



École Polytechnique Fédérale de Lausanne

Rotation estimation of detected satellites

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Bachelor Thesis

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# Abstract

I present an approach to extract the intensity along the satellite streaks that can be found in images that were obtained by OMEGACAM on the VLT Survey Telescope. In this project, I only analyzed the long tracks that were created by high velocity objects, to try to get an estimation about their rotation. I used Fourier transform to get the periodicity of the intensity along the tracks. Due to the lack of real data about rotation, I first had to create fake streaks with sinusoid intensity and different angles. I manually get the tracks to run my code on them, so ideally this project should be integrated with DetectSat [2] so that all the process can be done automatically. I found very great result with it and run the pipeline on real tracks. By making assumptions about the satellite, like their altitude, I get possible results, but since we don't have access to the real values, we can't say if it is close to reality or not.

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# Chapter 1

## Introduction

This project is about estimating the rotation of spatial objects detected by the OMEGA-CAM on the VLT Survey Telescope. In these times, we have a lot of flying objects, it can be debris, useful satellites or dead satellites, around the Earth and it becomes more and more necessary to have a clear view of all these objects. The project of Yann Bouquet "DetectSat"[2] was focused on detection of the satellite tracks on fits files, but with both, small (low velocity) and long (high velocity) streaks. To this project I had to focus on the long tracks, so that the variation of intensity along the line could be better detected. The rotation estimation can be used to determine which kind of object the streak is and, if it is a satellite, it may be used to better find out which one it is. Ideally, this should be run on the lines detected by DetectSat so all the process could be done with only the fits file as input.

The main challenge is that we don't have access to data about rotation of satellites, so I had to first create my own streaks with sinusoid intensity. This lack of resources is very restrictive, since even if I get a result, we cannot know the precision of this method and we cannot be sure that it is right or completely wrong. This is why I had to begin with creating lines that seem realistic and try with random angles.

The second part of this project was to retrieve the periodicity of the sinusoidal function along the streaks that I generated, and finally try this on real tracks. Since my project does not recognize the lines, I had to find the coordinates of the streaks by hand and give them as input to the pipeline. For this part, there are many factors that can

decrease the precision of the pipeline, e.g., if the line crosses a really bright star, it gives a peak of intensity and can lead to wrong result, or if the brightness of the streak is too small compared to the brightness of the background of the sky on the image.

# Chapter 2

## Inputs

In this project, we used fits images that were generated by the OMEGACAM on the VLT Survey Telescope, and we used a script to create a mosaic of 32 images, separated by NaN values, that are 4000x2000 pixels. So the image with all the mosaic's dimension is 16'000x16'000 pixels. A mosaic is approximately 1.5-2 GB of data.

I principally used the fits file "*OMEGA.2020-02-22T02:11:36.295\_fullfield.fits*", corresponding to the captured image February 22, 2020, at 02:11:36, to generate fake tracks and to test the pipeline.

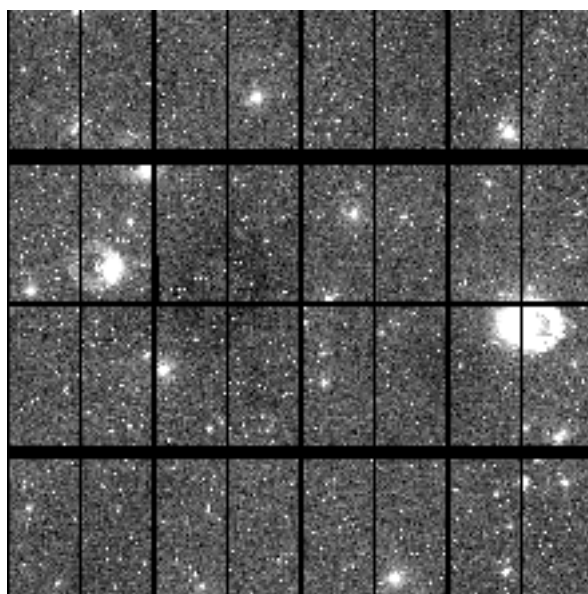


Figure 2.1 – Figure 2.1: The main FITS image that I used

The pipeline is being run on the entire file



# **Chapter 3**

## **Design**

Introduce and discuss the design decisions that you made during this project. Highlight why individual decisions are important and/or necessary. Discuss how the design fits together.

This section is usually 5-10 pages.

# **Chapter 4**

## **Implementation**

The implementation covers some of the implementation details of your project. This is not intended to be a low level description of every line of code that you wrote but covers the implementation aspects of the projects.

This section is usually 3-5 pages.

# **Chapter 5**

## **Evaluation**

In the evaluation you convince the reader that your design works as intended. Describe the evaluation setup, the designed experiments, and how the experiments showcase the individual points you want to prove.

This section is usually 5-10 pages.

# **Chapter 6**

## **Related Work**

The related work section covers closely related work. Here you can highlight the related work, how it solved the problem, and why it solved a different problem. Do not play down the importance of related work, all of these systems have been published and evaluated! Say what is different and how you overcome some of the weaknesses of related work by discussing the trade-offs. Stay positive!

This section is usually 3-5 pages.

# **Chapter 7**

## **Conclusion**

In the conclusion you repeat the main result and finalize the discussion of your project. Mention the core results and why as well as how your system advances the status quo.

# Bibliography

- [1] Mathias Payer. *Template for EPFL (BSc, MSc, or doctoral) theses and semester projects*. 2020. URL: [https://github.com/HexHive/thesis\\_template](https://github.com/HexHive/thesis_template) (visited on 2021).
- [2] Yann Bouquet. *Detectsat Repository*. 2020. URL: <https://github.com/YBouquet/detectsat> (visited on 2021).