

Updates on RFSoC / Attemplifier interface

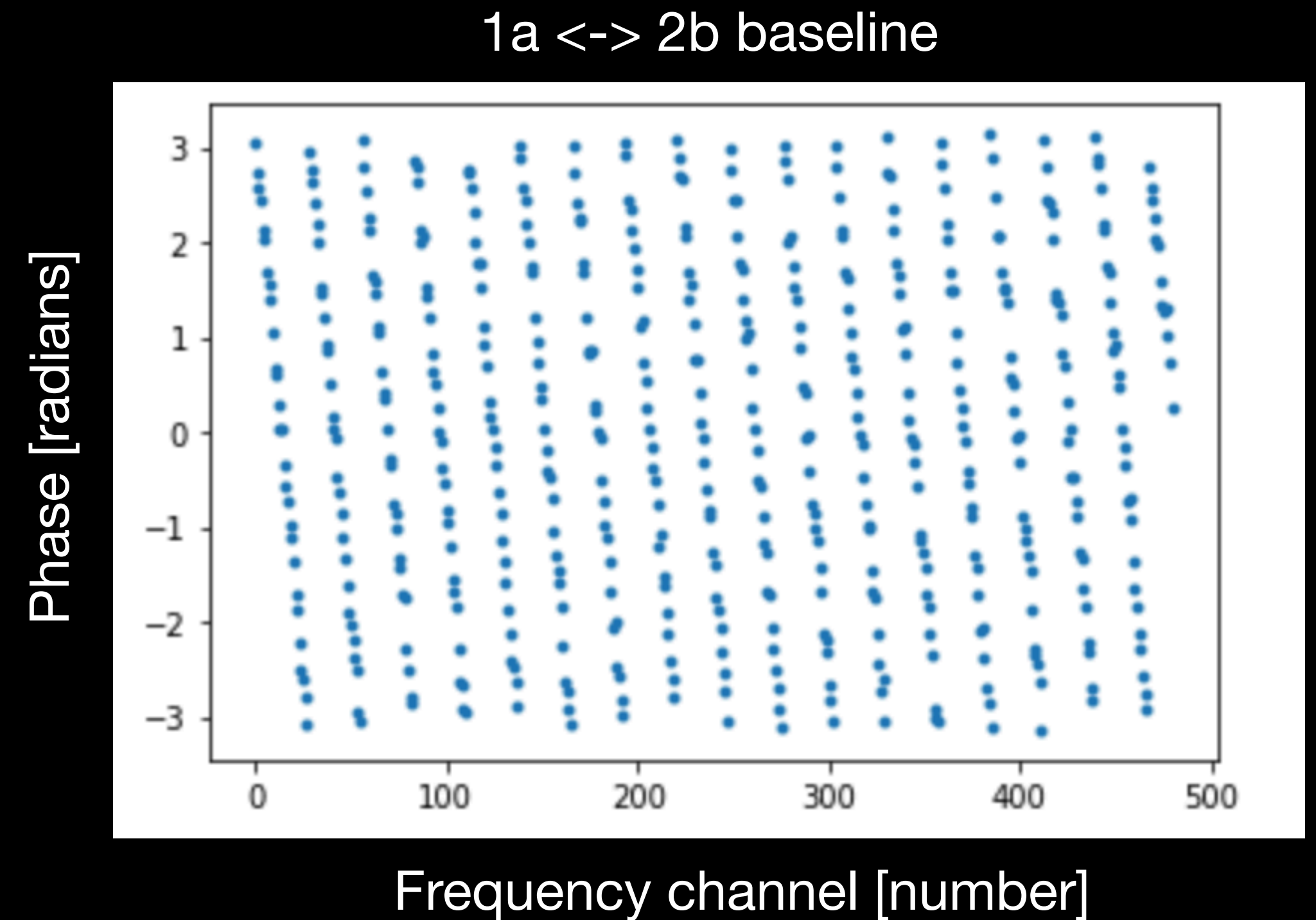
- “Spectrometer” (total power) mode compiled for 4x2 inputs + UDP streaming mode
- Packets are flowing through network.
- Tweaks to firmware are needed to:
 - Make sure UDP destination port is configurable for every output pair
 - Interface to control RFSoC boards as similar to SNAPs as possible
- Current backend receiver codes ingest data
- Connect new antennas to board
- attemplifiers <-> rfsoc <-> recorders mapping established.
 - + get-ing and set-ing attenuation values
- IF-balancing logic to deal with RFSoC + SNAP boards at the same time (testing changes to configuration files, + python modules, etc...)
- What's still needed:
- Tests listed in Jack's document underway:
 - Input panel <-> ADC <-> pipeline number mapping
 - + cross-talk measurement
- Collect data with backend:
 - perform on/off measurement
 - perform pulsar observation

Update on high-speed data capture

- Current voltage-streaming mode can ingest **4 GB/s** on each NIC
Would like to push it up to at least **8 GB/s** (64 gbits/s)
- Kernel-bypassing on Connect X5 NICs.
libvma + libibverbs
- Current status:
 - ping-pong between IBV builtin server and client
 - Flow creation within voltage recorder
 - “Work completion” (i.e. packet reception) errors out

Updates on Correlator

- Written “toy” python correlator.
- Collected “golden” data on bright quasar (3c84)
- xGPU + hashpipe:
 - Correlator compiled for 4-bit mode
 - Can run in “live” mode
- Testing 4-bit correlator:
 - Writing: disk \leftrightarrow RAM hashpipe threads to load and offload data easily
 - Investigate source of packet drops (turns out package was compiled with debug flag)
- Imaging software for faster display of data
- Live data flow + correlator
- GPU Complex multiplier for delay and phase compensation (currently happening offline post-correlation)

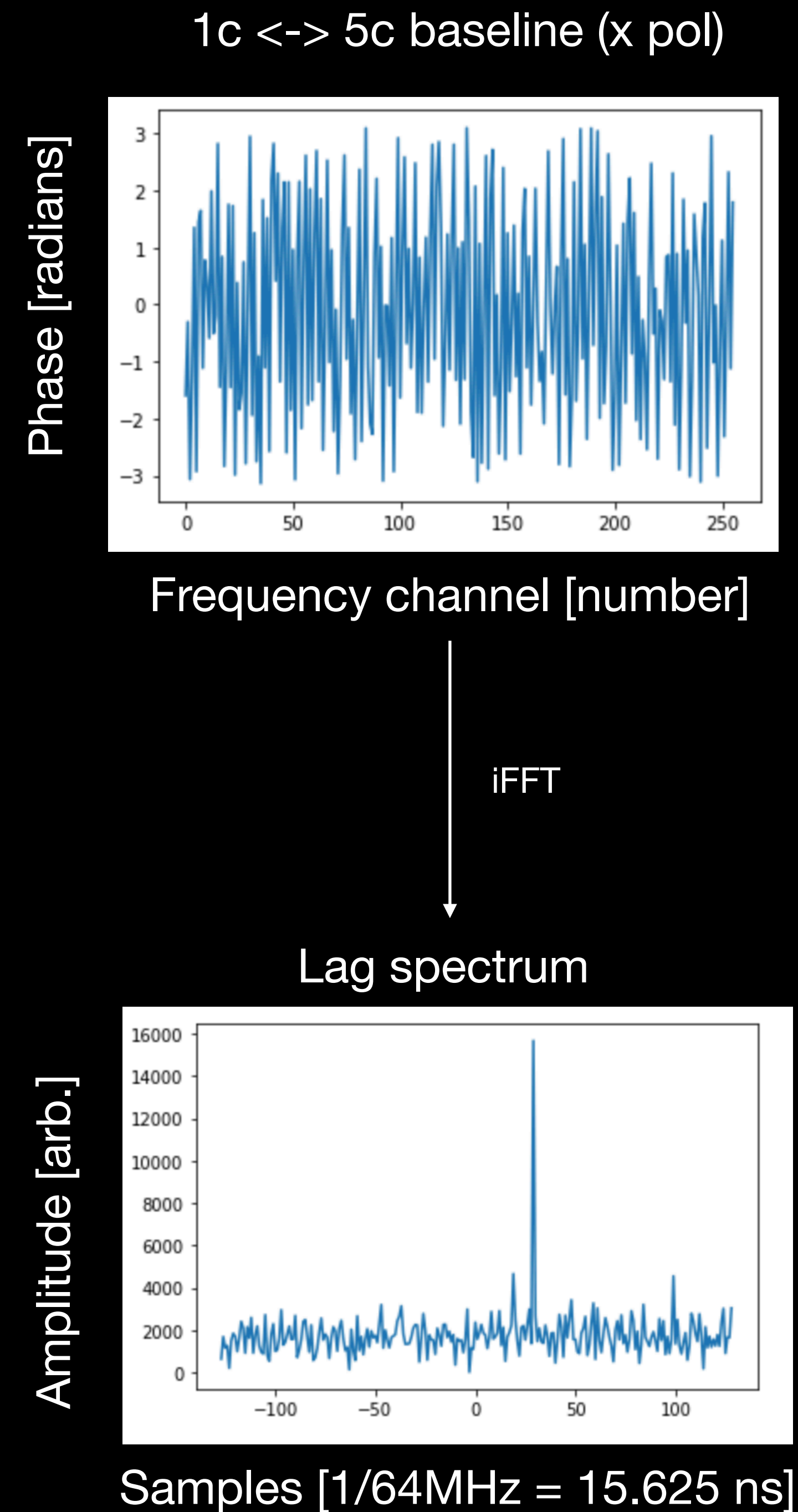


Prototype beam-former

- Written prototype python beam-former.
- Collected golden data on bright quasar (3c84) + Pulsar J0332+5434 in voltage mode:
- Observation:
 - 11 antennas (SNAP boards)
 - 64 MHz x 2 of bandwidth (256x2 channels)
 - 1.5 GHz sky frequency
 - 5 mins on each source
 - 1 + 1 TB of data
- Calibrate on bright quasar
- beam-form on pulsar

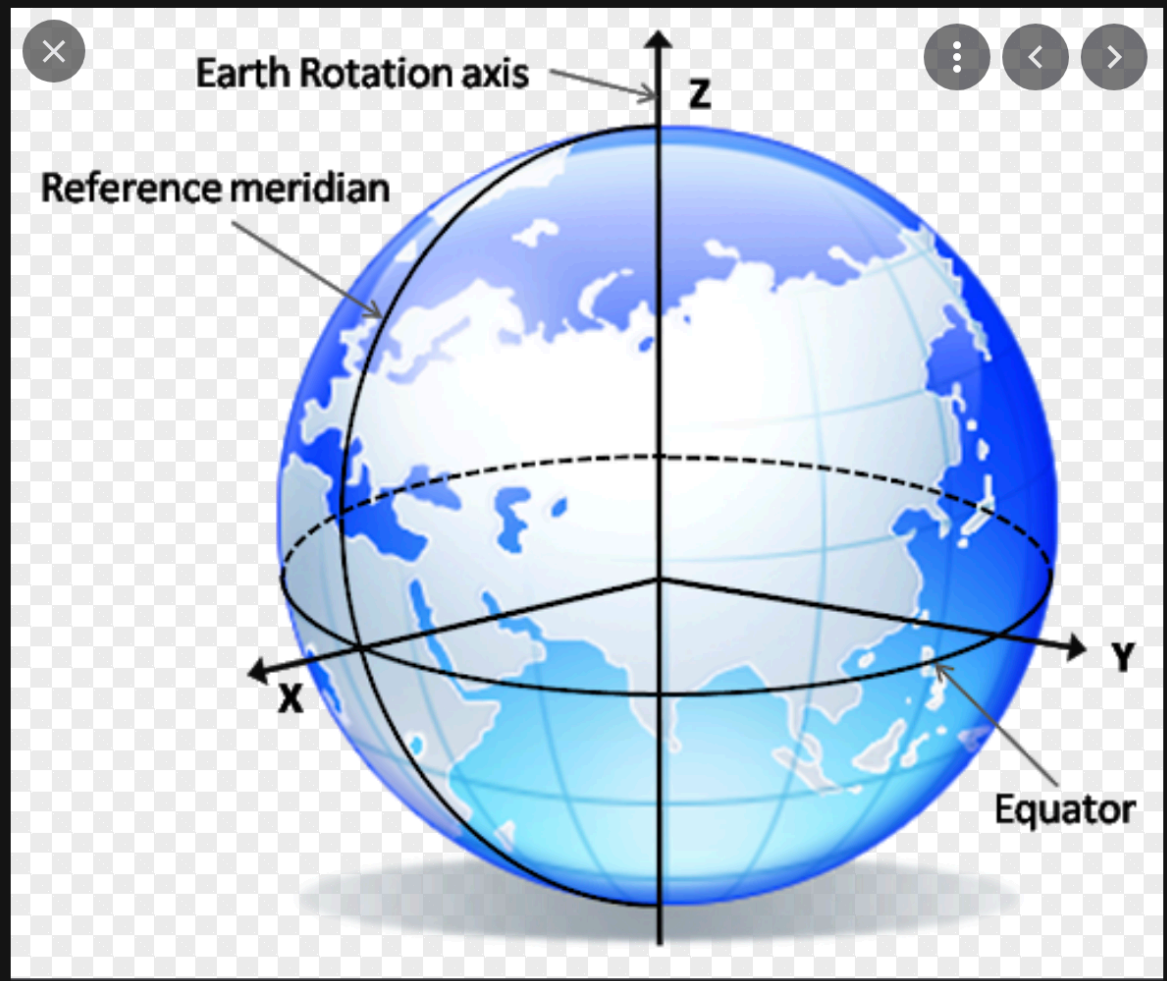
Calibration (delay+ phase) scheme

1. Cross-correlate antennas, display data + sanity checks



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- 1. Cross-correlate antennas, display data + sanity checks
- 2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)

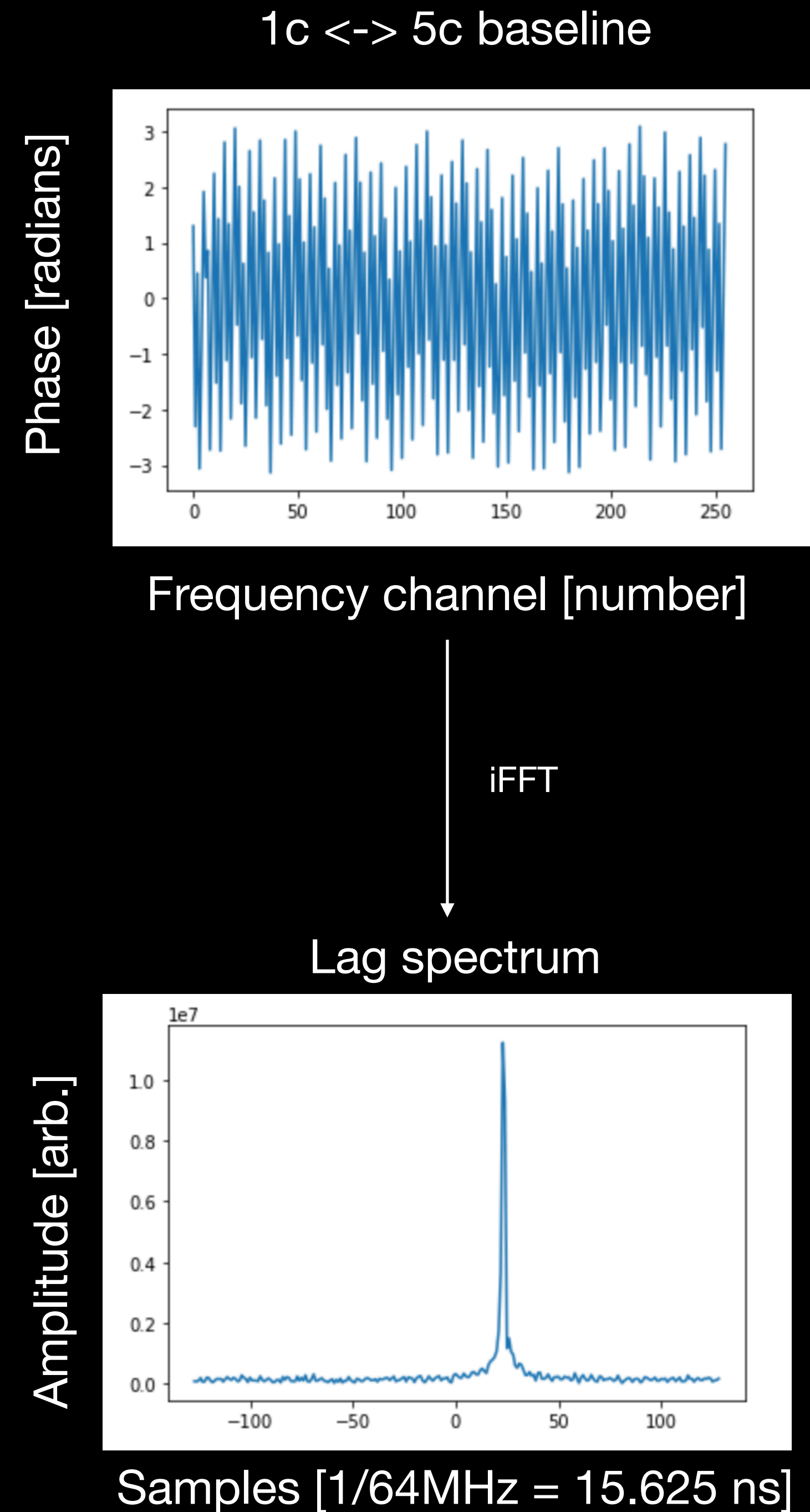
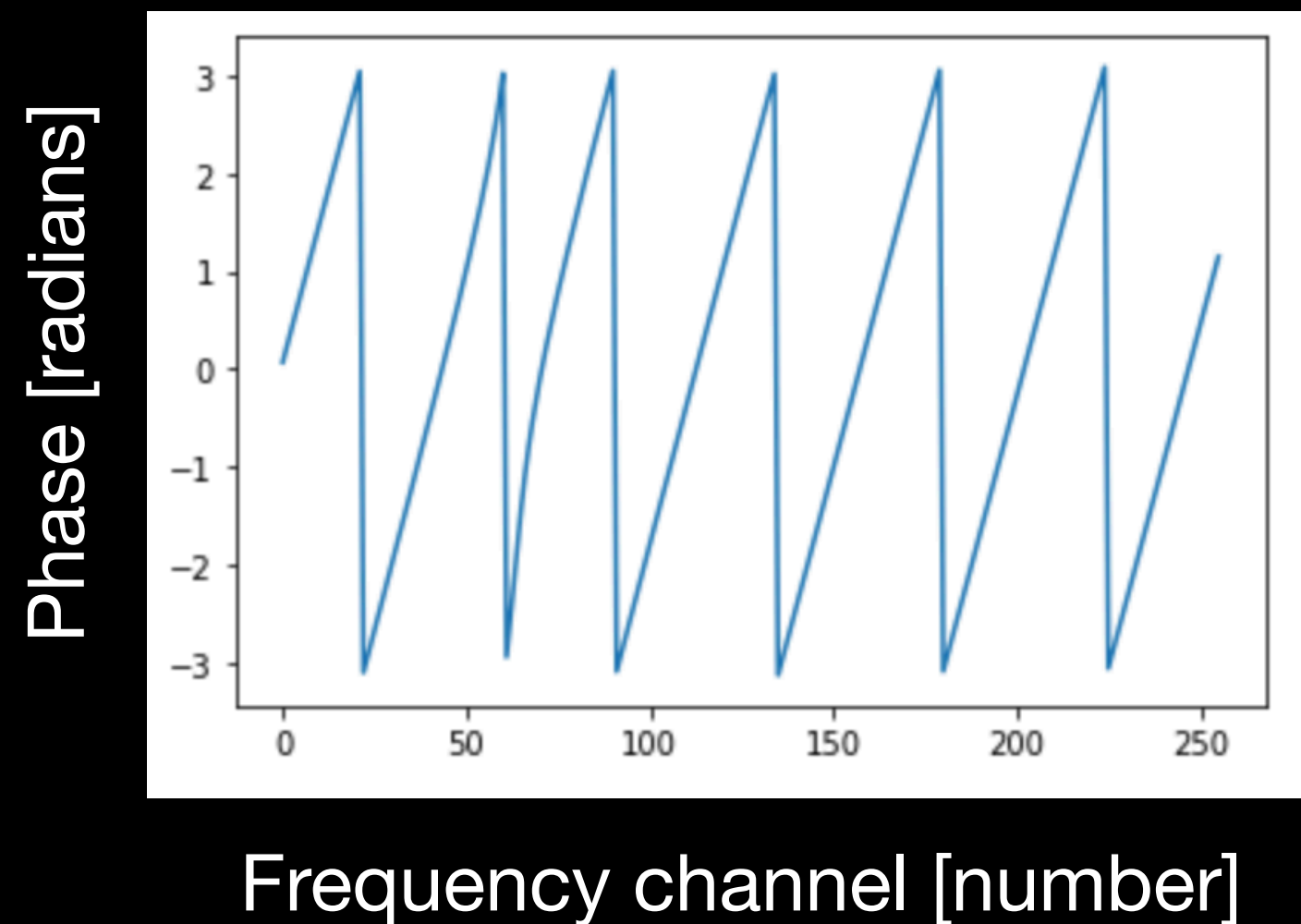


ITRF (ECEF; earth centered earth fixed)
antenna positions XYZ

	x	y	z
1A	-2.524136e+06	-4.123448e+06	4.147709e+06
1F	-2.524193e+06	-4.123464e+06	4.147658e+06
1C	-2.524141e+06	-4.123508e+06	4.147649e+06
2A	-2.524086e+06	-4.123417e+06	4.147769e+06
4J	-2.523998e+06	-4.123376e+06	4.147863e+06
2H	-2.524174e+06	-4.123388e+06	4.147745e+06
3D	-2.524045e+06	-4.123400e+06	4.147811e+06
5B	-2.523924e+06	-4.123448e+06	4.147837e+06
1K	-2.524156e+06	-4.123435e+06	4.147709e+06
5C	-2.523925e+06	-4.123461e+06	4.147824e+06
1H	-2.524203e+06	-4.123431e+06	4.147685e+06
2B	-2.524070e+06	-4.123435e+06	4.147761e+06

Calibration (delay+ phase) scheme

1. Cross-correlate antennas, display data + sanity checks
2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays:

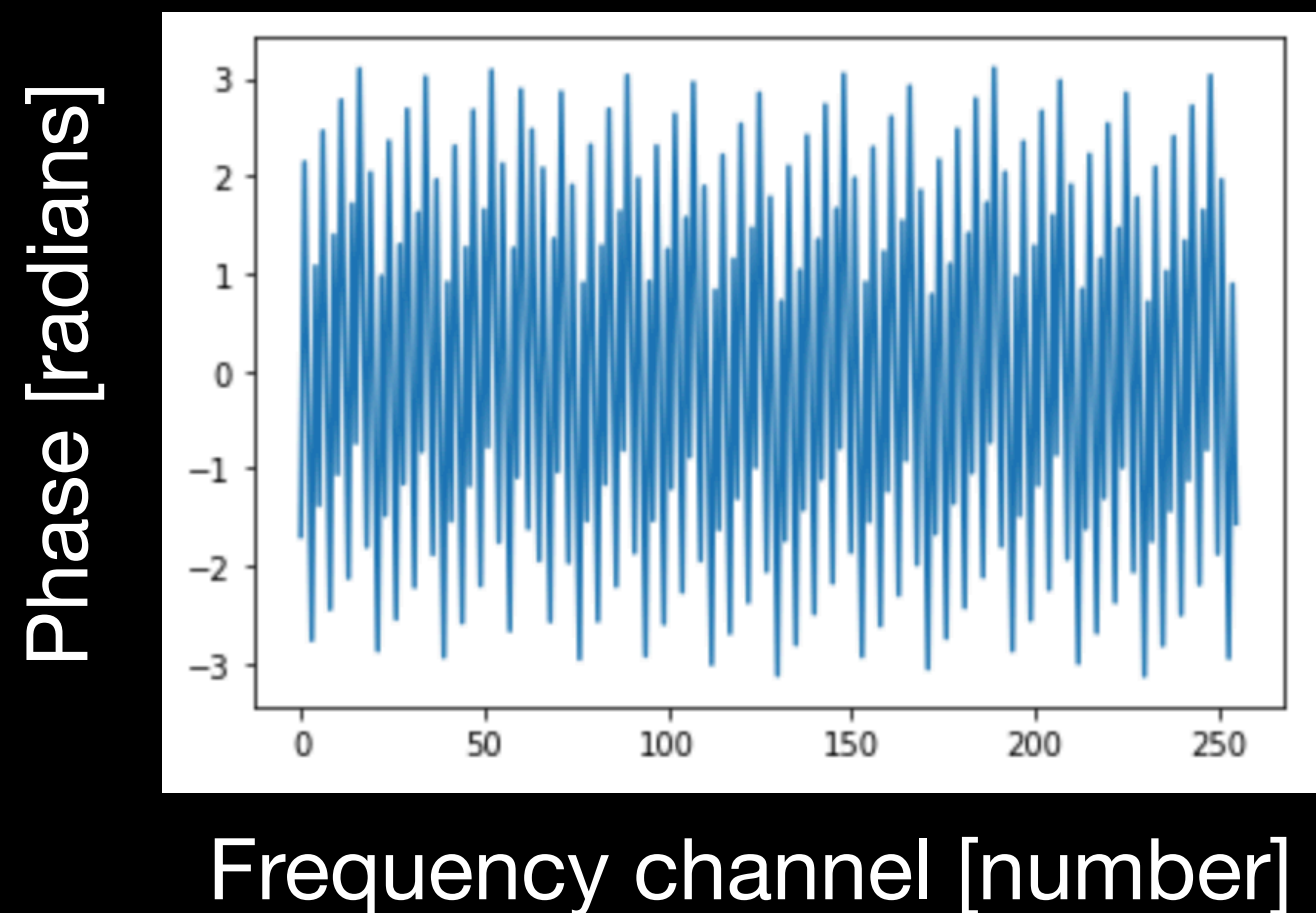


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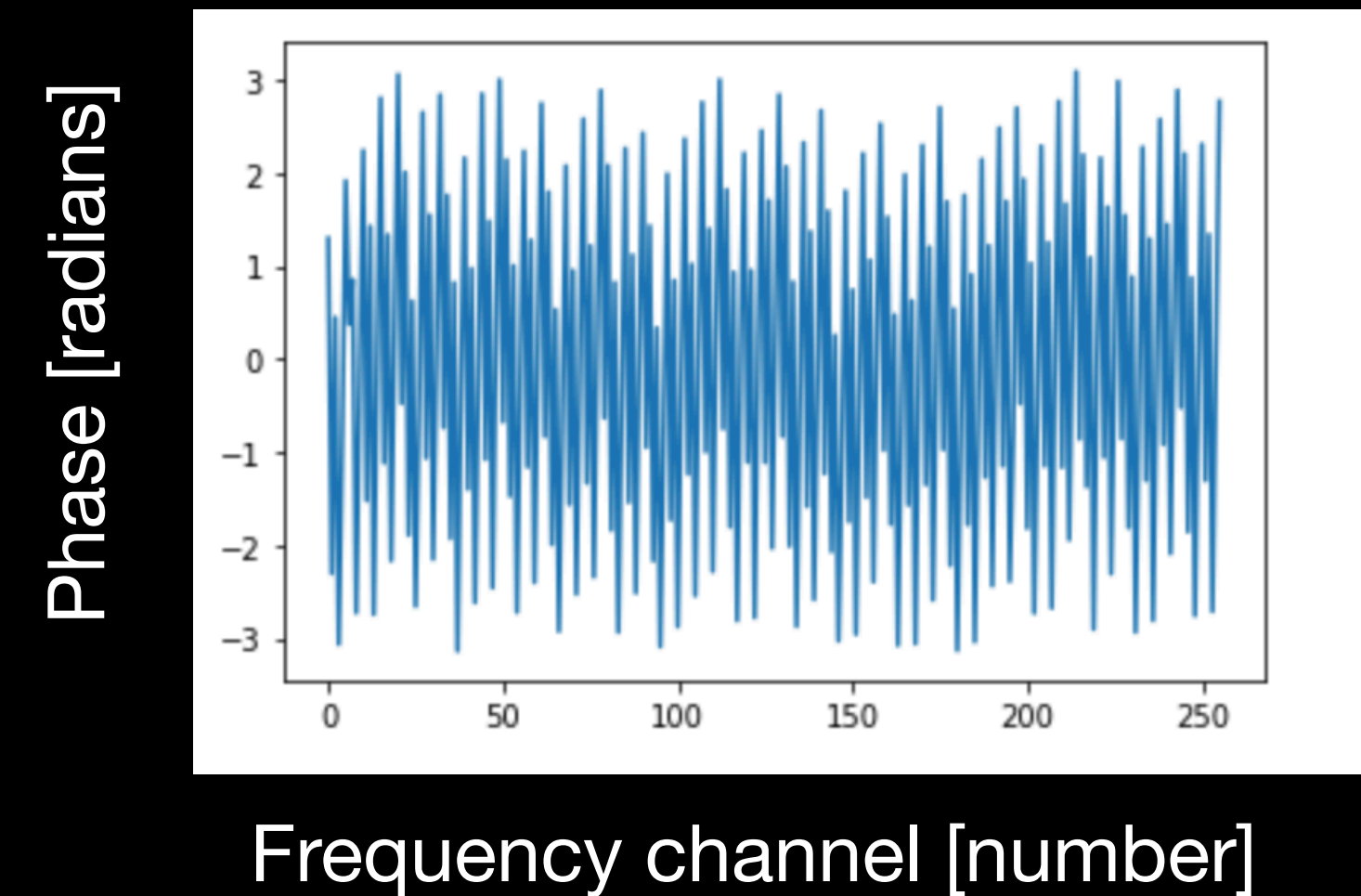
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4. Measure integer-sample fixed delays from lag spectrum.

22 samples = 343.75ns

Fixed delay phasor:

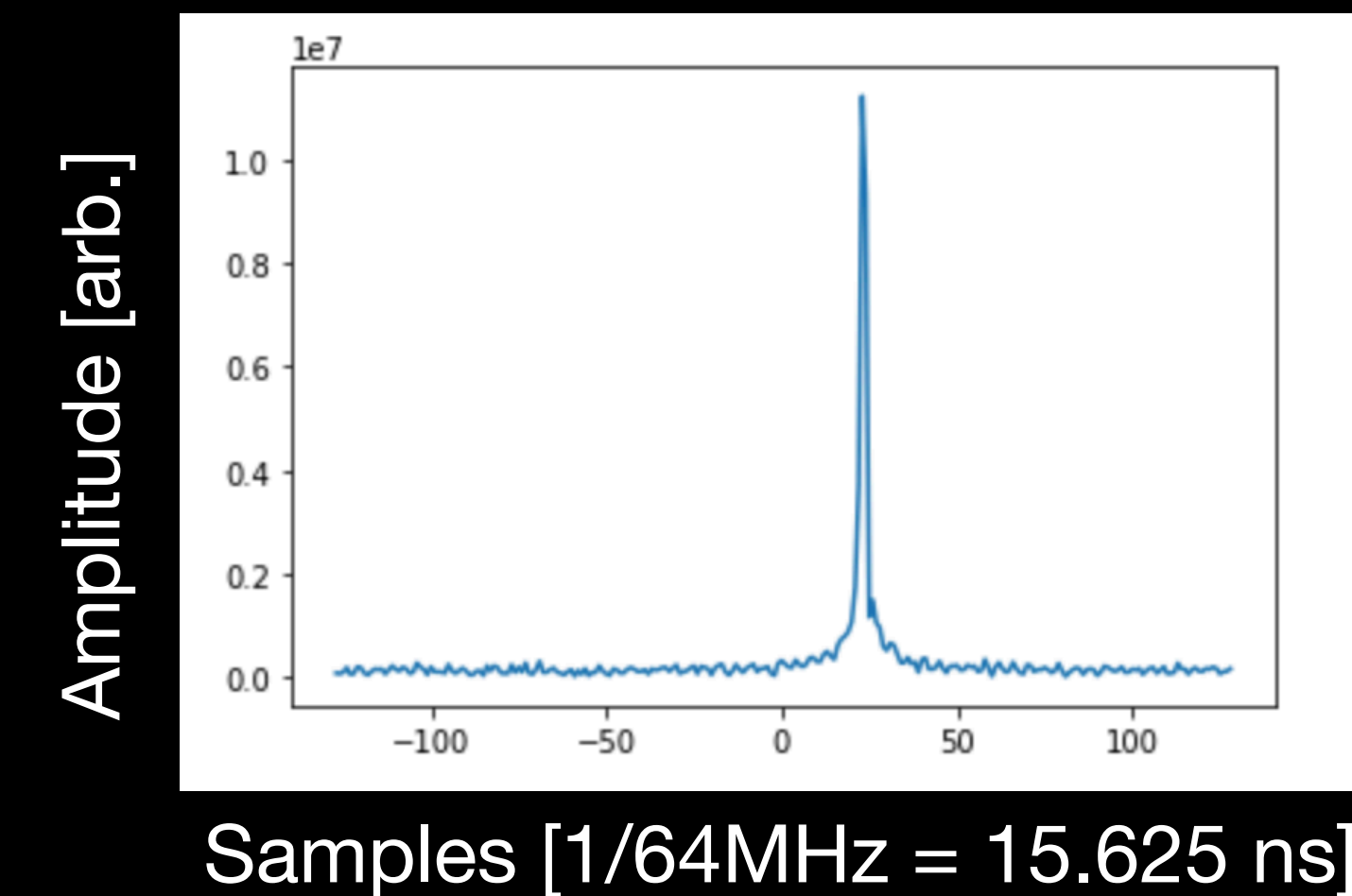


1c <-> 5c baseline



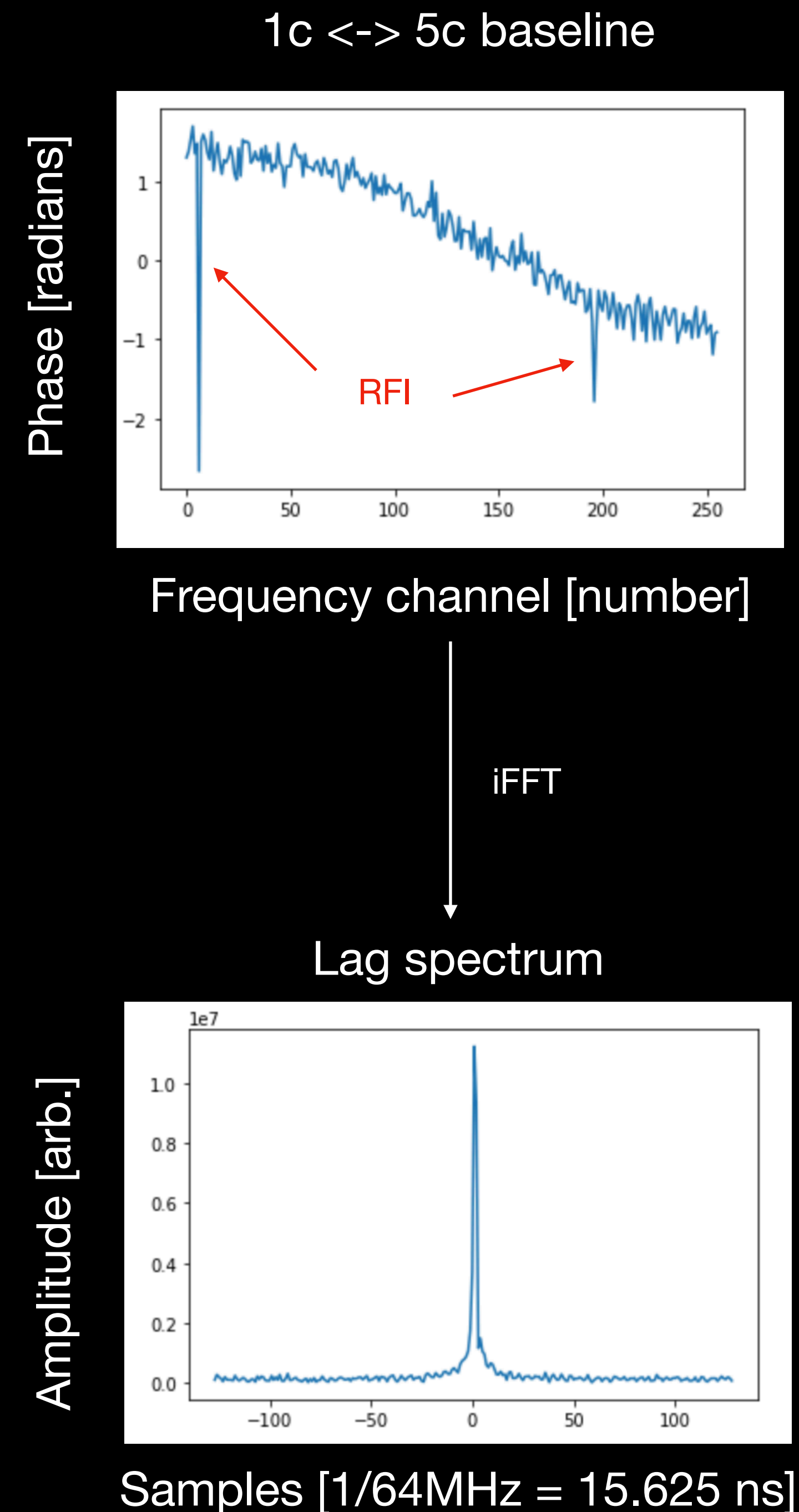
iFFT

Lag spectrum



Calibration (delay+ phase) scheme

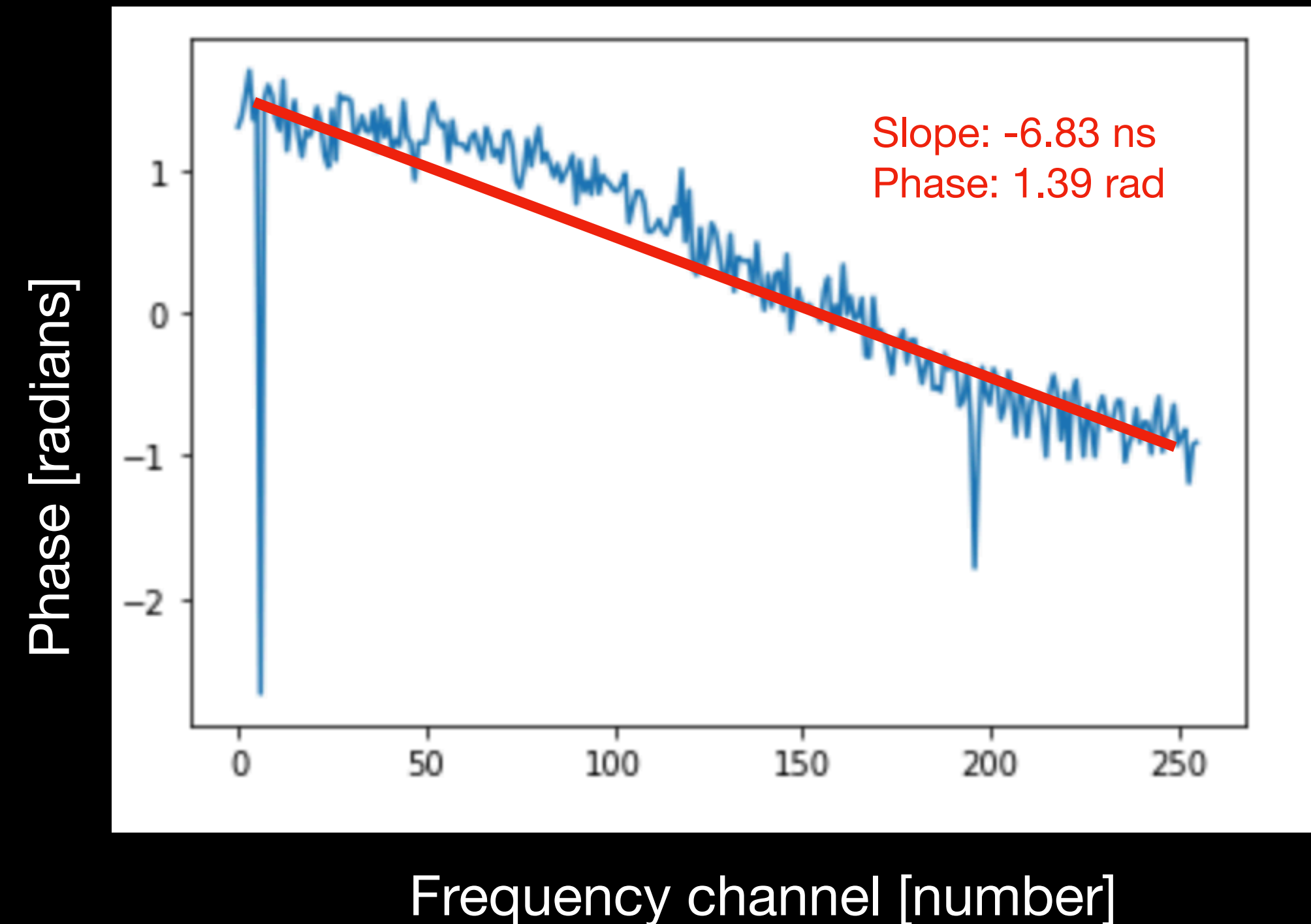
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5. Apply fixed sample delays:



Calibration (delay+ phase) scheme

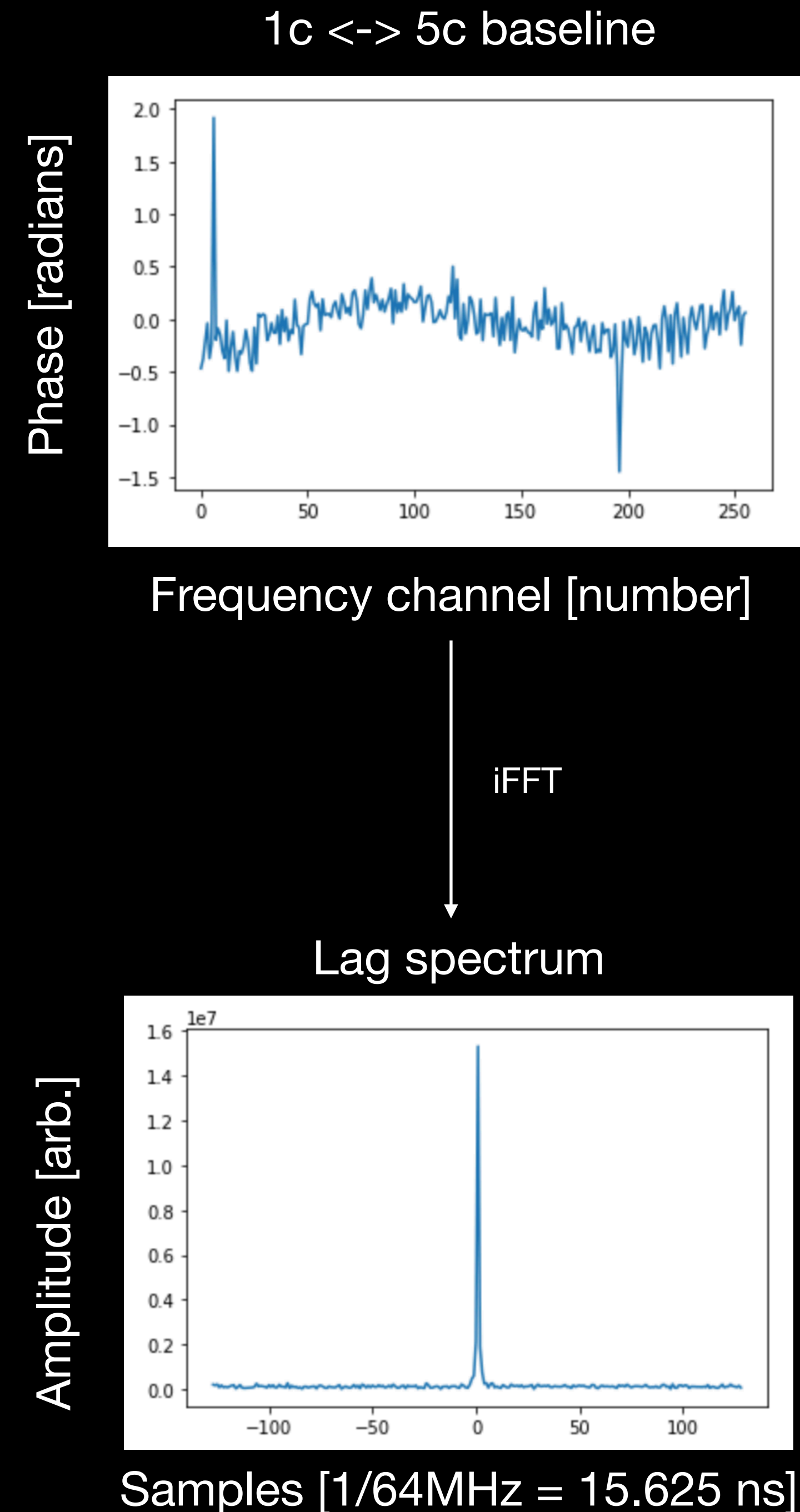
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6. Measure sub-sample delay (slope) and measure phase (intercept)

1c <-> 5c baseline



Calibration (delay+ phase) scheme

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2. Calculate geometric delays given the position of antennas + position of source (and a reference antenna position)
3. Apply geometric delays
4. Measure integer-sample fixed delays from lag spectrum.
5. Apply fixed sample delays
6. Measure sub-sample delay (slope) and measure phase (intercept)
7. Apply sub-sample delay + phase



Beamform on sky source

1. Apply geometric + fixed delays + phase offset

(no gain, nor bandpass calibration applied)

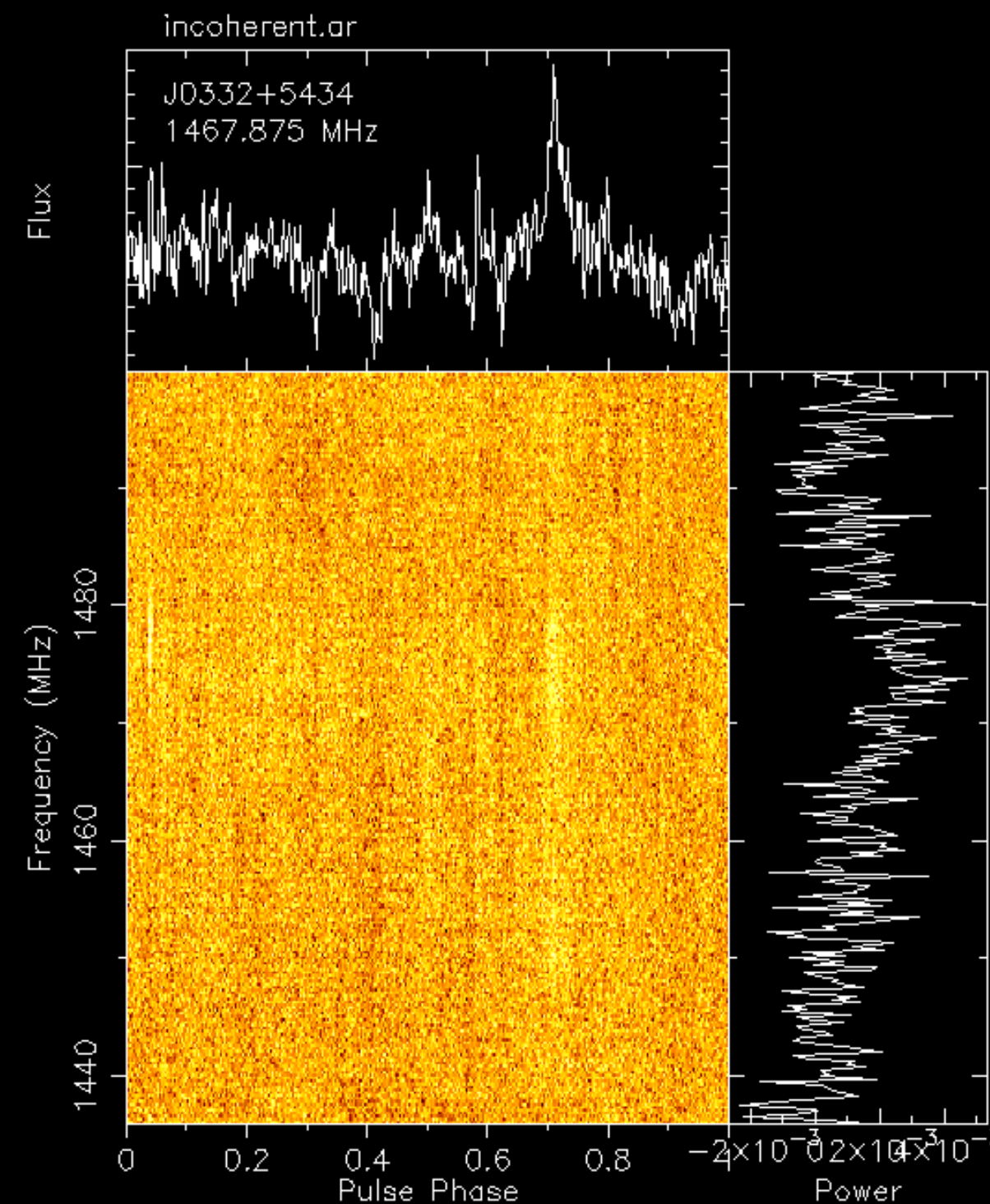
2. Coherent summation (in phase)

3. Incoherent sum (for comparison)

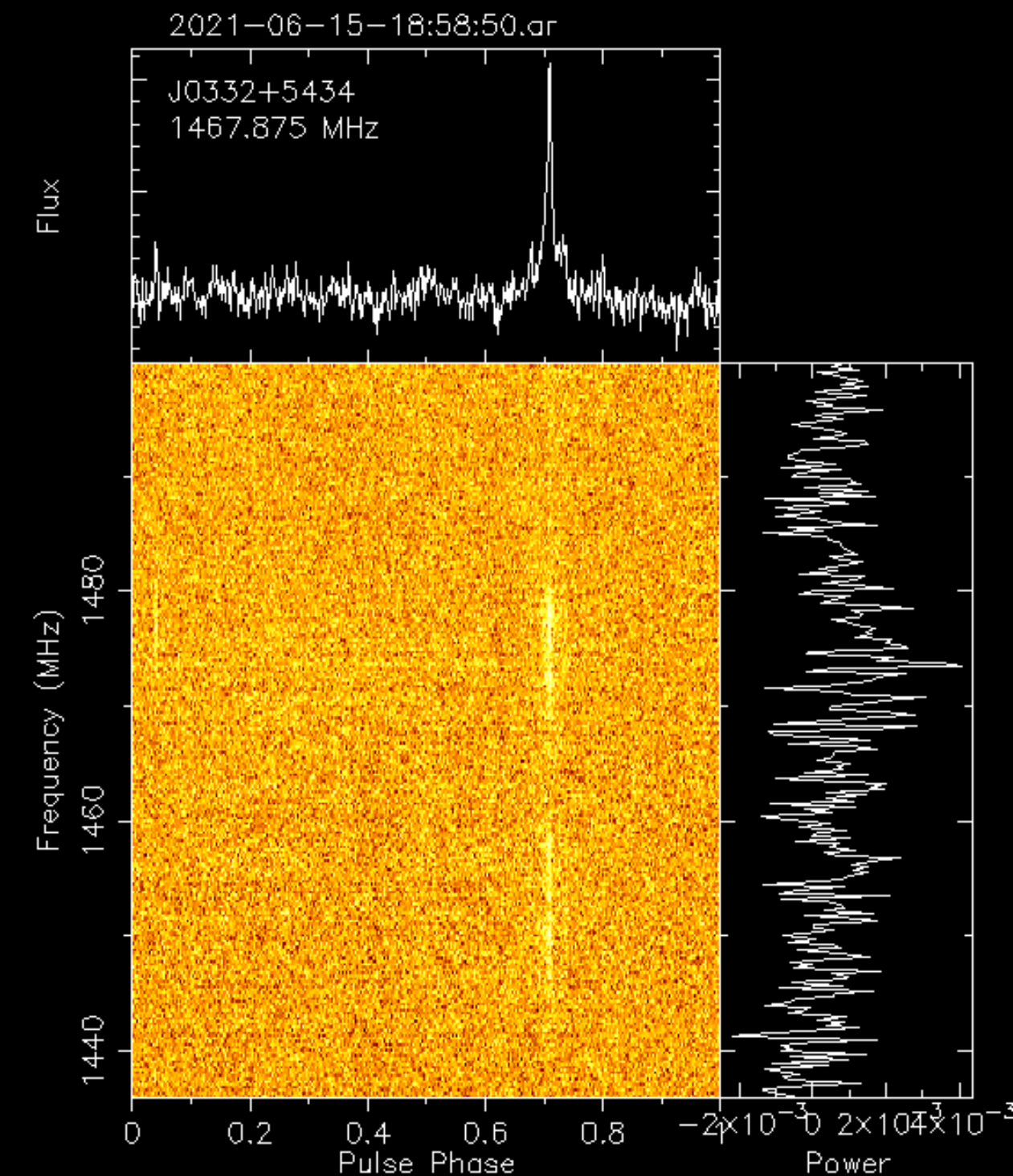
4. Pulsar processing

5. \sim factor of 2 improvement in S/N
(Theoretical improvement is $\sqrt{N_{\text{ants}}} = \sqrt{11} = 3.3$, with perfect beam forming efficiency)

Incoherent sum



coherent sum



Beamform on sky source

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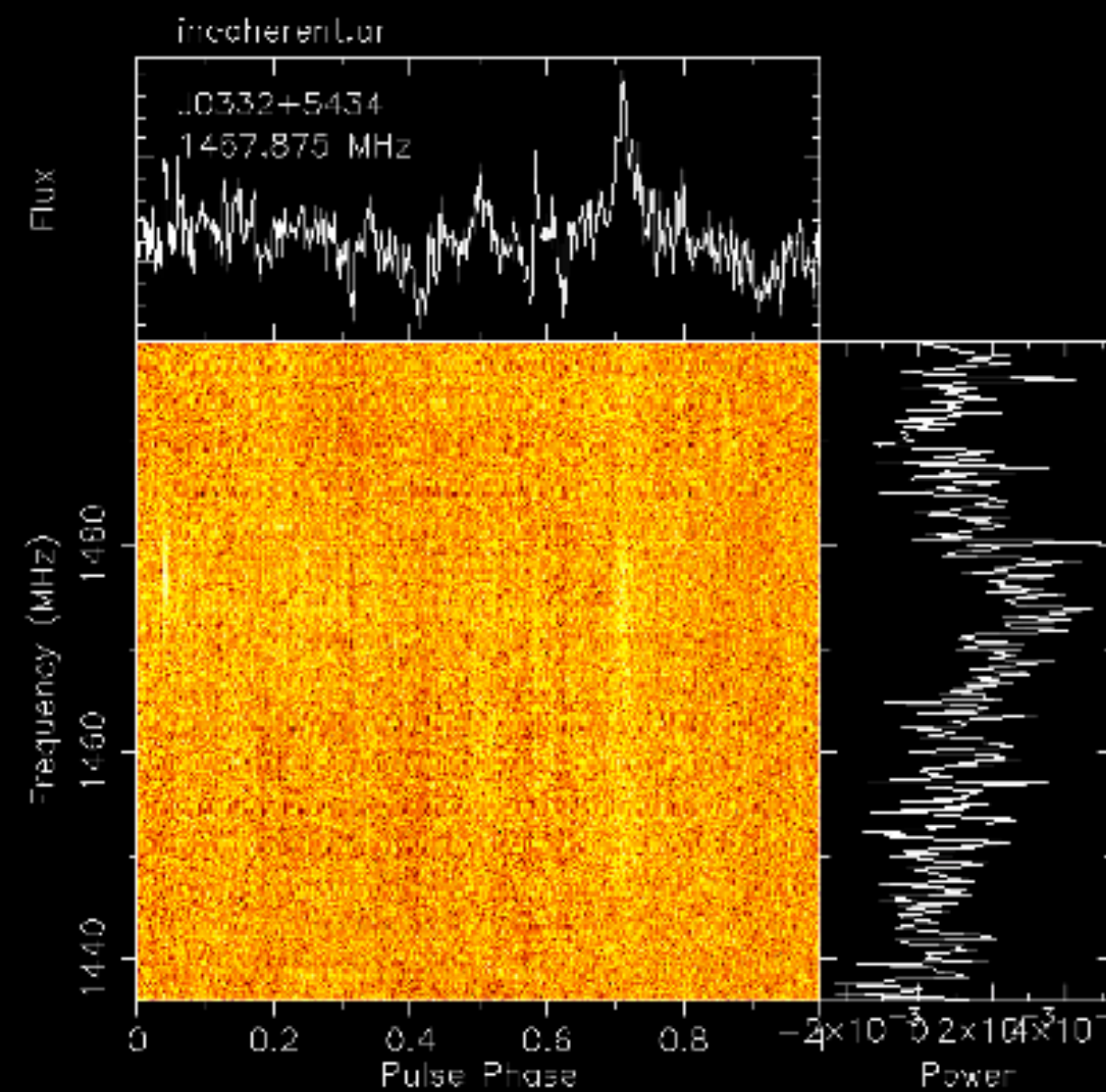
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4. Pulsar processing

5. Using polarisation-dependent phase solution. S/N scales correctly with number of antennas

Incoherent sum



coherent sum

