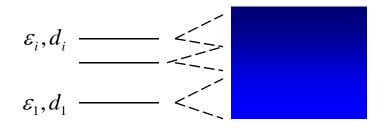
Lecture 20. Continuous Spectrum, the Density of States (Ch. 7), and Equipartition (Ch. 6)

Typically, it's easier to work with the integrals rather than the sums. Thus, whenever we consider an energy range which includes many levels (e.g., when $k_BT >>$ inter-level spacing), and, especially, when we are dealing with **continuous spectra**, we'll replace the sum over a discrete set of energy levels with an integral over a continuum. When doing this, we must replace the **level degeneracy** with a new variable, **the density of states**, $g(\varepsilon)$, which is defined as the number of states per unit energy integral:

$$P(\varepsilon_{i}) = \frac{d_{i}}{Z} \exp\left(-\frac{\varepsilon_{i}}{k_{B}T}\right) \qquad P(\varepsilon) d\varepsilon = \frac{g(\varepsilon)}{Z} \exp\left(-\frac{\varepsilon}{k_{B}T}\right) d\varepsilon$$



The units of $g(\varepsilon)$: (energy)⁻¹

The Density of States

The density of states, $g(\varepsilon)$, is the number of states per unit energy

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Let's consider $\textbf{\textit{G}}(\epsilon)$ - the number of states per unit volume with energy less than ϵ (for a one- and two-dimensional cases, it will be the number of states with energy less than ϵ per unit length or area, respectively). Then the density of states is its derivative:

$$g(\varepsilon) \equiv \frac{dG(\varepsilon)}{d\varepsilon}$$

The number of states per unit volume in the energy range $\epsilon - \epsilon + d\epsilon$:

$$g(\varepsilon)d\varepsilon$$

The probability that one of these states is occupied:

$$\frac{1}{Z}\exp(-\beta\varepsilon)$$

The probability that the particle has an energy between ε and ε +d ε :

$$P(\varepsilon)d\varepsilon = \frac{1}{Z}g(\varepsilon)\exp(-\beta\varepsilon)d\varepsilon$$

Since
$$\int_{0}^{\infty} P(\varepsilon) d\varepsilon = 1$$
 the partition function is

$$Z = \int_{0}^{\infty} g(\varepsilon) \exp(-\beta \varepsilon) d\varepsilon$$

The Density of States for a Single Free Particle in 1D

The energy spectrum for the translational motion of a molecule in free space is continuous. We will use a trick that is common in quantum mechanics: assume that a particle is in a large box (energy quantization) with zero potential energy (total energy= kin. energy). At the end of the calculation, we'll allow the size of the box to become infinite, so that the separation of the levels tends to zero. For any macroscopic **L**, the

 E_4 E_3 E_2 E_1

energy levels are very close to each other ($\Delta E \sim 1/L^2$), and the "continuous" description works well.

For a quantum particle confined in a 1D "box", the stationary solutions of Schrödinger equation are the standing waves with the wavelengths:

$$\lambda_n = \frac{2L}{n} = \frac{2\pi}{k}$$
, $n = 1, 2, \dots$ where $k = \frac{2\pi}{\lambda} = \frac{\pi n}{L}$

$$k = \frac{2\pi}{\lambda} = \frac{\pi n}{L}$$
 the

the wave number

The momentum in each of these states:
$$p_n = \frac{h}{\lambda_n} = \frac{hn}{2L} = \hbar k$$
 $\hbar = \frac{h}{2\pi}$

$$E_n = \frac{p_n^2}{2m} = \frac{h^2 n^2}{8mL^2} = \frac{(\hbar k)^2}{2m}$$
 - the energy (the non-relativistic case)

In the "k" (i.e., momentum) space, the states are equidistantly positioned (in contrast to the "energy" space).

The Density of States for a Single Free Particle in 1D (cont.)

The total number of states with the wave numbers < k is:

$$N(k) = \frac{k}{(\pi/L)} = \frac{kL}{\pi}$$

$$\Rightarrow G(k) = \frac{kL}{\pi}$$

 $\Rightarrow G(k) = \frac{kL}{\pi}$ G(k) - the number of states

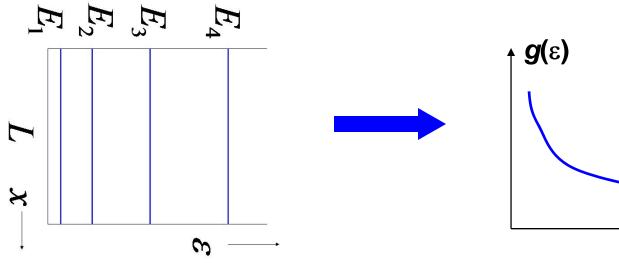
nbers
$$< \mathbf{k}$$
 is: $N(k) = \frac{k}{(\pi/L)} = \frac{kL}{\pi}$

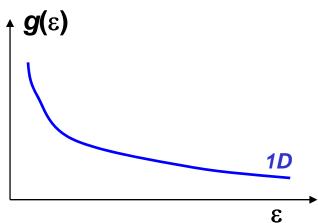
$$k = \frac{1}{\hbar} \sqrt{2m\varepsilon}$$

$$G(\varepsilon) = \frac{L\sqrt{2m\varepsilon}}{\pi\hbar}$$

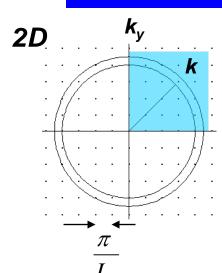
$$g^{1D}\left(\varepsilon\right) \equiv \frac{dG\left(\varepsilon\right)}{d\varepsilon} = \frac{L\left(2s+1\right)}{\pi} \sqrt{\frac{m}{2\varepsilon}}$$

additional degeneracy of the states (2**s**+1) because of **spin** [e.g., for electrons 2**s**+1=2]





The Density of States for a Particle in 2D and 3D



For quantum particles confined in a 2D "box" (e.g., electrons in FET):

$$k_{x} = \frac{\pi n_{x}}{L_{x}}$$
 $k_{y} = \frac{\pi n_{y}}{L_{y}}$ $k = \sqrt{k_{x}^{2} + k_{y}^{2}}$

$$N(k) = \frac{1}{4} \frac{\pi k^2}{\frac{\pi}{L_x} \times \frac{\pi}{L_y}}$$

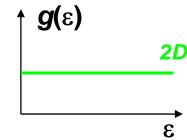
a circle of radius **k**

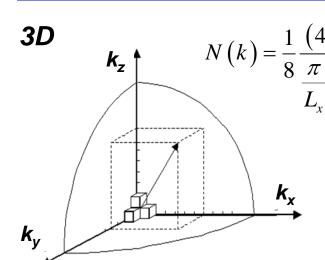
$$N(k) = \frac{1}{4} \frac{\pi}{\frac{\pi}{L_x}} \times \frac{\pi}{\frac{\pi}{L_y}} = \frac{k^2 (area)}{4\pi} G(k) = \frac{k^2}{4\pi} A G(\varepsilon) = \frac{A}{4\pi} \frac{2m\varepsilon}{\hbar^2}$$
states within ½ of a circle of radius k

$$g^{2D}(\varepsilon) = \frac{A(2s+1)m}{2\pi\hbar^2}$$

$$g^{2D}\left(\varepsilon\right) = \frac{A(2s+1)m}{2\pi\hbar^2}$$

- does not depend on ε





$$N(k) = \frac{1}{8} \frac{(4/3)\pi}{\frac{\pi}{L_x} \times \frac{\pi}{L_y} \times \frac{\pi}{L_z}} = \frac{k^3 \left(volume\right)}{6\pi^2} \quad G(k) = \frac{k^3 V}{6\pi^2} \quad G(\varepsilon) = \frac{V}{6\pi^2} \left(\frac{2m\varepsilon}{\hbar^2}\right)^{3/2}$$

$$g^{3D}(\varepsilon) = \frac{V(2s+1)}{4\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \varepsilon^{1/2}$$

Thus, for 3D electrons (2s+1=2):

$$g^{3D}\left(\varepsilon\right) = \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \varepsilon^{1/2}$$

Partition Function of a Free Particle in Three Dimensions

$$Z_{1} = V \int_{0}^{\infty} g(\varepsilon) \exp(-\beta \varepsilon) d\varepsilon = \left[g^{3D}(\varepsilon) = \frac{(2s+1)}{4\pi^{2}} \left(\frac{2m}{\hbar^{2}} \right)^{3/2} \varepsilon^{1/2} \right] = \frac{V(2s+1)}{4\pi^{2}} \left(\frac{2m}{\hbar^{2}} \right)^{3/2} \int_{0}^{\infty} \sqrt{\varepsilon} \exp(-\beta \varepsilon) d\varepsilon$$

 Z_1 – a reminder that this is the partition function for a **single** particle

The mean energy can be found without evaluating the integral:

$$\langle E \rangle = -\frac{\partial}{\partial \beta} \ln Z = \frac{\int_{0}^{\infty} \varepsilon^{3/2} \exp(-\beta \varepsilon) d\varepsilon}{\int_{0}^{\infty} \varepsilon^{1/2} \exp(-\beta \varepsilon) d\varepsilon}$$

The numerator can be integrated by parts:

$$\int_{0}^{\infty} \varepsilon^{3/2} \exp(-\beta \varepsilon) d\varepsilon = \frac{3}{2\beta} \int_{0}^{\infty} \varepsilon^{1/2} \exp(-\beta \varepsilon) d\varepsilon$$

(the integrand vanishes at both limits)

$$\langle E \rangle = \frac{3}{2} k_B T$$

- in agreement with the equipartition theorem

The Partition Function of a Free Particle (cont.)

$$Z_{1} = \frac{V}{4\pi^{2}} \left(\frac{2m}{\hbar^{2}}\right)^{3/2} \int_{0}^{\infty} \sqrt{\varepsilon} \exp(-\beta \varepsilon) d\varepsilon \qquad \text{(assume that s=0)}$$

$$\int_{0}^{\infty} \sqrt{\varepsilon} \exp(-\beta \varepsilon) d\varepsilon = (k_B T)^{3/2} \int_{0}^{\infty} \sqrt{x} \exp(-x) dx = (k_B T)^{3/2} \frac{\sqrt{\pi}}{2}$$

$$Z_1 = V \left(\frac{mk_B T}{2\pi \,\hbar^2}\right)^{3/2} = \frac{V}{V_Q}$$

$$Z_{1} = V \left(\frac{mk_{B}T}{2\pi\hbar^{2}}\right)^{3/2} = \frac{V}{V_{Q}}$$

$$V_{Q} = \left(\frac{2\pi\hbar^{2}}{mk_{B}T}\right)^{3/2} = \left(\frac{h^{2}\beta}{2\pi m}\right)^{3/2}$$

$$n_{Q} = \frac{1}{V_{Q}} = \left(\frac{mk_{B}T}{2\pi\hbar^{2}}\right)^{3/2}$$

$$n_Q \equiv \frac{1}{V_Q} = \left(\frac{mk_B T}{2\pi \,\hbar^2}\right)^{3/2}$$

For unit volume (
$$V=1$$
): $Z_1 = \frac{1}{V_Q} = n_Q$

- the quantum volume and density

$$V_Q = L_Q^3 \qquad L_Q \equiv \frac{h}{\sqrt{2\pi m k_B T}} \longleftarrow$$

 $V_Q = L_Q^3$ $L_Q \equiv \frac{h}{\sqrt{2\pi m k_B T}}$ the quantum length, coincides [within a factor of $(\pi)^{1/2}$] with the de Broglie length of a particle whose kinetic energy is k_BT .

 V_{o} is the volume of a box in which the ground state energy would be approximately equal to the thermal energy, and only the lowest few quantum states would be significantly populated. When $V \rightarrow V_Q$, the continuous approximation breaks down and we must take quantization of the states into account. We also need to consider quantum statistics (fermions and bosons)!

Partition Functions for Distinguishable Particles

There is nothing in the derivation of the partition function or the Boltzmann factor that restricts it to a *microsystem*. However, it is often convenient to express the partition function of a combined *macrosystem* in terms of the single-particle partition function. We restrict ourselves to the case of non-interacting particles (the model of *ideal gas*).

Recall the analogy between Ω and Z: $S(U,V,N) = k_B \ln \Omega$ $F(T,V,N) = -k_B T \ln Z$

For a *microcanonical* ensemble, we know the answer for Ω_{total} : $\Omega_{total} = \Omega_1 \times \Omega_2 \times ... \times \Omega_N$ How about the *canonical* ensemble?

For two non-interacting particles, $\boldsymbol{E_{total}} = \boldsymbol{E_1} + \boldsymbol{E_2}$: $Z_{total} = \sum \exp[-\beta(E_1(s) + E_2(s))]$

$$\Rightarrow Z_{total} = \sum_{s} \exp\left[-\beta E_{1}(s)\right] \times \exp\left[-\beta E_{2}(s)\right] = \left[\sum_{s_{1}} \sum_{s_{2}} \exp\left[-\beta E_{1}(s)\right] \times \exp\left[-\beta E_{2}(s)\right] = Z_{1} \times Z_{2}$$

all the states for the composite system

distinguishable

non-interacting distinguishable particles

$$Z_{total} = Z_1 \times Z_2 \times \dots \times Z_N$$

Example: two non-interacting distinguishable two-state particles, the states **0** and ε:

The partition function for a single particle:
$$Z_1 = \exp\left(-\frac{0}{k_B T}\right) + \exp\left(-\frac{\mathcal{E}}{k_B T}\right) = 1 + \exp\left(-\frac{\mathcal{E}}{k_B T}\right)$$

The states of this two-particle system are: (0,0), (ϵ,ϵ) , $(\epsilon,0)$, $(0,\epsilon)$ - **distinguishable**

$$Z_{total} = \sum_{1}^{4} \exp\left(-\frac{\varepsilon_{i}}{k_{B}T}\right) = 1 + 2 \exp\left(-\frac{\varepsilon}{k_{B}T}\right) + \exp\left(-\frac{2\varepsilon}{k_{B}T}\right) = \left[1 + \exp\left(-\frac{\varepsilon}{k_{B}T}\right)\right]^{2} = Z_{1}^{2}$$

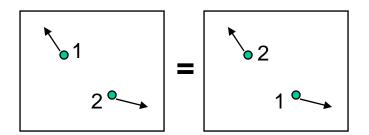
Partition Functions for Indistinguishable Particles (Low Density Limit)

If the particles are *indistinguishable*, the equation should be modified.

 $Z_{total} = Z_1 \times Z_2 \times \times Z_N$

Recall the multiplicity for IG:

 $\Omega_N = \frac{1}{N!} V^N \times \text{(the accessible momentum "volume")}$



If the states $(\varepsilon,0)$ and $(0,\varepsilon)$ of a system of two classical particles are indistinguishable:

classical particles are indistinguishable:
$$Z_{total} = \sum_{1}^{3} \exp\left(-\frac{\varepsilon_{i}}{k_{B}T}\right) = 1 + \exp\left(-\frac{\varepsilon}{k_{B}T}\right) + \exp\left(-\frac{2\varepsilon}{k_{B}T}\right) \neq Z_{1}^{2}$$

(For the *quantum indistinguishable* particles, there would be three microstates if the particles were bosons $[(0,0), (0,\varepsilon), (\varepsilon,\varepsilon)]$ and one microstate $[(0,\varepsilon)]$ if the particles were fermions).

$$Z_{total} \approx \frac{1}{2} \sum_{s_1} \sum_{s_2} \exp[-\beta E_1(s)] \times \exp[-\beta E_2(s)] = \frac{1}{2} Z_1 \times Z_2$$

Exceptions: the microstates where the particles are in the same state. If we consider particles in the phase space whose volume is the product of accessible volumes in the coordinate space and momentum space (e.g., an ideal gas), the states with the same position and momentum are very rare (unless the density is extremely high)

For a low-density system of classical indistinguishable particles:

$$Z_{total} = \frac{1}{N!} Z_1^N$$

Z and Thermodynamic Properties of an Ideal Gas

The partition function for an ideal gas:

$$Z = \frac{1}{N!} Z_1^N$$

$$Z = \frac{1}{N!} Z_1^N \qquad Z = \frac{1}{N!} V^N \left(\frac{m k_B T}{2\pi \hbar^2} \right)^{3N/2} = \frac{1}{N!} \left(\frac{V}{V_Q} \right)^N$$

Note that the derivation of Z for an ideal gas was much easier than that for Ω : we are not constrained by the energy conservation!

$$\ln Z = N \left[\ln V - \ln V_Q - \ln N + 1 \right] = N \left[\ln \left(\frac{V}{NV_Q} \right) + 1 \right] = N \left[\ln \left(\frac{n_Q}{n} \right) + 1 \right]$$

$$F(T,V,N) = -k_B T \ln Z = -Nk_B T \left| \ln \left(\frac{V}{NV_Q} \right) + 1 \right| = -Nk_B T \left[\ln \left(\frac{n_Q}{n} \right) + 1 \right]$$

$$P = -\left(\frac{\partial F}{\partial V}\right)_{TN} = Nk_B T \frac{\partial}{\partial V} \ln V = \frac{Nk_B T}{V} - \text{the "pressure" equation of state ("ideal gas" law)}$$

$$S = -\left(\frac{\partial F}{\partial T}\right)_{V,N} = Nk_B \frac{\partial}{\partial T} \left[T \left(\ln\left(\frac{n_Q}{n}\right) + 1 \right) \right] = Nk_B \left[\ln\left(\frac{n_Q}{n}\right) + T \frac{\partial}{\partial T} \ln T^{3/2} + 1 \right] = Nk_B \left[\ln\left(\frac{n_Q}{n}\right) + \frac{5}{2} \right]$$

$$U = -\frac{\partial}{\partial \beta} \ln Z = N \frac{\partial}{\partial \beta} \ln V_{\mathcal{Q}} = N \frac{1}{V_{\mathcal{Q}}} \frac{\partial V_{\mathcal{Q}}}{\partial \beta} = N \frac{3}{2} \frac{1}{\beta} = \frac{3}{2} N k_{B} T$$
 - the Sackur-Tetrode equation - the "energy" equation of state
$$U = F + TS = \frac{3}{2} N k_{B} T$$

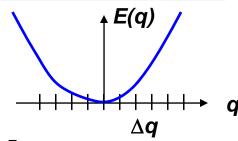
The Equipartition Theorem (continuous spectrum)

Equipartition Theorem: At temperature **T**, the average energy of any **quadratic** degree of freedom is $\frac{1}{2}k_{\rm B}T$.

We'll "one-dimensional" consider a system with just one degree of freedom:

$$E(q)=cq^2$$

q – a continuous variable, but we "discretize" it by considering small and countable Δq 's:



The partition function for this system:

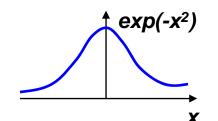
$$Z = \sum_{q} \exp \left[-\frac{E(q)}{k_B T} \right] = \sum_{q} \exp \left[-\beta c q^2 \right]$$

$$Z = \frac{1}{\Delta q} \sum_{q} \exp\left[-\beta c q^{2}\right] \Delta q = \frac{1}{\Delta q} \int_{-\infty}^{\infty} \exp\left[-\beta c q^{2}\right] dq$$

New variable: $x = \sqrt{\beta c} q$

$$x = \sqrt{\beta c} \ d$$

$$Z = \frac{1}{\Delta q} \frac{1}{\sqrt{\beta c}} \int_{-\infty}^{\infty} \exp\left[-x^2\right] dx = \left[\int_{-\infty}^{\infty} \exp\left[-x^2\right] dx = \sqrt{\pi}\right] = \frac{1}{\Delta q} \sqrt{\frac{\pi}{\beta c}}$$



The average energy:

$$\overline{E} = -\frac{1}{Z} \frac{\partial Z}{\partial \beta} = -\sqrt{\beta} \left(-\frac{1}{2} \beta^{-3/2} \right) = \frac{1}{2} \beta^{-1} = \frac{k_B T}{2}$$

Again, this result is valid only if the degree of freedom is "quadratic" (the limit of high temperatures, when the spacing between the energy levels of an individual system is << k_BT).

Equipartition Theorem and a Quantum Oscillator

Classical oscillator:
$$E = \frac{p^2}{2m} + \frac{kx^2}{2}$$
 $\overline{E} = \frac{k_B T}{2} + \frac{k_B T}{2} = k_B T$

Quantum oscillator:
$$\varepsilon_n = \left(n + \frac{1}{2}\right)\hbar\omega \quad \omega = \sqrt{\frac{k}{m}}$$

the equipartition function: $Z = \sum_{0}^{\infty} \exp(-\beta \varepsilon_{n}) = \exp\left(-\frac{\beta}{2}\hbar\omega\right) \sum_{0}^{\infty} \exp(-n\beta\hbar\omega) = \frac{\exp\left(-\frac{\beta}{2}\hbar\omega\right)}{1 - \exp(-\beta\hbar\omega)}$

$$1 + \exp(-\beta \hbar \omega) + \exp(-2\beta \hbar \omega) + \dots$$

$$\ln Z = -\frac{\beta}{2}\hbar\omega - \ln[1 - \exp(-\beta\hbar\omega)]$$

the average energy of the oscillator:

$$\overline{E} = -\frac{\partial}{\partial \beta} \ln Z = \frac{\hbar \omega}{2} + \frac{\exp(-\beta \hbar \omega) \times \hbar \omega}{1 - \exp(-\beta \hbar \omega)} = \hbar \omega \left[\frac{1}{2} + \frac{1}{\exp(\beta \hbar \omega) - 1} \right]$$

the limit of high
$$T$$
: $\frac{\hbar\omega}{k_BT} = \beta\hbar\omega <<1 \exp(\beta\hbar\omega) \approx 1 + \beta\hbar\omega \quad \overline{E} \approx \hbar\omega \left[\frac{1}{2} + \frac{1}{\beta\hbar\omega}\right] \approx \frac{1}{\beta} = k_BT$

the limit of low
$$T$$
: $\frac{\hbar\omega}{k_BT} = \beta\hbar\omega >> 1$ $\overline{E} \approx \frac{\hbar\omega}{2}$ (the ground state is by far the most probable state)

the assumption of a continuous energy spectrum was violated

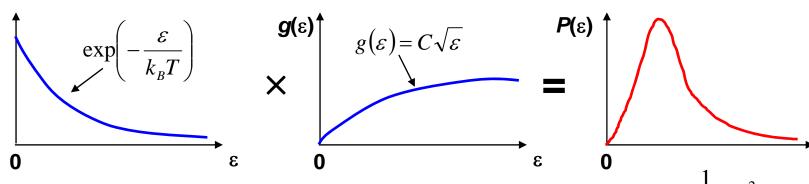
The Maxwell Speed Distribution (ideal gas)

The probability that a molecule has an energy between ε and $\varepsilon + d\varepsilon$ is:

$$P(\varepsilon)d\varepsilon = \frac{1}{Z_1}g(\varepsilon)\exp\left(-\frac{\varepsilon}{k_BT}\right)d\varepsilon = \frac{1}{n_Q}g(\varepsilon)\exp\left(-\frac{\varepsilon}{k_BT}\right)d\varepsilon$$

$$g^{3D}(\varepsilon) = \frac{1}{4\pi^2} \left(\frac{2m}{\hbar^2}\right)^{3/2} \varepsilon^{1/2} = \frac{2}{\sqrt{\pi}} \varepsilon^{1/2} \beta^{3/2} n_Q \qquad P(\varepsilon) = \frac{2}{\sqrt{\pi}} \beta^{3/2} \varepsilon^{1/2} \exp\left(-\frac{\varepsilon}{k_B T}\right)$$

$$P(\varepsilon) = \frac{2}{\sqrt{\pi}} \beta^{3/2} \varepsilon^{1/2} \exp\left(-\frac{\varepsilon}{k_B T}\right)$$



Now let's look at the **speed distribution** for these particles:

$$\mathcal{E} = \frac{mv^2}{2}$$

v, irrespective of the

The probability to find a particle with the speed between \mathbf{v} and $\mathbf{v}+\mathbf{dv}$, irrespective of the direction of its velocity, is the same as that finding it between ε and $\varepsilon + d\varepsilon$ where $d\varepsilon =$ mvdv:

$$P(v)dv = P(\varepsilon)d\varepsilon = P(\varepsilon)mvdv$$

$$P(v) = \left(\frac{2}{\pi}\right)^{1/2} (m\beta)^{3/2} v^2 \exp\left(-\frac{1}{2}\beta mv^2\right) = \left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{mv^2}{2k_B T}\right)$$
 Maxwell distributions

Note that Planck's constant has vanished from the equation – it is a classical result.

The Maxwell Speed Distribution (cont.)

$$P(v)dv = \left(\frac{m}{2\pi k_B T}\right)^{3/2} \exp\left(-\frac{mv^2}{2k_B T}\right) 4\pi v^2 dv$$

The structure of this equation is transparent: the Boltzmann factor is multiplied by the number of states between \mathbf{v} and $\mathbf{v}+\mathbf{dv}$. The constant can be found from the normalization:

$$\int_{0}^{\infty} P(v)dv = 1 \qquad C = \left(\frac{m}{2\pi k_{B}T}\right)^{3/2}$$

$$dN(\varepsilon) = NP(\varepsilon)d\varepsilon = N\exp\left(-\frac{\varepsilon}{k_B T}\right)d\varepsilon$$

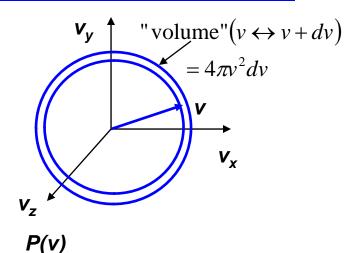
energy distribution, N - the total # of particles

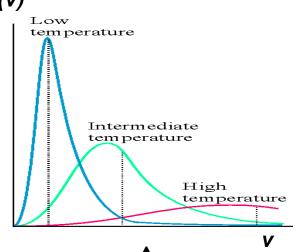
$$dN(v) = NP(v) dv = N\left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{mv^2}{2k_B T}\right) dv$$

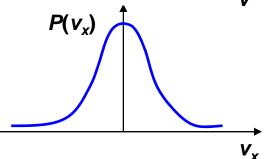
speed distribution

$$dN(v_x) = NP(v_x) dv = N\left(\frac{m}{2\pi k_B T}\right)^{1/2} \exp\left(-\frac{mv_x^2}{2k_B T}\right) dv_x$$

distribution for the projection of velocity, v_x







The Characteristic Values of Speed

$$P(v) = \left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{mv^2}{2k_B T}\right)$$

V

 $P(v) = \left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{mv^2}{2k_B T}\right)$ Because Maxwell distribution is skewed (not symmetric in \mathbf{v}), the root mean square speed is not equal to the most probable speed:

> The **root-mean-square speed** is proportional to the square root of the average energy:

$$\overline{E} = \frac{1}{2} m (v_{rms})^2$$
 $v_{rms} = \sqrt{\frac{2\overline{E}}{m}} = \sqrt{\frac{3k_B T}{m}}$

The most probable speed:

$$\left[\frac{dP(v)}{dv}\right]_{v=v} = 0 \quad \to \quad v_{\text{max}} = \sqrt{\frac{2k_BT}{m}}$$

The average speed:

$$\overline{v} = \int_{0}^{\infty} v \times P(v) dv = \left(\frac{m}{2\pi k_B T}\right)^{3/2} \int_{0}^{\infty} 4\pi v^3 \exp\left(-\frac{mv^2}{2k_B T}\right) dv = \sqrt{\frac{8k_B T}{\pi m}}$$

$$v_{\text{max}} : \overline{v} : v_{\text{rms}} = \sqrt{2} : \sqrt{8/\pi} : \sqrt{3} = 1 : 1.13 : 1.22$$

Problem (Maxwell distr.)

Consider a mixture of Hydrogen and Helium at T=300 K. Find the speed at which the Maxwell distributions for these gases have the same value.

$$P(v,T,m) = \left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{mv^2}{2k_B T}\right)$$

$$\left(\frac{m_1}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{m_1 v^2}{2k_B T}\right) = \left(\frac{m_2}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{m_2 v^2}{2k_B T}\right)$$

$$\frac{3}{2} \ln m_1 - \frac{m_1 v^2}{2k_B T} = \frac{3}{2} \ln m_2 - \frac{m_2 v^2}{2k_B T}$$

$$\frac{3}{2} \ln \frac{m_1}{m_2} = \frac{v^2}{2k_B T} \left(m_1 - m_2\right) \qquad v = \sqrt{\frac{3k_B T \ln \frac{m_1}{m_2}}{(m_1 - m_2)}} = \sqrt{\frac{3 \times 1.38 \cdot 10^{-23} \times 300 \times \ln 2}{2 \times 1.7 \cdot 10^{-27}}} = 1.6 \text{ km/s}$$

Problem (final 2005, MB speed distribution)

Consider an ideal gas of atoms with mass *m* at temperature *T*.

- (a) Using the Maxwell-Boltzmann distribution for the speed \mathbf{v} , find the corresponding distribution for the kinetic energy ϵ (don't forget to transform $d\mathbf{v}$ into $d\epsilon$).
- (b) Find the most probable value of the kinetic energy.
- (c) Does this value of energy correspond to the most probable value of speed? Explain.

(a)
$$dN(v) = NP(v) dv = N \left(\frac{m}{2\pi k_B T}\right)^{3/2} 4\pi v^2 \exp\left(-\frac{mv^2}{2k_B T}\right) dv$$

$$dN(\varepsilon) = NP(\varepsilon)\frac{dv}{d\varepsilon}d\varepsilon = \left[v = \sqrt{\frac{2\varepsilon}{m}} \quad dv = \frac{1}{2}\sqrt{\frac{2}{m\varepsilon}}d\varepsilon\right] = N\left(\frac{m}{2\pi k_B T}\right)^{3/2}\frac{8\pi\varepsilon}{m}\exp\left(-\frac{\varepsilon}{k_B T}\right)\frac{1}{2}\sqrt{\frac{2}{m\varepsilon}}d\varepsilon$$

$$P(\varepsilon) = \left(\frac{m}{2\pi k_B T}\right)^{3/2} \frac{4\sqrt{2\pi} \varepsilon^{1/2}}{m^{3/2}} \exp\left(-\frac{\varepsilon}{k_B T}\right)$$

(b)
$$\frac{\partial P(\varepsilon)}{\partial \varepsilon} = 0$$
 $\frac{1}{2} \varepsilon^{-1/2} \exp\left(-\frac{\varepsilon}{k_B T}\right) + \varepsilon^{1/2} \exp\left(-\frac{\varepsilon}{k_B T}\right) \times \left(-\frac{1}{k_B T}\right) = 0$ $\frac{1}{2} = \frac{\varepsilon_{\text{max}}}{k_B T}$ $\varepsilon_{\text{max}} = \frac{1}{2} k_B T$

(c) the most probable value of speed $\left[\frac{dP(v)}{dv}\right]_{v=v_{\text{max}}} = 0 \rightarrow v_{\text{max}} = \sqrt{\frac{2k_BT}{m}}$ doesn't correspond

the kin. energy that corresponds to the most probable value of speed

$$\varepsilon = k_B T$$