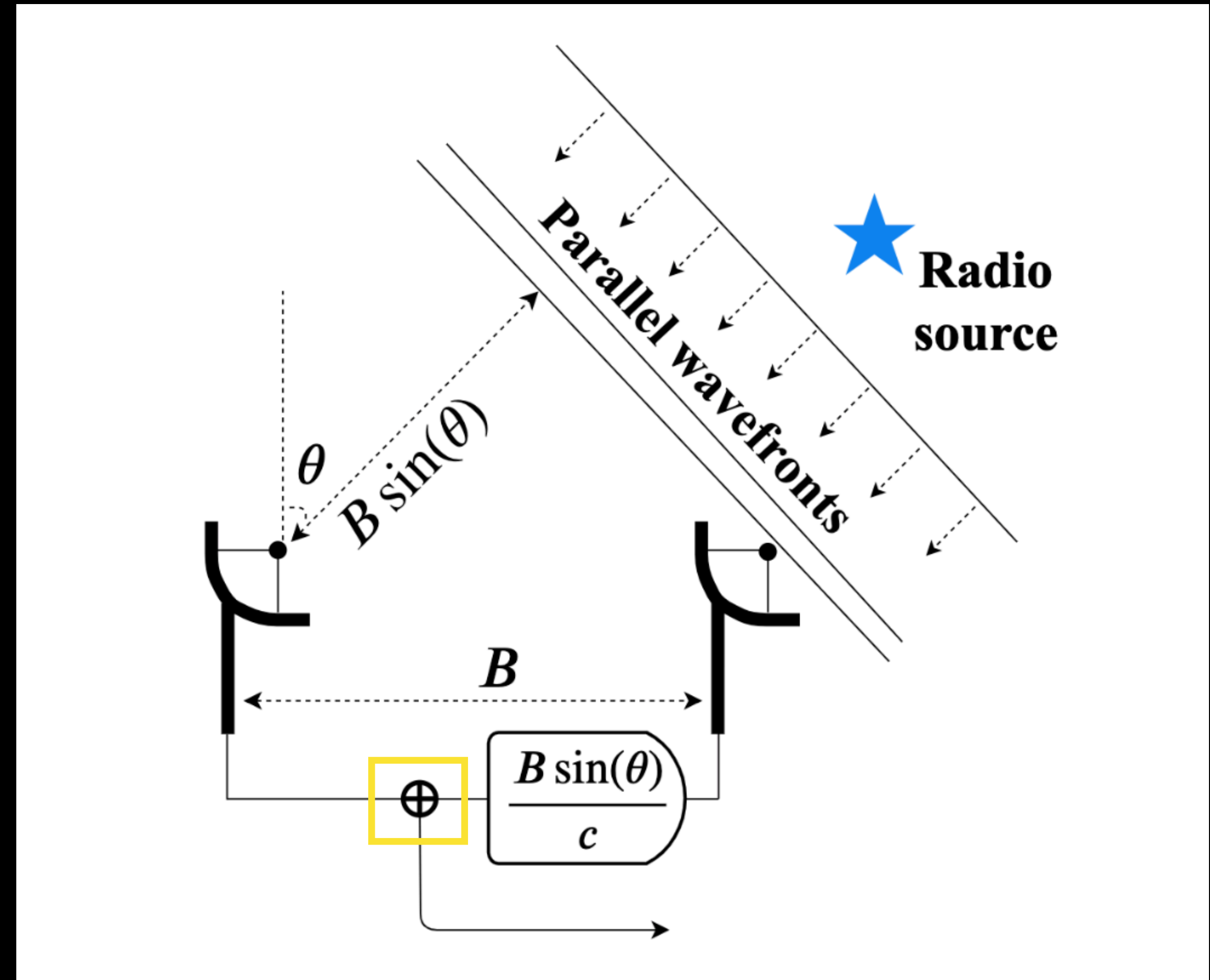


# Updates on ATA beam-forming calibration

Wael Farah - interferometric SETI meeting

# Simple Interferometer

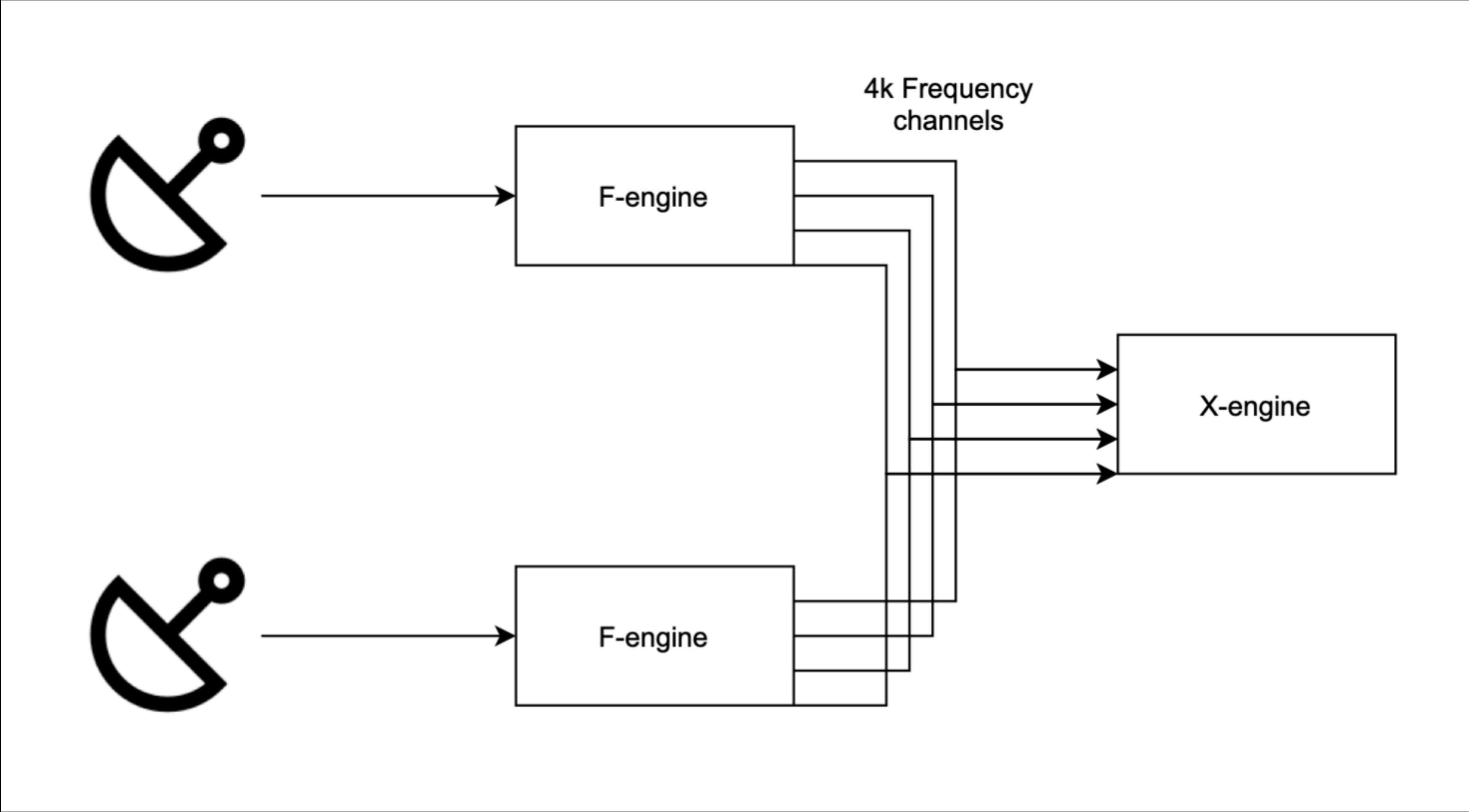
- 2 major **operations**:
  - Multiplication (aka correlation). Each antenna is multiplied by every other (including itself), forming baselines. 2D iFFT of this data product gives sky intensity distribution (i.e. image of the sky)
  - Addition (aka beamforming). Every antenna is compensated for the delay (with respect to a reference position) and instrumental phase, and then summed coherently to maximize the sensitivity of the array on a sky position



# Sources of delay

- Geometric delay: the fact that antennas are spread geometrically on earth
- Fixed delay: Due to fiber/copper, can change over time/temperature.
- Instrumental phase: electronics + atmosphere (concern at higher frequencies) + ...

# Backend Schematic



# Calibration (delay+ phase) scheme

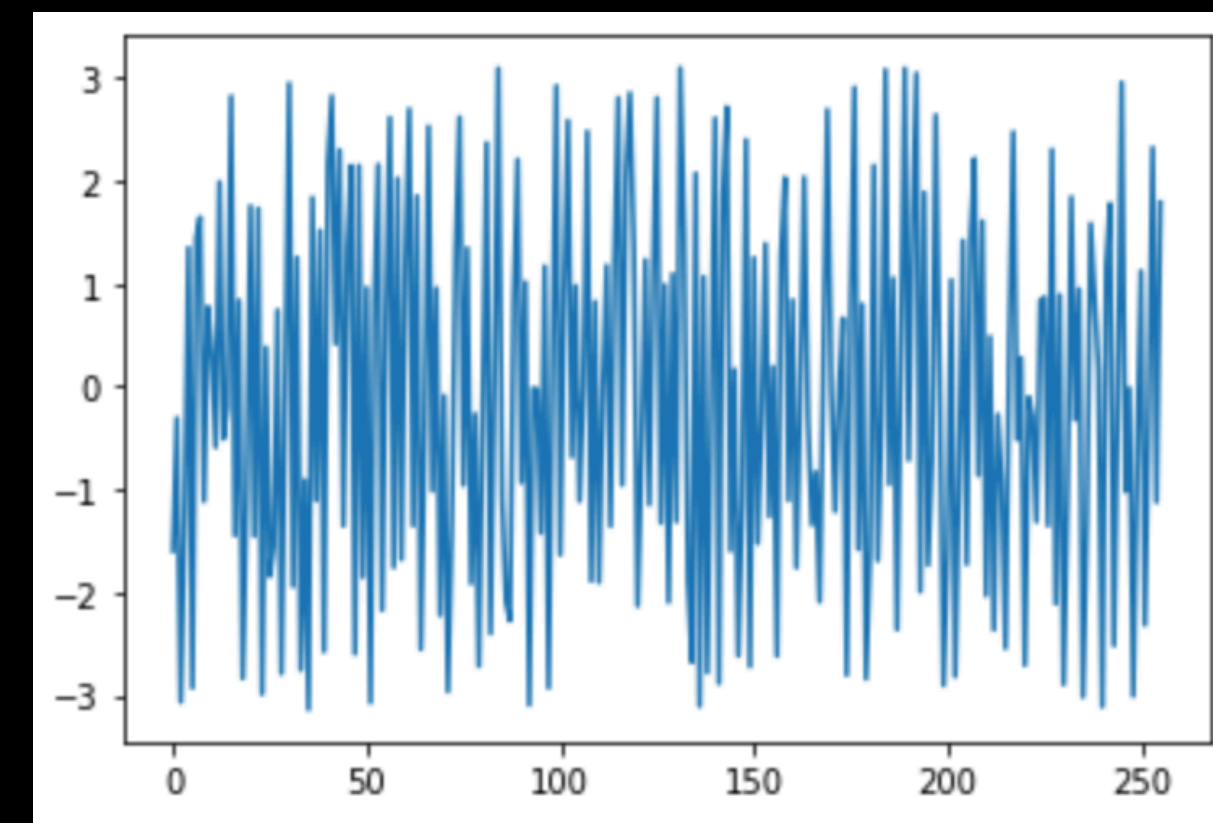
## 1. Cross-correlate antennas

$$F(\nu) = \sum_{t_i}^T V_a(t_i, \nu) \times V_b^*(t_i, \nu)$$

Where \* represents the complex conjugate,  
ν is frequency  
and “a” and “b” are antenna indices



Phase [radians]



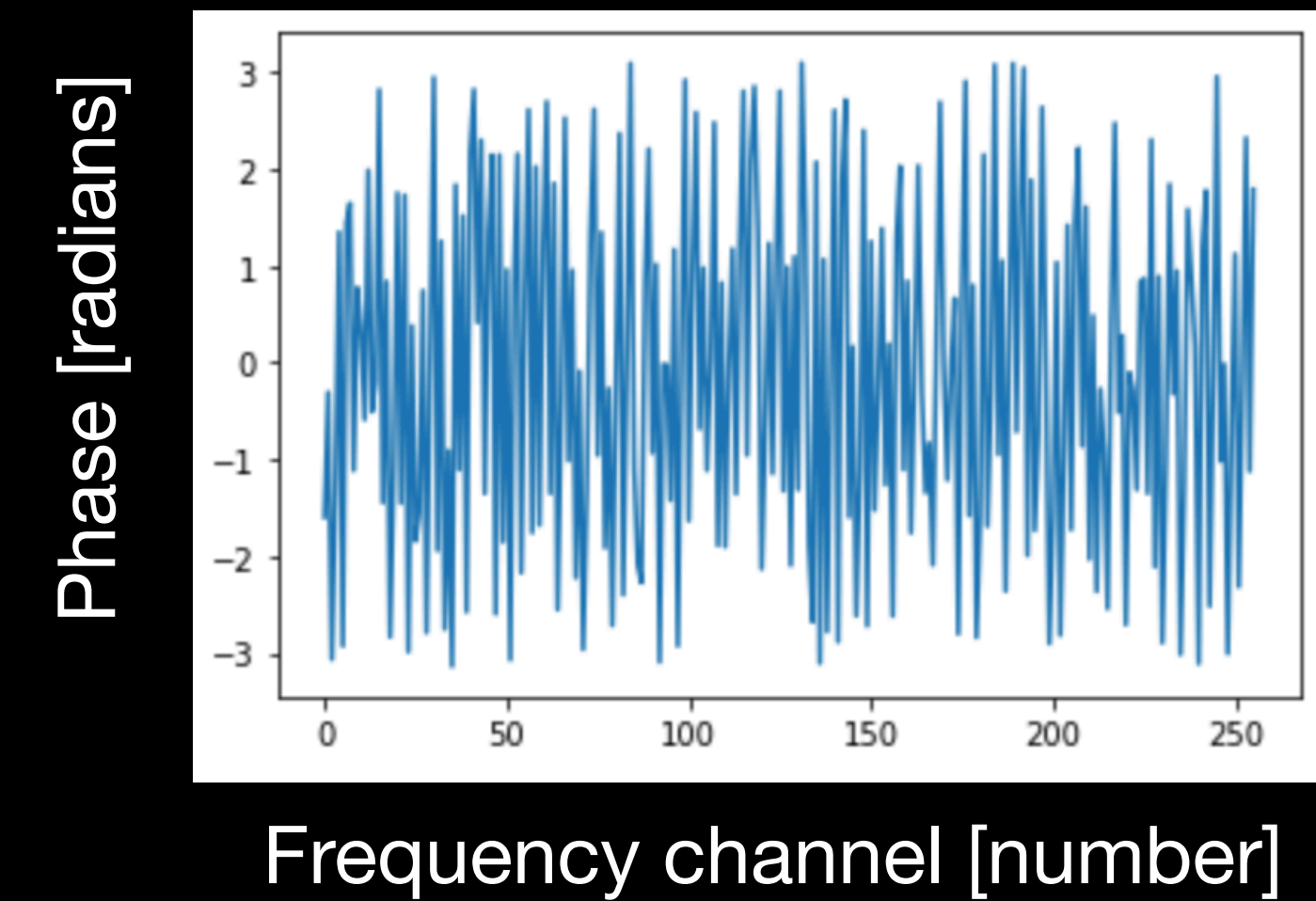
Frequency channel [number]

1c <-> 5c baseline (x pol)

# Calibration (delay+ phase) scheme

1. Cross-correlate antennas,  
display data + sanity checks

1c <-> 5c baseline (x pol)

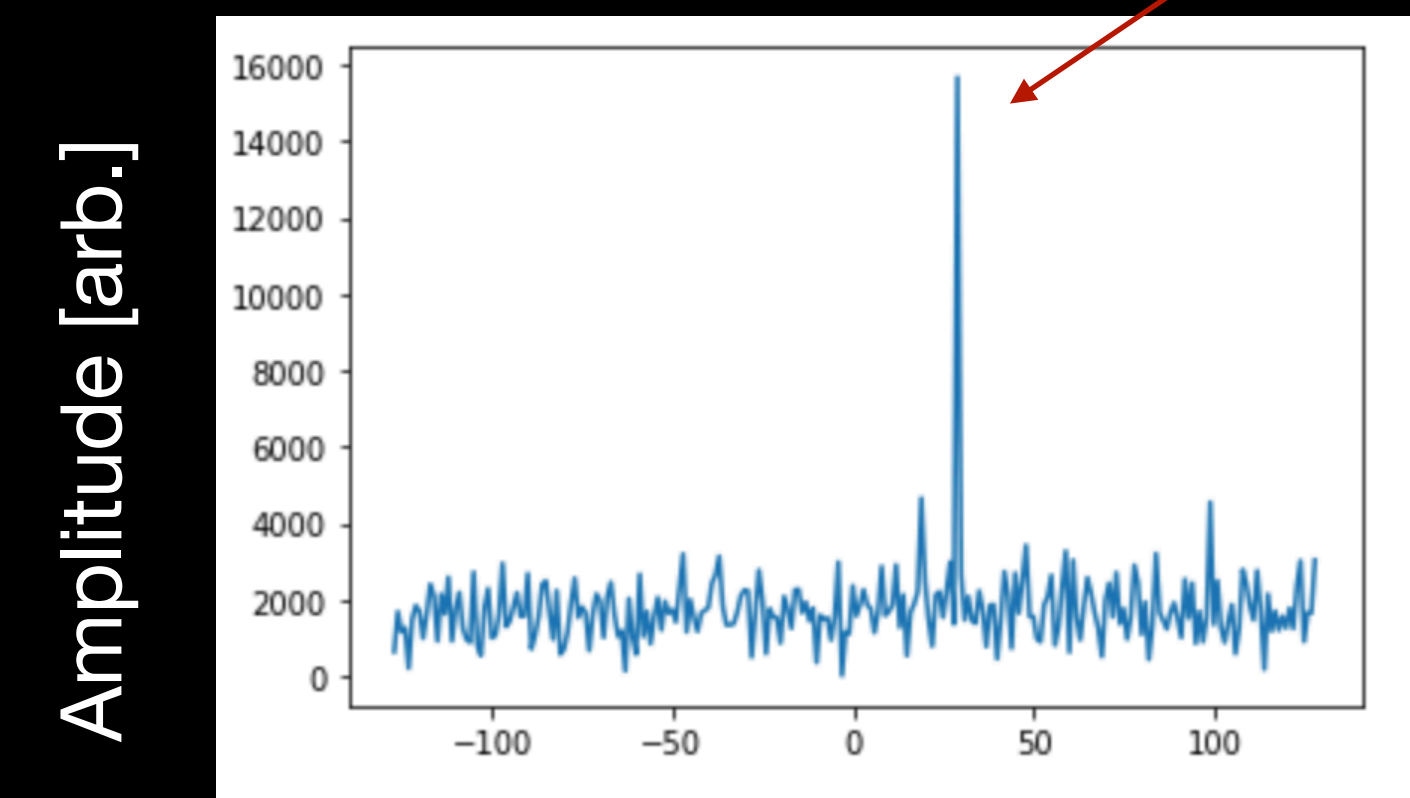


Lag in the time domain  
=  
fringe in the freq domain

iFFT

Lag spectrum

Spike in lag spectrum

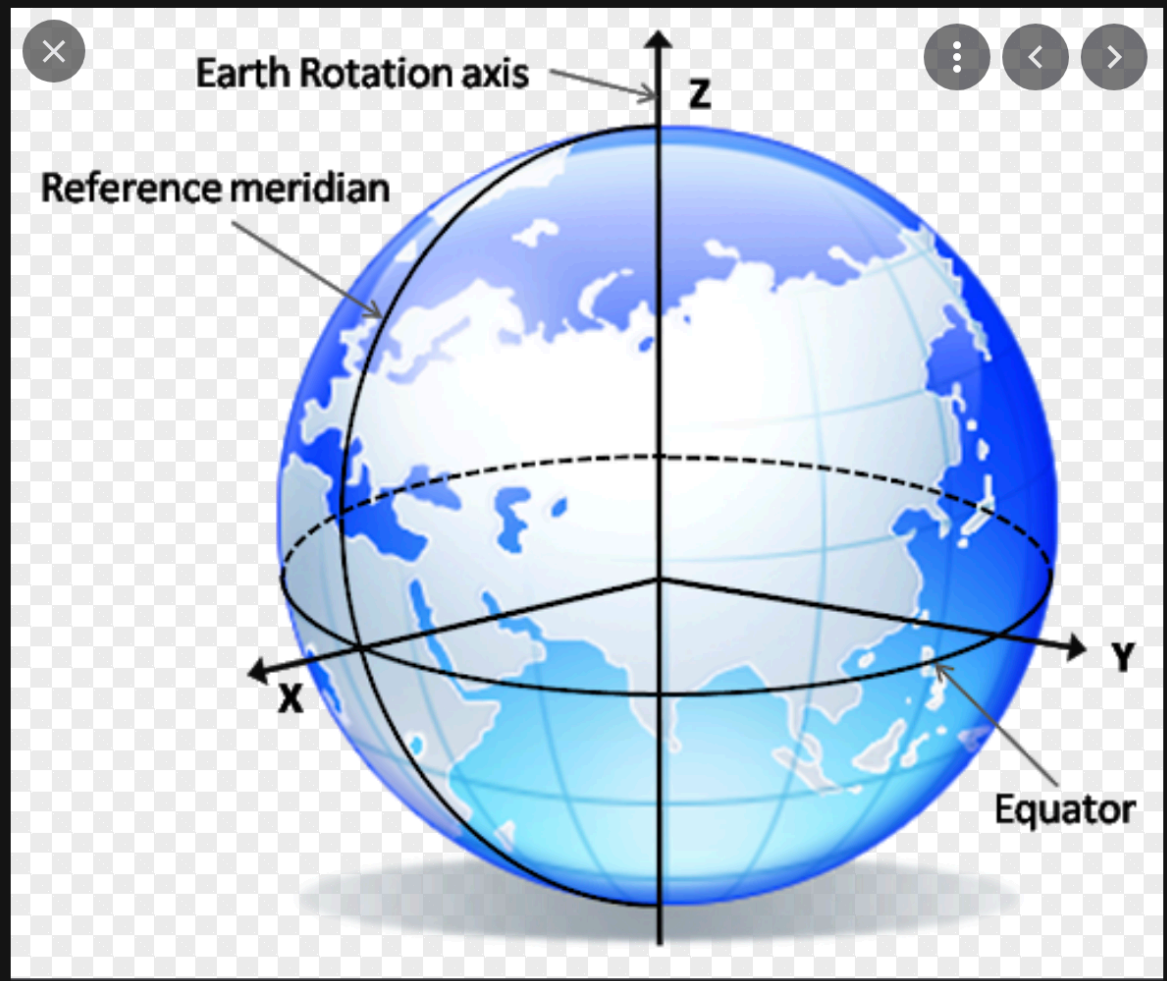


Samples [1/64MHz = 15.625 ns]



# Calibration (delay+ phase) scheme

- 1. Cross-correlate antennas, display data + sanity checks
- 2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)



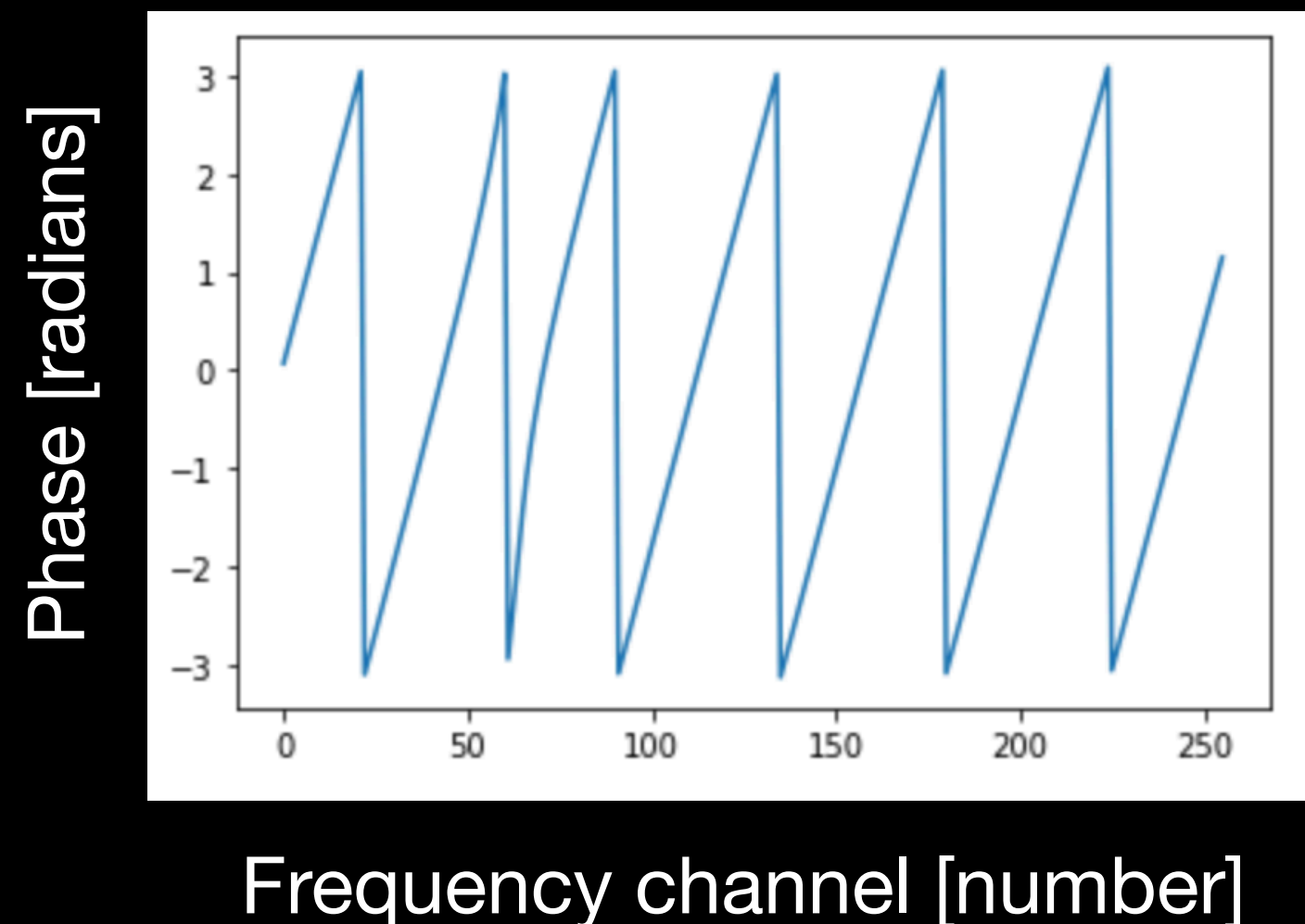
ITRF (ECEF; earth centered earth fixed)  
antenna positions XYZ

	x	y	z
1A	-2.524136e+06	-4.123448e+06	4.147709e+06
1F	-2.524193e+06	-4.123464e+06	4.147658e+06
1C	-2.524141e+06	-4.123508e+06	4.147649e+06
2A	-2.524086e+06	-4.123417e+06	4.147769e+06
4J	-2.523998e+06	-4.123376e+06	4.147863e+06
2H	-2.524174e+06	-4.123388e+06	4.147745e+06
3D	-2.524045e+06	-4.123400e+06	4.147811e+06
5B	-2.523924e+06	-4.123448e+06	4.147837e+06
1K	-2.524156e+06	-4.123435e+06	4.147709e+06
5C	-2.523925e+06	-4.123461e+06	4.147824e+06
1H	-2.524203e+06	-4.123431e+06	4.147685e+06
2B	-2.524070e+06	-4.123435e+06	4.147761e+06

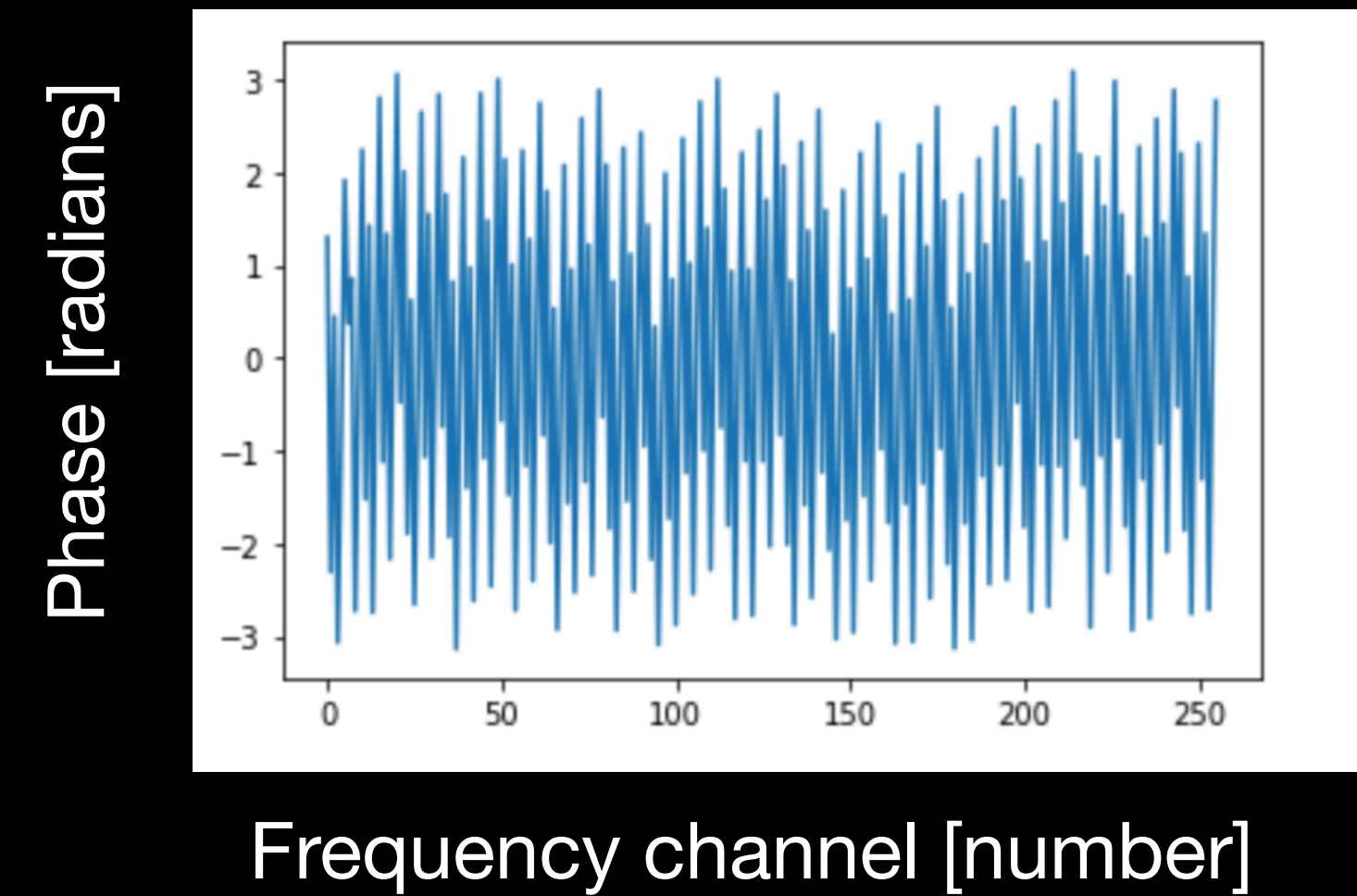
# Calibration (delay+ phase) scheme

1. Cross-correlate antennas, display data + sanity checks
2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays: ~6 samples  
~ 93.75 ns

Geometric delay phasor:

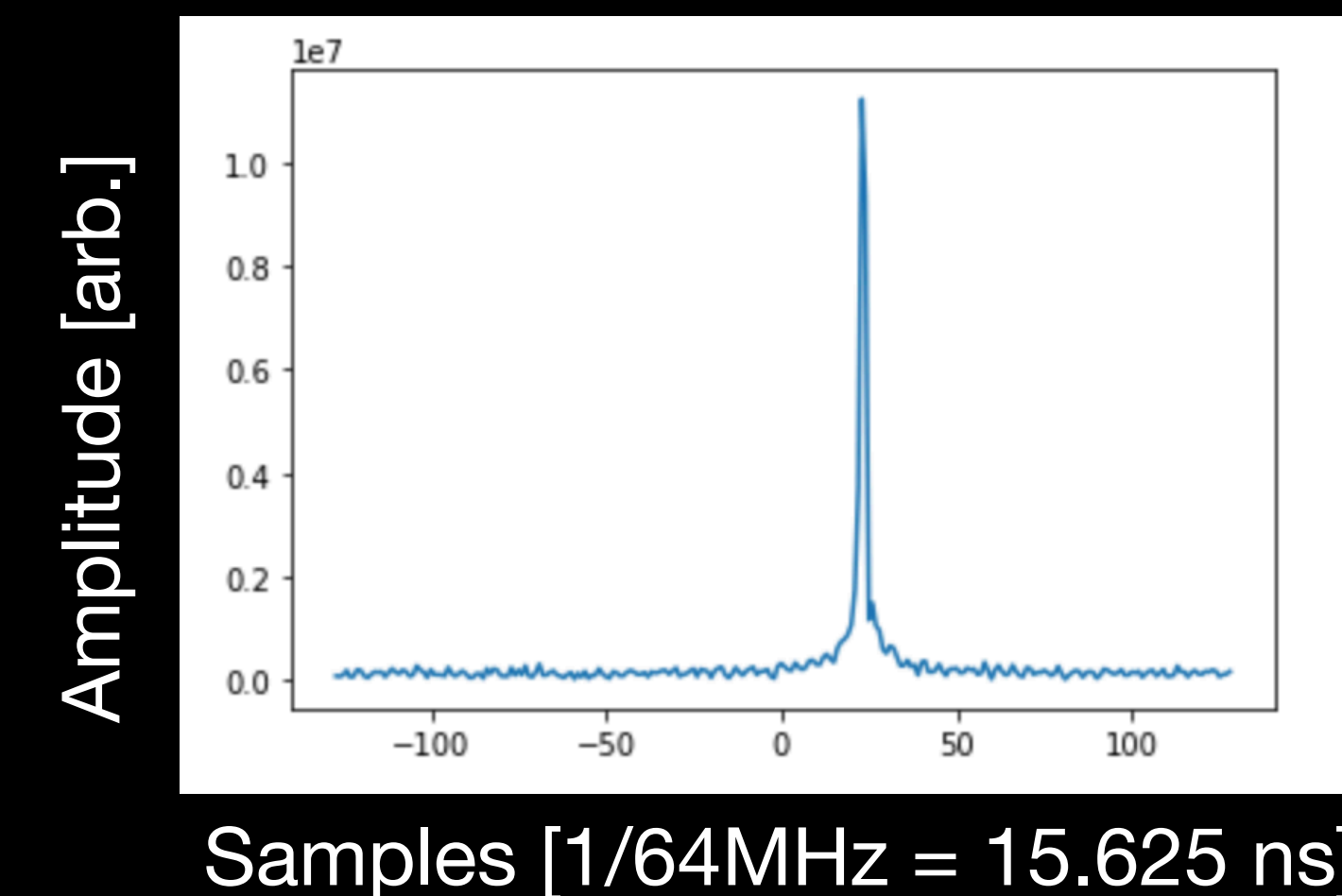


1c <-> 5c baseline



iFFT

Lag spectrum



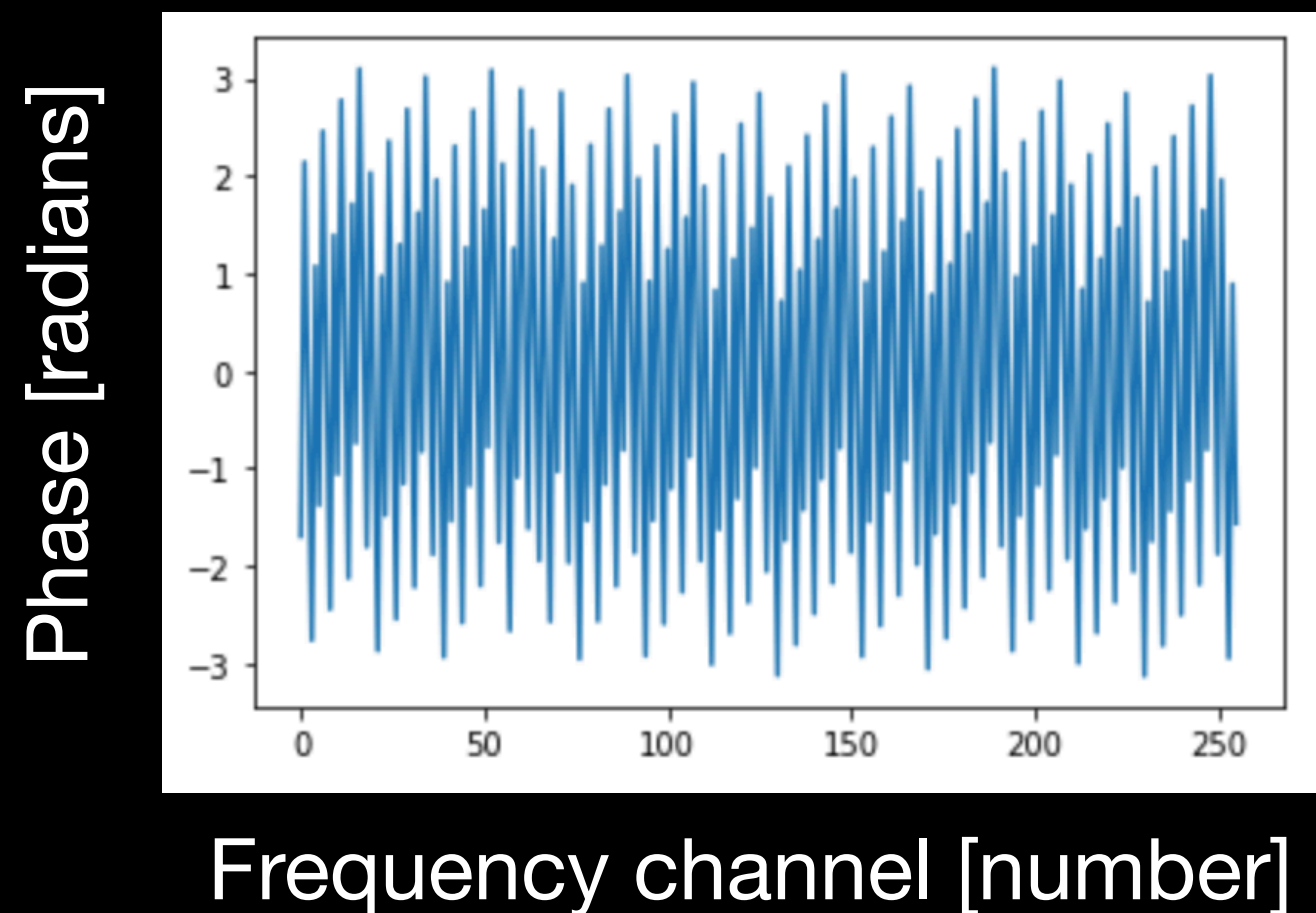


# Calibration (delay+ phase) scheme

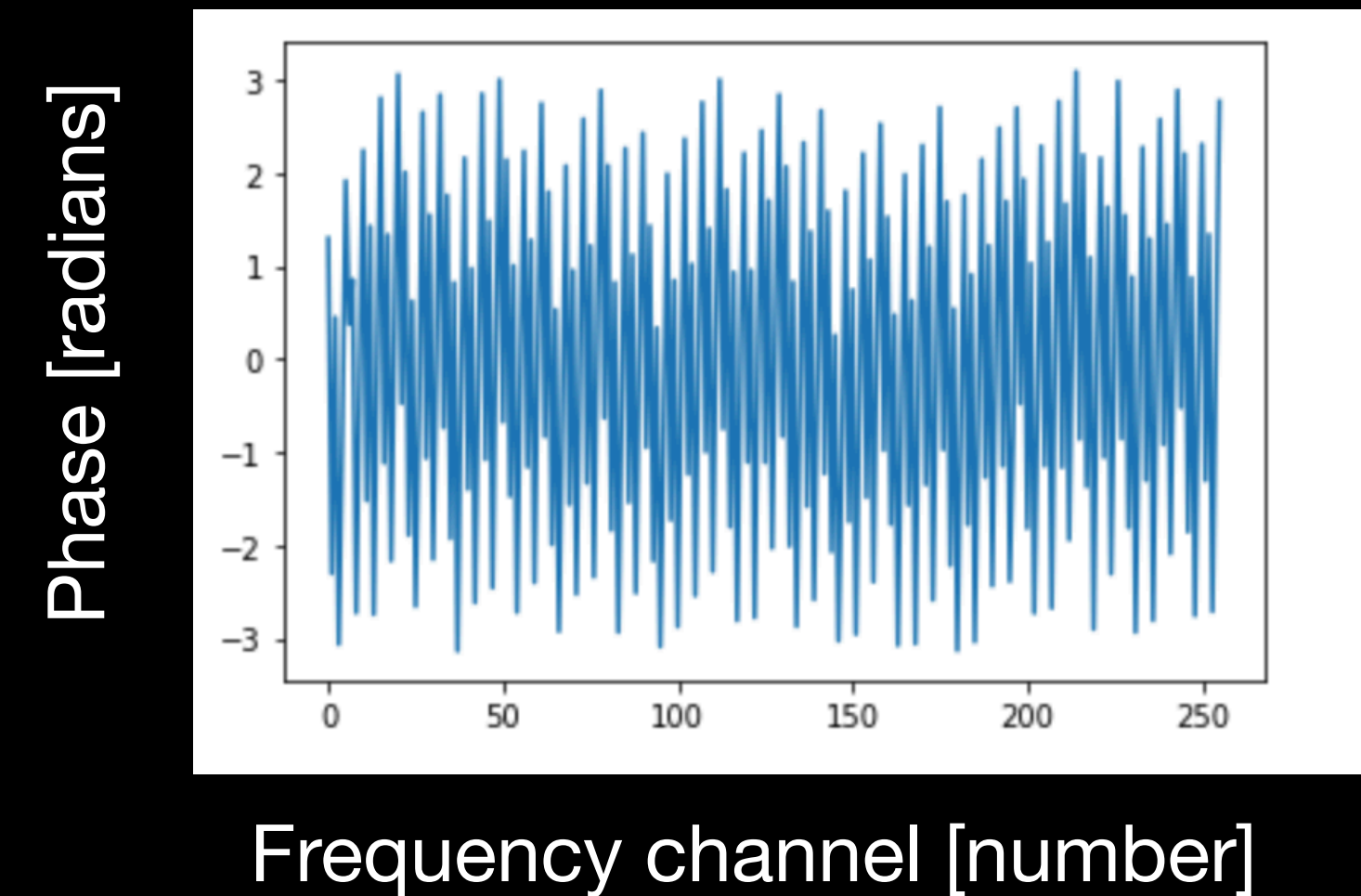
1. Cross-correlate antennas, display data + sanity checks
2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays
4. Measure integer-sample fixed delays from lag spectrum.

22 samples = 343.75ns

Fixed delay phasor:

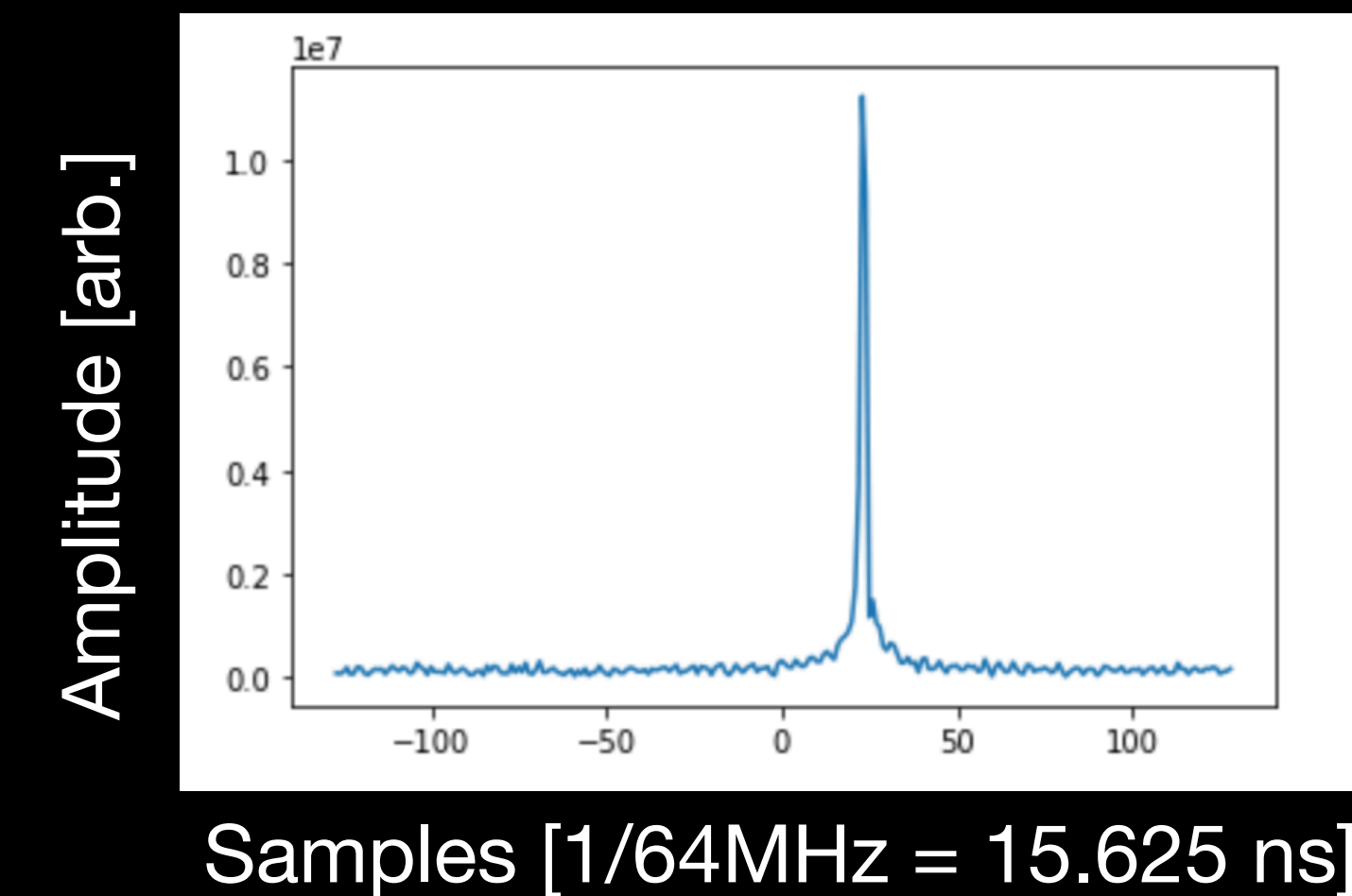


1c <-> 5c baseline



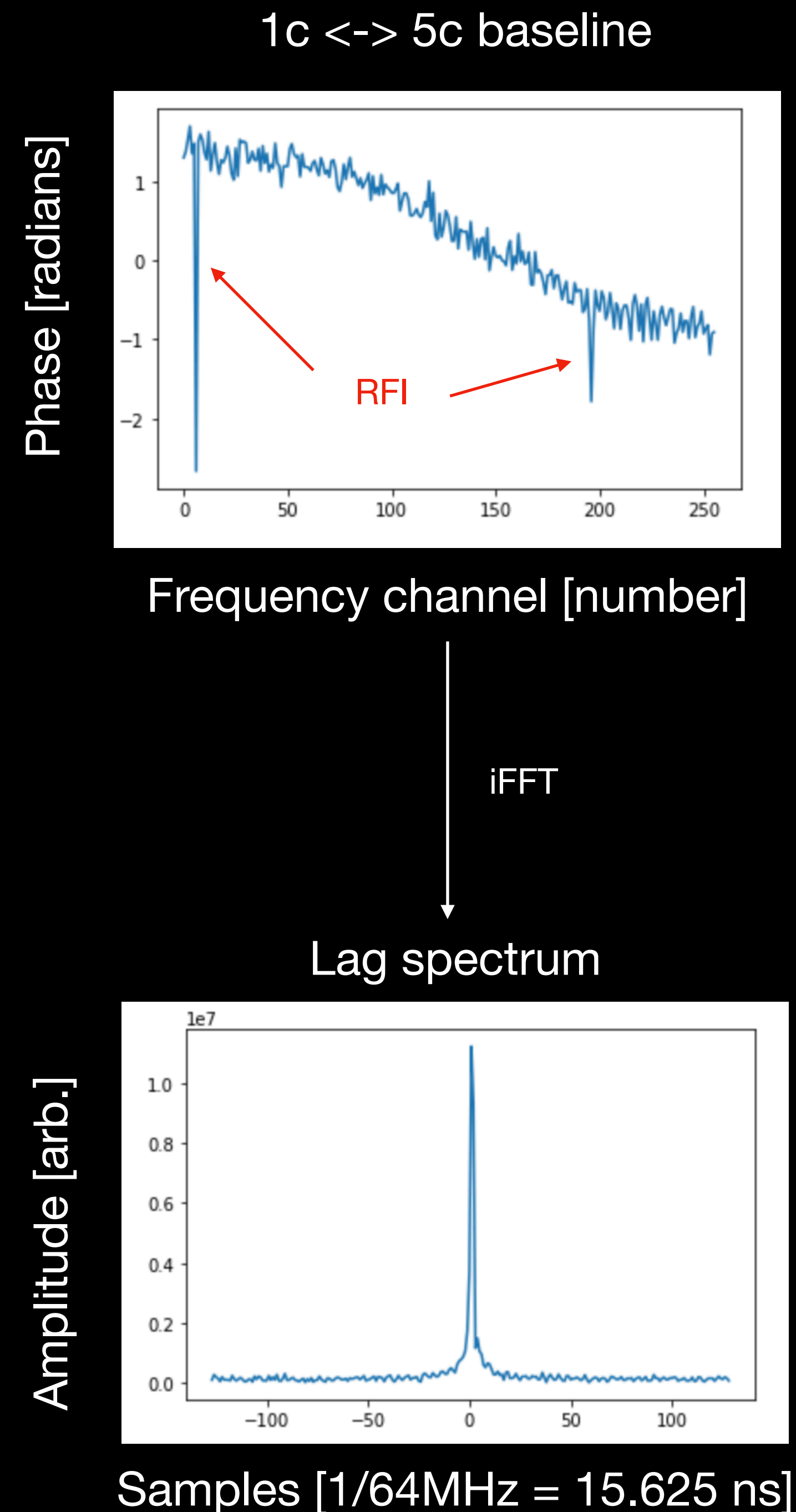
iFFT

Lag spectrum



# Calibration (delay+ phase) scheme

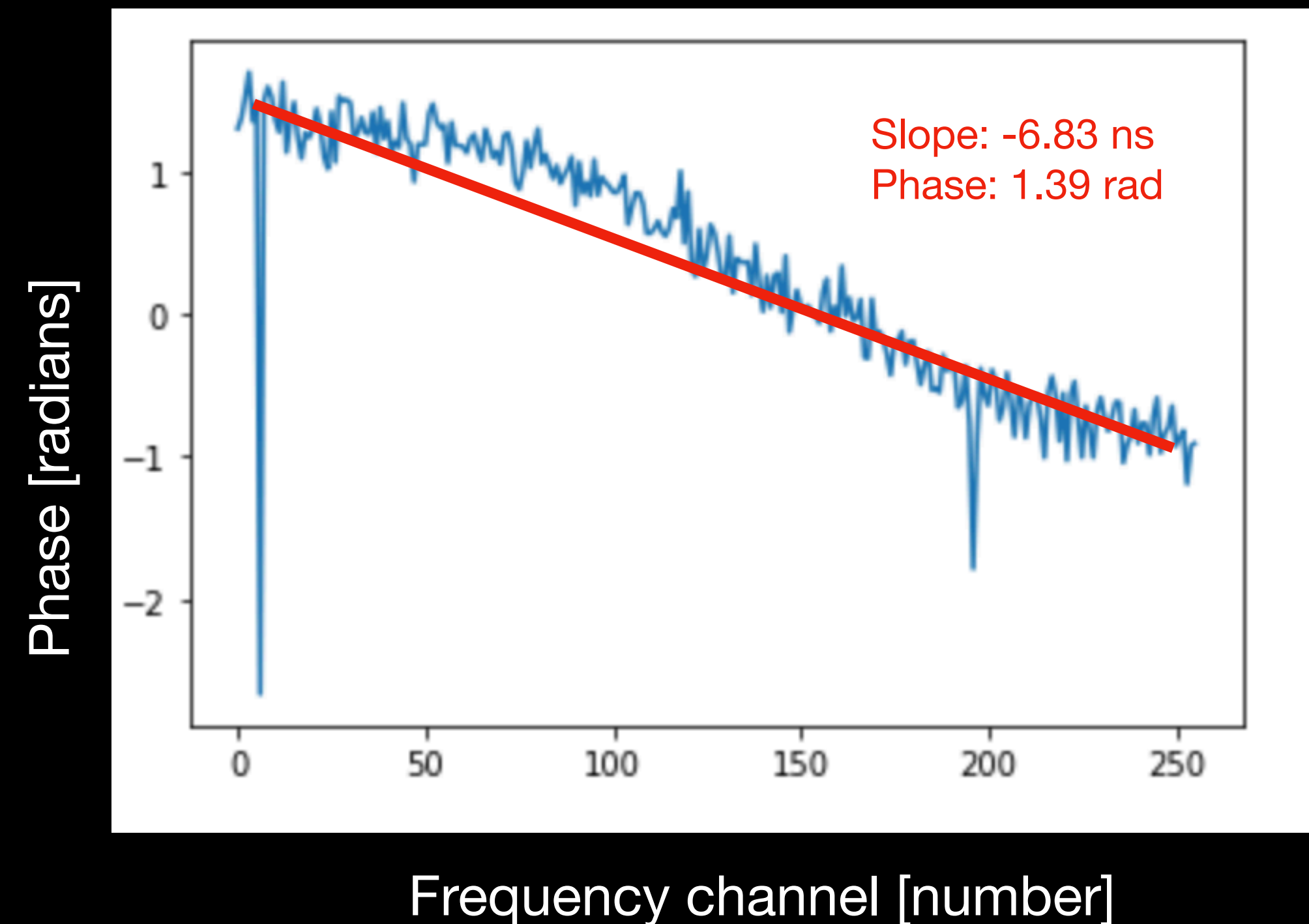
1. Cross-correlate antennas, display data + sanity checks
2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays
4. Measure integer-sample fixed delays from lag spectrum.
5. Apply fixed sample delays:



# Calibration (delay+ phase) scheme

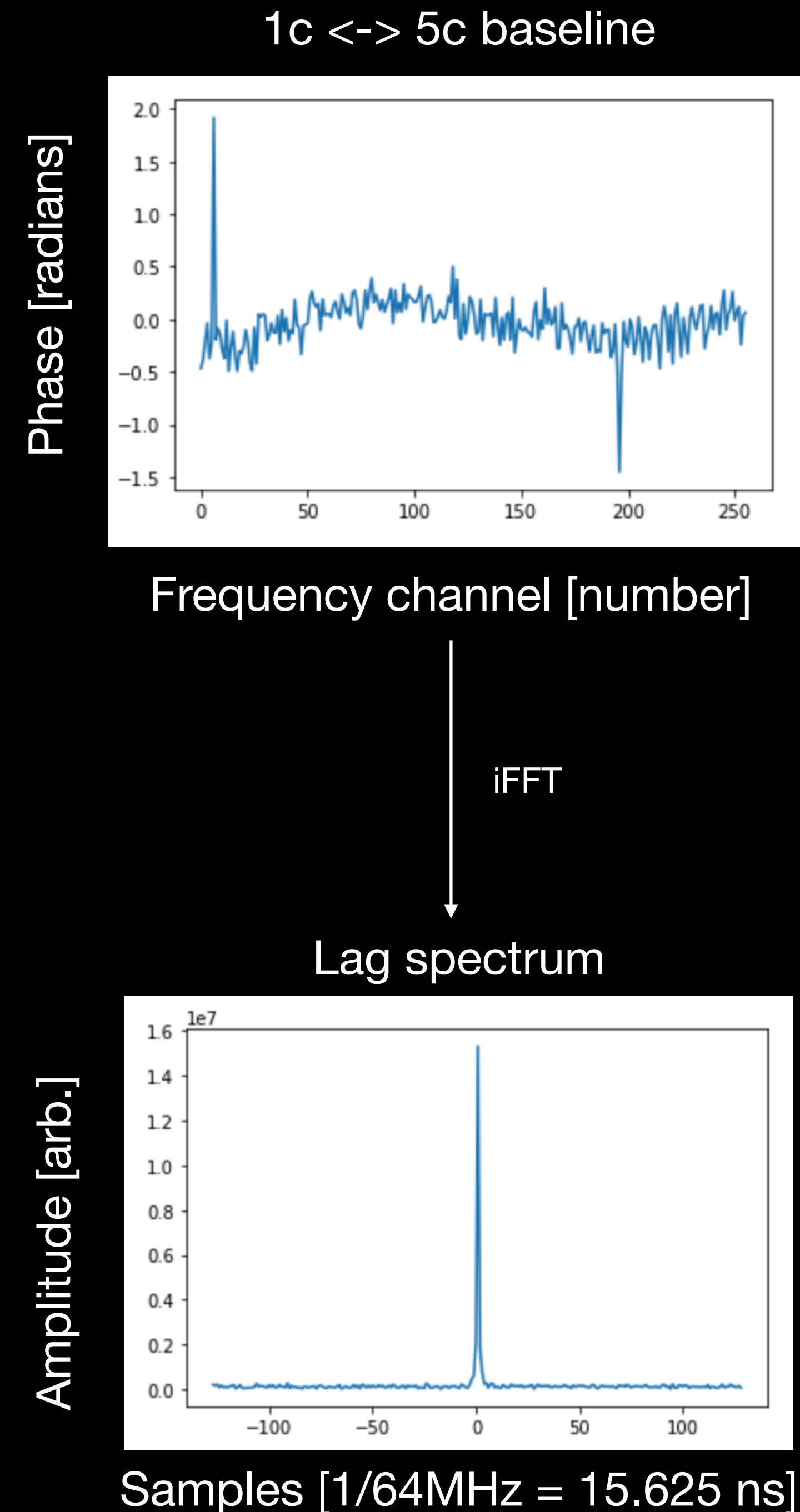
1. Cross-correlate antennas, display data + sanity checks
2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays
4. Measure integer-sample fixed delays from lag spectrum.
5. Apply fixed sample delays
6. Measure sub-sample delay (slope) and measure phase (intercept)

1c <-> 5c baseline

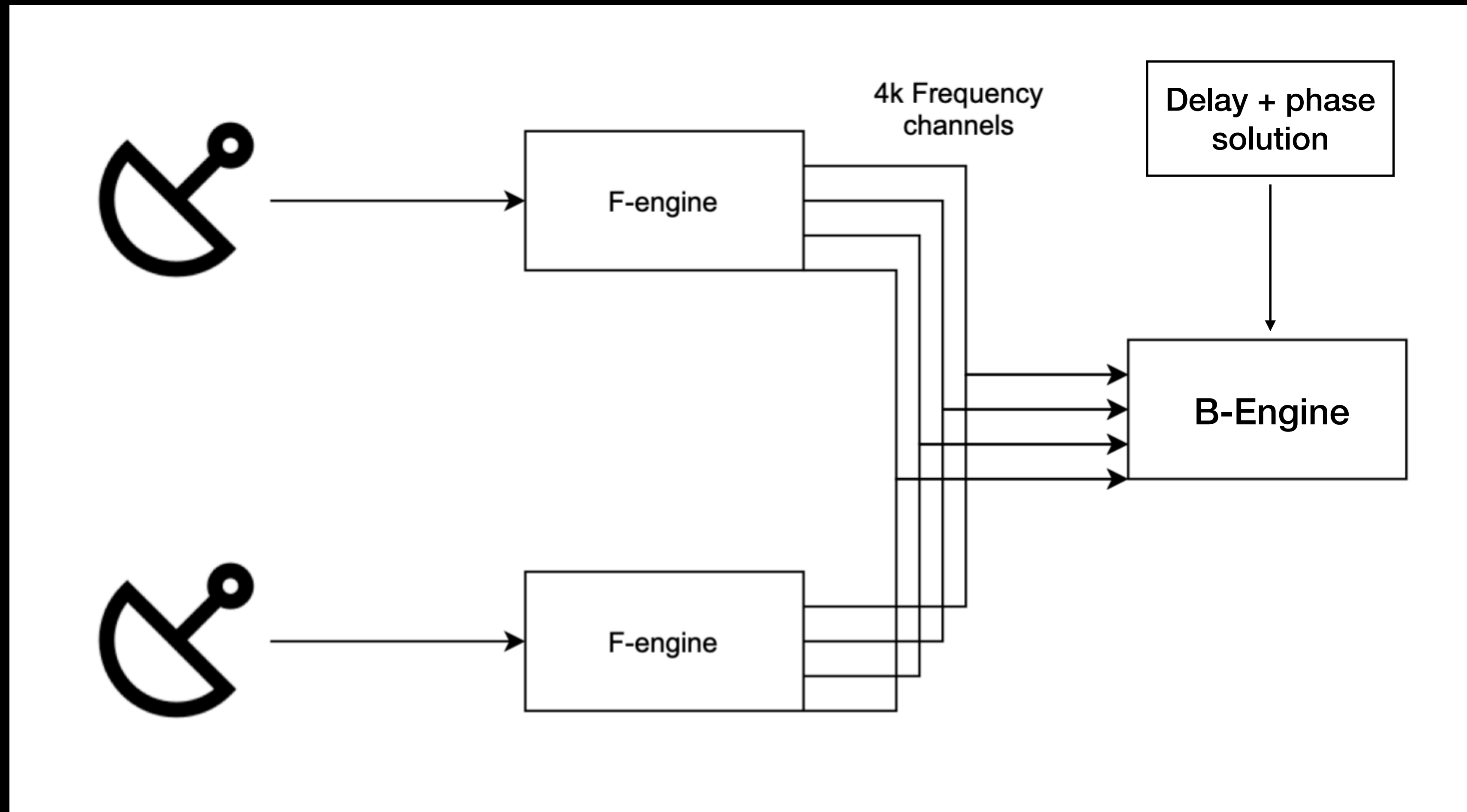


# Calibration (delay+ phase) scheme

1. Cross-correlate antennas, display data + sanity checks
2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays
4. Measure integer-sample fixed delays from lag spectrum.
5. Apply fixed sample delays
6. Measure sub-sample delay (slope) and measure phase (intercept)
7. Apply sub-sample delay + phase
8. Do the above for all other antennas



# Replace Correlator with Beamformer





# (Python/offline )Beamform on sky source

1. Apply geometric + fixed delays + phase offset

(no gain, nor bandpass calibration applied)

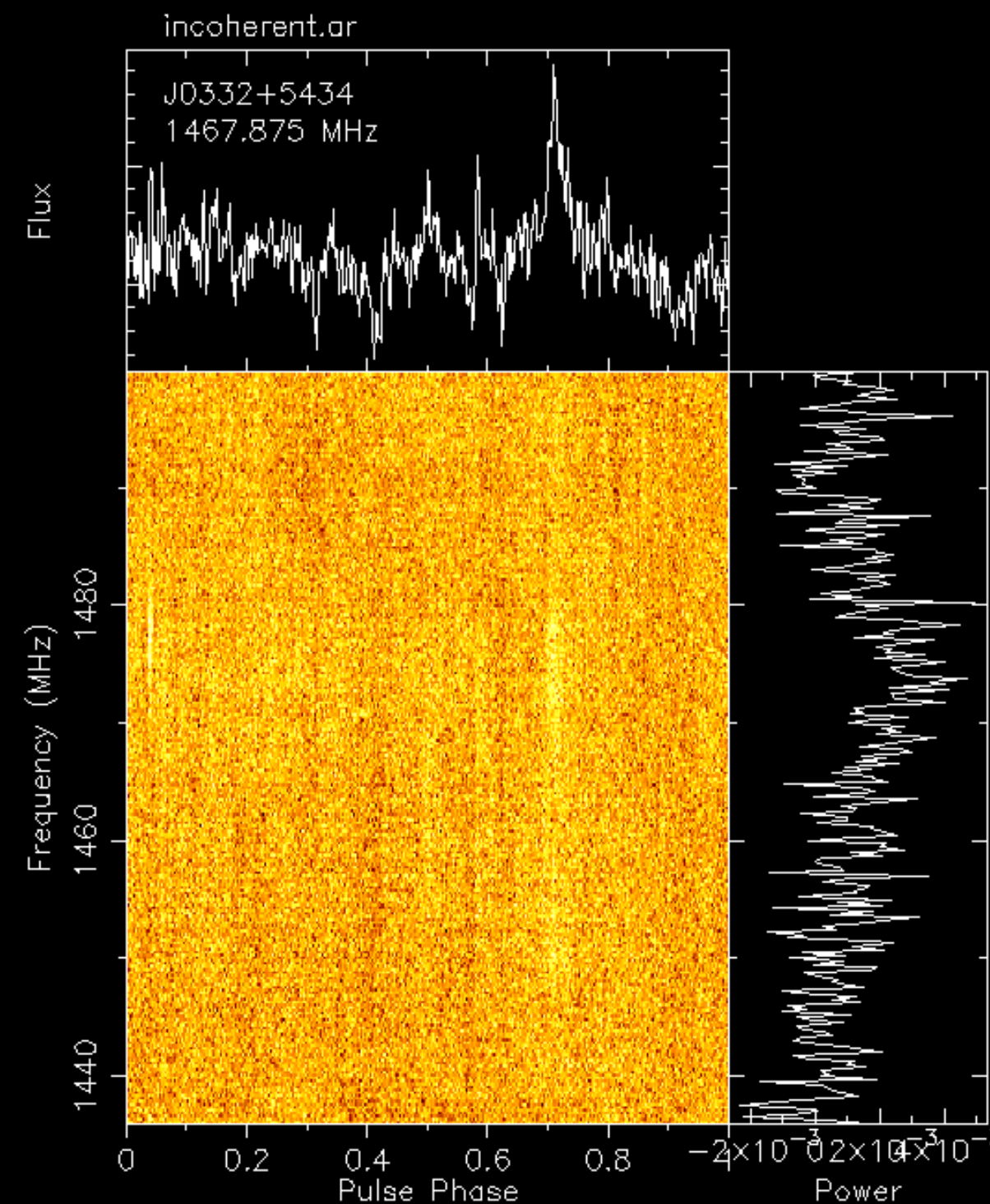
2. Coherent summation (in phase)

3. Incoherent sum (for comparison)

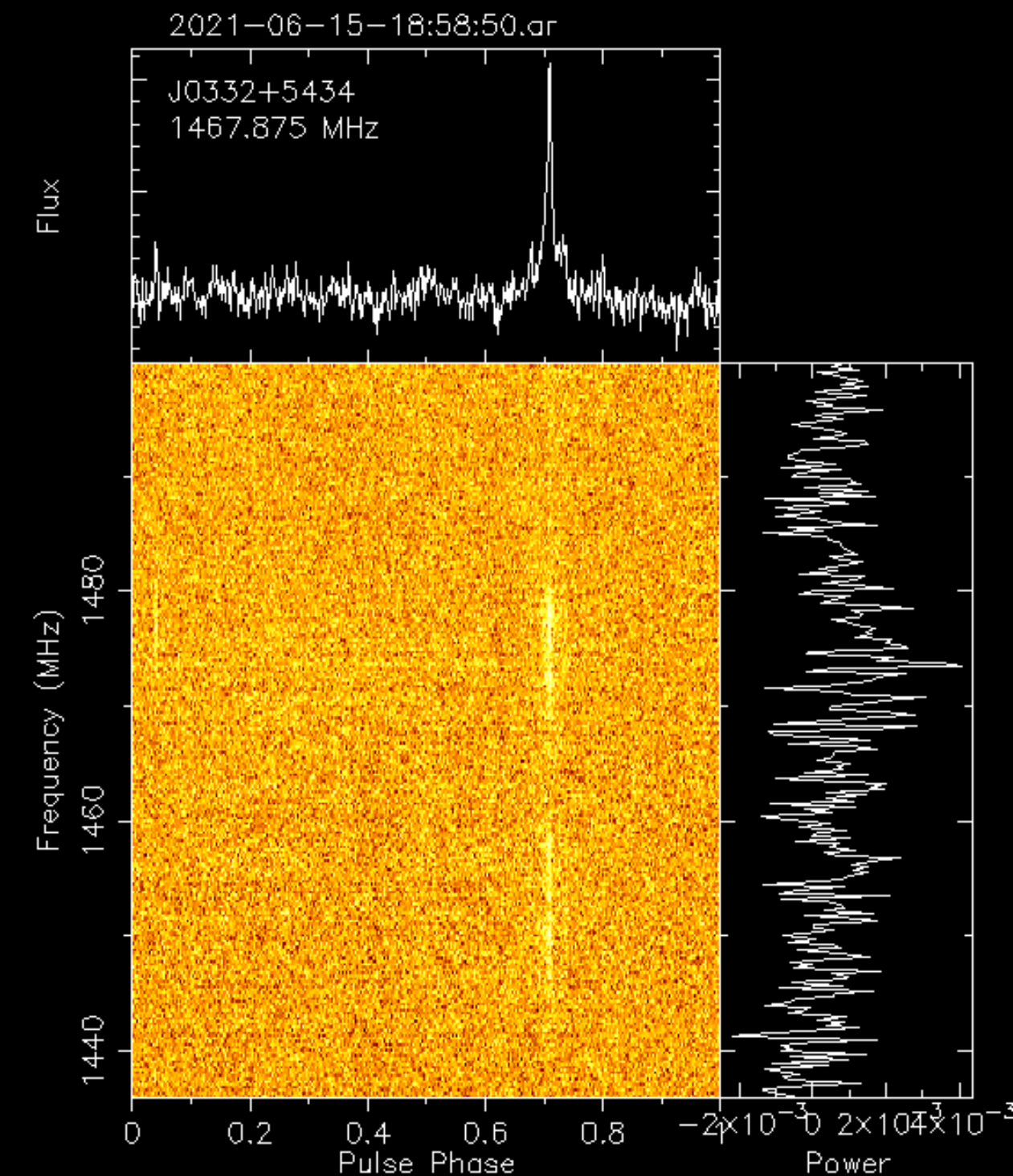
4. Pulsar processing (5 mins obs)

5.  $\sim$  factor of 2 improvement in S/N  
(Theoretical improvement is  $\sqrt{N_{\text{ants}}} = \sqrt{11} = 3.3$ , with perfect beam forming efficiency)

Incoherent sum



coherent sum





# Beamform on sky source (pol-dependent solution)

1. Apply geometric + fixed delays + phase offset

(no gain, nor bandpass calibration applied)

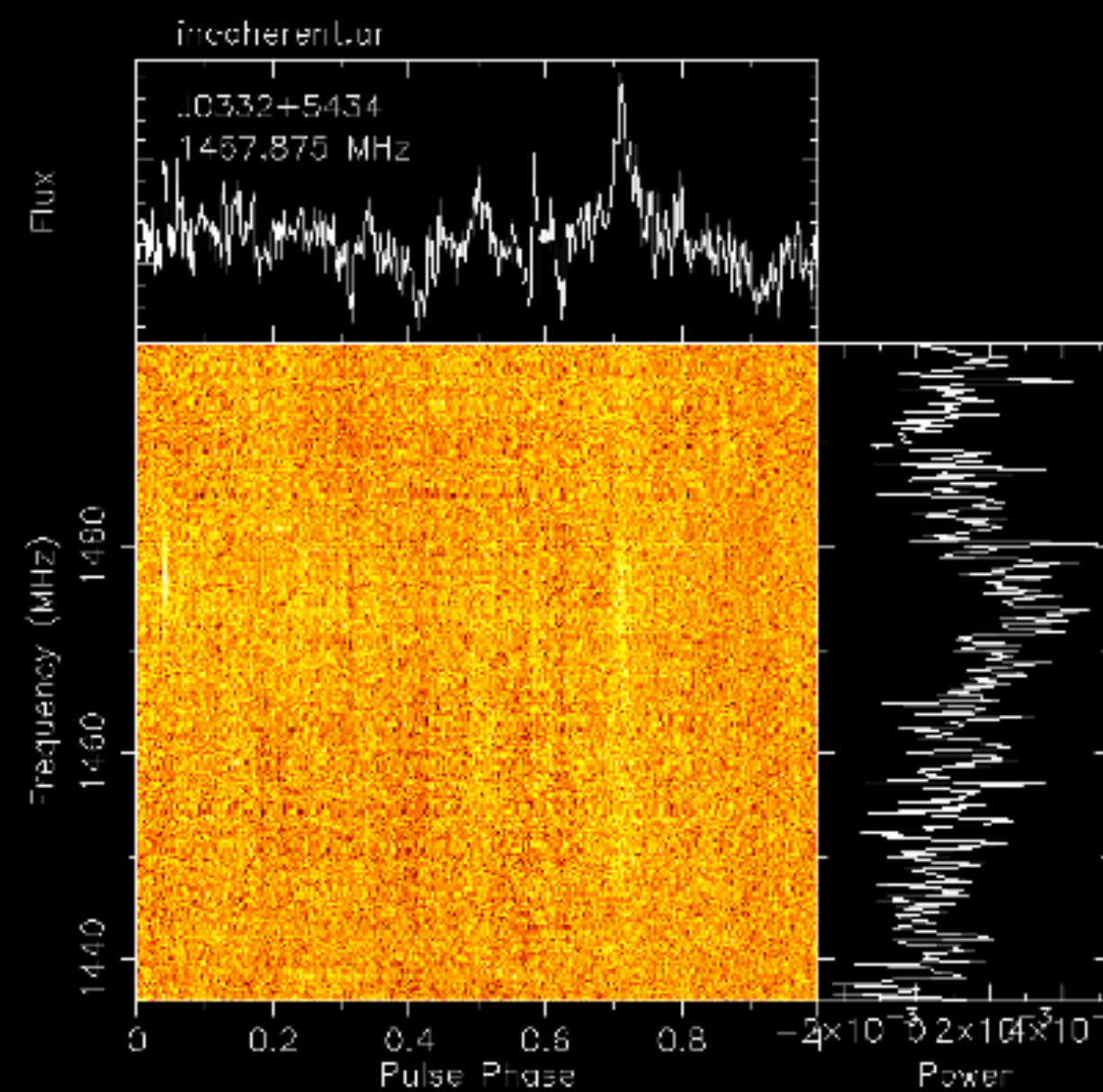
2. Coherent summation (in phase)

3. Incoherent sum (for comparison)

4. Pulsar processing (5 mins obs)

5. Using polarisation-dependent phase solution. S/N scales correctly with number of antpols

Incoherent sum



coherent sum

