

GRAVITATIONAL LENSING

13 – PLANET MICROLENSING (I)

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AA 2019-2020

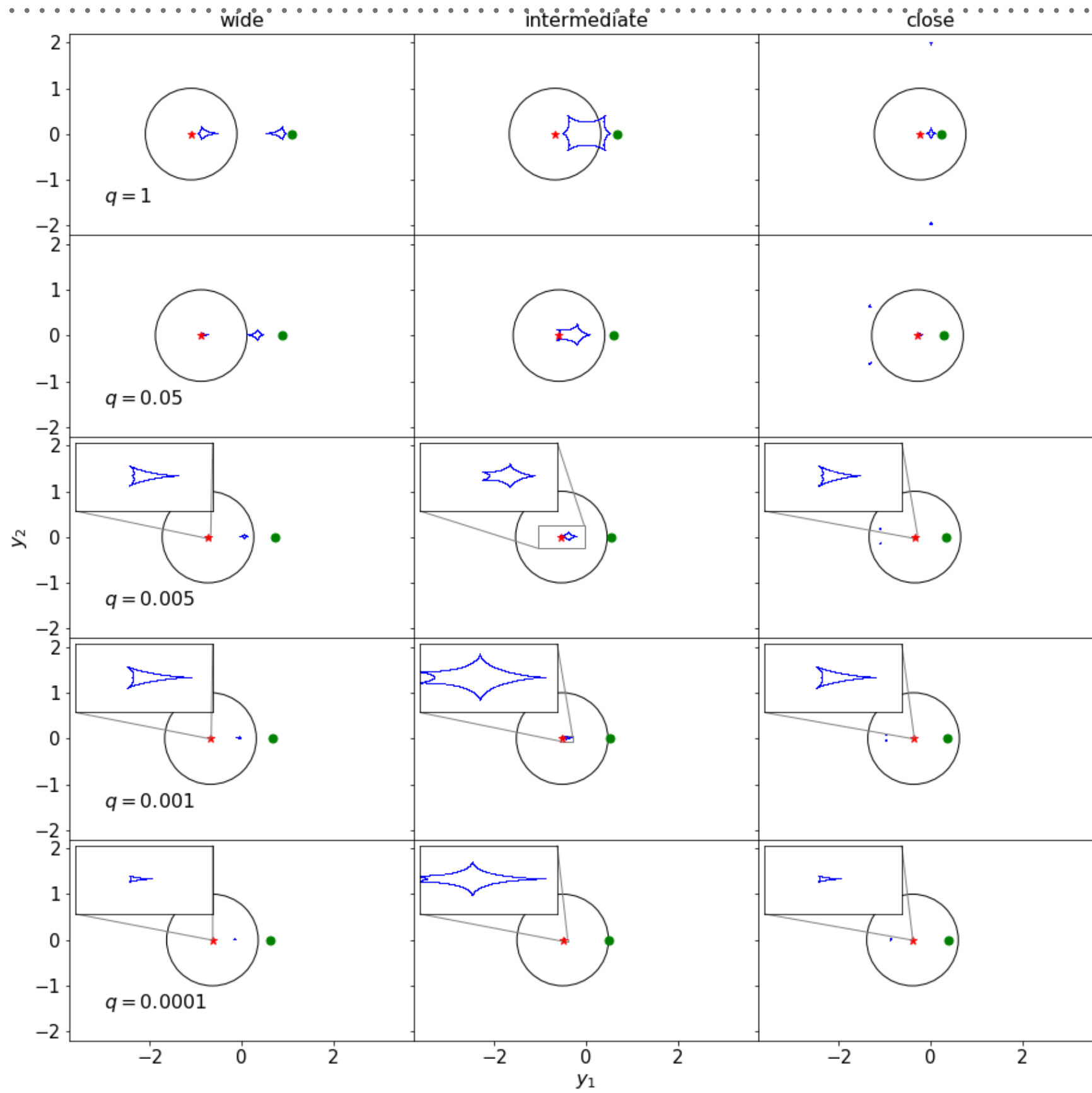
PLANETARY MICROLENSING

- Let us consider the system consisting of an host star and a planet orbiting around it.
- This is an example of **binary** lens
- The host star is of course much heavier than the planet!
 - example: for a Jupiter-like planet $q=0.001$ (solar mass star)
 - example: for a Earth-like planet $q=0.0000003$

WHAT KIND OF SIGNAL?

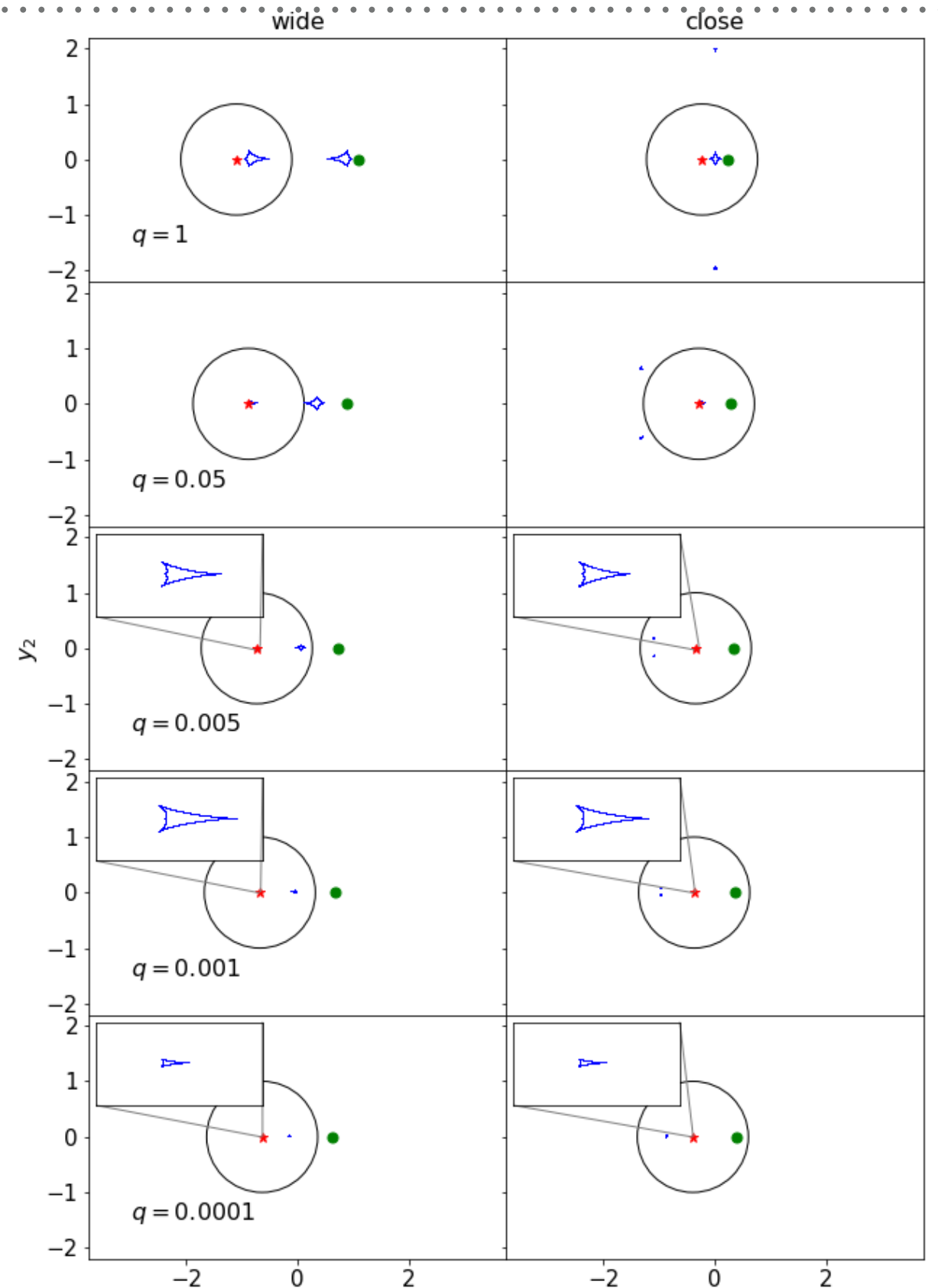
- The light curve is that of the star...
- The planet produces only a small perturbation to the magnification pattern, localized in a small region around the caustics
- Must cross one of these perturbed regions in order for the planet to be detected.
- The shape of the perturbation is determined by the caustic configuration...

CAUSTICS AS A FUNCTION OF $q = m_2/m_1$



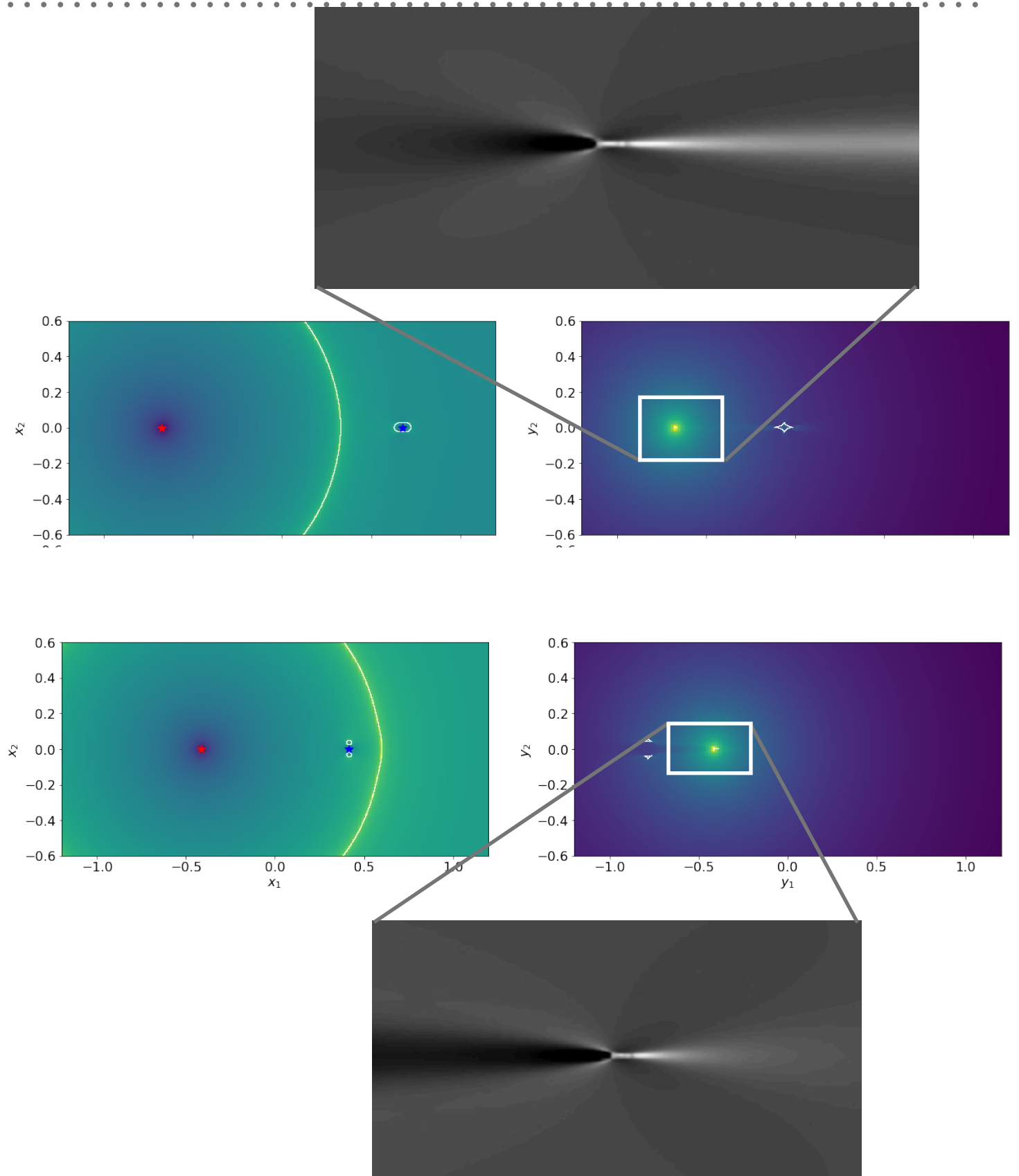
PERTURBATIONS OF THE CENTRAL CAUSTIC (WIDE/CLOSE SYSTEMS)

- As q decreases, we see that one caustic shrinks and approaches the primary lens (i.e. the star)
- This is what we call the “*central caustic*” in wide and close systems
- Four cusps and four folds
- One cusp is elongated towards the planet
- Three cusps on the back
- Different from point-like caustic of a point lens!

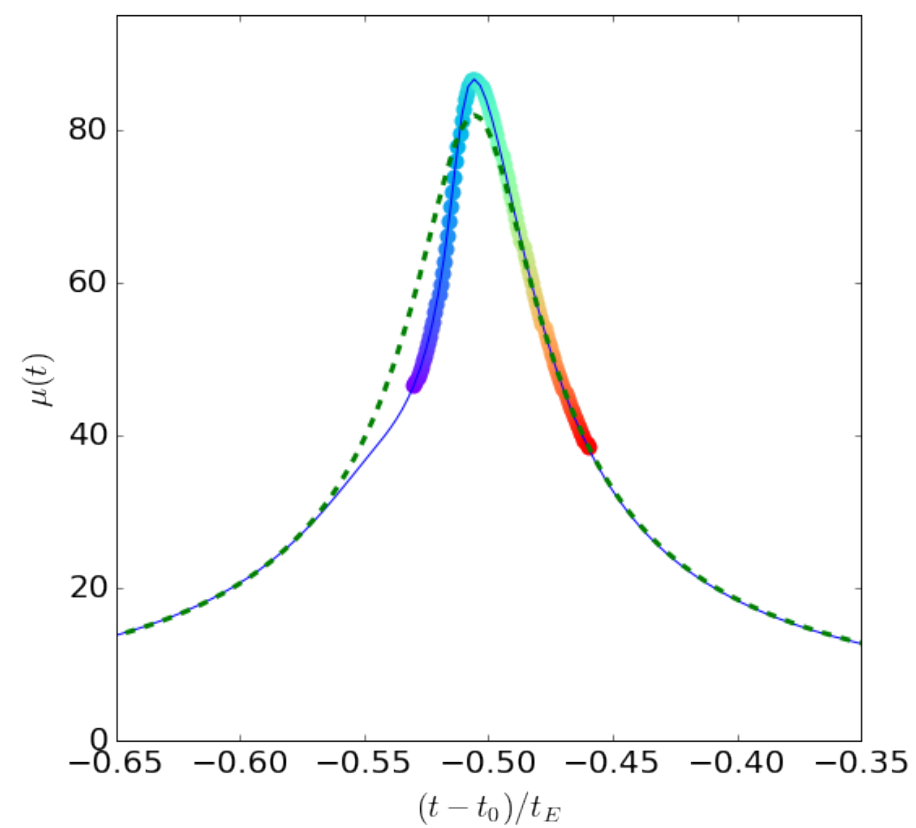
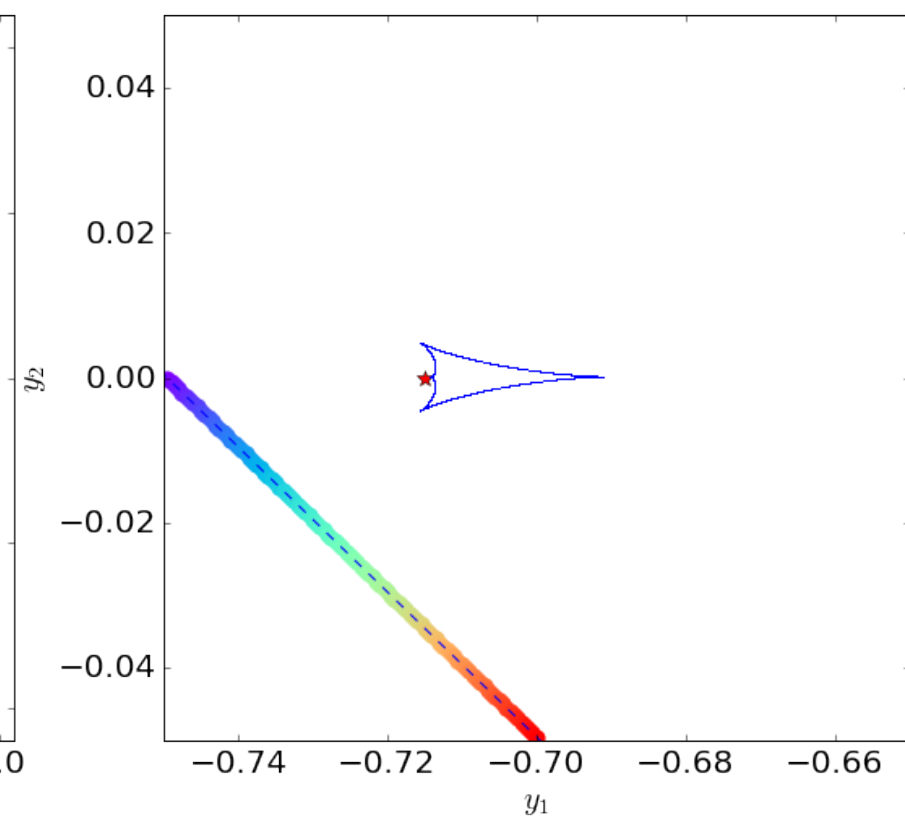
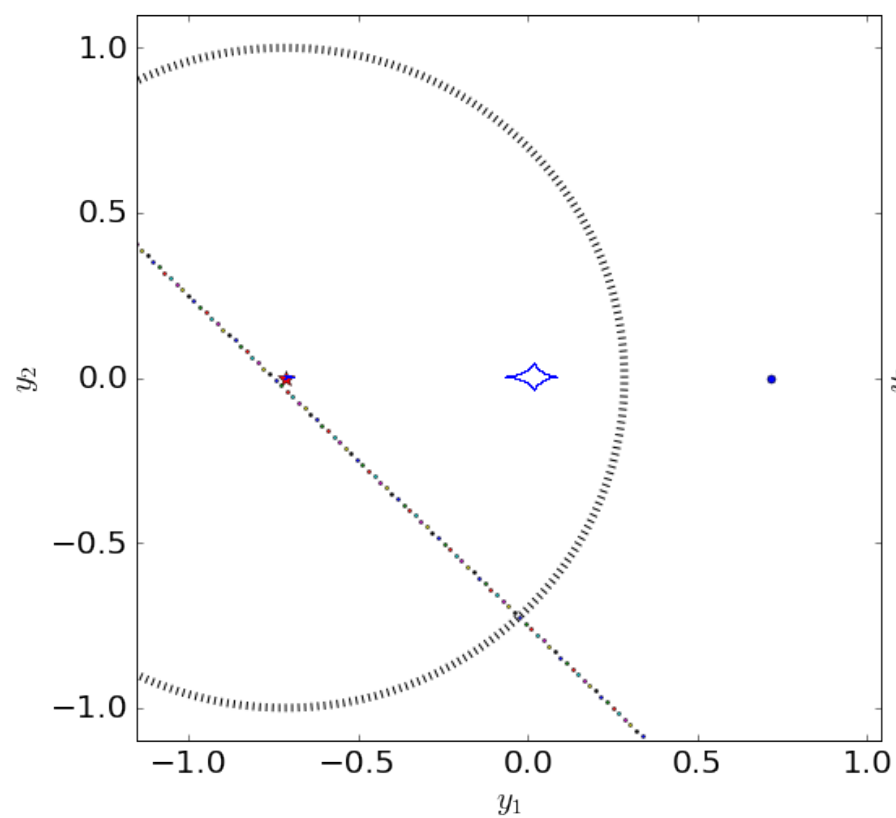
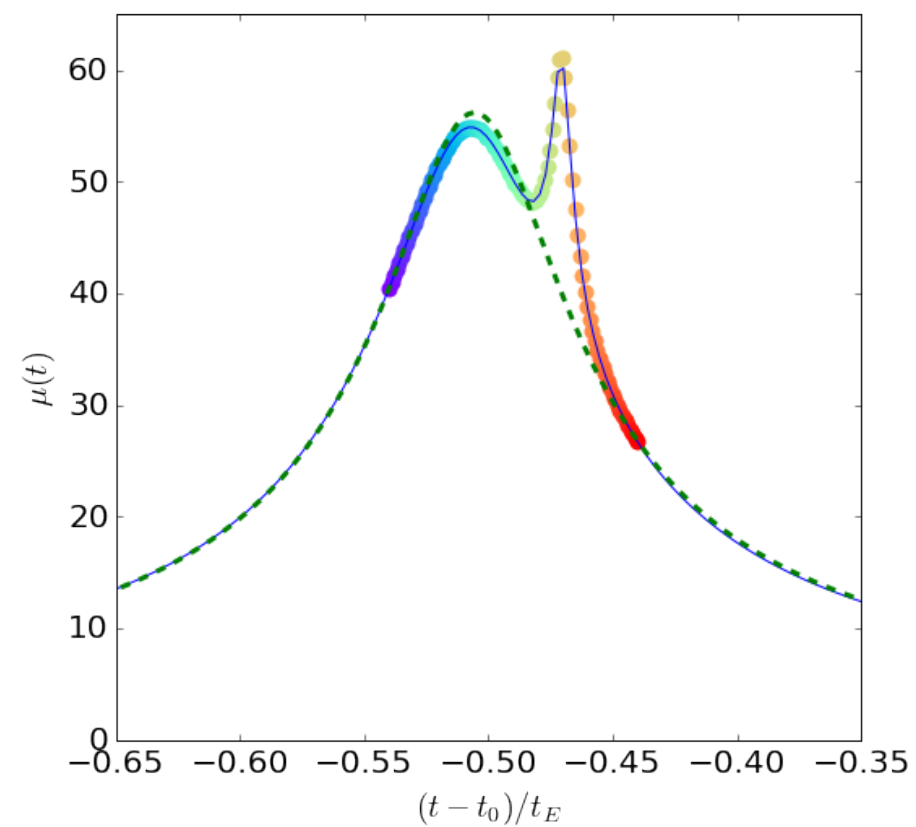
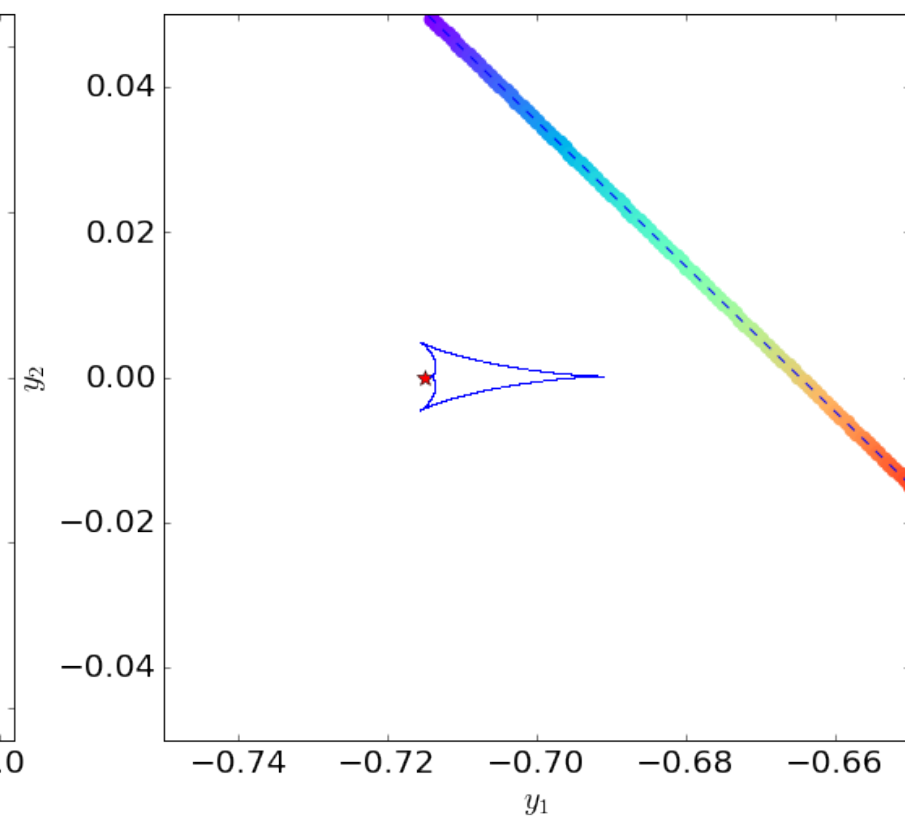
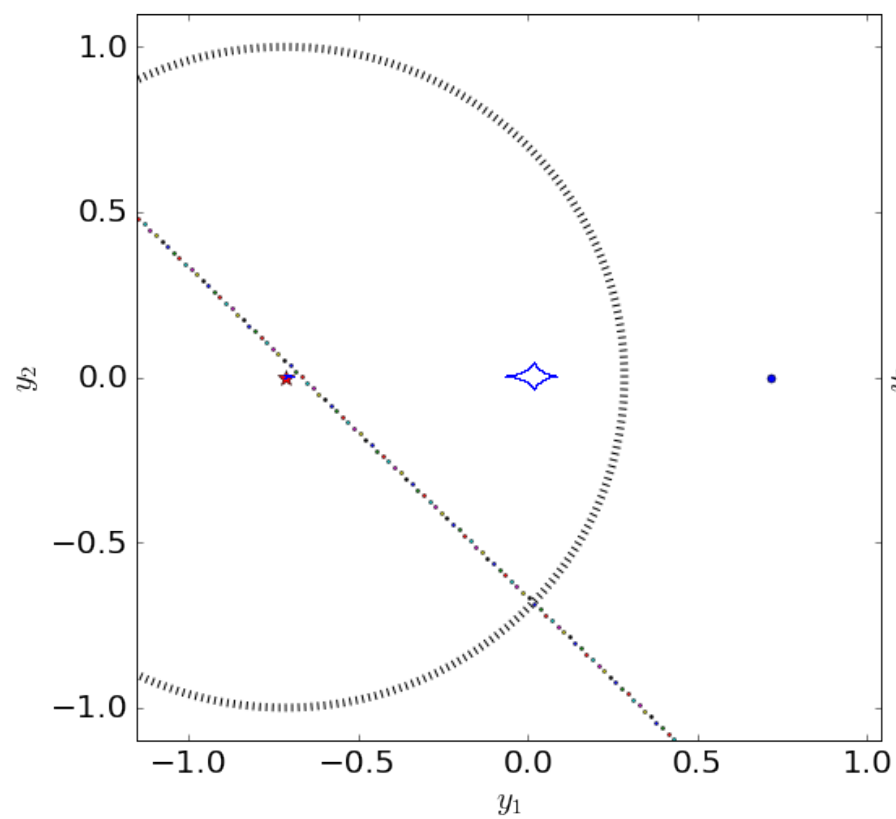


WHAT KIND OF SIGNATURES?

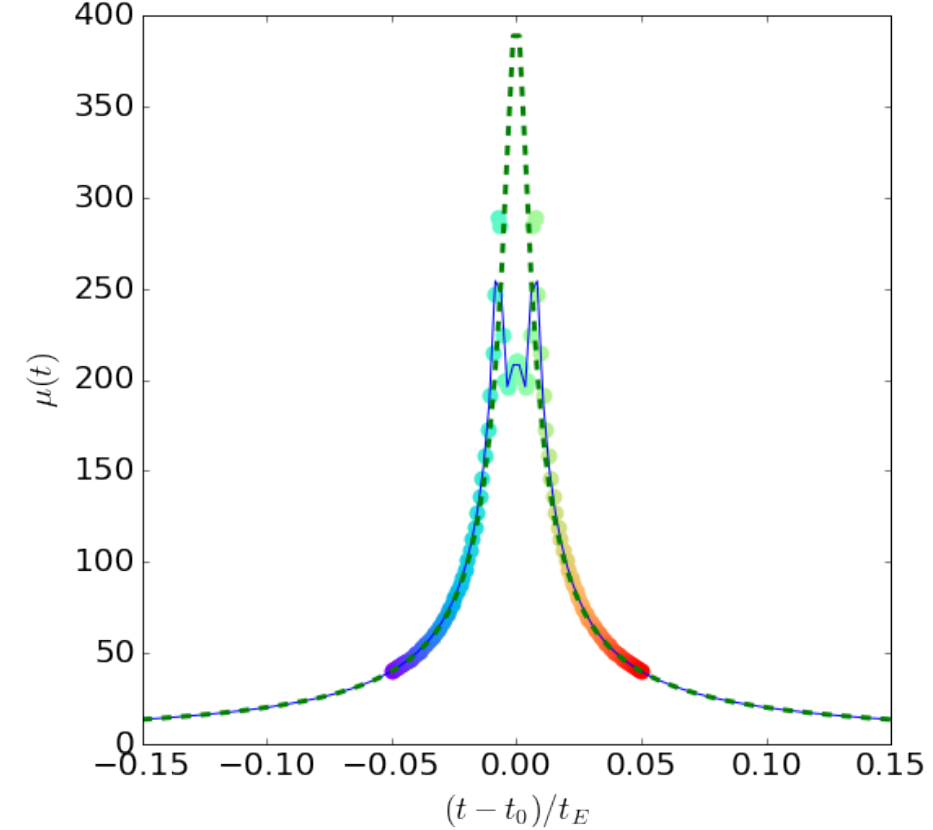
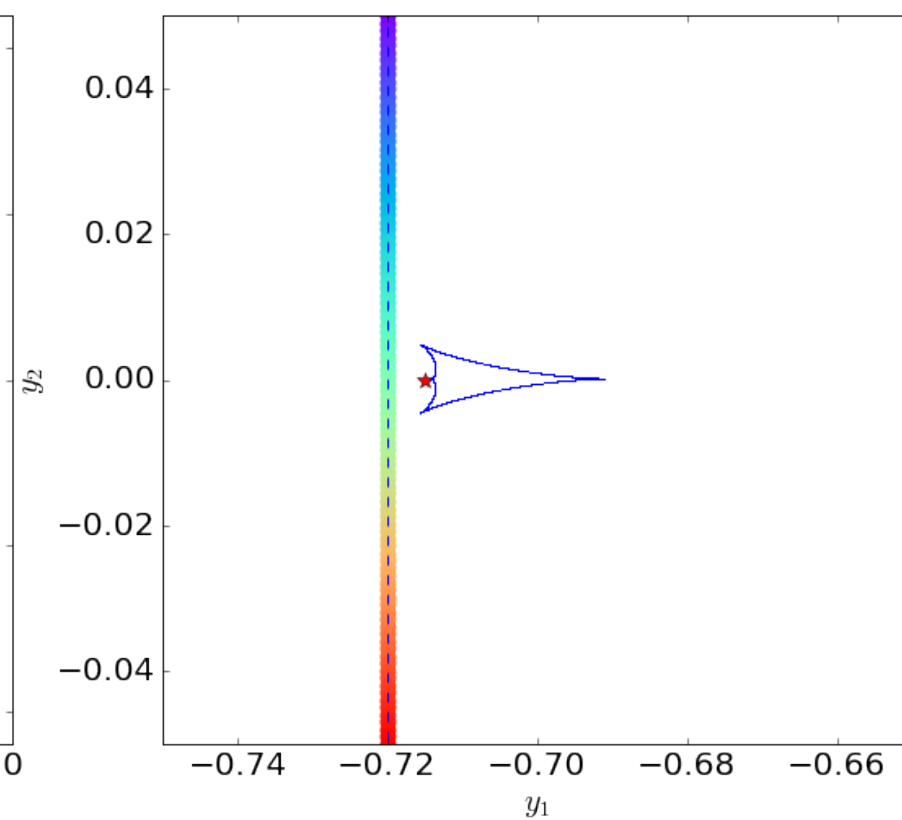
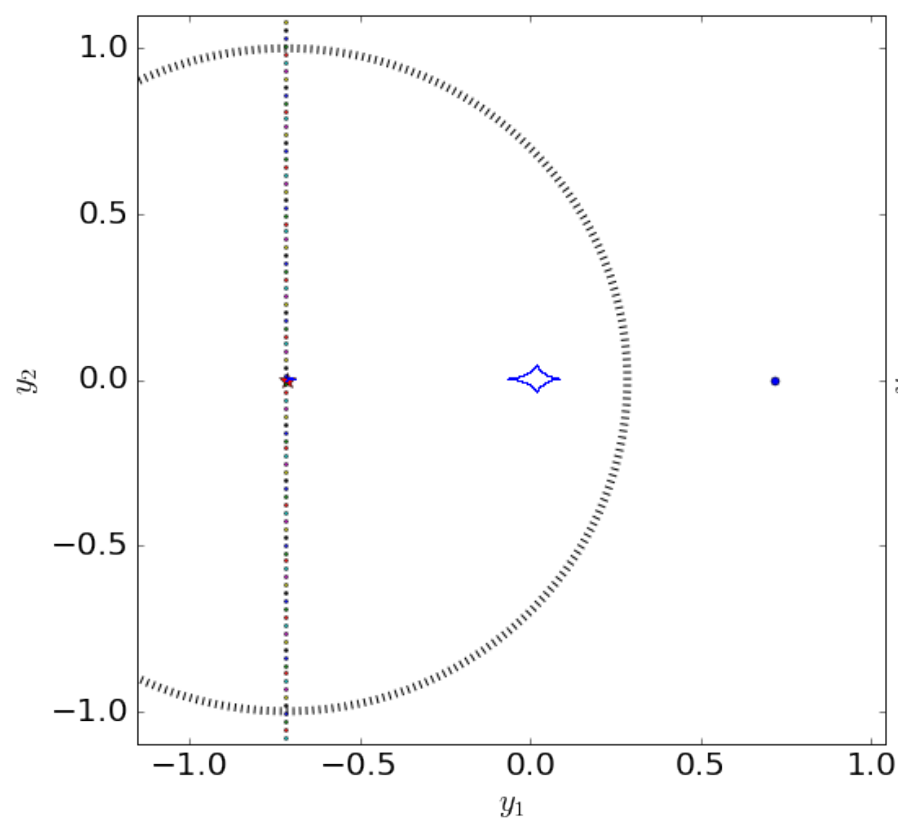
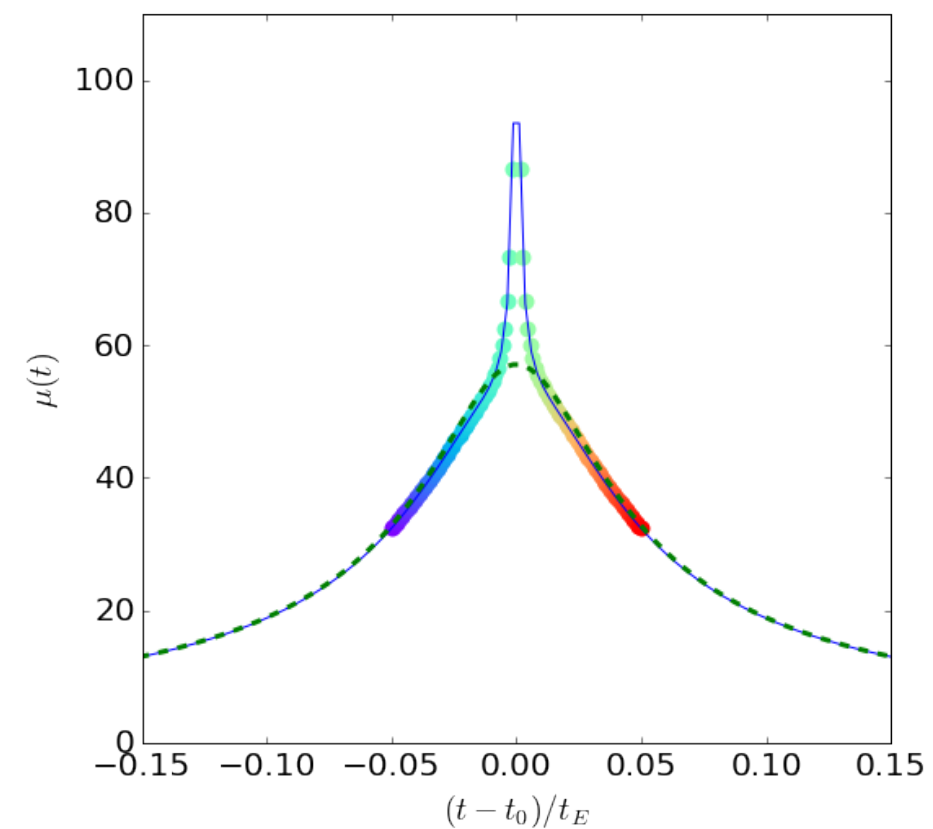
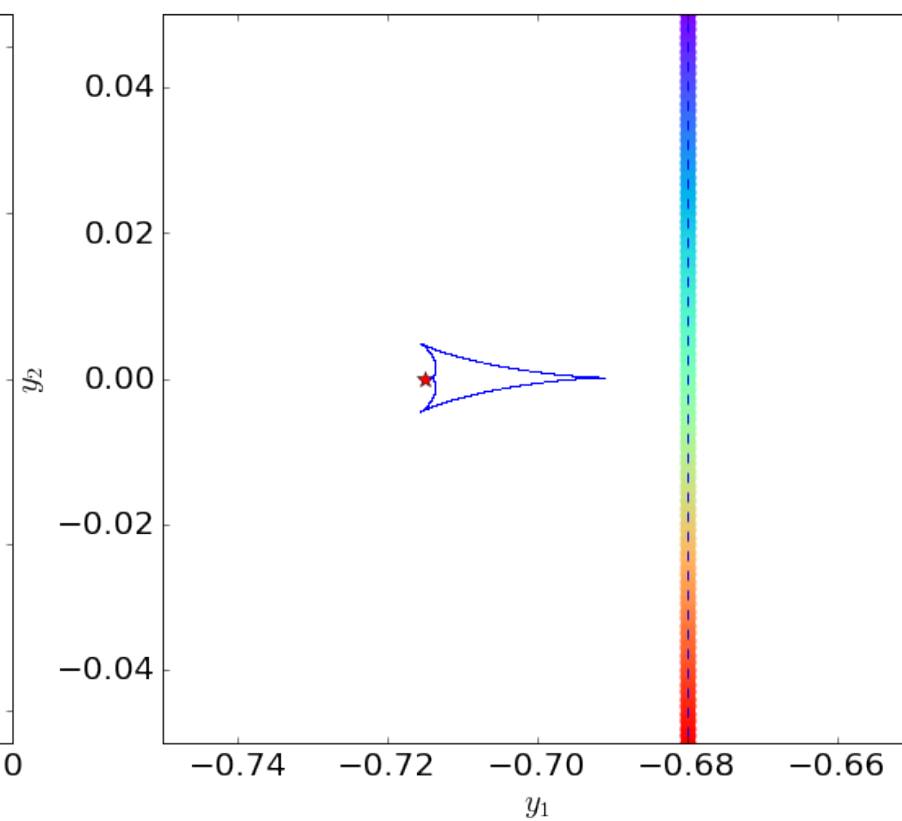
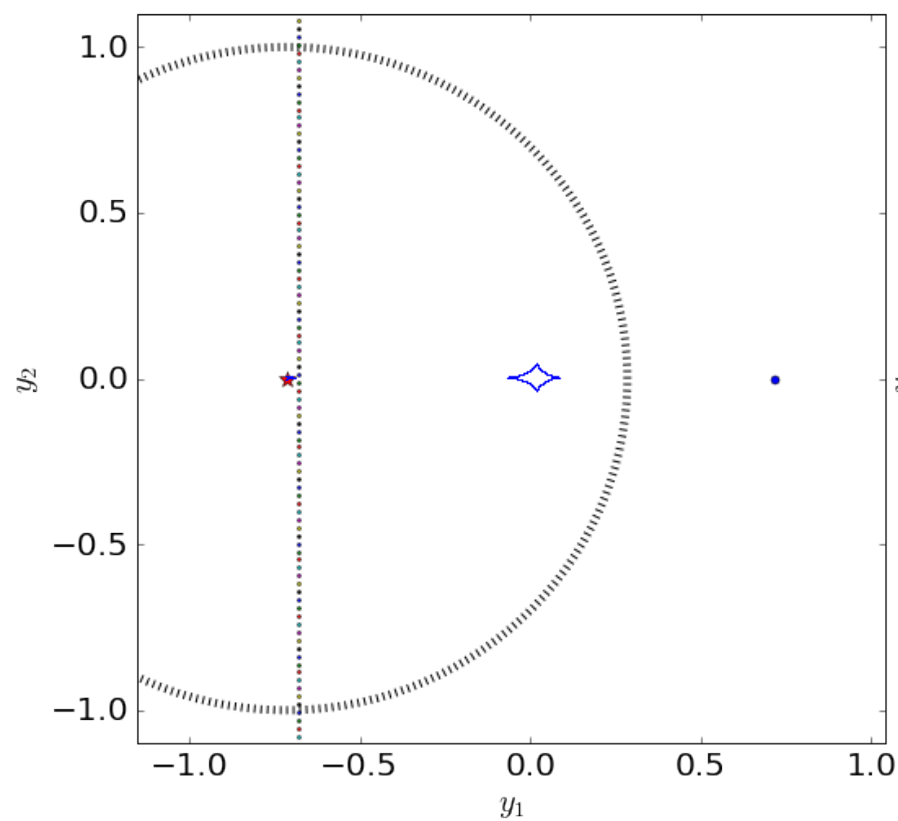
- Magnification if the source passes in between the planet and the star
- De-magnification if the source passes on the back of the caustic!
- Some examples...



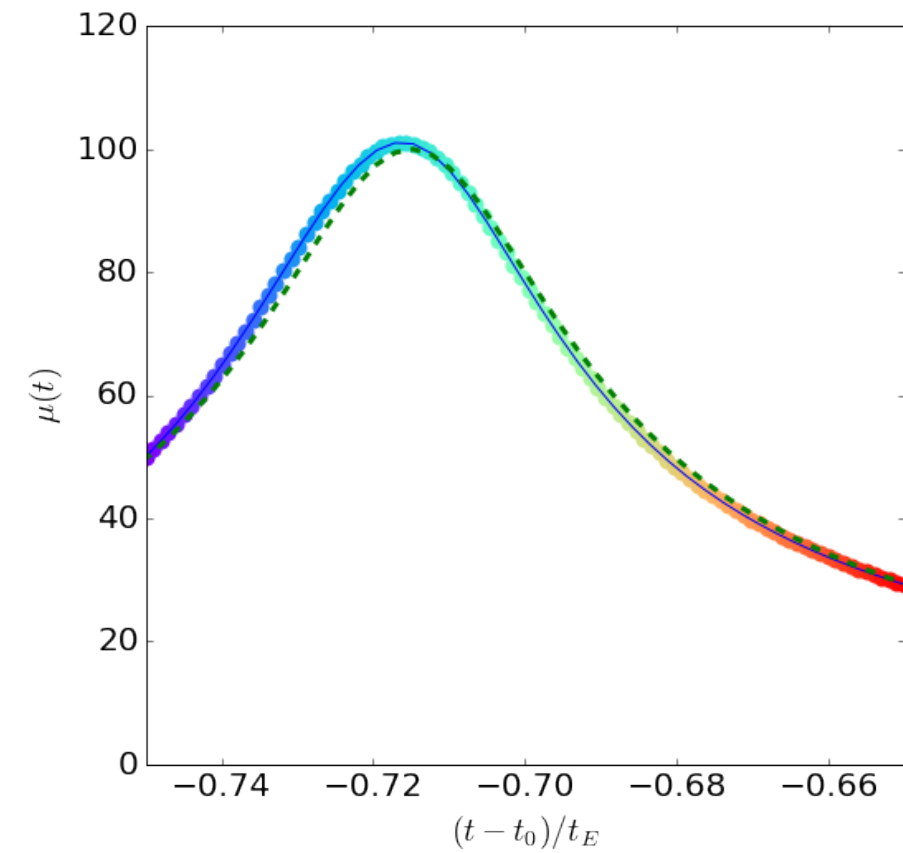
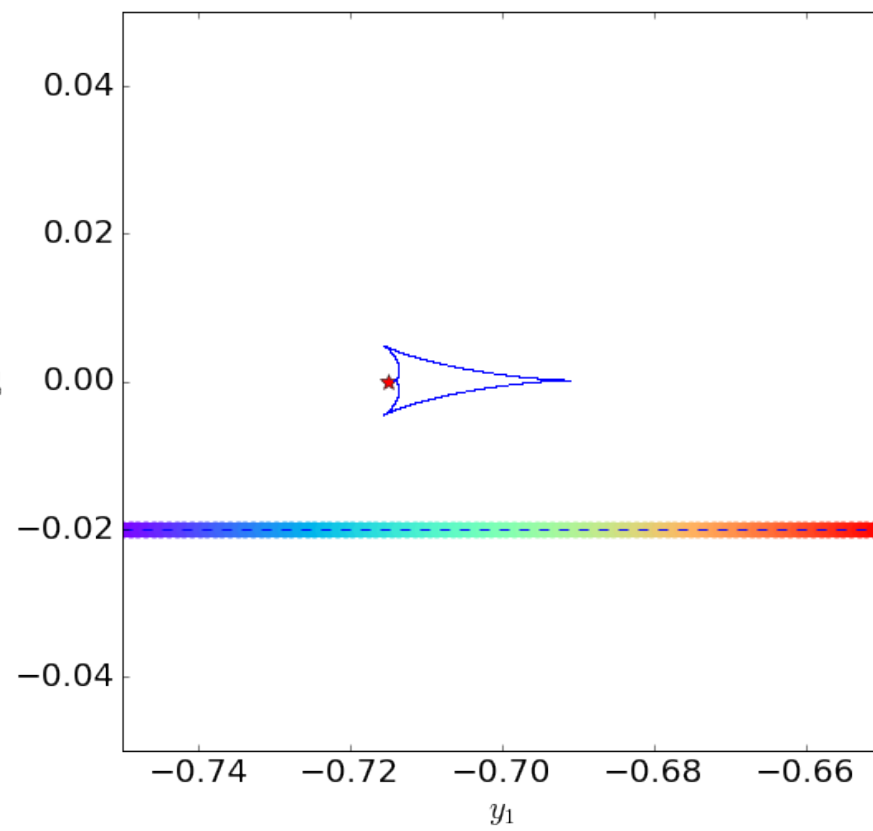
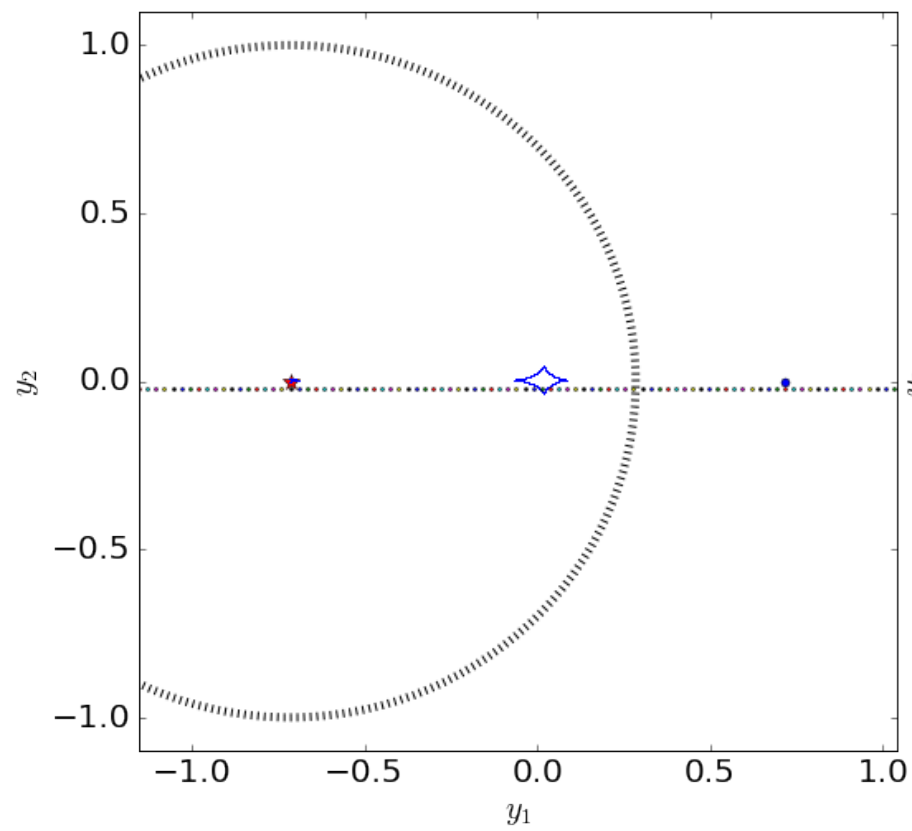
CENTRAL CAUSTIC PERTURBATIONS



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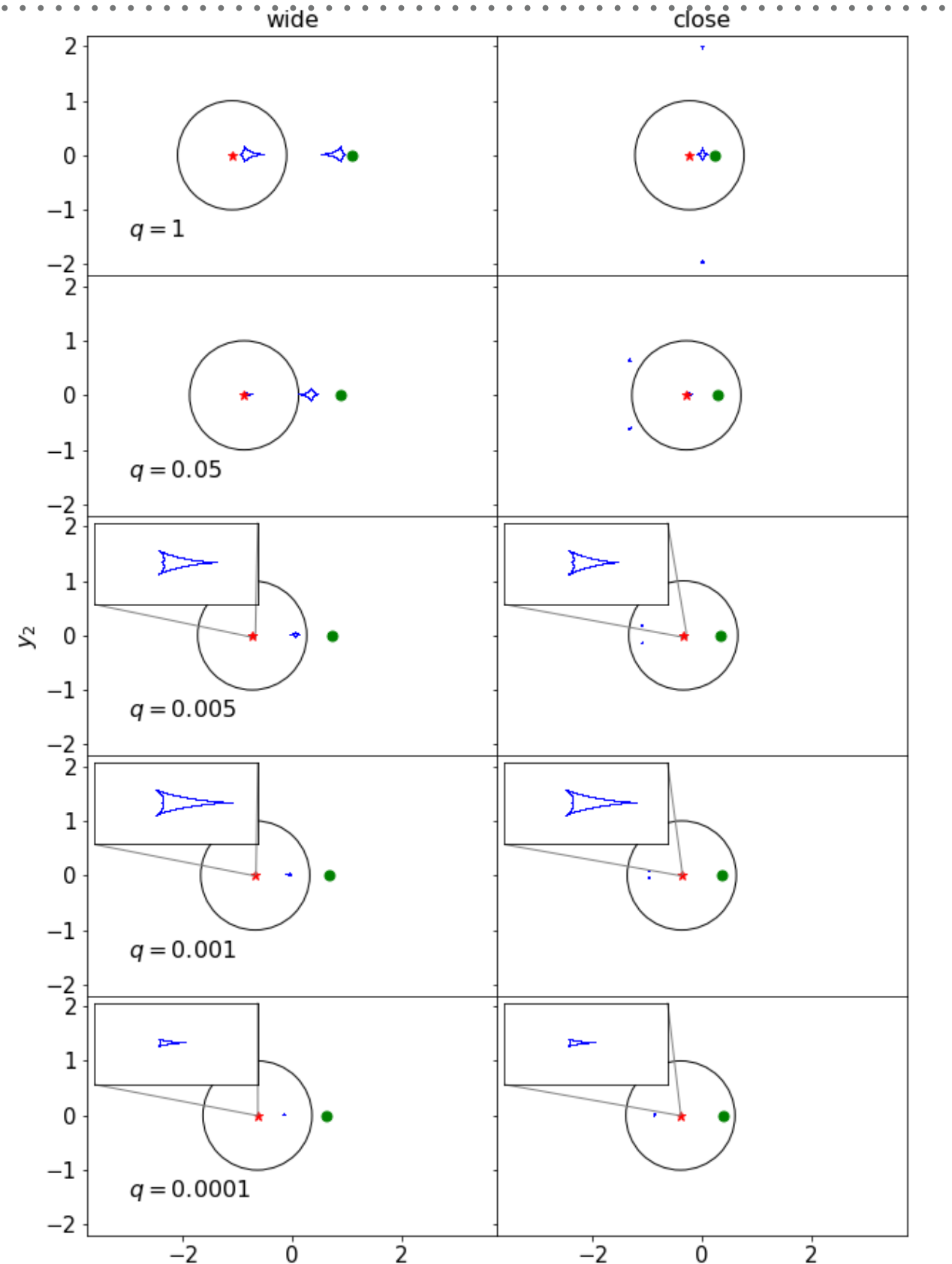
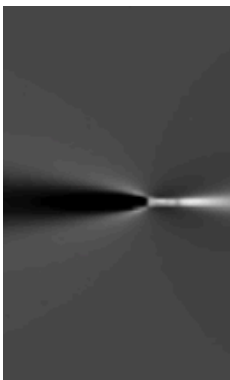
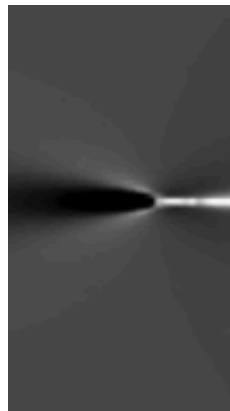


CENTRAL CAUSTIC PERTURBATIONS



WIDE-CLOSE DEGENERACY!

- Caustics and magnification patterns are identical in wide and close systems with $d_w = d_c^{-1}$



PLANET DETECTION THROUGH CENTRAL CAUSTICS PERTURBATIONS

- Only possible in the case of high magnification events (sources passing very close to the host stars)
- For this reason, they are rare events
- Advantages:
 - near the peak of the event
 - can sometimes be predicted in advance
 - high magnification makes possible to follow-up the events using small telescopes
 - more accurate photometry
- Disadvantages:
 - degeneracy wide-close topologies