流体力学笔记

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2021年11月19日

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Introduction

1.1 描述流体

密度 ρ , 压力p, 温度T, 速度 \vec{u}

1.2 稳态和非稳态流体

$$\frac{\partial}{\partial t} = 0$$

1.3 二维流动

two components (plane flow)

two coordinates (e.g. : $\partial/\partial\phi=0$)

1.4 流线

At a **fixed** time

$$\frac{dx}{u_x} = \frac{dy}{u_y} = \frac{dz}{u_z} = const$$

1.5 物质导数和迁移 material derivative and advection

$$\frac{\mathrm{d}f}{\mathrm{d}t} = \frac{\partial f}{\partial t} + \frac{\partial f}{\partial x} \frac{\partial x}{\partial t} + \frac{\partial f}{\partial y} \frac{\partial y}{\partial t} + \frac{\partial f}{\partial z} \frac{\partial z}{\partial t}
= \frac{\partial f}{\partial t} + u_x \frac{\partial f}{\partial x} + u_y \frac{\partial f}{\partial y} + u_z \frac{\partial f}{\partial z}
= \frac{\partial f}{\partial t} + (\vec{u} \cdot \vec{\nabla}) f$$
(1.1)

其中:

 $\frac{df}{dt}$ is material derivative

 $\frac{\partial f}{\partial t}$ is local rate

 $(\vec{u} \cdot \vec{\nabla})f$ is advection rate

1.6 拉格朗日描述和欧拉描述

If f is \vec{u} :

$$\frac{\mathrm{d}\vec{u}}{\mathrm{d}t} = \frac{\partial\vec{u}}{\partial t} + (\vec{u} \cdot \vec{\nabla})\vec{u}$$

左边为拉格朗日描述, 右边为欧拉描述

1.7 不可压缩流体(incompressible fluid)

$$\oint \vec{u} \cdot \vec{n} \ ds = 0$$

微分形式

$$\vec{\nabla} \cdot \vec{u} = 0$$

1.8 mass conservation 质量守恒

对于一个流体微团, 其质量变化率为:

$$\frac{\partial}{\partial t} \int \rho dV = -\oint \rho \ \vec{u} \cdot \vec{n} \ ds = -\int \nabla \cdot (\rho \vec{u}) dV$$

右边为流体微团质量的变化率; 右边为流体微团的面积微元的法向上, 物质流出/流入速率, 变成微分形式:

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{u}) = 0$$

等式左边第二项展开:

$$\frac{\partial \rho}{\partial t} + (\vec{u} \cdot \vec{\nabla})\rho + \rho(\vec{\nabla} \cdot \vec{u}) = 0$$

$$\frac{d\rho}{dt} + \rho(\vec{\nabla} \cdot \vec{u}) = 0$$
(1.2)

换个角度看质量守恒(连续方程),在拉格朗日视角,一个流体微元的体积为 $\delta\tau = \delta x \delta y \delta z$,质量为 $\delta m = \rho \delta \tau$;在运动过程中,该流体微元质量总是守恒的.

$$\frac{\mathrm{d}}{\mathrm{d}t}(\delta m) = \frac{\mathrm{d}}{\mathrm{d}t}(\rho \delta \tau) = 0$$

展开,得到:

$$\frac{1}{\rho} \frac{\mathrm{d}\rho}{\mathrm{d}t} + \frac{1}{\delta\tau} \frac{\mathrm{d}}{\mathrm{d}t} (\delta\tau) = 0$$

其中, $\frac{1}{\delta \tau} \frac{d}{dt} (\delta \tau)$ 为流体膨胀速率与原来体力比, 表征流体的体膨胀速度., 写开变为:

$$\frac{1}{\delta \tau} \frac{\mathrm{d}}{\mathrm{d}t} (\delta \tau) = \nabla \cdot u$$

整理得到连续方程. 对于不可压缩流体:

$$\rho(\vec{\nabla} \cdot \vec{u}) = 0$$

$$\frac{d\rho}{dt} = 0$$
(1.3)

1.9 Vorticity 涡度/旋度

$$\vec{\omega} = \vec{\nabla} \times \vec{u}$$

由斯托克斯公式 $\iint_s \nabla \times f \cdot n \mathrm{d}s = \oint_l f \cdot dl$ 对速度环量和涡度积分:

$$\oint \vec{u} d\vec{l} = \int \vec{\nabla} \times \vec{u} \ d\vec{s} = \int \vec{\omega} \ d\vec{s}$$

 Ω 为流体运动的角速度,积分得到:

$$\Omega R \cdot 2\pi R = \omega \cdot \pi R^2$$

即:

$$\omega=2\Omega$$

涡度是局地旋转角速度的两倍。

1.10 potential flow 势场

 $\overset{\omega}{\exists}\vec{\omega}=0$:

$$\vec{u} = \vec{\nabla} \psi$$

1.11 算符:梯度,散度和旋度

Inviscid fluid

2.1 Euler equation

momentum equation:

$$\frac{\partial}{\partial t} \int \rho u_i dV = -\oint \rho u_i (\boldsymbol{u} \cdot \boldsymbol{n}) ds - \oint p n_i d\boldsymbol{s} + \int \rho f_i dV$$

$$= -\int \boldsymbol{\nabla} (\rho u_i \boldsymbol{u}) dV - \int \frac{\partial p}{\partial x_i} dV + \int \rho f_i dV$$
(2.1)

这个式子运用了雷诺输运方程. 左边的式子是一个流体微团在 u_i 方向上的动量随着时间的变化率。计算可知左边第一项为动量的量纲。右边第一项为流体的流出率,表征流体微团因为流进流出导致的动量变化。 $(u\cdot n)ds$ 表示在一个很小的面元ds上的流体速度u在其法向上的大小,表征流体微团在方向n的扩散速度,可以表征流体的动量守恒.

有:

$$\frac{\partial}{\partial t}(\rho u_i) + \nabla(\rho u_i \boldsymbol{u}) = -\frac{\partial p}{\partial x_i} + \rho f_i$$

其中:

$$LHS = \rho \frac{\partial u_i}{\partial t} + u_i \frac{\partial \rho}{\partial t} + \rho (\boldsymbol{u} \cdot \boldsymbol{\nabla}) u_i + u_i \boldsymbol{\nabla} (\rho \boldsymbol{u})$$
$$= \rho \left(\frac{\partial u_i}{\partial t} + (\boldsymbol{u} \cdot \boldsymbol{\nabla}) u_i \right)$$
$$= \rho \frac{du_i}{dt}$$

对所有分量:

$$\frac{d\boldsymbol{u}}{dt} = -\frac{1}{\rho}\boldsymbol{\nabla}p + \boldsymbol{f}$$

2.2 Close the equation

流体方程的参数: $\rho \vec{u} p$ 一共五个参数。

连续方程:

$$\frac{d\rho}{dt} + \rho(\vec{\nabla} \cdot \vec{u}) = 0$$

欧拉方程:

$$\frac{d\boldsymbol{u}}{dt} = -\frac{1}{\rho}\boldsymbol{\nabla}p + \boldsymbol{f}$$

$$p = \rho RT$$

2.3 Bernoulli theorem

steady flow : $\partial u/\partial t = 0, \rho = \text{const.}$

$$\vec{u} \cdot \vec{\nabla} u = -\frac{1}{\rho} \vec{\nabla} p + \vec{g} \tag{2.2}$$

$$\vec{u} \cdot \vec{\nabla} u = \vec{\nabla} (u^2/2) - \vec{u} \times \vec{\omega} \tag{2.3}$$

$$-\nabla(p/\rho + u^2/2 + gz) + u \times \omega = 0 \tag{2.4}$$

 $i \exists H = p/\rho + u^2/2 + gz:$

$$-\nabla H + u \times \omega = 0$$

两边同时乘u:

$$u \cdot \nabla H = 0 \partial H / \partial s = 0$$

因此

$$H = \frac{u^2}{2} + \frac{p}{\rho} + gz = const$$

2.4 Vorticity equation

$$\frac{\partial u}{\partial t} = -\nabla H + u \times \omega$$

$$\frac{\partial \omega}{\partial t} = \nabla \times (u \times \omega) = (\omega \cdot \nabla)u - (u \cdot \nabla)\omega \frac{d\omega}{dt} = (\omega \cdot \nabla)u$$

涡度形变:

$$(\omega \cdot \nabla)u \simeq \omega_z \frac{\partial u_z}{\partial z}$$

2.5 Kelvin theorems

$$\Gamma = \oint u \cdot dl = \int \omega \cdot ds$$

$$\frac{d\Gamma}{dt} = \frac{d}{dt} \int_{s(t)} \omega \cdot ds = \int_{s(t)} \left[\frac{\partial \omega}{\partial t} - \nabla \times (u \times \omega) \right] ds = 0$$

$$\frac{d}{dt}(d\vec{l}) = \vec{u}(\vec{x} + d\vec{l}) - \vec{u}(\vec{x}) = (d\vec{l} \cdot \nabla)\vec{u}(\vec{x})$$

$$\frac{d\omega}{dt} = (\omega \cdot \nabla)u$$

如果 ω 方向与l方向平行, 那之后将会一直平行. 开尔文定理:

$$\frac{d\Gamma}{dt} = 0$$

也就是说涡度强度不随时间变化,或者说无外力下无旋流体不可能获得涡度.

Viscous fluid

3.1 Velocity decomposition

$$d\vec{u} = (d\vec{r} \cdot \vec{\nabla})\vec{u}$$

$$du_i = dx \frac{\partial u_i}{\partial x} + dy \frac{\partial u_i}{\partial y} + dz \frac{\partial u_i}{\partial z}$$
(3.1)

$$=\sum_{i=1}^{3} \frac{\partial u_i}{\partial x_j} dx_j \tag{3.2}$$

$$= \frac{\partial u_i}{\partial x_j} dx_j \tag{3.3}$$

Einstein summation 爱因斯坦求和约定:相同指标求和

一个张量可以写成对称部分和反称部分

$$du_i = \frac{\partial u_i}{\partial x_j} dx_j \tag{3.4}$$

$$= \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) dx_j + \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} - \frac{\partial u_j}{\partial x_i} \right) dx_j \tag{3.5}$$

其中对称部分和反称部分记为:

$$e_{ij} = \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$$

$$\xi_{ij} = \frac{1}{2} \left(\frac{\partial u_i}{\partial x_j} - \frac{\partial u_j}{\partial x_i} \right)$$

 e_{ij} : strain tensor, symmetric about i, j. (对称的)

 ξ_{ij} : rotation tensor, anti-symmetric about i, j. (反称的) 我们发现这刚好和涡度对应上.

因此:

$$du_i = e_{ij}dx_j + \xi_{ij}dx_j$$

对法向形变求和:

$$e_{ii} = e_{xx} + e_{yy} + e_{zz}$$

$$= \frac{1}{2} \left(\frac{\partial u_x}{\partial x} + \frac{\partial u_x}{\partial x} \right) + \cdots$$

$$= \vec{\nabla} \cdot \vec{u}$$

形变张量的主对角元是速度散度。散度实际上就是面元的膨胀率. 除了法向形变, 还有剪切形变. 公式[$\mathbf{3.1}$]中当 $i \neq j$ 时为剪切形变. 表征度量流体畸变程度的物理量.

 ξ_{ij} 表征旋转量,是反称的。

$$\xi_{ij} = \frac{1}{2} \begin{bmatrix} 0 & -\omega_z & \omega_y \\ \omega_z & 0 & -\omega_x \\ -\omega_y & \omega_x & 0 \end{bmatrix}$$

3.2 Constitutive relation

 σ_{ij} : stress tensor; 从无粘流体到粘性流体:

$$\oint -p\vec{n}\cdot ds \longrightarrow \oint \sigma \cdot \hat{n} \ ds$$

而 σ_{ij} 分解为:

$$\sigma_{ij} = -p\delta_{ij} + \tau_{ij}$$

应力张量由压力和剪切组成。对于无粘流体,剪切不起作用:

$$(\sigma \cdot \hat{n})_i = -p\delta_{ij}n_j = -pn_i$$

对于粘性流体:

$$\frac{\mathrm{d}u}{\mathrm{d}t} = -\frac{1}{\rho}\nabla p + \frac{1}{\rho}\nabla \tau + f$$

其中, 广义牛顿粘性假设告诉我们:

$$\tau_{ij} = 2\mu e_{ij} + \mu'(\nabla \cdot u)\delta_{ij}\mu' = -\frac{2}{3}\mu$$

3.3 Navier-Stokes Equation

$$\frac{\mathrm{d}\vec{u}}{\mathrm{d}t} = -\frac{1}{\rho}\nabla p + \nu\nabla^2 u + (\nu + \nu')\nabla(\nabla \cdot u) + f$$

其中, $\nu = \frac{\mu}{\rho}$ 为运动学粘性系数, 不可压缩流体:

$$\frac{\mathrm{d}\vec{u}}{\mathrm{d}t} = -\frac{1}{\rho}\nabla p + \nu\nabla^2 u + f$$

对于广义牛顿粘性假设的流体:

$$\frac{\mathrm{d}\vec{u}}{\mathrm{d}t} = -\frac{1}{\rho}\nabla p + \nu\nabla^2 \vec{u} + \frac{1}{3}\nu\nabla(\nabla \cdot \vec{u}) + \vec{f}$$

3.4 Boundary Condition

1. fixed boundary 对于固定边界, 法向速度分量为: $\vec{u} \cdot \hat{n}$; 切向速度分量为: $\vec{u} - \vec{u} \cdot \hat{n}$. 对于无粘流体:

$$\vec{u} \cdot \hat{n} = 0$$

对于粘性流体,流体固定于边界时:

$$\vec{u} = 0$$
 and
$$\begin{cases} \vec{u} \cdot \hat{n} = 0 \\ \vec{u} - \vec{u} \cdot \hat{n} = 0 \end{cases}$$

流体对边界没有应力时:

$$\begin{cases} \vec{u} \cdot \hat{n} = 0 \\ \frac{\partial}{\partial n} (\vec{u} - \vec{u} \cdot \hat{n}) = 0 \end{cases}$$
 (3.6)

2. free boundary 运动学边界:

$$u_z = \frac{d\eta}{dt} = \frac{\partial \eta}{\partial t} + u_x \frac{\partial \eta}{\partial x} + u_y \frac{\partial \eta}{\partial y}$$

动力学边界:

$$p_- = p_+$$

即在边界上得压力相等。

3.5 Normalisation

无量纲化方程

对于无外力的不可压缩的N-S方程:

$$\frac{\partial \vec{u}}{\partial t} + u \cdot \nabla u = -\frac{1}{\rho} \nabla p + \nu \nabla^2 u$$

距离尺度: l_0 ; 时间尺度 t_0 ; 速度尺度: l_0/t_0 ; 压力尺度: $(l_0/t_0)^2\rho$. 把尺度因子代入[3.5]:

$$\frac{l_0/t_0}{t_0} \frac{\partial \tilde{u}}{\partial \tilde{t}} + \frac{\left(l_0/t_0\right)^2}{l_0} \tilde{u} \cdot \tilde{\nabla} \tilde{u} = -\frac{1}{\rho} \frac{\left(l_0/t_0\right)^2 \rho}{l_0} \tilde{\nabla} \tilde{p} + \nu \frac{l_0/t_0}{\ell_0^2} \tilde{\nabla}^2 \tilde{u}$$

整理得:

$$\frac{\partial \tilde{u}}{\partial \tilde{t}} + \tilde{u} \cdot \nabla \tilde{u} = -\tilde{\nabla} \tilde{p} + \frac{\nu t_0}{\ell_0^2} \tilde{\nabla}^2 \tilde{u}$$

其中:

$$\frac{\nu t_0}{l_0^2} = \frac{\nu}{l_0 u_0} = \frac{1}{R_e}$$

 R_e 称为雷诺数。[3.5]公式则化简为:

$$\frac{\partial \tilde{u}}{\partial \tilde{t}} + \tilde{u} \cdot \nabla \tilde{u} = -\tilde{\nabla} \tilde{p} + \frac{1}{R_e} \tilde{\nabla}^2 \tilde{u}$$

从这个方程我们可以这样看:

$$R_e \simeq \frac{|u \cdot \nabla u|}{|\nu \nabla^2 u|} \simeq \frac{u_0^2/l_0}{\nu u_0/l_0^2} = \frac{l_0 u_0}{\nu}$$

雷诺数实际上表征惯性力和粘滞力之比.

3.6 Kinetic Energy Equation

u点乘不可压缩流体的N-S方程,并对该流体微元体积分:

$$\frac{\mathrm{d}}{\mathrm{d}t} \int \frac{u^2}{2} \mathrm{d}V = -\int u \cdot (u \cdot \nabla u) dV - \frac{1}{\rho} \int u \cdot \nabla p \mathrm{d}V + \int u \cdot (\nabla \cdot \tau) \mathrm{d}V$$

由于:

$$\int u \cdot (u \cdot \nabla u) dV = \int u_i u_j \frac{\partial u_i}{\partial x_j} dV$$

$$= \int u_j \frac{\partial}{\partial x_j} (\frac{u^2}{2}) dV$$

$$= \int \frac{\partial}{\partial x_j} (u_j u^2 / 2) dV$$

$$= \oint \frac{u^2}{2} u_j n_j dS$$

$$= 0$$
(3.7)

因此:

$$\frac{\mathrm{d}}{\mathrm{d}t} \int \frac{u^2}{2} \mathrm{d}V = -2\nu \int e_{ij} e_{ij} \mathrm{d}V$$

能量耗散与粘性有关。

3.7 Vorticity Equation

对不可压缩的N-S方程[3.3],两边取速度旋度:

$$\frac{\partial \omega}{\partial t} = \nabla \times (u \times \omega) + \nu \nabla^2 \omega$$

因为由粘性项,与Kelvin定理[2.5]不一致,涡线重联。

3.8 Microscopic Physics

有概率密度函数f(r, v, t)

$$N = \int f \mathrm{d}^3 r \mathrm{d}^3 v$$

玻尔兹曼分布律:

$$\frac{\partial f}{\partial t} + v \cdot \nabla f + a \cdot \nabla_v f = C$$

其中

$$\rho(r,t) = \int m f \mathrm{d}^3 v$$

$$\rho < v > = \int m f v \mathrm{d}^3 v$$

公式[3.8]对时间求偏导,得到质量守恒.能量守恒和动量守恒也是这个方式.

Applications

4.1 Pressure Driven Flow

plan posieuille flow

对于稳态流体du/dt=0, 只受到压力,两边平板固定, $p_1>p_2$ 边界条件:

$$u_x = u; u_y = u_z = 0$$

考虑不可压流体:

$$\nabla \cdot u = 0 \longrightarrow \frac{\partial u}{\partial x} = 0 \longrightarrow u(y, z, t) = u(y)$$

流体速度与x,z方向无关,只与y有关. 得到:

$$-\frac{1}{\rho}\frac{\partial p}{\partial x} + \nu \frac{\partial^2 u}{\partial y^2} = 0 - \frac{1}{\rho}\frac{\partial p}{\partial y} = 0 - \frac{1}{\rho}\frac{\partial p}{\partial z} = 0$$

得到:

$$\nu \frac{d^2 u}{dy^2} = \frac{1}{\rho} \frac{dp}{dx}$$

设两平板之间的高度差为h, 解得:

$$u = -\frac{h^2}{2\mu} \frac{dp}{dx} \frac{y}{h} (1 - \frac{y}{h})$$

容易发现速度最大值在y = h/2上.

4.2 Plane Couette Flow

Boundary driven flow

边界条件:

$$u(h) = Uu(0) = 0$$

由于两边压强为0,即dp/dx=0,代入N-S方程:

$$\mu \frac{d^2 u}{du^2} = \frac{dp}{dx} \mu \frac{d^2 u}{du^2} = 0$$

解得:

$$u = U \frac{y}{h}$$

当两边加上压力时:

$$u = U \left[\frac{y}{h} - \frac{h^2}{2\mu} \frac{dp}{dx} \frac{y}{h} (1 - \frac{y}{h}) \right]$$

4.3 Taylor-Couette Flow

在一个稳态不可压缩的流体在旋转的圆柱上,边界条件为:

$$u_R = 0$$
, $u_z = 0$, $u_\theta = u(R)$, $p = p(R)$

在柱坐标上,压力梯度与离心势平衡:

$$\frac{\mathrm{d}p}{\mathrm{d}R} = \rho \frac{u^2}{R}$$

代入柱坐标下的拉普拉斯算子 $\nabla^2 = \frac{\partial^2}{\partial r^2} + \frac{1}{r} \frac{\partial}{\partial r} + \frac{1}{r^2} \frac{\partial}{\partial \theta^2} + \frac{\partial}{\partial z^2}$:

$$\frac{1}{R}\frac{d}{dR}(r\frac{du}{dR}) - \frac{u}{R^2} = 0$$

假设 $u \propto R^{\alpha}$, 当 $\alpha = \pm 1$ 时,

$$u = c_1 R + \frac{c_2}{R}$$

$$\Omega = c_1 + \frac{c_2}{R^2}$$

对于Kepler运动: $\Omega \propto R^{-1.5}$

4.4 Rotating Flow

对于旋转流动,在随动坐标中展开:

$$\frac{du}{dt} \longrightarrow \frac{du}{dt} + \Omega \times (\Omega \times r) + 2\Omega \times u$$

其中右边第二项为离心项, 第三项为科里奥利力项

$$\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \vec{\nabla} \vec{u} = -\frac{1}{\rho} \vec{\nabla} p + \vec{g} - \vec{\Omega} \times (\vec{\Omega} \times \vec{r}) + 2\vec{u} \times \vec{\Omega}$$

其中离心势写为:

$$\vec{\Omega} \times (\vec{\Omega} \times \vec{r}) = \nabla (\frac{1}{2} |\vec{\Omega} \times \vec{r}|^2)$$

方程则变为:

$$\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \vec{\nabla} \vec{u} = -\frac{1}{\rho} \vec{\nabla} p + \vec{g} - \nabla (\frac{1}{2} |\vec{\Omega} \times \vec{r}|^2) + 2\vec{u} \times \vec{\Omega}$$

对比左边的第二项和右边的最后一项

$$R_0 = \frac{|\vec{u} \cdot \vec{\nabla} \vec{u}|}{|2\vec{u} \times \vec{\Omega}|} \simeq \frac{u_0^2/l_0}{2\Omega u_0} = \frac{u_0}{2\Omega l_0} R_0 = \frac{u_0}{2\Omega l_0}$$

也就是惯性力项/科氏力项,这个数称为Rossby数.在地球上,假设空气是不可压缩流体,有势函数:

$$-\frac{1}{\rho}\vec{\nabla}p + \vec{g} - \vec{\Omega} \times (\vec{\Omega} \times \vec{r}) = \vec{\nabla}\psi$$

认为地球大气是准静态的,地球上的 $R_0 << 1$,有du/dt = 0,则有准地转平衡:

$$\nabla \psi + 2u \times \Omega = 0$$

由于势函数是无旋的,我们对方程[4.4]取旋度:

$$\nabla \times (2\vec{u} \times \vec{\Omega}) = 0$$

即:

$$2\vec{\Omega} \cdot \nabla u = 0 \frac{\partial \vec{u}}{\partial l} = 0$$

4.5 Accretion Disk

吸积盘结构:黑洞吸积盘, AGN, 原行星盘, 行星环, 双星系统盘的洛希极限, X射线源

1. 估计重力势能

$$E_p = \frac{GMm}{a}/(mc^2) = \frac{GM}{ac^2}$$

2. 薄盘近似 在柱坐标中: (R, ϕ, z) , 薄盘即为 $h \ll R_d$, 因此物理量为: $(\rho, p, T, v_R, v_\phi, v_z)$. 其中:

$$v_z \simeq 0 v_\phi \gg v_R \frac{\partial}{\partial \phi} = 0$$

也就是说在盘上垂直速度很小, 且以6方向速度为主, 相比起来径向速度非常小. 这里引入引力势函数

$$\Psi = -GM(R^2 + z^2)^{-1/2}$$

在R方向和z方向的引力为:

$$F_R = -\partial \Psi/\partial R = -GM(R^2 + z^2)^{-3/2} \simeq -GMR^{-2}$$

$$F_z = -\partial \Psi / \partial z = -GMz(R^2 + z^2)^{-3/2} \simeq -zGMR^{-3}$$

由于是稳态场, $\partial/\partial t=0$, 在径向和垂直方向的运动方程为:

$$v_R \frac{\partial v_R}{\partial R} - \frac{v_\phi^2}{R} = -\frac{1}{\rho_g} \frac{\partial p}{\partial R} - GM/R^2$$

$$0 = -\frac{1}{\rho} \frac{\partial p}{\partial z} - \frac{GMz}{R^3}$$
(4.1)

对于理想气体状态方程:

$$p = \frac{\rho_g}{m_H} k_B T$$

密度 ρ_q 的分布为一个高斯分布:

$$\rho_g \propto \exp\left(-\frac{GMm_H}{2k_BTR^3}z^2\right)$$

等温过程为:

$$-\frac{1}{\rho_g}\frac{k_BT}{m_H}\frac{d\rho_g}{dz} - \frac{GM}{R^3}z = 0$$

其中:

$$H_z = \left(2k_BTR^3/GMm_H\right)^{1/2}$$
 scale height
$$\rho_g \propto \exp\left(-z^2/H_z^2\right) \tag{4.2}$$

通过垂直方向的运动方程,对压力p的估计为:

$$\frac{1}{\rho_g} \frac{p}{h} \simeq \frac{GMh}{R_d^3} \Rightarrow p \simeq \frac{GM\rho_g h^2}{R_d^3}$$

有了上面的公式, 我们对上面的公式进行量级估计:

$$v_R \ll v_{\varphi}; \quad v_R \frac{\partial v_R}{\partial R} \ll v_{\varphi}^2 / R$$

得到:

$$\left(\frac{1}{\rho_g}\frac{p}{R_d}\right)/\left(\frac{GM}{R_d^2}\right)\simeq \left(\frac{1}{\rho_g}\frac{1}{R_d}\frac{GM\left(gh^2\right)}{R_d^3}\right)/\left(\frac{GM}{R_d^2}\right)=h^2/R_d^2\ll 1$$

因此在薄盘中, 径向压力梯度远远小于重力, 压力梯度不是主要的, 但在厚盘中需要考虑. 离心项写为:

$$\frac{v_{\varphi}^2}{R} \simeq \frac{GM}{R^2} \longrightarrow v_{\varphi} = (GM/R)^{1/2}$$

在开普勒盘中, 角速度为:

$$\Omega = v_{\varphi}/R = (GM/R^3)^{1/2}$$

考虑角动量. 吸积盘面密度为

$$\sum = \int \rho dz$$

连续方程写为:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho u) = 0$$

稳态, 在柱坐标中, r方向的连续方程为:

$$\frac{1}{R}\frac{\partial}{\partial R}(R\rho v_R) = 0$$

因此:

$$\int \frac{1}{R} \frac{\partial}{\partial R} (R\rho v_R) dz = 0$$

由于密度取决于z, 但是速度不取决于z, 改写为面密度:

$$\frac{d}{dR}(R\Sigma v_R) = 0$$

吸积率可以表示为单位面密度成径向速度的和, 因此吸积率为:

$$\dot{m} = -2\pi R \Sigma v_R$$

因此:

$$R\Sigma v_R = -\dot{m}/2\pi$$

计算角动量, 这时候的粘性力是重要的. 单位面积上的角动量为:

$$J = \Sigma R^2 \Omega$$

单位面积的粘滞力矩为:

$$\frac{\partial}{\partial t} \left(\sum R^2 \Omega \right) + \frac{1}{R} \frac{\partial}{\partial R} \left(R \cdot \sum R^2 \Omega \cdot V_R \right) = G \qquad \text{(viscous torgue per unit area)}$$

力矩等于 $\partial L/\partial t$; 第二项为散度算子在柱坐标中的展开. 其中角动量守恒:

$$\frac{\partial}{\partial t} \left(\sum R^2 \Omega \right) = 0$$

因此力矩为:

$$G \cdot 2\pi R \cdot dR = d\Gamma \quad (\Gamma : \text{ viscous torgue})$$

$$G = \frac{1}{2\pi R} \frac{d\Gamma}{dR}$$
(4.3)

有速度差就有速度剪切. 因此:

$$\frac{dv_{\varphi}}{dR} = \frac{d}{dR}(\Omega R) = \Omega + R\frac{d\Omega}{dR}$$

viscous force =
$$\mu R \frac{d\Omega}{dR}$$

因此粘滞力矩可以写为:

$$\Gamma(R) = \iint \left(\mu R \frac{d\Omega}{dR}\right) \cdot (Rd\varphi dz) \cdot R$$

$$= R^{3} \frac{d\Omega}{dR} \cdot 2\pi \cdot \int \rho \nu dz$$

$$= 2\pi \nu \sum R^{3} \frac{d\Omega}{dR}$$
(4.4)

因此力矩可以写为:

$$G(R) = \frac{1}{R} \frac{d}{dR} \left(\nu \sum R^3 d\Omega / dR \right)$$

$$\frac{1}{R}\frac{d}{dR}\left(\sum R^3\Omega V_R\right) = \frac{1}{R}\frac{d}{dR}\left(\nu\Sigma R^3\frac{d\Omega}{dR}\right)$$

$$\sum R^3 \Omega v_R = \nu \sum R^3 \frac{d\Omega}{dR} + C$$

当没有剪切的时候, $\frac{d\Omega}{dR} = 0$, 确定系数C为:

$$C = \sum R_*^3 \Omega_* v_{R_*} = -\dot{m}/2\pi \cdot R_*^2 \Omega_*$$

代回去,得到:

$$-\dot{m}/2\pi \cdot R^2\Omega = \nu \sum R^3 \frac{d\Omega}{dR} - \dot{m}/2\pi \cdot R_*^2\Omega_*$$

$$\Omega = \left(GM/R^3\right)^{1/2}$$

解得:

$$\nu \Sigma = \frac{\dot{m}}{3\pi} \left(1 - \left(\frac{R_*}{R} \right)^{1/2} \right)$$

因此粘滞系数正比于吸积率. 接下来估计能量和角动量的变化率.

能量变化率为:

$$\begin{split} \frac{dE}{dt} &= -\int_{\text{(per unit area)}} \mu (Rd\Omega/dR)^2 dz \\ &= -\nu \sum_{k} R^2 (d\Omega/dR)^2 \\ &= -\frac{3GM\dot{m}}{4\pi R^3} \left(1 - \left(\frac{R_*}{R}\right)^{1/2}\right) \end{split} \tag{4.5}$$

角动量为:

$$\begin{split} L &= \int_{R_*}^{\infty} -\frac{dE}{dt} 2\pi R dR \\ &= GM\dot{m}/2R_* \\ &= \text{ half of } \frac{GM}{R_*} \dot{m} \end{split} \tag{4.6}$$

从上面的分析我们知道

$$L \longrightarrow \dot{m} \longrightarrow \nu$$

因此吸积率越大, 粘滞力越大, 角动量越大, 盘越不稳定.

3. 薄盘演化 综合上面的式子, 得到diffusion equation:

$$\frac{\partial \Sigma}{\partial t} = \frac{3}{R} \frac{\partial}{\partial R} \left(R^{1/2} \frac{\partial}{\partial R} \left(\nu \Sigma R^{1/2} \right) \right)$$

$$\sum \propto \exp\left(-\frac{R^2}{12\nu t}\right)$$

这就是不稳定薄盘的演化方程.

4.6 Stellar Structure

compressible fluid and spherical symmetric

$$-\vec{\nabla}p + \rho\vec{g} = 0 - \frac{dp}{dr} - \rho g = 0$$

重力g表示为:

$$g = \frac{G}{r^2} \int_0^r 4\pi r'^2 \rho dr'$$

也就是说对球壳dr'积分,得到重力g.

在球坐标中表示:

$$\frac{dp}{dr} + \rho \frac{G}{r^2} \int_0^r 4\pi r'^2 dr' = 0$$

对r求导:

$$\frac{d}{dr}\left(\frac{dp}{dr}r^2\frac{1}{\rho}\right) + 4\pi G\rho r^2 = 0$$

整理得:

$$\frac{1}{r^2}\frac{d}{dr}\left(r^2\frac{1}{\rho}\frac{dp}{dr}\right) = -4\pi G\rho$$

another approach

$$-\vec{\nabla}p + \rho\vec{g} = 0 \quad , \quad \vec{g} = -\vec{\nabla}\Phi$$

对两边求微分,有:

$$\vec{\nabla}(\frac{1}{\rho}\vec{\nabla}p) = \vec{\nabla}\vec{g} = -\vec{\nabla}^2\Phi$$

引力势能与引力的关系:

$$\oint \vec{g} d\vec{s} = \int -\nabla^2 \phi dV$$

积分得到:

$$\frac{G\rho V}{r^2} 4\pi r^2 = \nabla^2 \phi V$$

得到:

$$\nabla^2 \phi = 4\pi G \rho$$

polytropic gas:

多方关系:

$$p = \kappa \rho^{\gamma} = \kappa \rho^{(1+1/n)}$$

其中n为多方指数。对于每单位体积的能量:

$$u = np$$

对于恒星对流区,以H原子为主,因此:

$$\gamma = 5/3; \quad n = 1.5$$

对于辐射区以光子为主:

$$\gamma = 4/3; \quad n = 3$$

边界条件为: isothermal边界

$$p \propto \rho; \quad n \to \infty$$

incompressible边界

$$p \propto \rho, \quad n \to 0$$

方程为:

$$K\left(1+\frac{1}{n}\right)\frac{1}{r^2}\frac{d}{dr}\left[r^2\rho^{(-1+1/n)}\frac{d\rho}{dr}\right] + 4\pi G\rho = 0$$

设:

$$a = \frac{4\pi G}{K(1+n)} \rho_c^{1-1/n} z = aru = (\rho/\rho_c)^{1/n}$$

得到:

$$\frac{1}{z^2}\frac{d}{dz}(z^2\frac{du}{dz}) + u^n = 0$$

4.7 Boundary Layer

边界层是把靠近边界的地区考虑粘滞力,

$$\partial u_{x}/\partial x + \partial u_{y}/\partial y = 0$$

$$u_{x}\frac{\partial u_{x}}{\partial x} + u_{y}\frac{\partial u_{x}}{\partial y} = -\frac{1}{\rho}\frac{\partial p}{\partial x} + \nu\left(\frac{\partial^{2}U_{x}}{\partial x^{2}} + \frac{\partial^{2}u_{x}}{\partial y^{2}}\right)$$

$$u_{x}\frac{\partial u_{y}}{\partial x} + u_{y}\frac{\partial u_{y}}{\partial y} = -\frac{1}{\rho}\frac{\partial p}{\partial y} + \nu\left(\frac{\partial^{2}u_{y}}{\partial x^{2}} + \frac{\partial^{2}U_{y}}{\partial y^{2}}\right)$$

$$(4.7)$$

设边界层为 δ , 平板长度为l, 化简得到:

$$\delta \simeq \left(\frac{\nu l}{u_x}\right)^{1/2} \simeq Re^{-1/2}l$$

当雷诺数越大时, 边界层越小. 同时也可以知道:

$$Re^{-1/2} \ll 1$$

对 u_y 方向也同理. 最后化简得到的方程为:

$$\begin{cases} \frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} = 0\\ u_x \frac{\partial u_x}{\partial x} + u_y \frac{\partial u_x}{\partial y} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \nu \frac{\partial^2 u_x}{\partial y^2}\\ 0 = -\frac{1}{\rho} \frac{\partial p}{\partial y} \end{cases}$$

Prandtl's equation

Gas Dynamics

5.1 thermodynamics

energy per unit volume:

$$u = np$$

pdytropic index:

$$\gamma = 1 + \frac{1}{n}$$

energy per unit mass

$$\varepsilon = u/p = c_V T$$

 c_V : specific heat per unit mass

$$p = Nk_BT$$

N: number desity, $\rho=m_pN$

$$p = \frac{\rho}{m_p} k_B T = \rho \frac{k_B N_A}{m_p N_A} T$$

there: $k_B N_A = R = 8.314$; $m_p N_A = \mu$ mass per mole

$$p = \rho \frac{R}{\mu} T$$

$$c_V T = u/\rho = np/\rho = n\frac{R}{\mu}T$$

有:

$$c_V = nR/\mu = \frac{1}{\gamma - 1} \frac{R}{\mu}$$

带入压力方程:

$$p = (\gamma - 1)\rho c_V T$$

the first law of thermodynamics:

$$d\varepsilon = Tds + pd(\frac{1}{\rho})$$

带入p:

$$c_V dT = T ds + (\gamma - 1)c_V T \frac{1}{\rho} d\rho d \ln T = \frac{1}{c_V} ds + (\gamma - 1)d \ln \rho$$

得到:

$$ds = c_V d \ln \frac{T}{\rho^{\gamma}} s = c_V \ln \frac{p}{\rho^{\gamma}} + s_0$$

adiabatic: ds = 0

5.2 Sound Wave

由于

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho u) = 0 \frac{\partial u}{\partial t} + u \cdot \nabla = -\frac{1}{\rho} \nabla p$$

微扰:

$$\rho = \rho_0 + \rho_1$$
 $p = p_0 + p_1$ $u = 0 + u$

代入线性微扰:

$$\frac{\partial \rho_1}{\partial t} + \rho_0 \nabla \cdot u_1 = 0 \frac{\partial u}{\partial t} = -\frac{1}{\rho_0} \nabla p_1$$

解得:

$$\frac{\partial^2 \rho_1}{\partial t^2} - \nabla^2 p_1 = 0$$

ਪੋਟੀ $dp/d
ho=C_s^2,\, p_1=dp/d
ho imes
ho_1$:

$$\frac{\mathrm{d}^2 \rho_1}{\mathrm{d}t^2} - C_s^2 \nabla^2 \rho_1 = 0$$

代入波动解:

$$\rho_1 = \hat{\rho}e^{i(kx - \omega t)}$$

解得:

$$-\omega^2 + C_s^2 k^2 = 0 \quad \to \quad \omega = \pm C_s k$$

相速度:

$$\vec{c}_p = \frac{\omega}{k} = \pm c_s \hat{k}$$

群速度:

$$\vec{c}_g = \vec{\nabla}_k \omega = \pm c_s \hat{k}$$

将波动解代入[5.2]:

$$-\mathrm{i}\omega\hat{u} = \frac{1}{\rho_0}\mathrm{i}\vec{k}\hat{p}$$

则: $\hat{u} = \hat{p}$,两个方向平行

多方过程:

$$p=\kappa\rho^{\gamma}$$

得到:

$$c_s = (\gamma \frac{p}{\rho})^{1/2} = (\gamma \frac{R}{\mu} T)^{1/2}$$

当 $\gamma = 1$, 即为等温过程时:

$$c_s = (\frac{p}{\rho})^{1/2}$$

- 5.3 Shock
- 5.4 Blast Wave
- 5.5 Galatic Jet
- 5.6 Sphere Wind and Accretion

Waves

6.1 water wave

unperturbed:

$$-\nabla p_0 + \rho g = 0$$

浮力定律:

$$p_0 = p_a + \rho g(h - y)$$

微扰欧拉方程:

$$\frac{\partial u_i}{\partial t} + u_i \nabla u_i = -\frac{1}{\rho} \nabla p_1$$

其中 $u_i \nabla u_i = 0$ (微扰), 有

$$\frac{\partial u_i}{\partial t} = -\frac{1}{\rho} \nabla p_1$$

取旋:

$$\frac{\partial \omega_1}{\partial t} = 0$$

速度分解: $u_1 = v_1 + v_2$,

$$\nabla \times v_1 = 0$$
 $\frac{\partial v_1}{\partial t} \neq 0 \nabla \times v_2 = 0$ $\frac{\partial v_2}{\partial t} = 0$ (no waves)

对于上面的公式, v_1 旋度为0,有波动行为,有速度势 ϕ

$$v_1 = \nabla \phi$$

$$\nabla v_1 = 0$$

$$\nabla^2 \phi = 0$$

kinematic B. C. : $\partial \phi/\partial y = d\xi/dt = \partial \xi/\partial t + \partial \phi/\partial x \cdot \partial \xi/\partial x$

dynamic B. C. : $p_0 + p_1 = p_a$ and $p_a = \rho g \xi$

At y = h:

$$\frac{\partial^2 \phi}{\partial t} + g \frac{\partial \phi}{\partial y} = 0$$

波动解的形式为: $\phi = \hat{\phi} \exp i(kx - \omega t)$:

解得

$$\omega^2 = gk \tanh(kh)$$

当为深水波时: $kh \gg 1$, 有:

$$\omega^{2} = gk$$

$$c_{p} = \frac{\omega}{k} = \sqrt{g/k}$$

$$c_{g} = \frac{\partial \omega}{\partial k} = \frac{1}{2}\sqrt{g/k}$$
(6.1)

其中, c_p 为相速度, c_g 为群速度. 记波动速度为c, 一个波动周期为 γ , 则有: $\omega=2\pi/\gamma=kc_p$; 一个周期内有多少波动周期称为波数, 波数为: $k=2\pi/\lambda$; 波的移动速度又称为相速度, 而当不同频率的波叠加时的波包移动速度即为群速度 c_g .

6.2 Internal wave

1. Internal Gravity wave (g mode) 流体密度是线性的:

$$\rho_0 = \bar{\rho} + \frac{d\rho_0}{dz}z$$

对于线性部分的连续方程:

$$\frac{d\rho_0}{dt} = 0; \quad \nabla \cdot u = 0$$

$$\frac{\partial \rho}{\partial t} + u \cdot \nabla \rho + \rho \cdot \nabla u = 0$$

欧拉方程:

$$\bar{\rho}\frac{du}{dt} = -\nabla p + \rho g$$

对于扰动部分展开,我们注意到 $u_1 \cdot \nabla u_1$ 为二阶微扰,因此连续方程和欧拉方程可写为:

$$\frac{\partial \rho_1}{\partial t} + u_z \frac{d\rho_0}{dz} = 0, \qquad \nabla \cdot u = 0 \bar{\rho} \frac{\partial u_1}{\partial t} = -\nabla p_1 + \rho_1 g, \quad -\nabla p_0 + \rho_0 g = 0$$

联立得到方程:

$$\frac{\partial^2}{\partial t^2} \nabla^2 u_z^2 + N^2 \nabla_\perp^2 u_z = 0 \\ N^2 = -\frac{g}{\bar{\rho}} \frac{d\rho_0}{dz}$$

其中 $\nabla_{\perp} = \frac{\partial}{\partial x} + \frac{\partial}{\partial y}$, 设波动解为:

$$u_z = \hat{u}_z \exp i(k \cdot x - \omega t)$$

解得:

$$\omega = \pm N \frac{k_{\perp}}{k}$$

其中 $k^2=k_x^2+k_y^2+k_z^2=k_\perp^2+k_z^2$

群速度和相速度为:

$$c_p = \frac{\omega}{k}\hat{k}c_g = \nabla_k\omega c_g \cdot c_p = 0$$

也就是说群速度垂直于相速度.

2. Internal Inertial Wave

6.3 Planetary Wave

0. Shallow Water h << L and $\nabla u = 0$, in other word : $u_z/u_x \approx h/L << 1$ so u_x and u_y is independent of z

对于大气,海洋:浅水波,有 $h \ll L$,对于不可压流体:

$$\nabla \cdot u = 0$$

$$\frac{u_z}{u_x} \simeq \frac{h}{L} \ll 1$$

$$h = h_0 + h_1(x, y, t)$$

假设 u_x, u_y 与z无关, 非扰动项:

$$p_0 = p_a + \rho g(h_0 - z)$$

扰动项有:

$$p_1 = \rho g h_1$$

$$\frac{\partial p_1}{\partial x} = \rho g \frac{\partial h_1}{\partial x} \frac{\partial p_1}{\partial y} = \rho g \frac{\partial h_1}{\partial y}$$

因此

$$u_z = \frac{dh}{dt} = \frac{\partial h}{\partial t} + u_x \frac{\partial h}{\partial x} + u_y \frac{\partial h}{\partial y}$$

由于与z无关,有:

$$\frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} = -\frac{\partial u_z}{\partial z}$$

两边对z积分,得到:

$$u_z = -h\nabla_{//}u$$

$$\frac{dh}{dt} - u_z = 0 \frac{dh}{dt} + h\nabla \cdot u_{//} = 0$$

展开,得到:

$$\frac{\partial h}{\partial t} + \nabla \cdot (hu_{//}) = 0$$

1. Kelvin Wave mean flow : $u_{x_0} = U$

$$\begin{cases} \frac{\partial u_x}{\partial t} + U \frac{\partial u_x}{\partial x} = -\frac{1}{\rho} \frac{\partial p_1}{\partial x} = -g \frac{\partial h_1}{\partial x} \\ \frac{\partial u_y}{\partial t} + U \frac{\partial u_y}{\partial x} = -\frac{1}{\rho} \frac{\partial p_1}{\partial y} = -g \frac{\partial h_1}{\partial y} \\ \frac{\partial h_1}{\partial t} + h_0 \left(\frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} \right) + U \frac{\partial h_1}{\partial x} = 0 \end{cases}$$

有形式解:

$$u_{x} = \hat{u}_{x}e^{i(k_{x}x + k_{y}y - \omega t)}u_{y} = \hat{u}_{y}e^{i(k_{x}x + k_{y}y - \omega t)}h_{1} = \hat{h}e^{i(k_{x}k + k_{y}y - \omega t)}$$

即:

$$A \begin{pmatrix} \hat{u}_x \\ \hat{u}_y \\ \hat{u}_z \end{pmatrix} = 0$$

$$A = \begin{pmatrix} -i\omega + iUk_x & 0 & igk_x \\ 0 & -i\omega + i0k_x & igk_y \\ ih_0k_x & ih_0k_y & -i\omega + ivk_x \end{pmatrix}$$

当该方程组有解时, det(A) = 0

解得:

$$\omega = Uk_x \pm (gh_0(k_x^2 + k_y^2))^{1/2}$$

2. Poincare Wave

$$\vec{\Omega} = \Omega[\cos(\theta \hat{e}_y) + \sin(\theta \hat{e}_z)]$$

and

$$2u \times \Omega = \begin{vmatrix} \hat{e}_x & \hat{e}_y & \hat{e}_z \\ 2u_x & 2u_y & 2u_z \\ 0 & \Omega \cos \theta & \Omega \sin \theta \end{vmatrix} = \begin{bmatrix} 2\Omega \sin(\theta u_y) - 2\Omega \cos(\theta u_z) \\ -2\Omega \sin(\theta u_x) \\ 2\Omega \cos(\theta u_x) \end{bmatrix}$$
(6.2)

Neglect $\cos \theta$ terms on high latitude region :

$$2u \times \Omega = 2\Omega \sin \theta_y \hat{e}_x - 2\Omega \sin \theta u_x \hat{e}_y$$

= $f u_y \hat{e}_x - f u_x \hat{e}_y$ (6.3)

Corilis number: $f = 2\Omega \sin \theta$

$$\begin{cases} \frac{\partial u_x}{\partial t} = -g \frac{\partial h_1}{\partial x} + f u_y \\ \frac{\partial u_y}{\partial t} = -g \frac{\partial h_1}{\partial x} - f u_x \\ \frac{\partial h_1}{\partial t} + h_0 \left(\frac{\partial u_x}{\partial x} + \frac{\partial u_y}{\partial y} \right) = 0 \end{cases}$$

we assume the form of the solution:

$$\begin{cases} u_x = \hat{u}_x \exp i(k_x x + k_y y - \omega t) \\ u_y = \hat{u}_y \exp i(k_x x + k_y y - \omega t) \\ h_1 = \hat{h}_1 \exp i(k_x x + k_y y - \omega t) \end{cases}$$

带入上述方程组:

$$A = \begin{bmatrix} -i\omega & -f & igk_x \\ f & -i\omega & igk_y \\ ih_0k_x & ih_0k_y & -i\omega \end{bmatrix}$$

即:

$$A \begin{bmatrix} \hat{u}_x \\ \hat{u}_y \\ \hat{h}_1 \end{bmatrix} = 0$$

由于det(A) = 0,有:

$$\begin{vmatrix} -i\omega & -f & igk_x \\ f & -i\omega & igk_y \\ ih_0k_x & ih_0k_y & -i\omega \end{vmatrix}$$

$$= -i\omega \begin{vmatrix} -i\omega & igk_y \\ ih_0k_y & -i\omega \end{vmatrix} + f \begin{vmatrix} f & igk_y \\ ih_0k_x & -i\omega \end{vmatrix} + igk_x \begin{vmatrix} f & -i\omega \\ ih_0k_x & ih_0k_y \end{vmatrix}$$

$$= i\omega^3 - i\omega f^2 - i\omega gh_0 k$$

$$= 0$$

$$(6.4)$$

解得:

$$\omega^2 = f^2 + gh_0k$$

即:

$$\omega = \pm \sqrt{f^2 + gh_0k}$$

3. Rossby Wave 科氏力:

$$f=2\Omega\sin\theta$$

and

$$Rd\theta = dy$$

 β 平面近似:

$$f = f_0 + \beta y$$

其中:

$$f_0 = 2\Omega \sin \theta \beta = \partial f / \partial y$$

综合上以上式子,

$$\beta = 2\Omega\cos\theta/R$$

因此

$$\frac{\partial U_x}{\partial t} = -g \frac{\partial h_1}{\partial x} + f u_y : \partial/\partial y
\omega^3 - (g h_0 k^2 + f_0^2) \omega - \beta g h_0 k_x = 0$$
(6.5)

对于运动缓慢的大气长波:

$$\omega \simeq -\frac{\beta k_x}{k^2 + f_0^2/gh_0}$$

x方向的相速度:

$$c_{px} = \frac{\omega}{k} \frac{k_x}{k} < 0$$

因此Rossy wave是向西方向传播的.

Instability

7.1 Gravitational Instability

微扰方程:

$$\begin{cases} \frac{\partial \rho_1}{\partial t} + \rho_0 \nabla u_1 = 0\\ \rho_0 \frac{\partial u_1}{\partial t} = -\nabla p_1 - \rho_0 \nabla \phi_1\\ \nabla^2 \phi = 4\pi G \rho_1 \end{cases}$$
(7.1)

代入
$$p_1=c_s^2\rho_1$$
以及 $c_s^2=dp/d\rho$

- 7.2 Rotating Flow Instability
- 7.3 Convective Instability

Turbulence

8.1 Concepts from phenomenology

动量方程:

$$\frac{\partial \vec{u}}{\partial t} + u \cdot \nabla u = -\frac{1}{\rho} \nabla p + \nu \nabla^2 u$$

涡度方程:

$$\frac{\partial \vec{\omega}}{\partial t} + u \cdot \nabla \omega = \omega \cdot \nabla u + \nu \nabla^2 \omega$$

其中 $\omega \cdot \nabla u$ 为拉伸项。

8.2 Reynolds Average and Turbulent Viscosity

设:

$$u = \bar{u} + u'$$

动量方程:

$$\frac{\partial u_i}{\partial t} + u_j \frac{\partial u_i}{\partial x_j} = -\frac{1}{\rho} \frac{\partial p}{\partial x_i} + \nu \frac{\partial}{\partial x_j} (\frac{\partial u_i}{\partial x_j})$$

代入微扰假设:

$$\frac{\partial \bar{u}_i}{\partial t} + \bar{u}_j \frac{\partial \bar{u}_i}{\partial x_j} = \overline{-u_j' \frac{\partial u_i'}{\partial x_j}} - \frac{1}{\rho} \frac{\partial \bar{p}}{\partial x_i} + \nu \frac{\partial}{\partial x_j} (\frac{\partial \bar{u}_i}{\partial x_j})$$

8.3 Cascade and Scaling Law

Largest Scale : l , τ_l , u_l

Any Other Scale : λ , τ_{λ} , u_{λ}

各个尺度下能量传输速率相等:

$$\varepsilon \sim u_l^2/\tau_l \sim u_\lambda^2/\tau_\lambda$$

由于 $\tau_l \simeq u_l/l$, 代入上式,整理得:

$$u_{\lambda}/u_{l} \sim (\lambda/l)^{1/3} \tau_{\lambda}/\tau_{\lambda} \sim (\lambda/l)^{2/3}$$

粘滯系数得量纲为: $[m^2s^{-1}]$, 即 $\nu_l \sim u_l \cdot l$:

$$\nu_{\lambda}/\nu_{l} \sim \frac{u_{\lambda}\lambda}{u_{l}l} \sim (\frac{\lambda}{l})^{4/3}$$

湍流粘滞系数 ν_t 为:

$$u_t \sim rac{\int_0^l
u_\lambda d\lambda}{\int_0^l d\lambda} \sim
u_l$$

若外力强迫是时间周期尺度,产生共振:

$$\nu_t \sim \nu_\lambda|_{\tau_\lambda = T}$$

由于:

$$\nu_{\lambda}/\nu_{l} \sim (\lambda/l)^{4/3} \sim (\tau_{\lambda}/\tau_{l})^{2}$$

当 $\lambda \sim T$:

$$\nu_t/\nu_l \sim (T/\tau_l)^2$$

得出:

$$\nu_t \sim \nu_l (T/\tau_l)^2$$

tidal turbulence $\nu_t << \nu_l$, 对于 smallest scale : $\eta, \tau_{\eta}, u_{\eta}$:

$$\varepsilon \sim \frac{u_l^2}{\tau_l} \sim \frac{u_\eta^2}{\tau_\eta} \sim \nu \tau_\eta^{-2} \quad and \quad [\varepsilon] = m^2 s^{-3}$$

$$\frac{u_{\eta}^2}{\tau_n} \sim \nu \tau_{\eta}^{-2} \to u_{\eta} \cdot \eta / \nu \sim 1$$

$$u_l^2/\tau_l \sim \nu \tau_\eta^{-2} \to \eta/l \sim (u_l \cdot l/\nu)^{-3/4} \sim \text{Re}_l^{-3/4}$$

因此:

$$u_{\eta}/u_{l} \sim \operatorname{Re}_{l}^{-1/4} \tau_{\eta}/\tau_{l} \sim \operatorname{Re}_{l}^{-1/2}$$

Example: 设有云朵: $l \sim 10^3 \text{m}, u_l \sim 1 \text{m/s}, \nu \sim 10^{-5} \text{m}^2 \text{s}^{-1}$

$$\begin{aligned} \text{Re}_l &\sim \frac{u_l \cdot l}{\nu} \sim 1 \times 10^3 / 10^{-5} = 10^8 \\ \eta &\sim \text{Re}_l^{-3/4} \cdot l = (10^8)^{-3/4} \times 10^3 \text{m} = 10^{-3} \text{m} \\ u_{\eta} &\sim \text{Re}_l^{-1/4} \cdot u_l = (10^8)^{-1/4} \times 1 \text{m/s} = 10^{-2} \text{m/s} \end{aligned} \tag{8.1}$$

8.4 Energy Spectrum

Mangeto-hydrodynamics (MHD)

Introduction to MHD

9.1 Equation

在低速流动: u << c

流体参考系和实验室参考系的关系:

$$E' = E + U \times BB' = B$$

法拉第定律:

$$\nabla \times E = \frac{\partial B}{\partial t} \to B_0/t_0 \sim E_0/l_0 \to E_0/U_0$$

$$\mu \epsilon E/t_0 / B/l_0 \sim (u_0/c)^2 << 1$$

位移电流小, 去掉

欧姆定律:

$$j' = \sigma E' = \sigma (E + U \times B)$$

$$\nabla \times B = \mu j$$

$$\partial B/\partial t = -\nabla \times E$$

η: 磁扩散系数

磁感应项+磁扩散项

$$R_m = \frac{\nabla \times (U \times B)}{\eta \nabla^2 B} = \frac{U_0 B / l_0}{\eta B / l_0^2} = u_0 l_0 / \eta$$

不可压缩流体:速度散度=0

 $\eta = 0$: 阿尔文磁冻结定理 $(R_m \to \infty)$

磁力线冻结-->阿尔文波 1942

$$\rho(\frac{\partial u}{\partial t} + u \cdot \nabla u) = -\nabla p$$

洛伦兹项:应力张量,正+剪切

切向投影s

法向投影n-->张力

理想磁流体:没有粘性和磁扩散,

不可压, 微扰

磁力线与群速度方向一样

群速度: v_a 阿尔文速度

u/v_a=M_a: 磁马赫数

可压缩: 磁声波 $c_g = \sqrt{c_s^2 + v_a^2}$