Classical statistical inference

Part 1

Associated notebook:

03-Basic_statistics_and_proba_concepts/Basic-statistics_01.ipynb

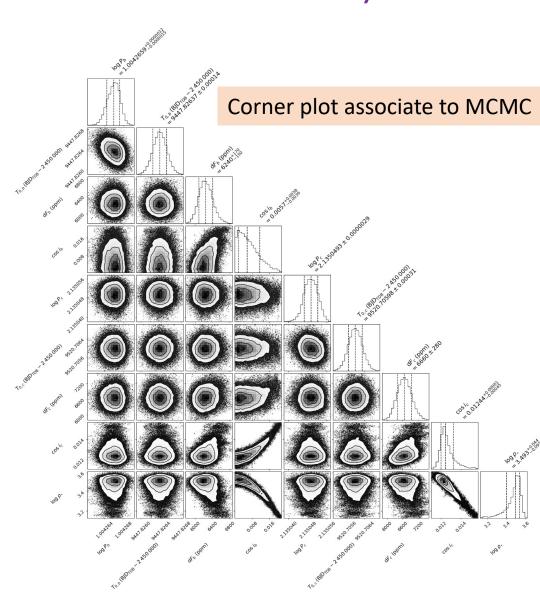
Why some statistics? (Concrete Motivation)

Two temperate super-Earths transiting a nearby late-type M dwarf*

L. Delrez 1, 2, C.A. Murray 3, 4, F.J. Pozuelos 1, 2, 5, N. Narita 6, 7, 8, E. Ducrot 9, **, M. Timmermans¹, N. Watanabe 10, A.J. Burgasser 11, T. Hirano 7, 12, B.V. Rackham 13, 14, ***, K.G. Stassun 15, V. Van Grootel 2, C. Aganze 11, M. Cointepas 16, 17, S. Howell 18, L. Kaltenegger 19, P. Niraula 13, D. Sebastian 20, J.M. Almenara 6, K. Barkaoui 1, 13, 8, T.A. Baycroft 20, X. Bonfils 16, F. Bouchy 17, A. Burdanov 13, D.A. Caldwell 18, 21, D. Charbonneau 22, D.R. Ciardi 23, K.A. Collins 2, T. Daylan 14, 24, 25, et al. (2014) relation between these observables and [Fe/H]. We used a Monte Carlo approach to calculate the uncertainty in our estimate, drawing 10 samples from normal distributions given by the means and standard deviations of the measurements. We calculated the mean and standard deviation of the resulting values and, adding in quadrature the systematic uncertainty of the

We performed a fit using NextGen stellar atmosphere models, with the free parameters being the effective temperature $(T_{\rm eff})$ and metallicity ([Fe/H]). The remaining free parameter is the extinction A_V , which we fixed at zero due to the star's proximity. The resulting fit (Fig. 6) has a reduced χ^2 of 1.7, with best fit $T_{\rm eff} = 2850 \pm 75$ K and [Fe/H] = 0.0 ± 0.5 . Integrating the model SED gives the bolometric flux at Earth, $F_{\rm bol} = 4.41 \pm 0.15 \times 10^{-11}$ erg s⁻¹ cm⁻². Taking the $F_{\rm bol}$ and $T_{\rm eff}$, together with the Gaia EDR3 parallax, gives the stellar radius, $R_{\star} = 0.1556 \pm 0.0086$ R_{\odot} .

relation (0.07), we arrived at our final metallicity estimate of



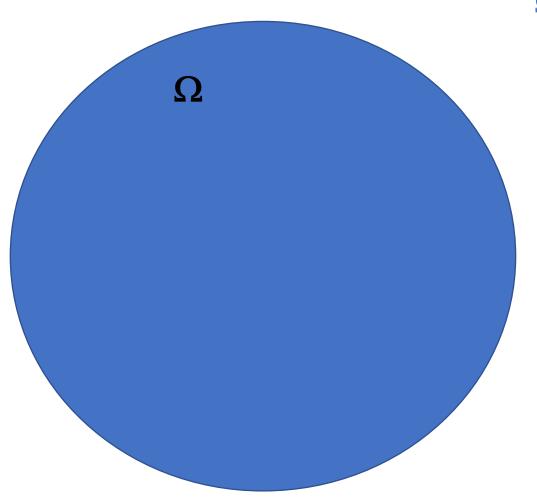
Why some statistics?

- Python for data (observation / numerical simulations) manipulation
- Data most often contain a stochastic component: observational device, numerical noise, simulation of stochastic process, ...

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⇒Data ≈ Random variable (RV)
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- Statistics is the tool needed to manipulate RV. Python allows one to have a more practical understanding of statistics.
- Main Goals:
 - Uncertainty calculation (no, this is not black magic)
 - Make prediction based on data modelling (first step towards machine learning)

http://www.wooclap.com/RSLBIT



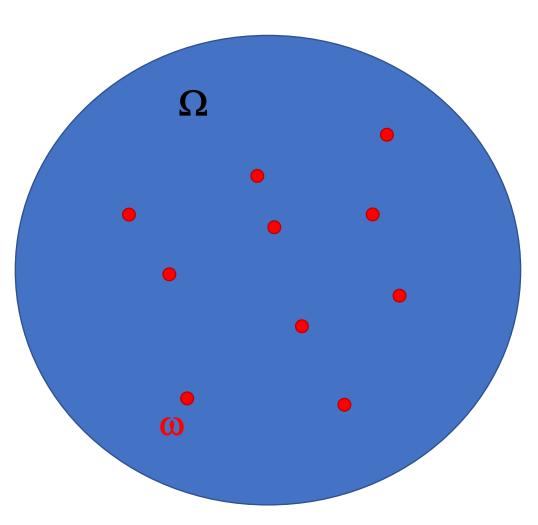
 Ω : Sample space \equiv all possible outcome of an experiment

e.g. of experiment

- I measure the magnitude of a star (in a binary system, for a transit, ...)
- I count galaxies for different L at a given z
- I obtain the spectrum of a candidate SN
- I measure a GW signal

- ...

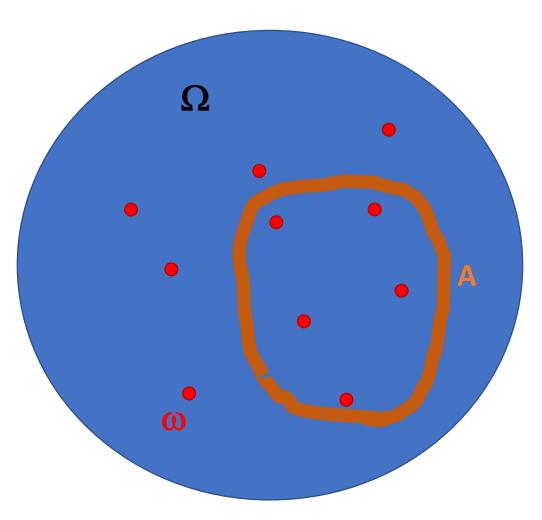
This is an abstract space. For the mag of a star, $\Omega \equiv \mathbb{R}$



 Ω : Sample space \equiv all possible outcome of an experiment

Realisations of the experiment

E.g. There have been 10 measurements of the magnitude of a star. Each measurement is a different realisation



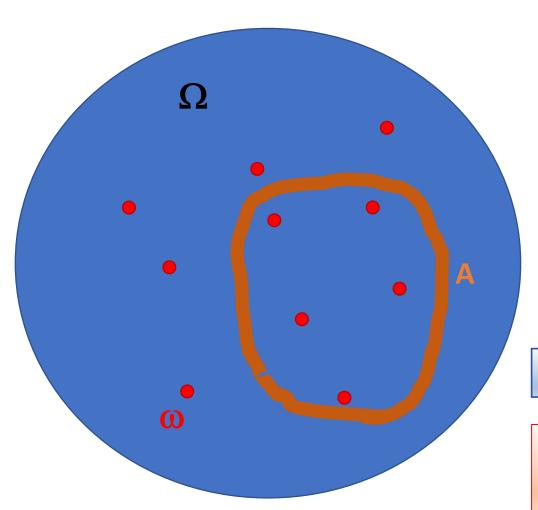
 Ω : Sample space \equiv all possible outcome of an experiment

Realisations of the experiment

A: Event \equiv a subsample of $\omega \cong$ Your data set

E.g. You have obtained and are working on 5 measurements of the magnitude of the star.

But an event can be a bit more convoluted quantity, e.g. all measurements you've done that have m < 15 mag



 Ω : Sample space \equiv all possible outcome of an experiment

o: Realisations of the experiment

A: Event \equiv a subsample of $\omega \cong$ Your data set

p(A): Probability of an event / value to be in [x-dx, x+dx]

e.g. probability that m < 15 mag

What means p(A) in frequentist/classical inference?

Relative frequency of an event if experiment is repeated an infinite number of times

Random variable

A random variable is a variable whose value results from the measurement of a quantity that is subject to random variations

In Python: np.random

- np.random.choice(array): choice at random in an array
- np.random.seed(value): sets the seed of the rnd generator
- np.random.rand(shape): random floats drawn from uniform distribution
- np.random.randint(low, high, shape): rnd integers btw low and high

Go to: Sect. I.2. of the notebook

Conditional probability p(A | B)

$$p(A \mid B) = \frac{p(A \cap B)}{p(B)} = \text{fraction of times that A occurs given that B occurred}$$
Reads "Probability of A given B"

- The calculation of **p(A | B)** follows **Bayes** theorem

$$p(A \mid B) = \frac{p(B \mid A) p(A)}{p(B)}$$

 The probability to have a flu given that you have fever is different from the probability to have fever given that you have a flu

$$p(A|B) \neq p(B|A)$$

Bayes theorem

$$p(A \mid B) = \frac{p(B \mid A) p(A)}{p(B)}$$

Question:

A: rare disease that affects 0.1 % of the population.

B: test that is efficient at 99 % (i.e. 1 % False positive rate).

If you have a positive test (B), what is the probability for you to be affected by this disease (A)?

NB: Efficiency is NOT sensitivity (sensitivity generally means fraction of true positive).

Bayes theorem

$$p(A \mid B) = \frac{p(B \mid A) p(A)}{p(B)}$$

Question

A: rare disease that affects 0.1 % of the population.

B: test that is efficient at 99 % (i.e. 1 % False positive rate).

If you have a positive test (B), what is the probability for you to be affected by this disease (A)?

Solution: (See Sect. I.3. of the notebook)

Among 1000 persons, 1 has the disease (it touches 0.1% of the population = p(A)).

The test has 99% efficiency (=p(B | A)). Which means that 1% of the people will be tested positive while not being sick. => 10 people will be positive while healthy. You should also have \approx 1 being positive while being effectively sick. p(B)=0.01 + 0.001 =0.011

 $=> p(disease | +) \approx 1/11 = 9 \%$

BEWARE

RARE events common in astronomy Conditional probabilities are often implicit

Contingency Table

E\H	Disease = H	No Disease = !H	Σ
Test + = E	P(E H) p(H)	P(E !H) p(!H)	P(E)
Test - = !E	P(!E H) p(H)	P(!E !H) p(!H)	P(!E)
Σ	P(H)	P(!H)	1

E \ H	Disease = H	No Disease = !H	Σ
Test + = E			
Test - = !E			
Σ	100	999000	100000

Fill this row first

Contingency Table

E\H	Disease = H	No Disease = !H	Σ
Test + = E	P(E H) p(H)	P(E !H) p(!H)	P(E)
Test - = !E	P(!E H) p(H)	P(!E !H) p(!H)	P(!E)
Σ	P(H)	P(!H)	1

E\H	Disease = H	No Disease = !H	Σ
Test + = E	0.99 * 100 = 99		
Test - = !E	(1-0.99) * 100 = 1		
Σ	100	999000	100000

The test is efficient at 99%: $P(E \mid H) = 0.99$

Contingency Table

E\H	Disease = H	No Disease = !H	Σ
Test + = E	P(E H) p(H)	P(E !H) p(!H)	P(E)
Test - = !E	P(!E H) p(H)	P(!E !H) p(!H)	P(!E)
Σ	P(H)	P(!H)	1

E\H	Disease = H	No Disease = !H	Σ
Test + = E	0.99 * 100 = 99	(1 - 0.99) * 999000 = 999	1098
Test - = !E	(1-0.99) * 100 = 1		
Σ	100	999000	100000

 $P(E \mid !H) = proba$. To be positive without being sick = (1-0.99)

Contingency Table

E \ H	Disease d = H	No Disease = !H	Σ
Test + = E	P(E H) p(H)	P(E !H) p(!H)	P(E)
Test - = !E	P(!E H) p(H)	P(!E !H) p(!H)	P(!E)
Σ	P(H)	P(!H)	1

$$P(H \mid E) = \frac{P(E \mid H) P(H)}{P(E)}$$

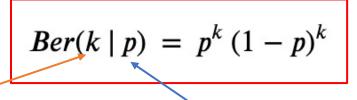
E\H	Disease d = H	No Disease = !H	Σ
Test + = E	0.99 * 100 = 99	(1 - 0.99) * 999000 = 999	<mark>1098</mark>
Test - = !E	(1-0.99) * 100 = 1	98901	100000- 1098=98902
Σ	100	999000	100000

$$P(d|+) = \frac{P(+|d)P(d)}{P(+)} = \frac{99}{1098}$$

$$T_{+}(+) = \frac{1098}{1098}$$

Probability density / mass function

Coin Toss (Bernouilli PMF): The PDF is the normalised histogram we had obtained



 $k \text{ in } \{0, 1\} \equiv \{\text{failure, success}\}\$

parameter (success rate)

Uniform PDF:

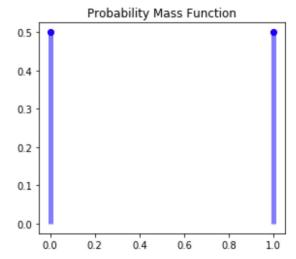
$$h(x) = \frac{1}{b-a} \text{ if } a \leq x \leq b$$

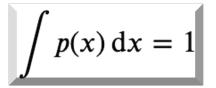
$$h(x) = 0$$
 otherwise

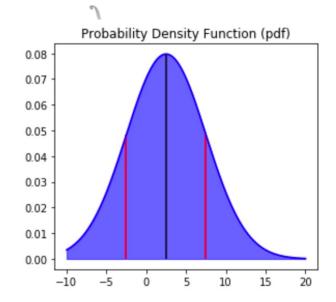
Gaussian PDF:

$$h(x; \mu, \sigma) = \frac{1}{\sqrt{2\pi} \sigma} \exp\left(-\frac{1}{2} \frac{(x - \mu)^2}{\sigma^2}\right)$$

PDF In Python: go to Sect. I.4 of the notebook

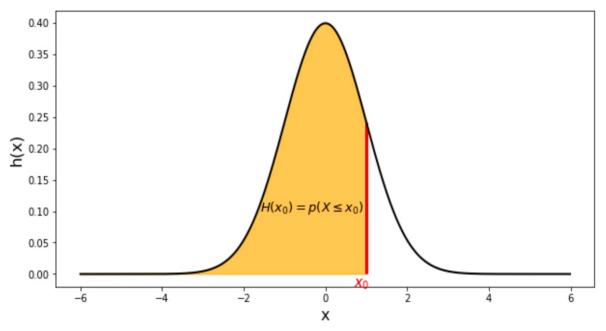


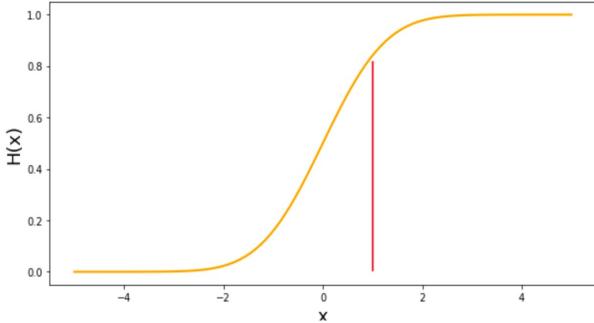




Cumulative distribution function

This is the integral of the PDF:
$$p(X \le x) = H(x) = \int_{-\infty}^{x} h(x') \, \mathrm{d}x'$$





$$H(x \mid \mu, \sigma) = \frac{1}{\sqrt{2\pi} \sigma} \int_{-\infty}^{x} \exp\left(-\frac{1}{2} \frac{(x' - \mu)^2}{\sigma^2}\right) dx'.$$

CDF In Python: go to Sect. I.5 of the notebook

Probability enclosed between 1-2-3 σ for N(μ , σ)

See supplementary exercise of Sect. I.6.2 of the notebook Basic statistics 02.ipynb

