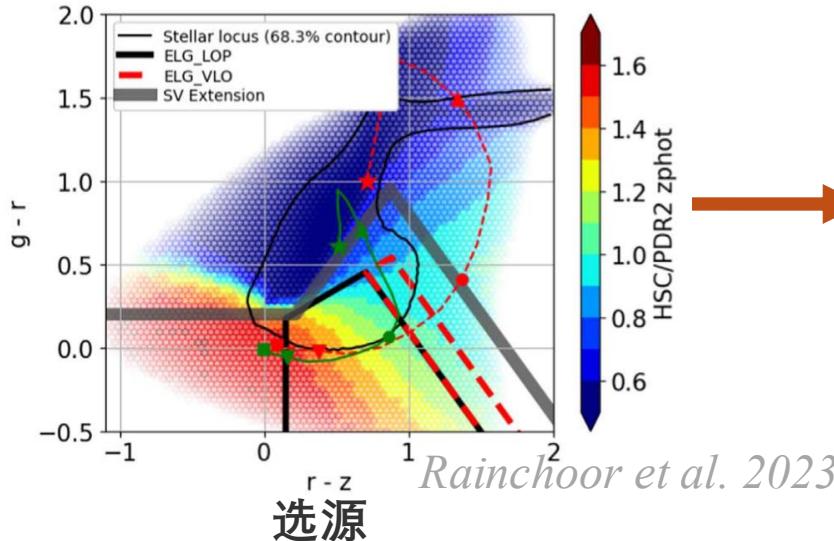
2024 -
5千万



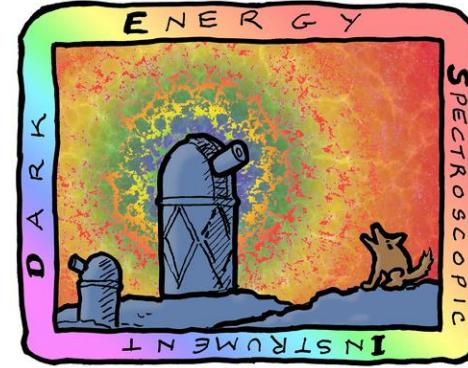
Dark Energy Spectroscopic Instrument



DESI Legacy Imaging Surveys



观测



Survey Validation

DESI - 5 year project

2020

2021

2022

2023

2024

2025

2026

测光

DR9

DR10 (28亿)

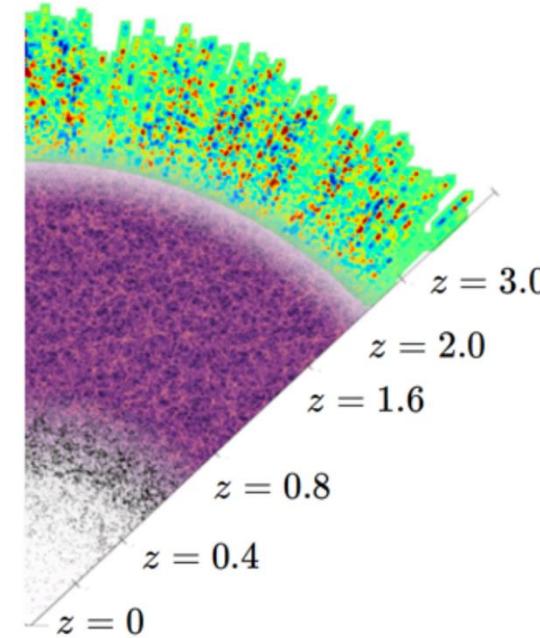
光谱

EDR
(180万)

DR1
(1870万)



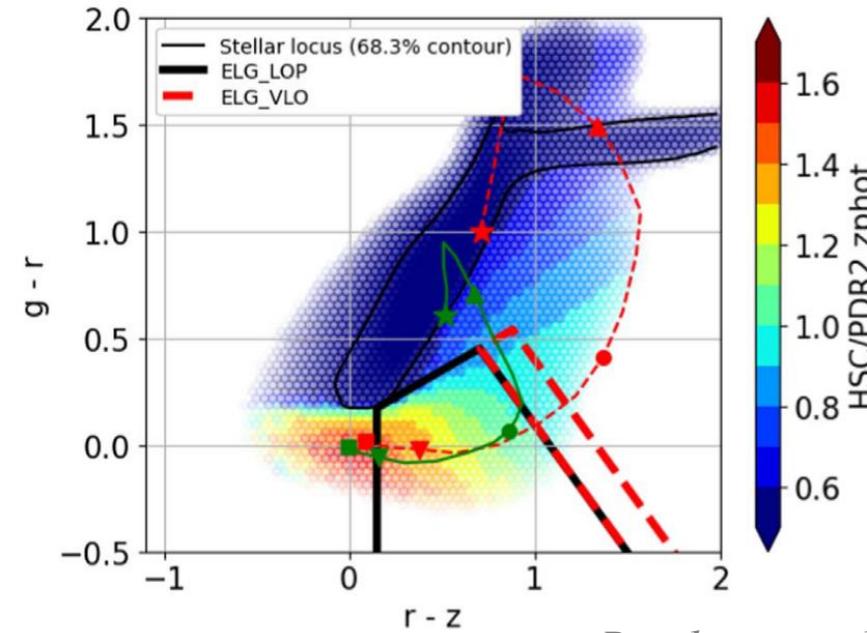
QSO: 3M (SDSS: 500k)
 $\text{Ly}\alpha$ $1.8 < z$
 Tracers $0.8 < z < 2.1$
 ELG: 16M (SDSS: 200k)
 $0.6 < z < 1.6$
 LRG: 8M (SDSS: 1M)
 $0.4 < z < 0.8$
 Bright Galaxies: 14M
 (SDSS: 600k)
 $0 < z < 0.4$



Credit: Pauline Zarrouk

ELG对于理解星系形成演化、化学成分以及宇宙大尺度结构至关重要

提升ELGs的测光红移预测准确性有助于光谱观测选源



Raichoor et al. 2023

- ✓ 观测目标的成功率
- ✓ 选源的完备性
- ✓ 减少非ELG的污染

发射线星系

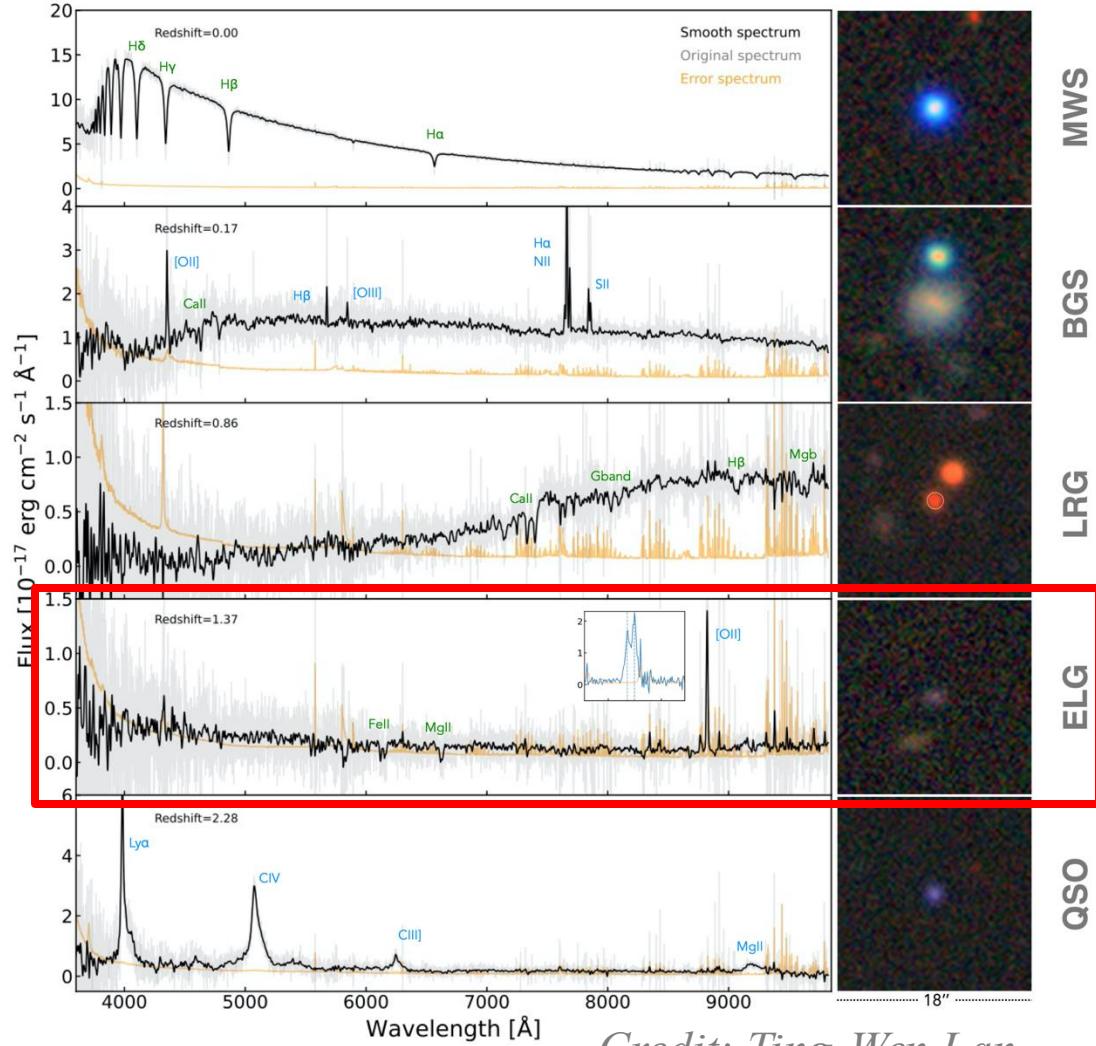


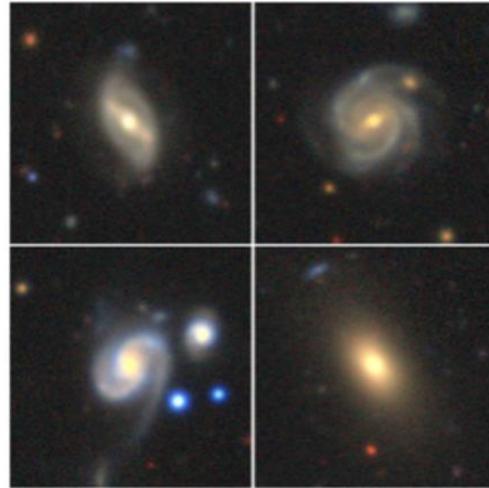
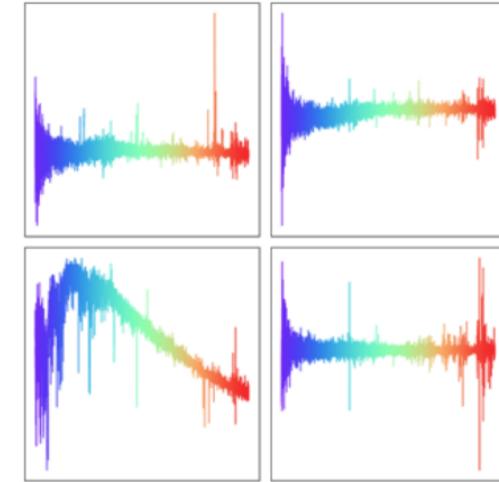
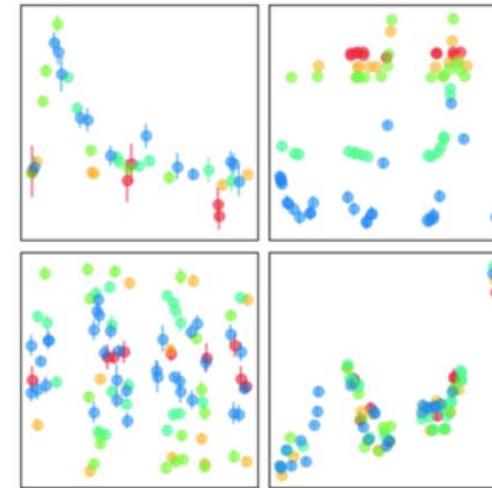
Table 12. The performance comparison of our models with different samples from DESI EDR to the work of Zhou et al. (2023).

Sample	Method	σ_{NMAD}	$\sigma_{\text{NMAD}}(z < 21\text{mag})$	$O(0.10)(\%)$	$O(0.10)(z < 21\text{mag})(\%)$
DESI SV1 BGS	Zhou et al. (2023)	0.021	0.021	2.0	2.0
	Our work	0.013	0.013	0.3	0.3
DESI SV1 LRG	Zhou et al. (2023)	0.028	0.025	3.3	2.7
	Our work	0.019	0.017	3.4	2.6
DESI SV1 ELG	Zhou et al. (2023)	0.095	0.041	34.7	14.4
	Our work	0.073	0.038	28.3	13.4
DESI SV3 BGS	Zhou et al. (2023)	0.019	0.019	1.3	1.3
	Our work	0.012	0.012	0.2	0.2
DESI SV3 LRG	Zhou et al. (2023)	0.025	0.024	1.9	1.6
	Our work	0.012	0.012	0.9	0.8

Credit: Li et al. 2024

Targets	BGS				LRG					
	Metrics	η	σ_{NMAD}	$\bar{\Delta z}$	N_{training}	Metrics	η	σ_{NMAD}	$\bar{\Delta z}$	N_{training}
Our work (Separately)	0.14%	0.020	-0.0031	0.3 m	0.68%	0.030	0.0111	0.1 m		
Our work (Bayesian)	0.14%	0.018	-0.0015	0.3 m	0.45%	0.026	0.0022	0.1 m		
Our work (Collectively)	0.83%	0.029	0.0160	-	1.07%	0.033	0.0151	-		
Zou et al. (2019)	0.19%	0.013	0.0	1.2 m	0.18%	0.016	0.0013	0.8 m		
Targets	ELG				NON					
	Metrics	η	σ_{NMAD}	$\bar{\Delta z}$	N_{training}	Metrics	η	σ_{NMAD}	$\bar{\Delta z}$	N_{training}
Our work (Separately)	16.65%	0.112	0.0257	0.1 m	9.44%	0.058	0.0014	0.2 m		
Our work (Bayesian)	16.07%	0.107	0.0184	0.1 m	7.87%	0.052	0.0140	0.2 m		
Our work (Collectively)	15.78%	0.108	0.0154	-	8.15%	0.053	0.0304	-		
Zou et al. (2019)	6.23%	0.053	0.0067	0.02 m	2.46%	0.024	0.0022	0.6 m		

Credit: Zhou et al. 2024

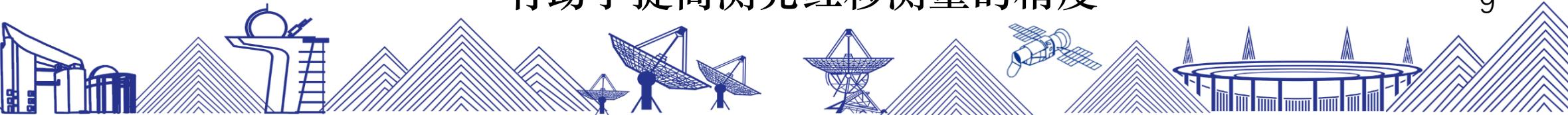
The Multimodal Universe Collaboration et al. 2024图片
形态信息光谱
特征线信息光变曲线
时序信息*Rizhko et al. 2024*

Feature	Description
mean_vmag	Mean magnitude in the visible band
phot_g_mean_mag	Gaia G-band mean magnitude
e_phot_g_mean_mag	Uncertainty in Gaia G-band mean magnitude
phot_bp_mean_mag	Gaia BP band mean magnitude
e_phot_bp_mean_mag	Uncertainty in Gaia BP band mean magnitude
phot_rp_mean_mag	Gaia RP band mean magnitude
e_phot_rp_mean_mag	Uncertainty in Gaia RP band mean magnitude
bp_rp	BP mean magnitude minus RP mean magnitude
parallax	Gaia DR3 Parallax measurement
parallax_error	Uncertainty in parallax measurement
parallax_over_error	Signal-to-noise ratio for parallax measurement

星表
测光、自行信息

多模态信息互补，丰富特征

有助于提高测光红移测量的精度



数据集构建

ELG样本

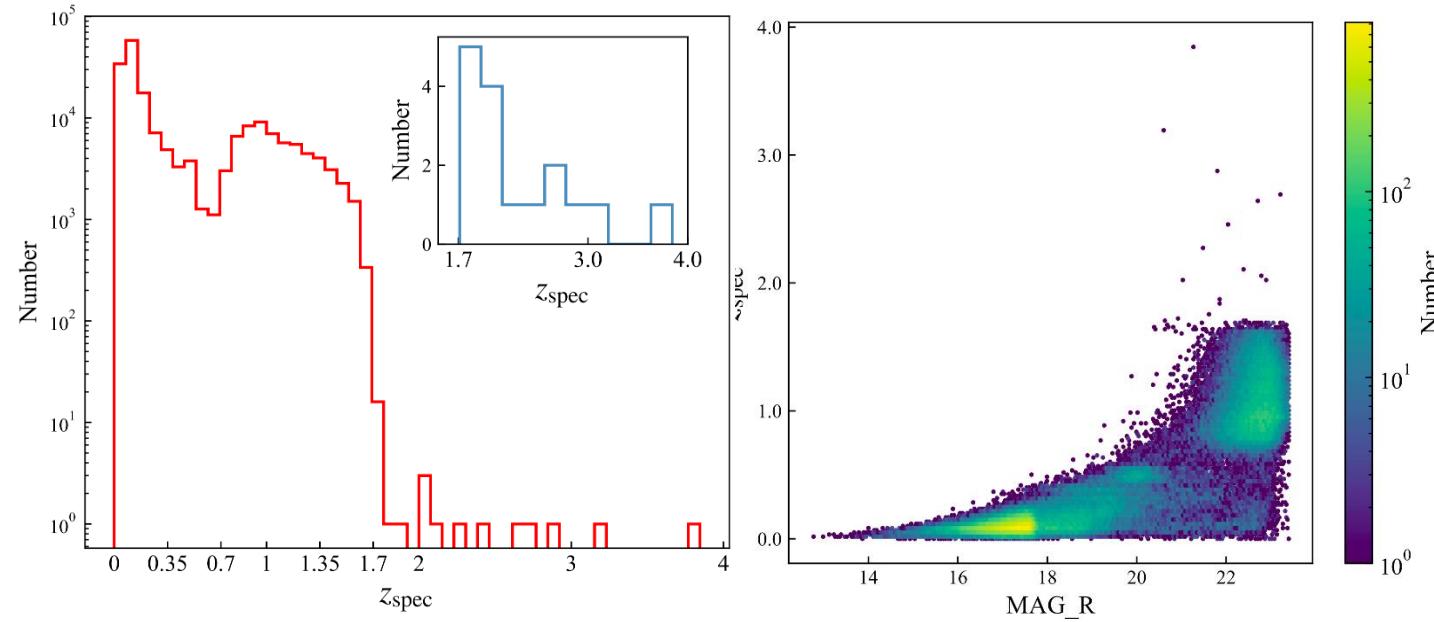
Table 1. Information about the various spectroscopic surveys included in the dataset. \bar{z}_{spec} denotes the median of z_{spec} . $\bar{r}(\text{mag})$ denotes the median of r -band magnitude.

Survey	No. of sources	Quality criterion	\bar{z}_{spec}	$\bar{r}(\text{mag})$
SDSS	122 452	$\text{zWarning}=0$	0.11	17.48
DESI SV3	52 491	$\text{zWarning}=0, \text{ZERR}<0.01$	1.04	22.73
DESI SV1	10 782	$\text{zWarning}=0, \text{ZERR}<0.01$	0.94	22.63
dFGRS	3 220	$Q=3,4$	0.06	15.57
PRIMUS	909	$Z\text{QUALITY}=3,4$	0.56	21.37
zCOSMOS	716	$5 \geq CC \geq 3$	0.68	21.67
LAMOST	612	$Z_ERR \leq 0.01$	0.12	16.98
GAMA	442	$NQ=4$	0.19	18.78
VIPERS	274	$z\text{flag}=3,4$	0.79	21.83
WiggleZ	118	$qz=4,5$	0.63	20.90
VVDS	109	$\text{SIMBAD: AGN}/\text{EmG}$ $Z\text{FLAGS}=4$	0.74	22.13
fCOSMOS	103	$Z\text{FLAG} \geq 3$	1.48	22.73
dFGRS	88	$Q=4$	0.11	17.38
OzDES	28	$qop=4$	0.46	20.46
C3R2	19	$\text{Qual} \geq 3$	0.82	22.43
DEEP2	12	$Z\text{QUALITY} \geq 3$	0.82	22.67
Total No.	192 375		0.16	18.41

STARFORMING / STARBURST / AGN /
BROADLINE

ELG

训练：验证：测试
144663: 9237: 38475



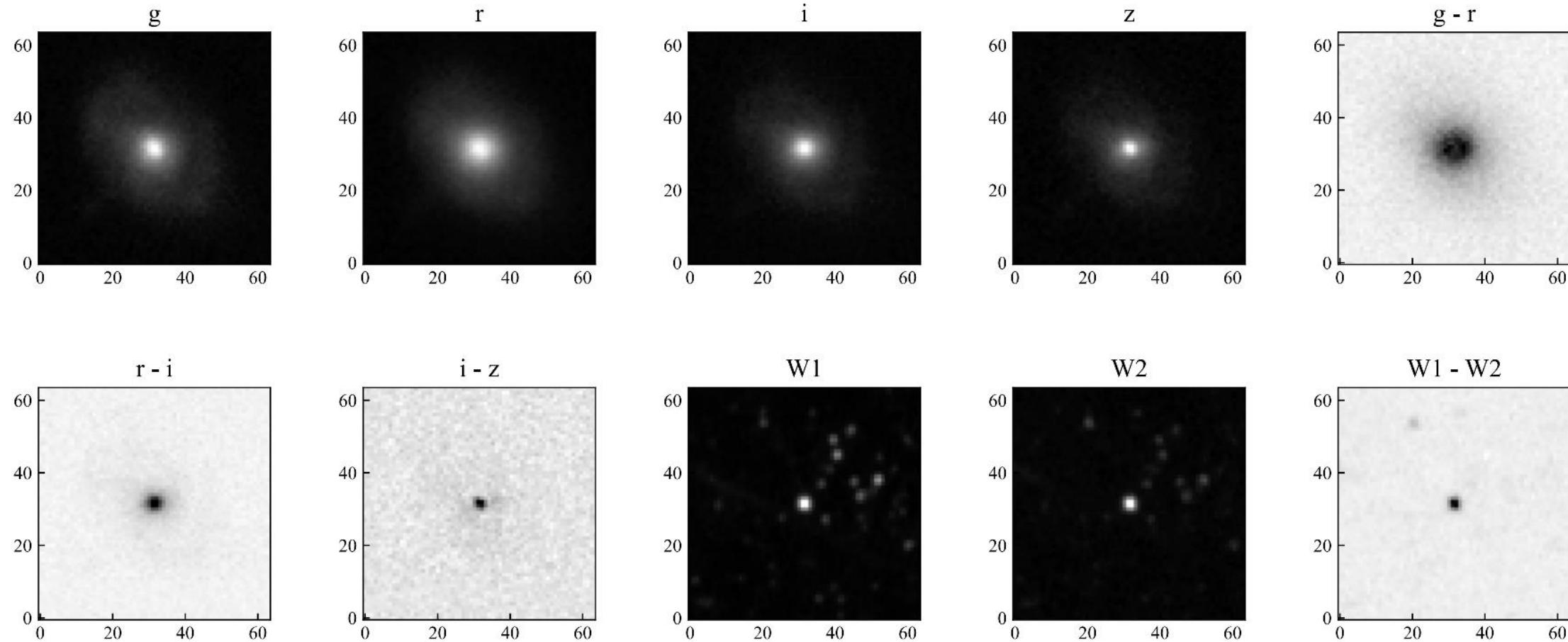


Figure 4. Example multi-band images of a single ELG source with a z_{spec} of 0.156 from the dataset. The 64×64 imaging data consist of 10 channels, with the bands ordered from optical to infrared. Notable resolution discrepancies are observed between the optical and infrared images.

Table 2. Photometric data corresponding to the known ELG dataset.

Name	Definition	Catalogue	Waveband
MAG_G	Model magnitude in <i>g</i> band	DESI	Optical band
MAG_R	Model magnitude in <i>r</i> band	DESI	Optical band
MAG_I	Model magnitude in <i>i</i> band	DESI	Optical band
MAG_Z	Model magnitude in <i>z</i> band	DESI	Optical band
MAG_W1	Model magnitude in <i>W1</i> band	DESI	Infrared band
MAG_W2	Model magnitude in <i>W2</i> band	DESI	Infrared band
APMAG_G_(1~8) ^a	Aperture magnitude in <i>g</i> band	DESI	Optical band
APMAG_R_(1~8)	Aperture magnitude in <i>r</i> band	DESI	Optical band
APMAG_I_(1~8)	Aperture magnitude in <i>i</i> band	DESI	Optical band
APMAG_Z_(1~8)	Aperture magnitude in <i>z</i> band	DESI	Optical band
APMAG_W1_(1~5) ^b	Aperture magnitude in <i>W1</i> band	DESI	Infrared band
APMAG_W2_(1~5)	Aperture magnitude in <i>W2</i> band	DESI	Infrared band
ebv	Galactic extinction E(B-V) reddening	DESI	
z_spec	Spectroscopic redshift	Known sample	

^aOptical aperture magnitude suffixes 1~8 correspond to fluxes measured within circular apertures of radii [0.5, 0.75, 1.0, 1.5, 2.0, 3.5, 5.0, 7.0] arcsec.

^bInfrared aperture magnitude suffixes 1~5 correspond to apertures with radii [3, 5, 7, 9, 11] arcsec.

模型、孔径星等

模型星等颜色: MAG_G - MAG_R

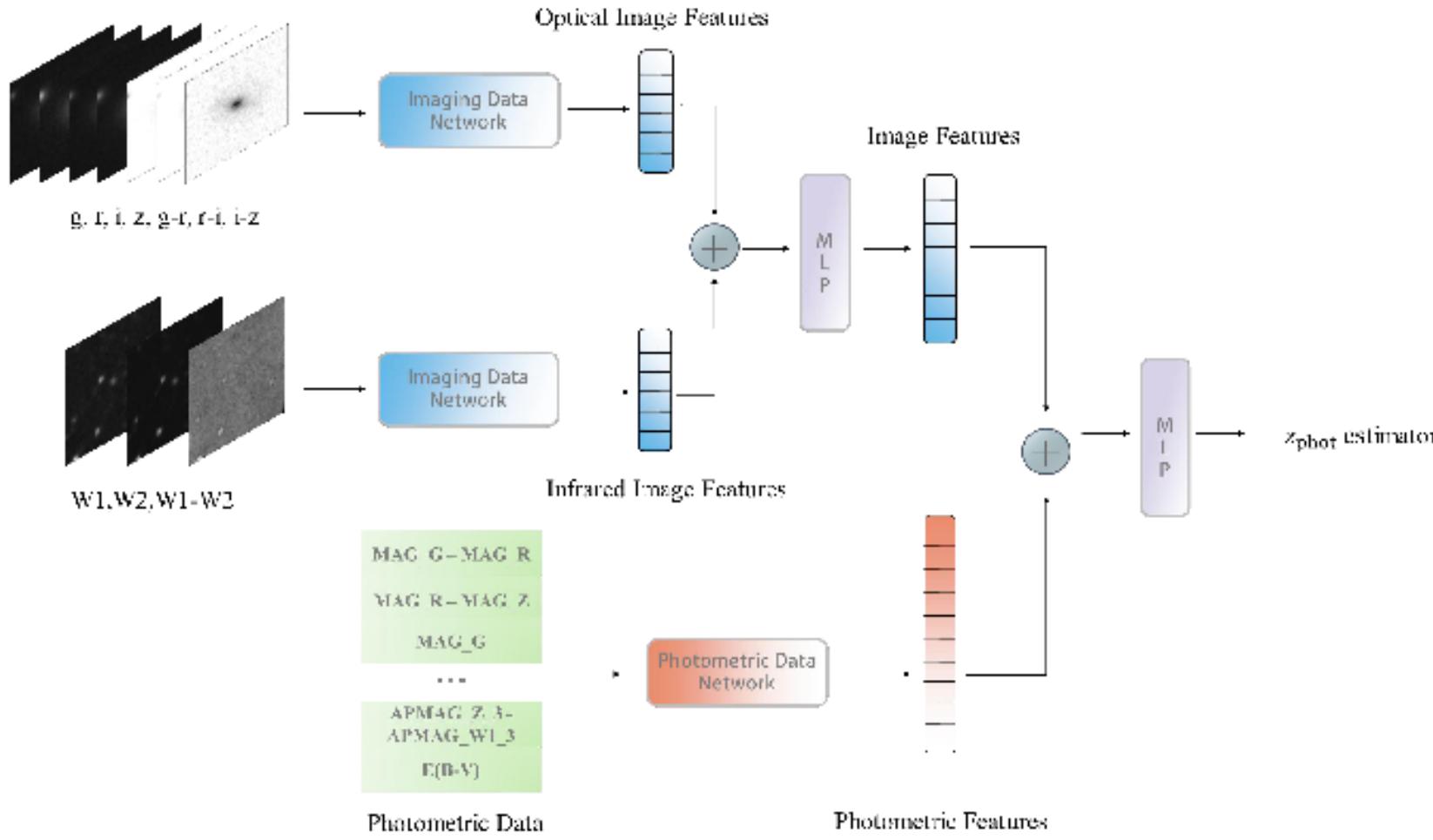
孔径星等颜色:

APMAG_G_1 - APMAG_R_1

Appendix A.2. List of selected photometric features

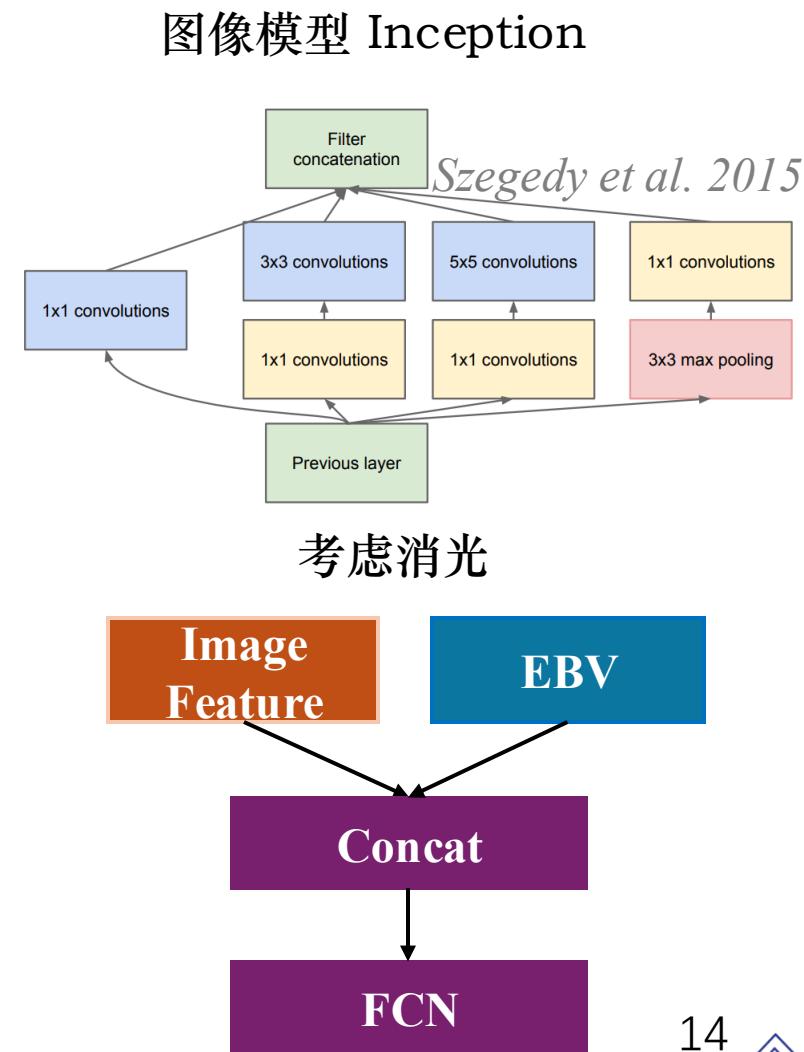
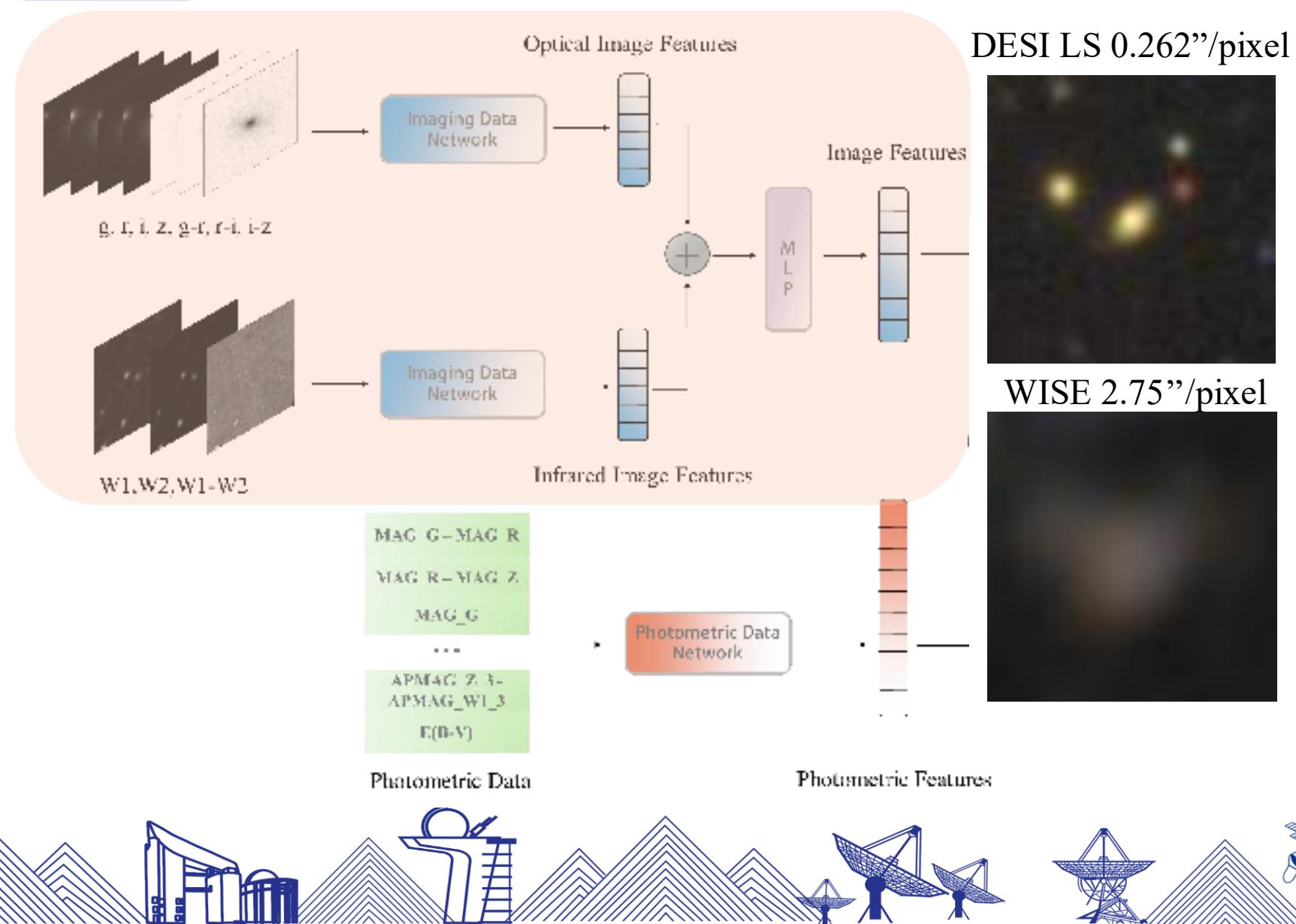
The 85 photometric features from the DESI LS10 catalogue ranked by PFI, used as the input to the model are as follows: MAG_R - MAG_W2, MAG_I - MAG_W1, MAG_R - MAG_W1, MAG_I - MAG_W2, MAG_Z - MAG_W1, -MAG_R - MAG_I, MAG_G - MAG_W2, MAG_Z - MAG_W2, MAG_G - MAG_R, MAG_I - MAG_Z, MAG_G - MAG_I, APMAG_W1_2 - APMAG_W2_2, MAG_R - APMAG_R_1, MAG_R - MAG_Z, MAG_G - APMAG_G_1, APMAG_Z_5 - APMAG_W1_5, MAG_G, MAG_G - MAG_W1, MAG_I - APMAG_I_1, APMAG_W1_3 - APMAG_W2_3, MAG_I, MAG_Z, APMAG_I_7 - APMAG_Z_7, MAG_Z - APMAG_Z_1, APMAG_R_6 - APMAG_I_6, APMAG_I_6 - APMAG_Z_6,

模型构建 多模态模型

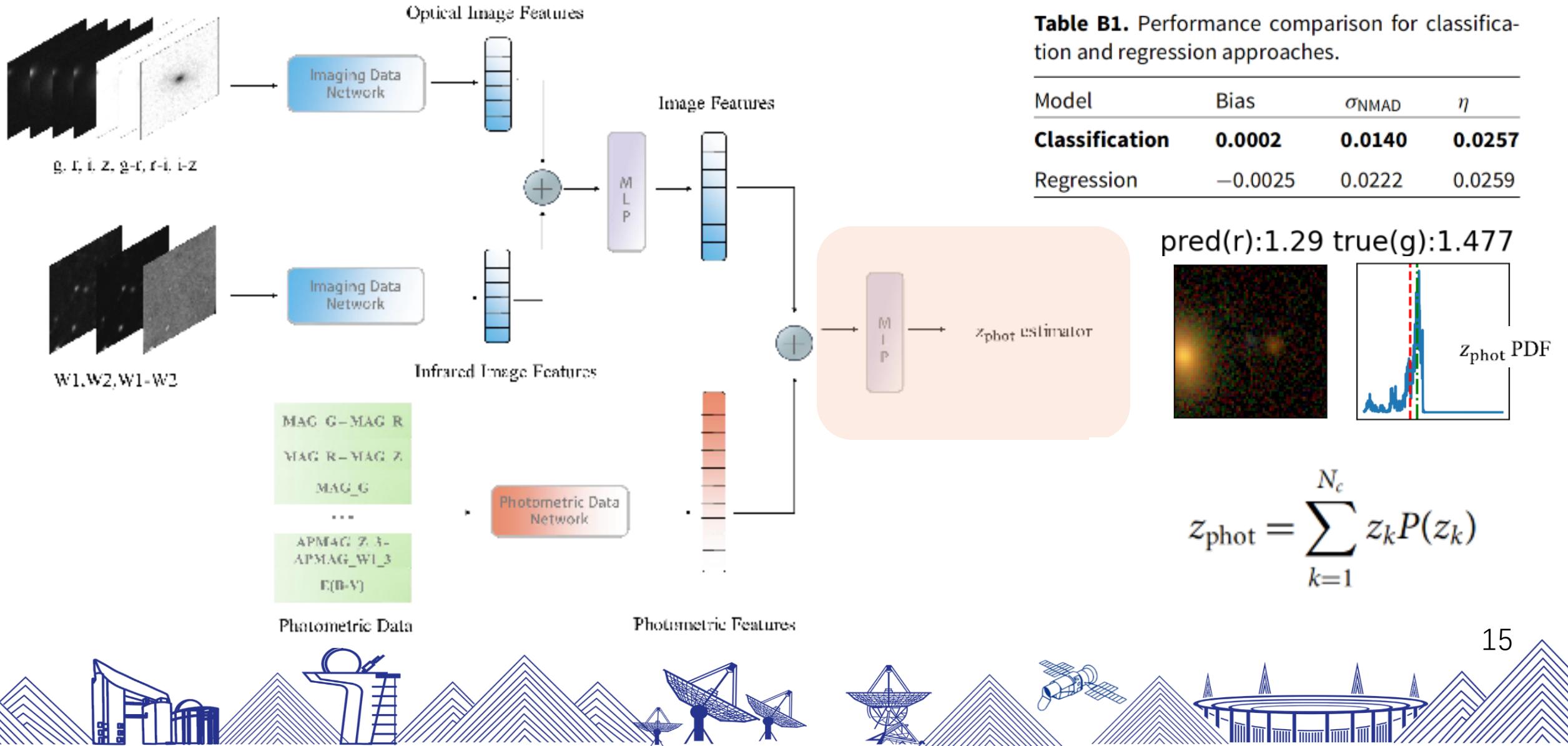


CNN – MLP
多波段图像：Inception
星表：MLP
模态融合：特征拼接
红移估计：MLP

模型构建 图像模块



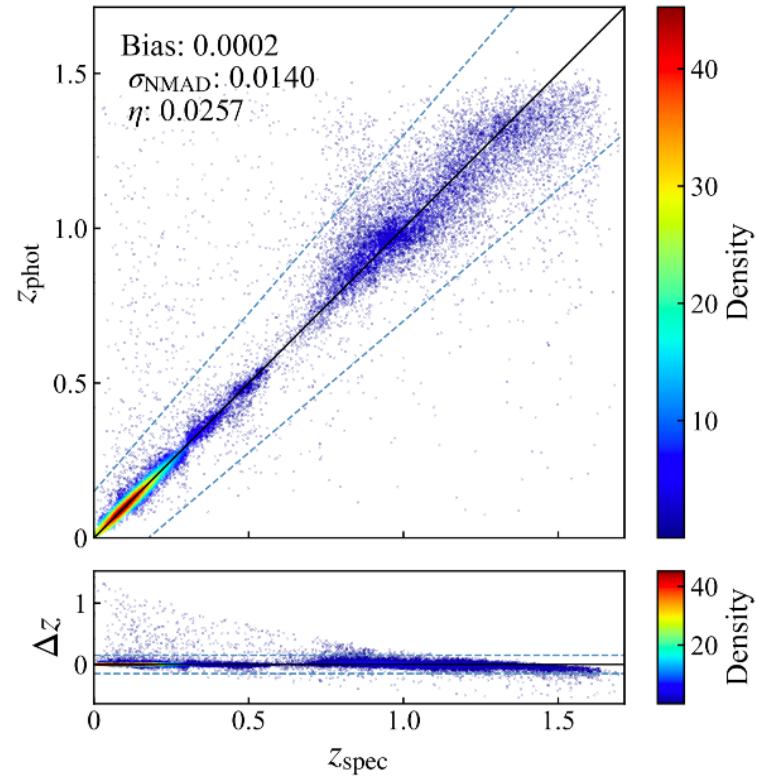
模型构建 红移估计模块



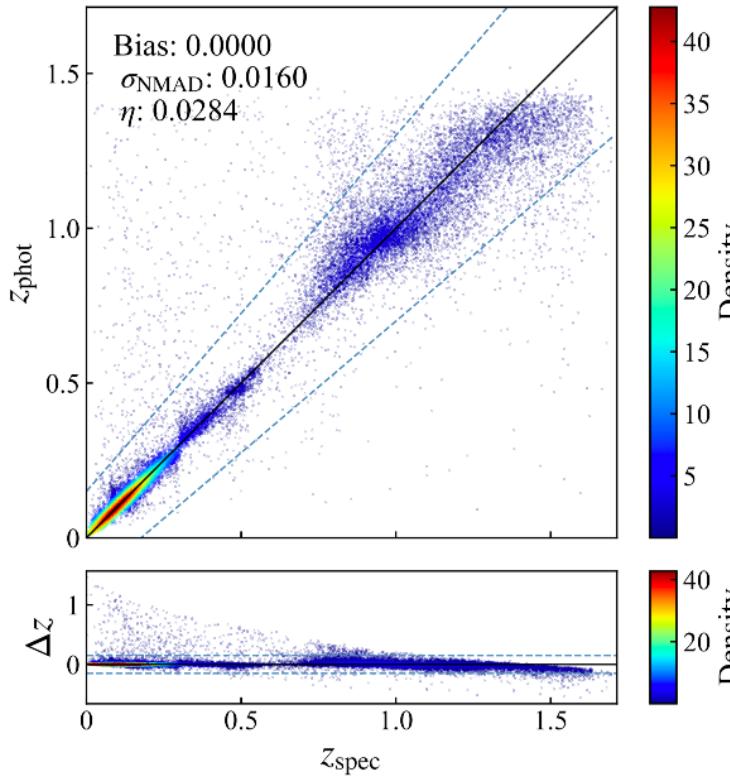
实验结果

结合**多模态**数据提供的互补信息有助于ELG测光红移估计效果的**提升**

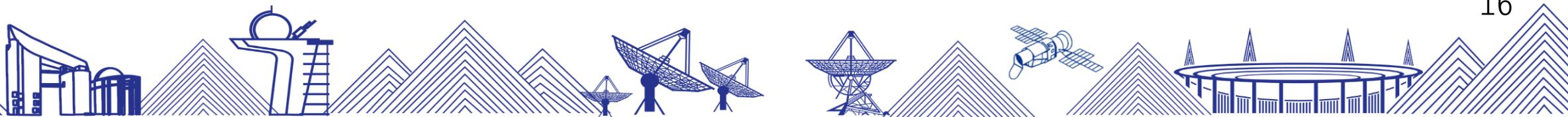
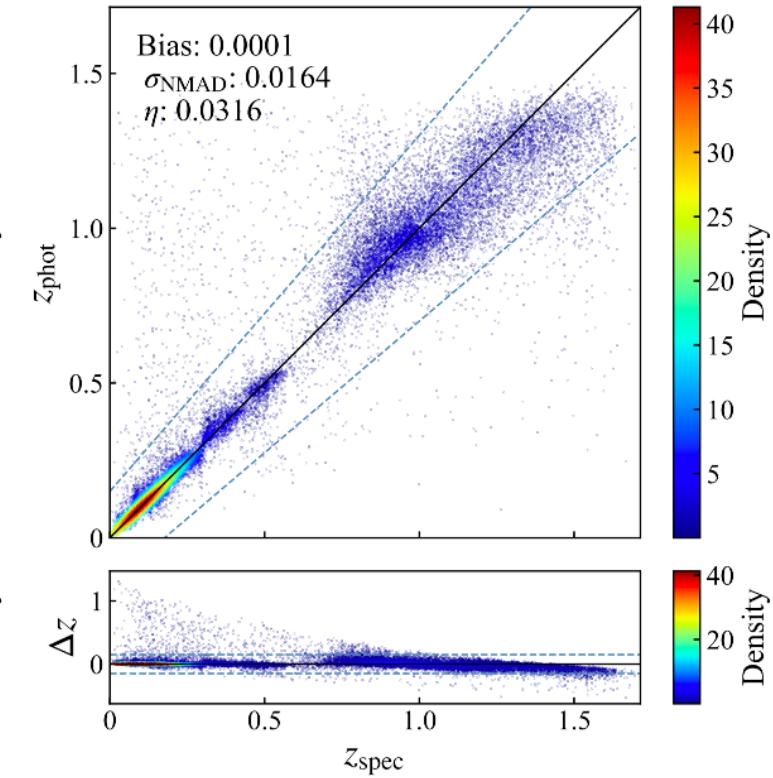
图像+星表



只使用星表



只使用图像



实验结果

模型在低红移和亮源上对于测光红移估计的准确度较高

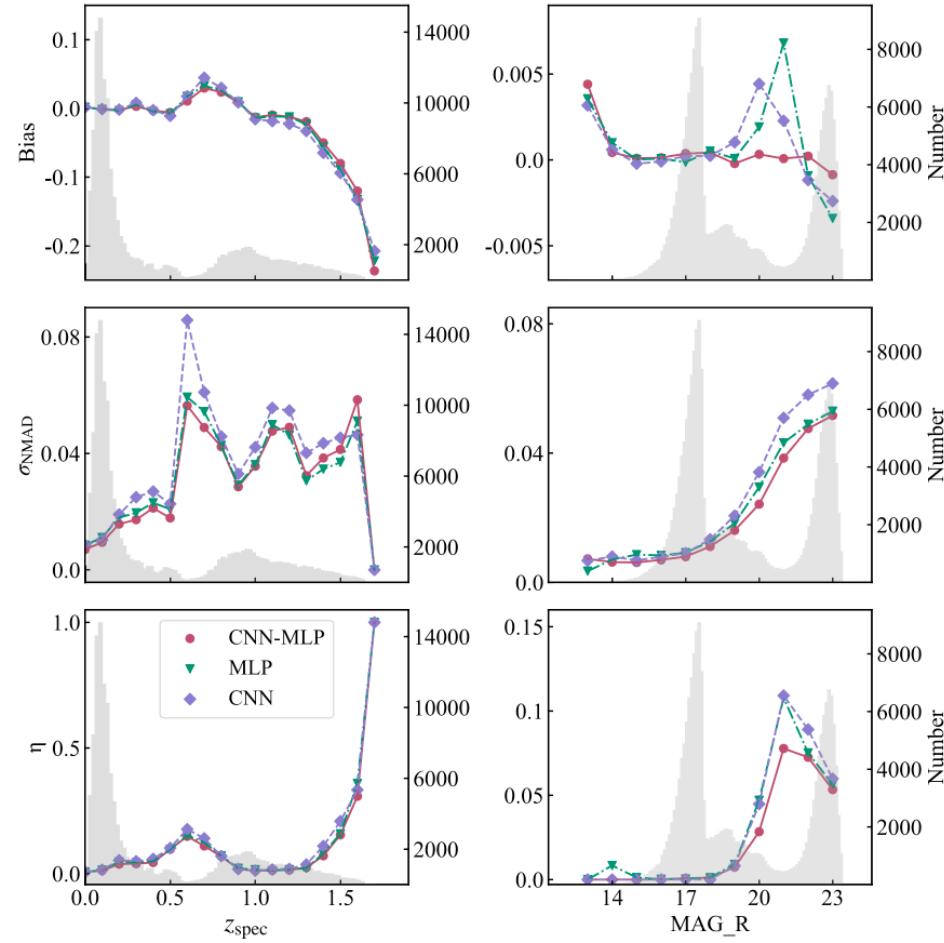
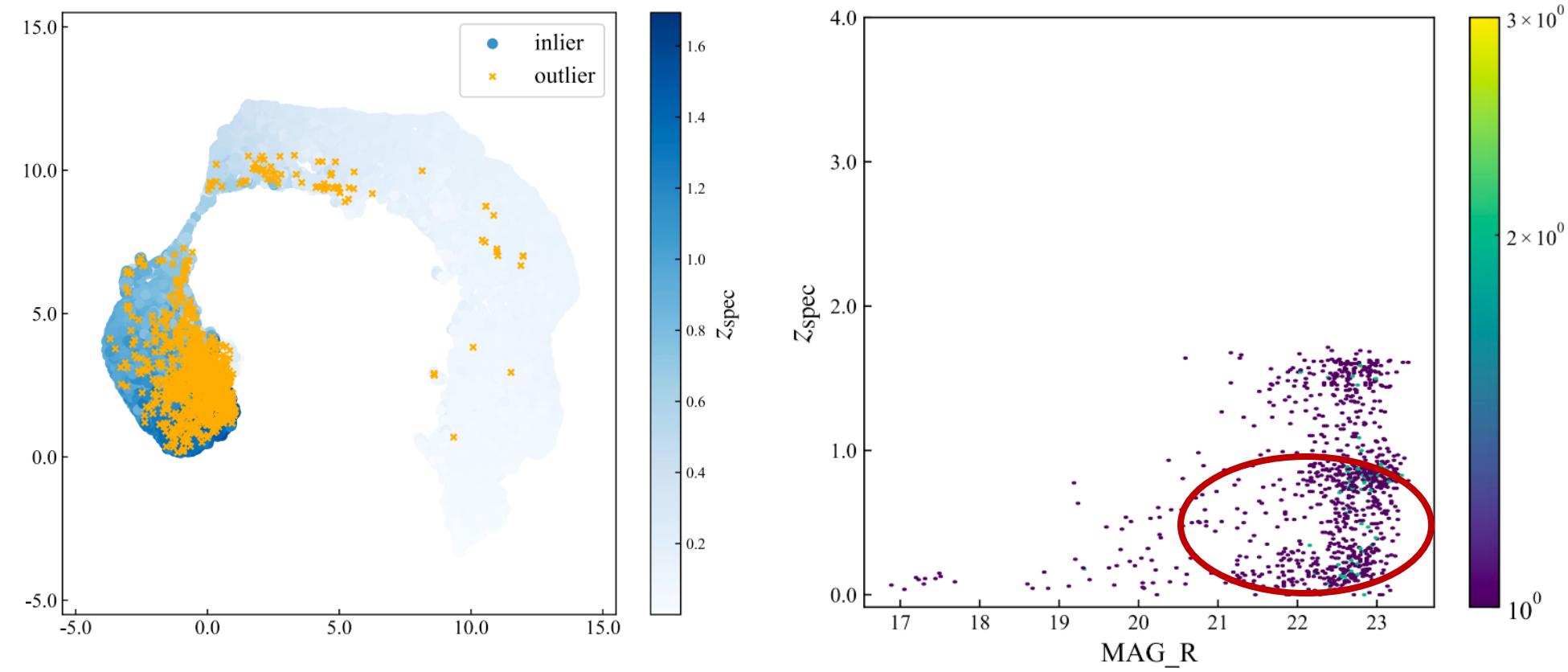


Table 5. Performance of bright and faint sources using the two-part model and one-part model.

Model	Sources	Bias	σ_{NMAD}	η
Two-part model	$MAG_R < 21.5$	0.0000	0.0089	0.0044
	$MAG_R > 21.5$	0.0010	0.0494	0.0715
	All	0.0001	0.0139	0.0270
One-part model	$MAG_R < 21.5$	0.0003	0.0090	0.0041
	$MAG_R > 21.5$	0.0000	0.0479	0.0691
	All	0.0002	0.0140	0.0257

部分离群点来自低红移和暗源的简并性

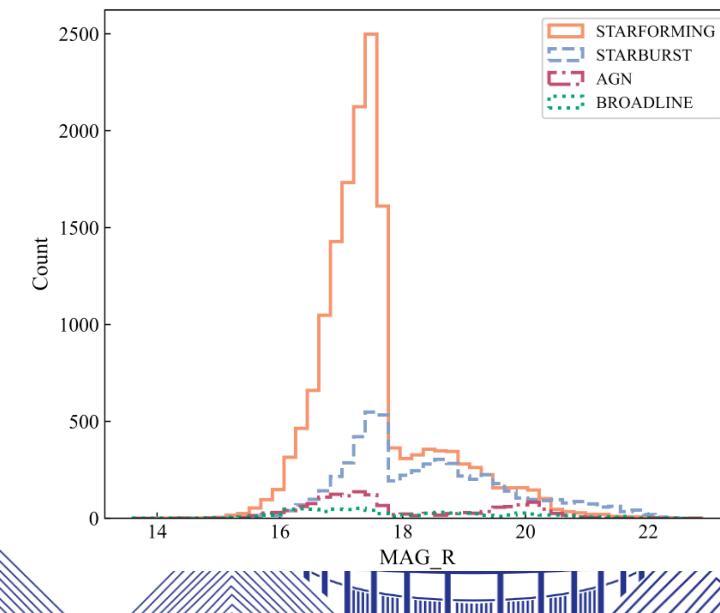
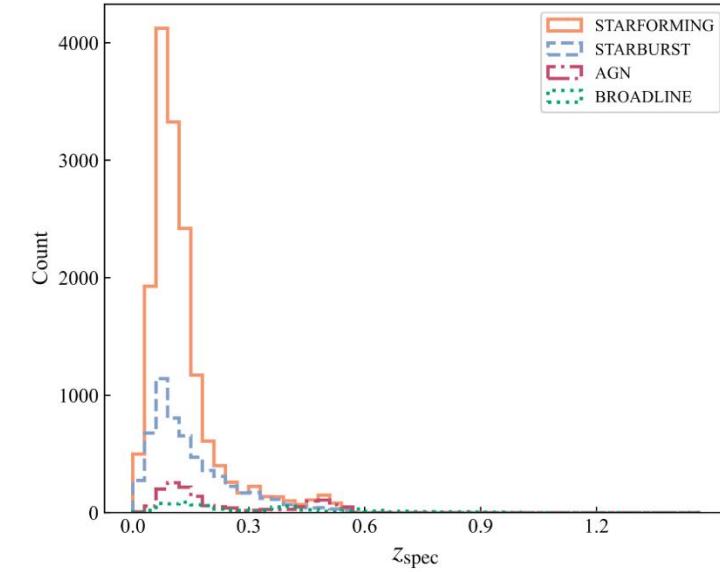


实验结果

发射线星系类型对于模型效果无较大差异

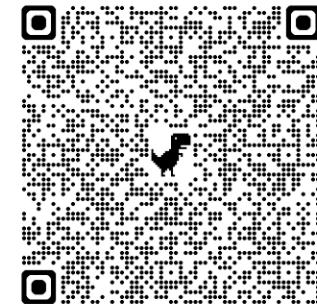
Table 6. Performance for different galaxy types.

Galaxy type	No. of sources	Bias	σ_{NMAD}	η
Starforming	15 944	0.0003	0.0083	0.0023
Starburst	5 738	0.0003	0.0101	0.0052
AGN	1 512	0.0030	0.0134	0.0040
Broadline	803	-0.0027	0.0103	0.0050
All	23 997	0.0003	0.0090	0.0032



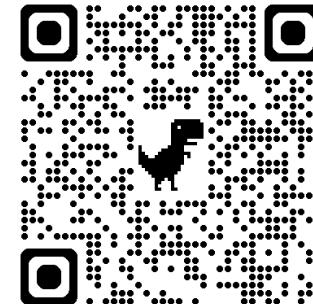
- ✓ 构建CNN-MLP模型整合多波段测光图像和测光星表多模态天文数据
- ✓ 提高DESI Legacy Imaging Surveys 发射线星系的测光红移预测精度
- ✓ 为DESI及后续大型巡天项目ELG选源和测光红移测量提供参考

论文



Wei et al. 2025, PASA
arXiv:2505.24175

代码



<https://github.com/RuiNov1st/Zphot4ELG>

