

Nonlinear evolution of the Kelvin-Helmholtz instability

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- the Kelvin-Helmholtz instability in compressible HD
- Linear (2D) results
 - ⇒ HD versus MHD: parallel and anti-parallel B configurations
- Non-linear (2D) behaviour
- 3D simulations: A case study



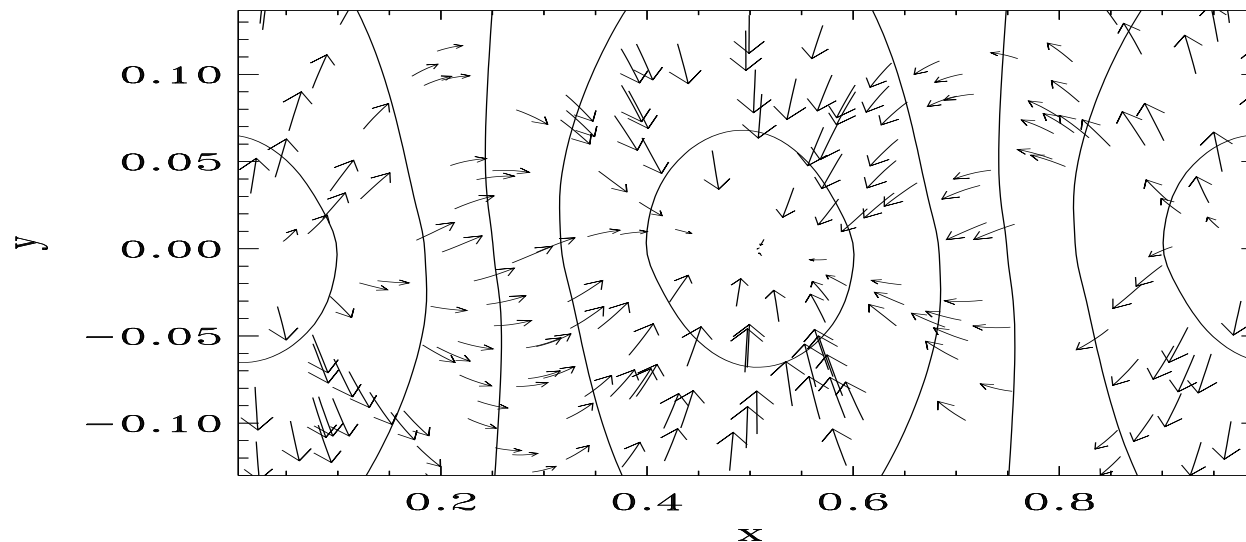
Kelvin-Helmholtz instability

- 2D (x, y) -plane, compressible HD, constant ρ and p
- shear $\tanh(y/a)$ horizontal v_x flow profile, width $2a$
- perturb with vertical flow $\delta v_y \sim \sin(k_x x)$

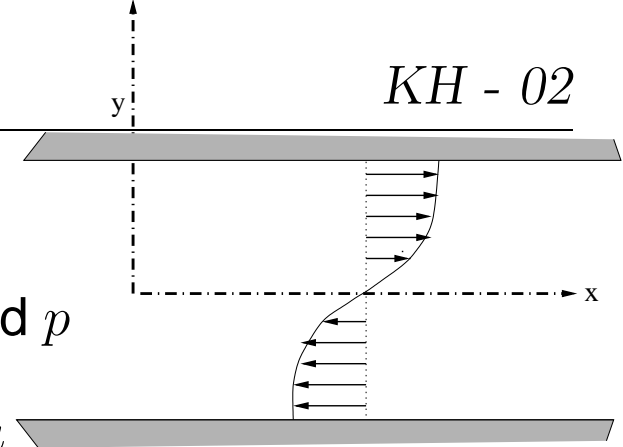
⇒ within shear layer: $-\rho v'_x \delta v_y$ horizontal force

⇒ density ($\propto p$) redistributed in compression-depression pattern

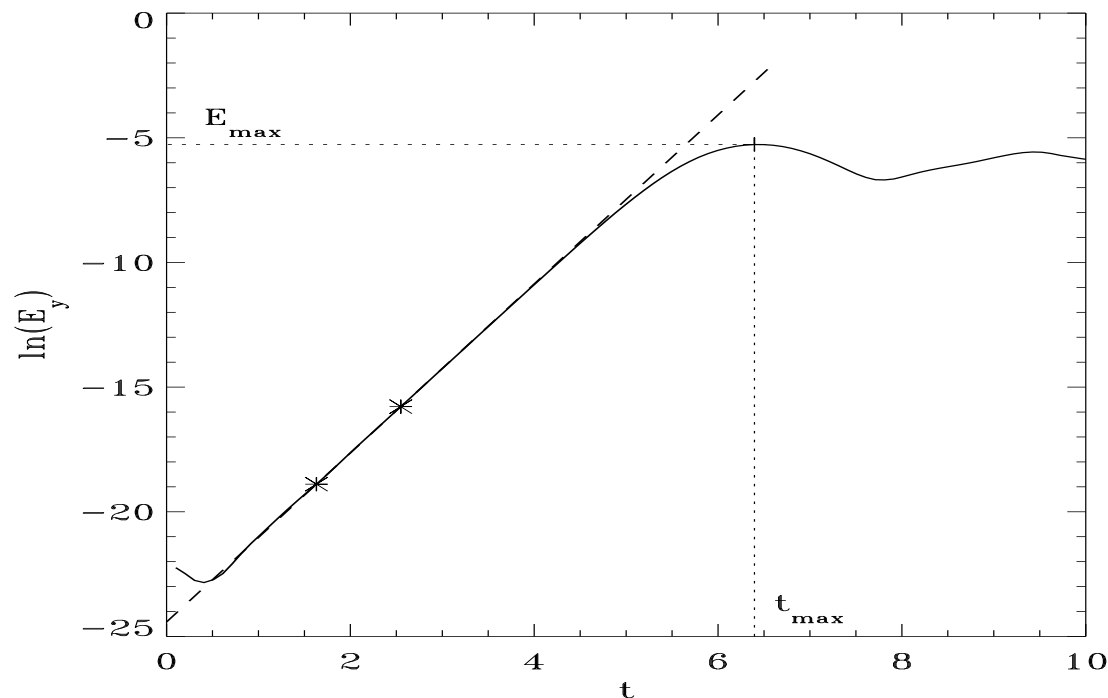
⇒ pressure perturbation → pressure gradient force field



⇒ amplifies the vertical velocity perturbation: unstable!

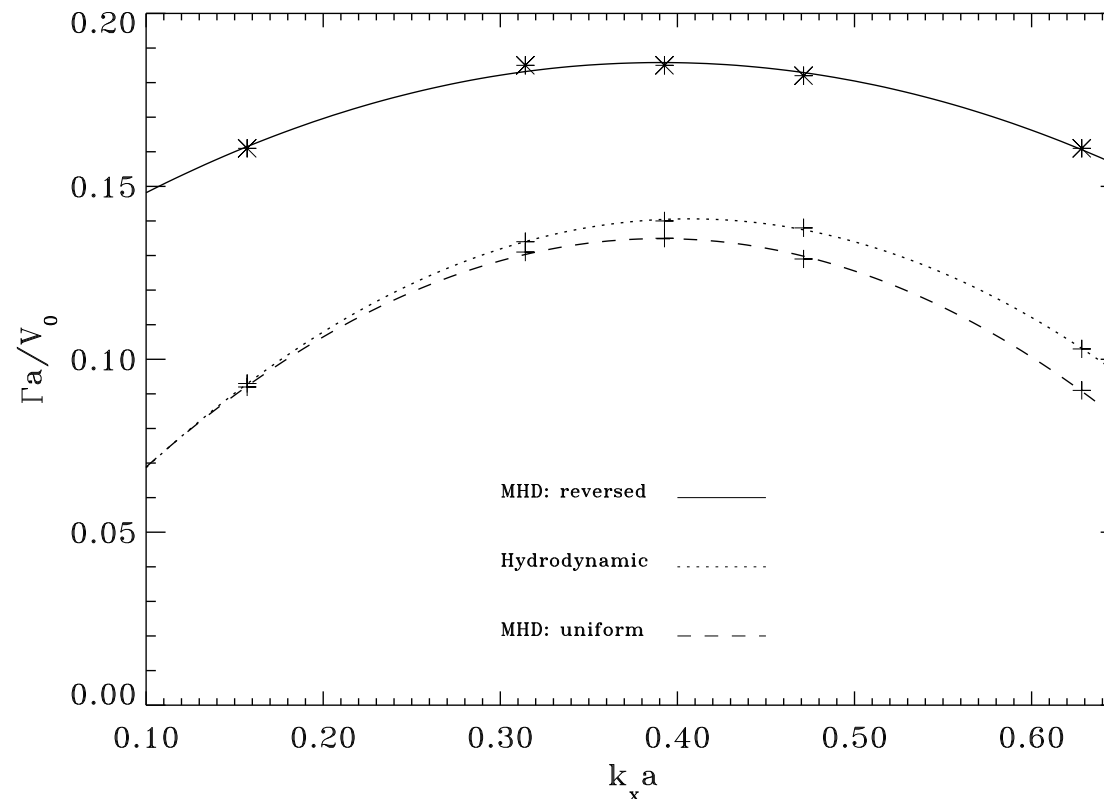


- 2D studies: HD and MHD uniform horizontal (\parallel flow) and reversed B
 \Rightarrow reversed field: uniform strength, direction flips at $y = 0$
- linear and non-linear behaviour: growth & saturation
- monitor vertical kinetic energy: deduce growth rate and E_{max}



Linear (2D) results

- reference case with $V_0 = 0.645$, $a = 0.05$, $k_x = 2\pi$, $B_0 = 0.129$
 - \Rightarrow KH unstable for HD, initially weak B field in MHD
 - \Rightarrow dimensionless $k_x a = 0.314$, $M_s = 0.5$, $M_a = 5$, $\beta = 120$
 - \Rightarrow growth rates agree with LEDAFLOW results!
- dependence of growth Γ on wavelength

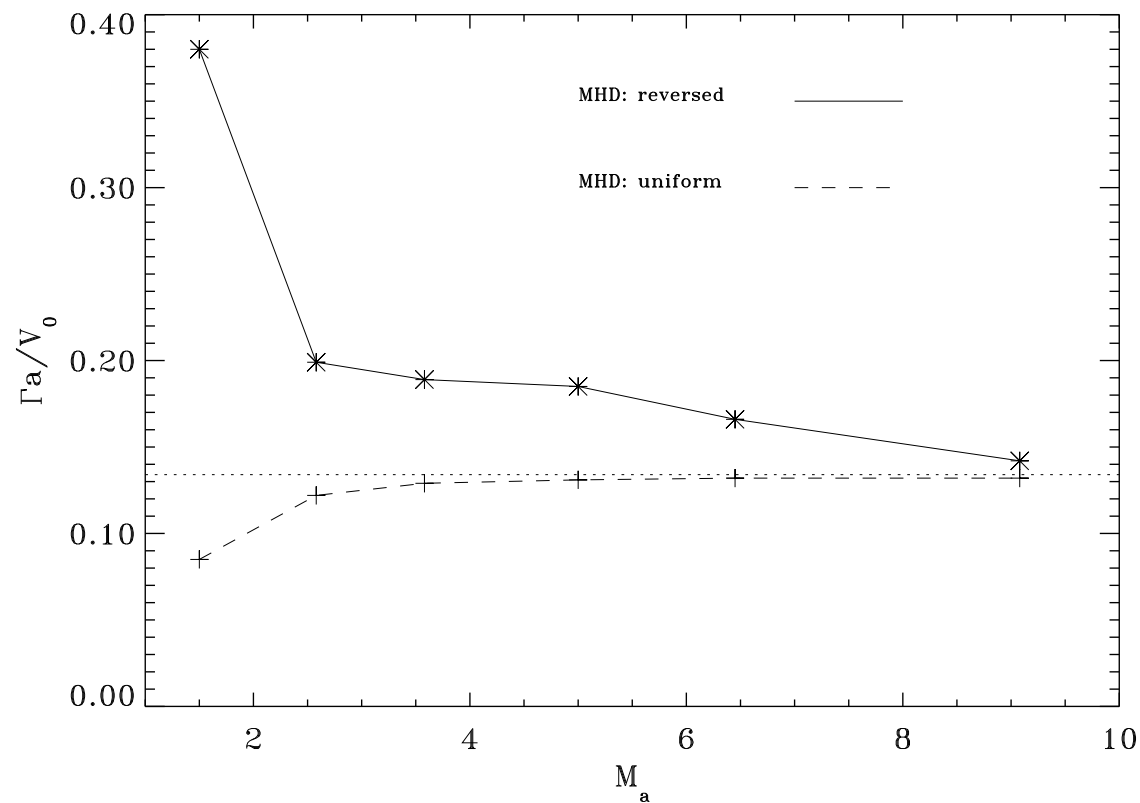


- stabilizing influence of uniform B

⇒ tension in field lines

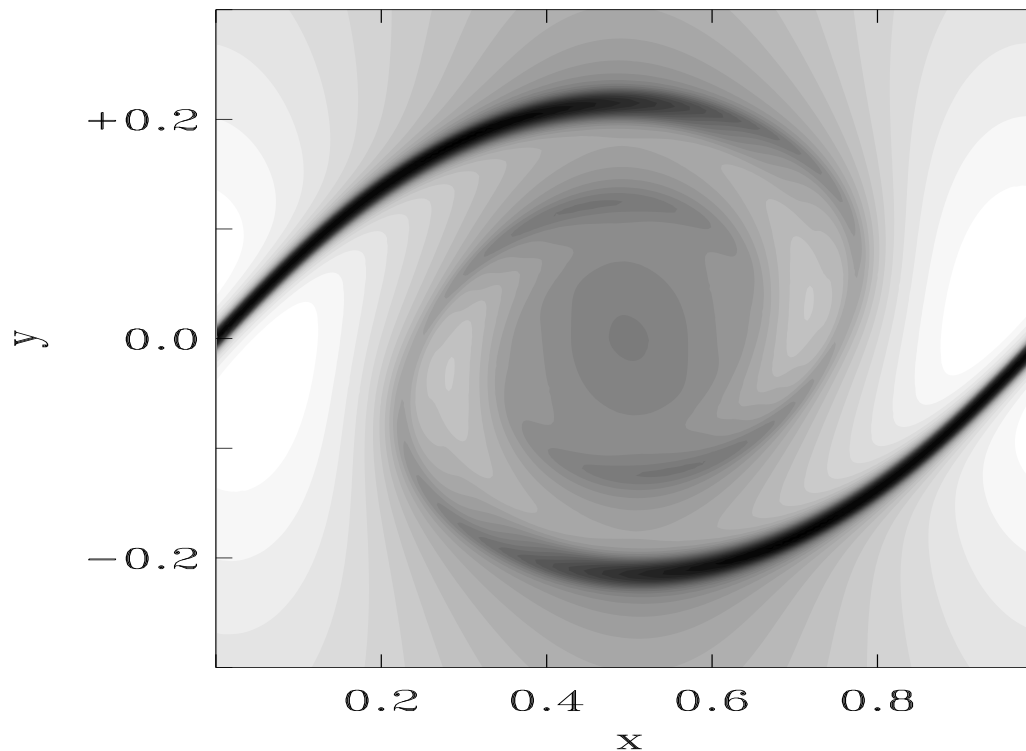
- destabilizing influence of reversed B stronger for B ↗

⇒ growth rate as function of Alfvén Mach M_a

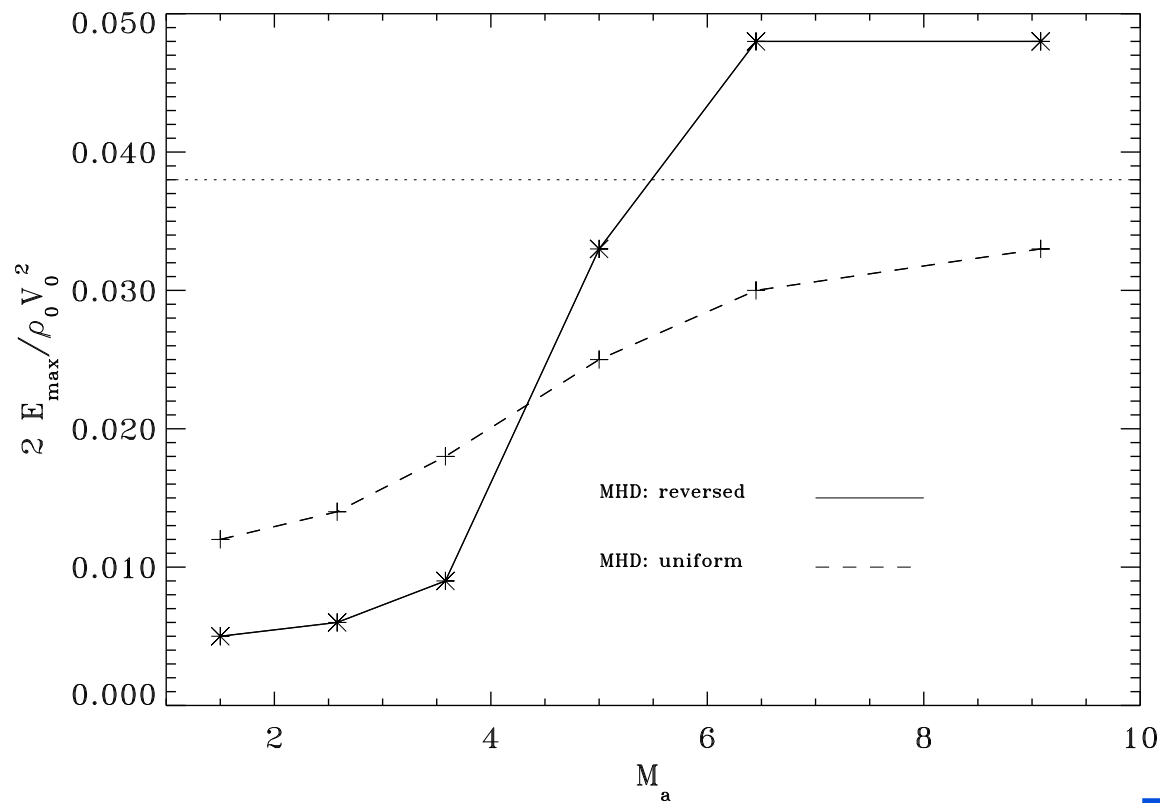


Nonlinear results

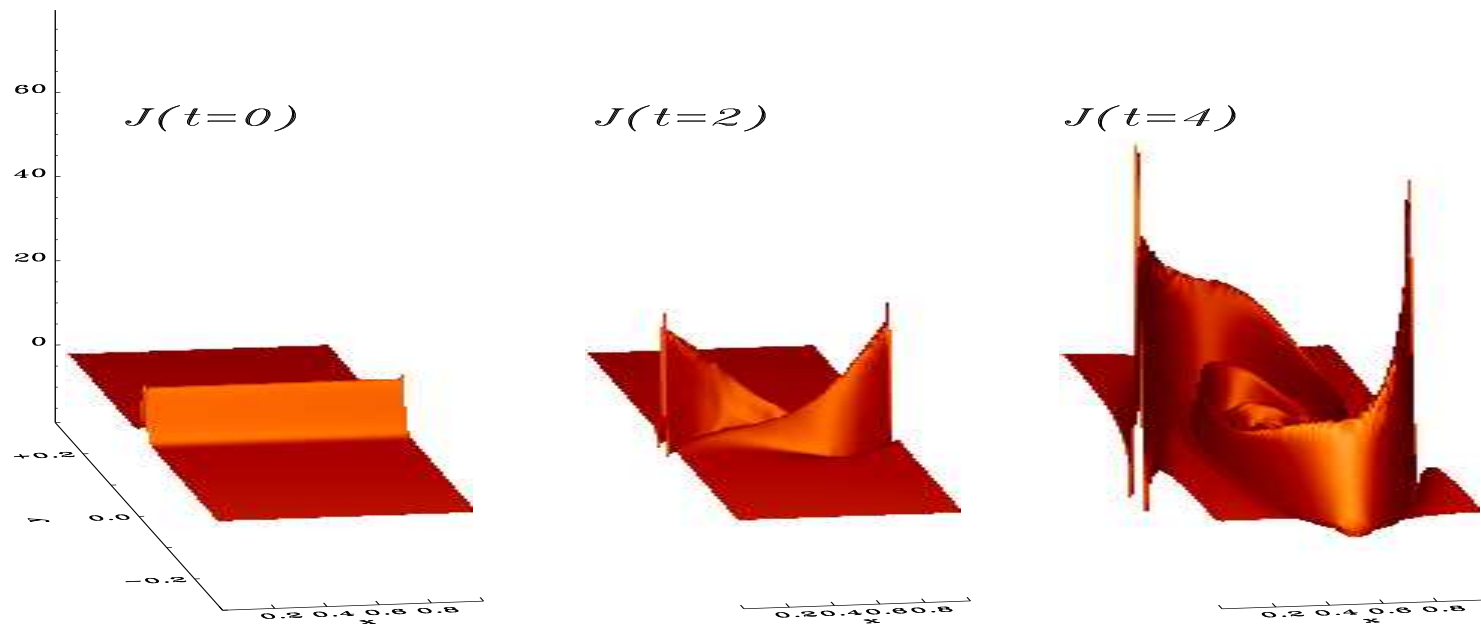
- KH instability for reference parameter values
 - ⇒ weak initial \mathbf{B} ($\beta = 120$): vortical flow pattern sets in
 - ⇒ magnetic field dragged into spiral configuration
 - ⇒ eventually gets amplified in thin spiral sheetlike structure
 - ⇒ \mathbf{B} locally dynamically important: halts further winding
 - ⇒ where \mathbf{B} is amplified in sheets: cospatial low density lanes



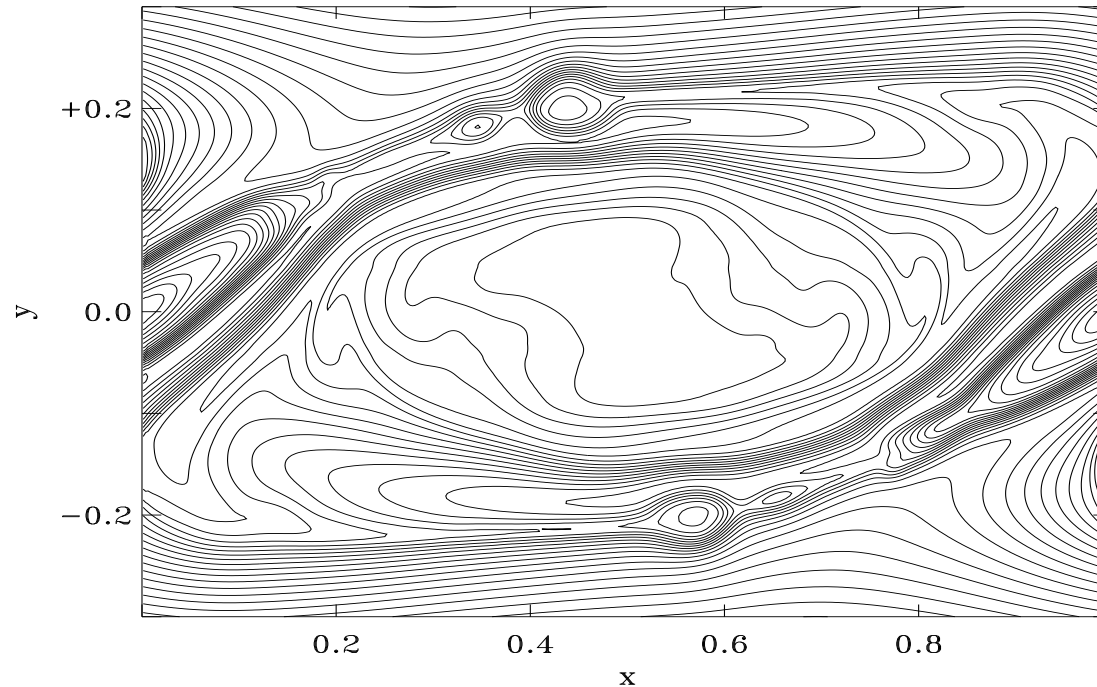
- study of saturation energy level E_{max} versus $k_x a$, M_s , and M_a
 - ⇒ indication of non-linear saturation
- dependence on Alfvén Mach number (initial field strength)
 - ⇒ reversed case may saturate above/intermediate/below HD value



- reversed case: initial $\mathbf{J} = \nabla \times \mathbf{B}$ infinitely thin sheet
 - \Rightarrow limit $b \rightarrow 0$ with $B_x = \tanh(y/b)$ [then $p(y)$]
 - \Rightarrow additional pinching mode (with reconnection) accessible
- current sheet gets amplified by vortex flow
 - \Rightarrow strong current sheet \Rightarrow resistive dissipation important
 - \Rightarrow systematic study in resistive MHD for $\eta \rightarrow 0$ (finer grid)
 - \Rightarrow extra magnetic energy tapped

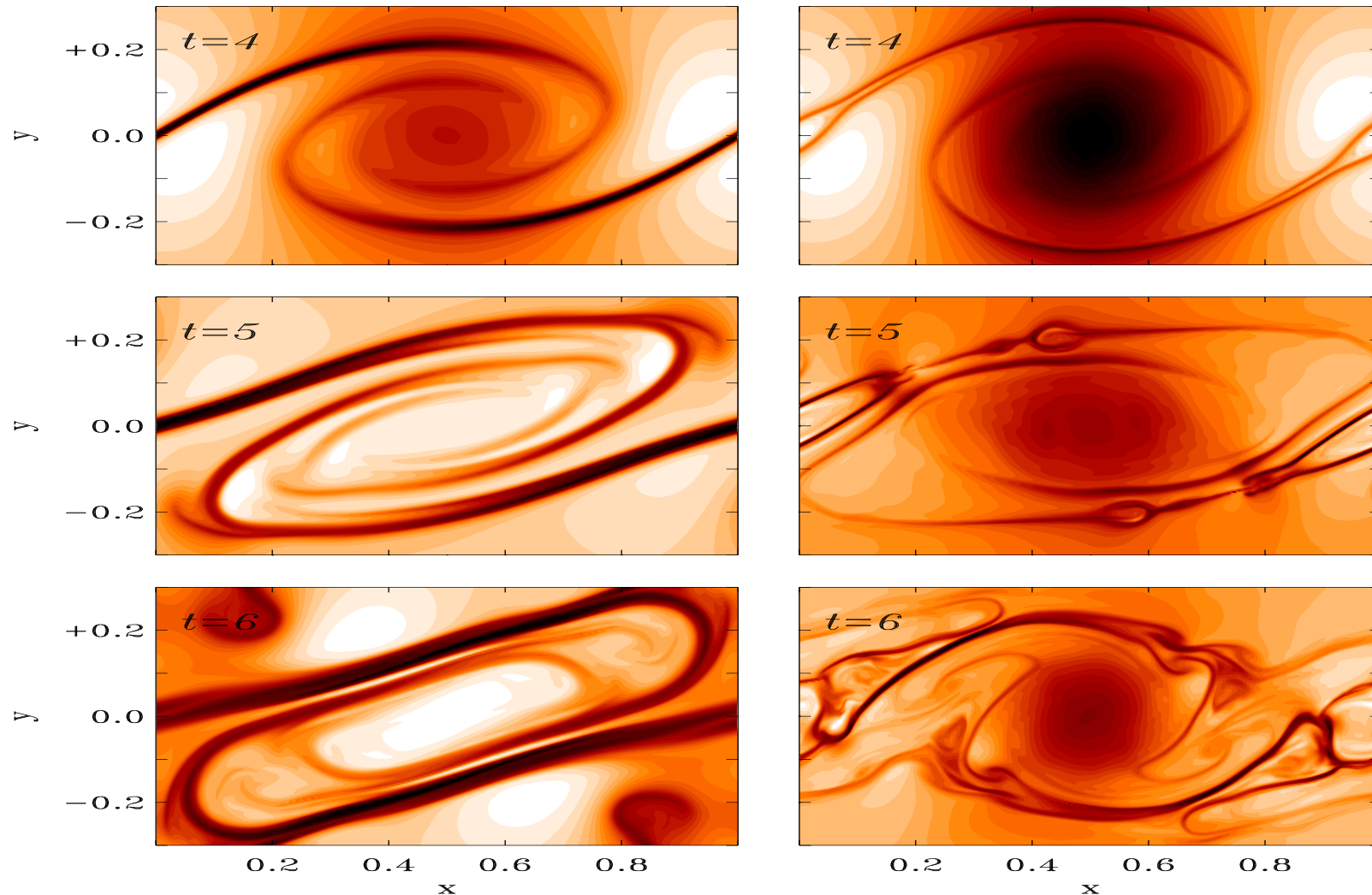


- anti-parallel field lines pushed together
⇒ magnetic islands form (tearing unstable)



- reconnection plays role sooner than for uniform case
- turbulent state sets in fast, complicating saturation behaviour
- compressibility: density deviations up to 40 %

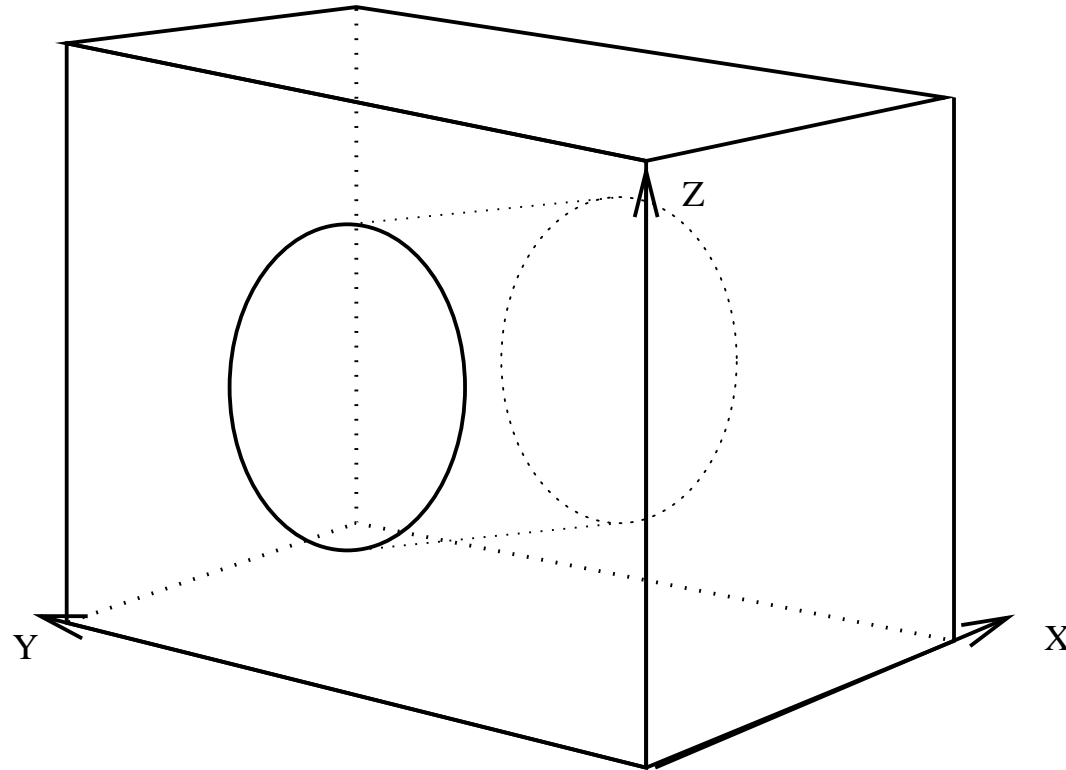
- density evolution for case without and with J-sheet



- ⇒ induced island formation ('tearing') when J-sheet
- ⇒ 2D current-vortex transits to magnetically modified turbulence
- ⇒ KH unstable shear flow triggers small-scale reconnection events

3D KH case study

- Model [astrophysical] jet of radius R_{jet}



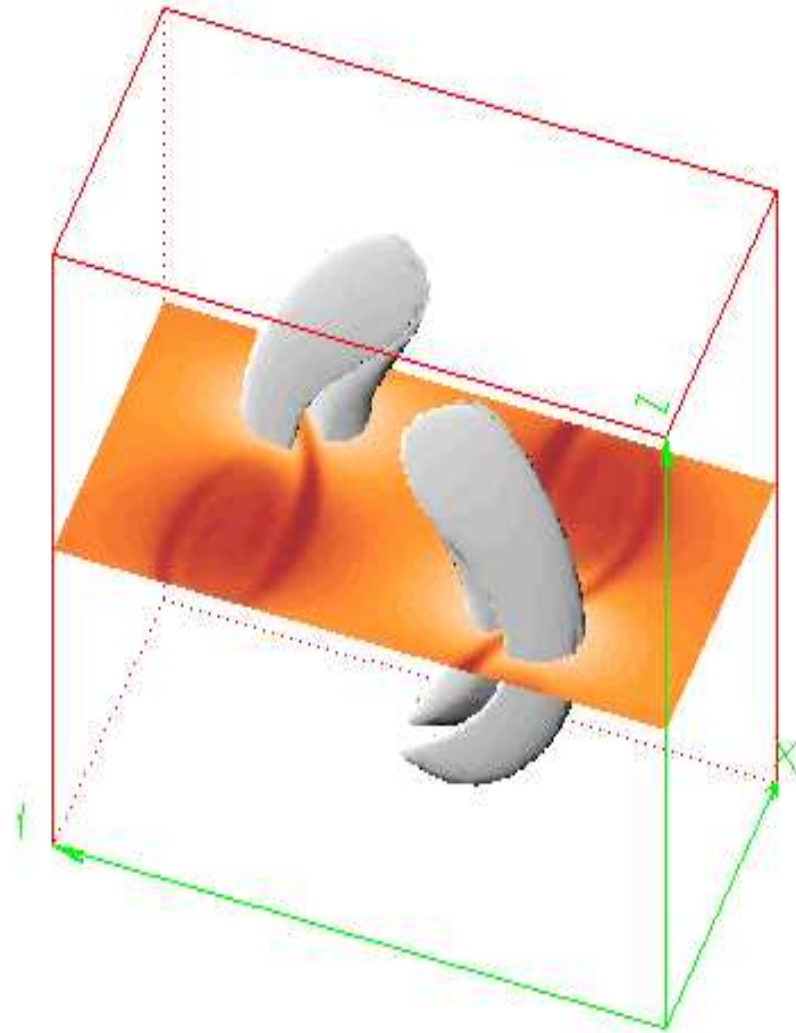
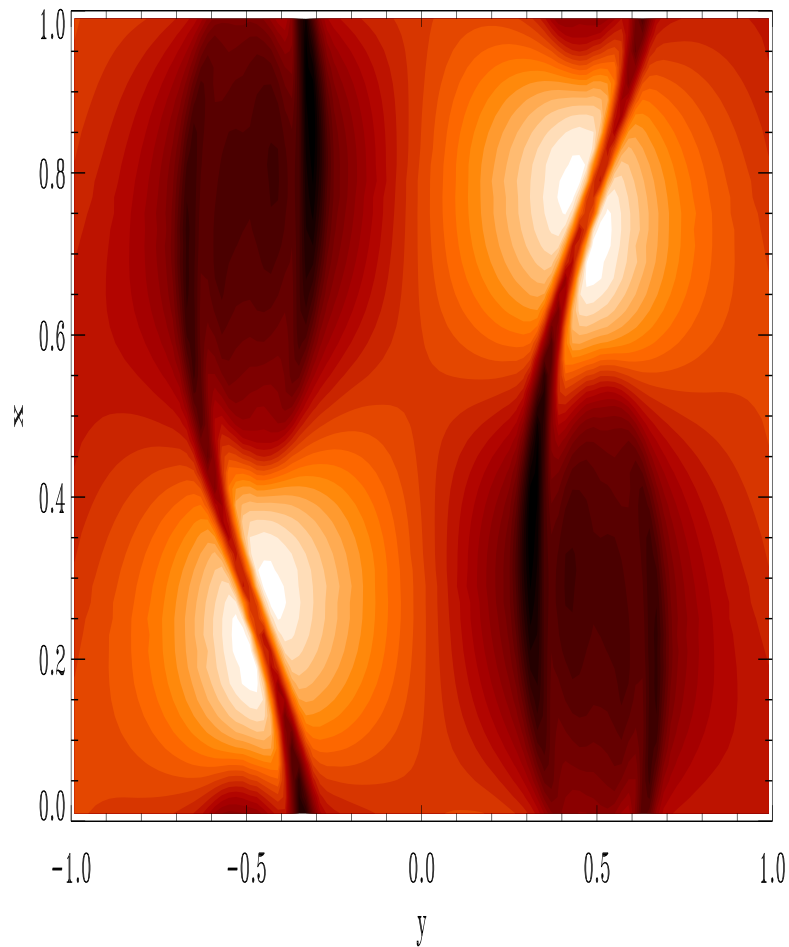
⇒ shear flow (width $2a$) across its circumference

⇒ $t = 0$ parallel uniform \mathbf{B} , reference $V_0 = 0.645$, $a = 0.05$, $B_0 = 0.129$

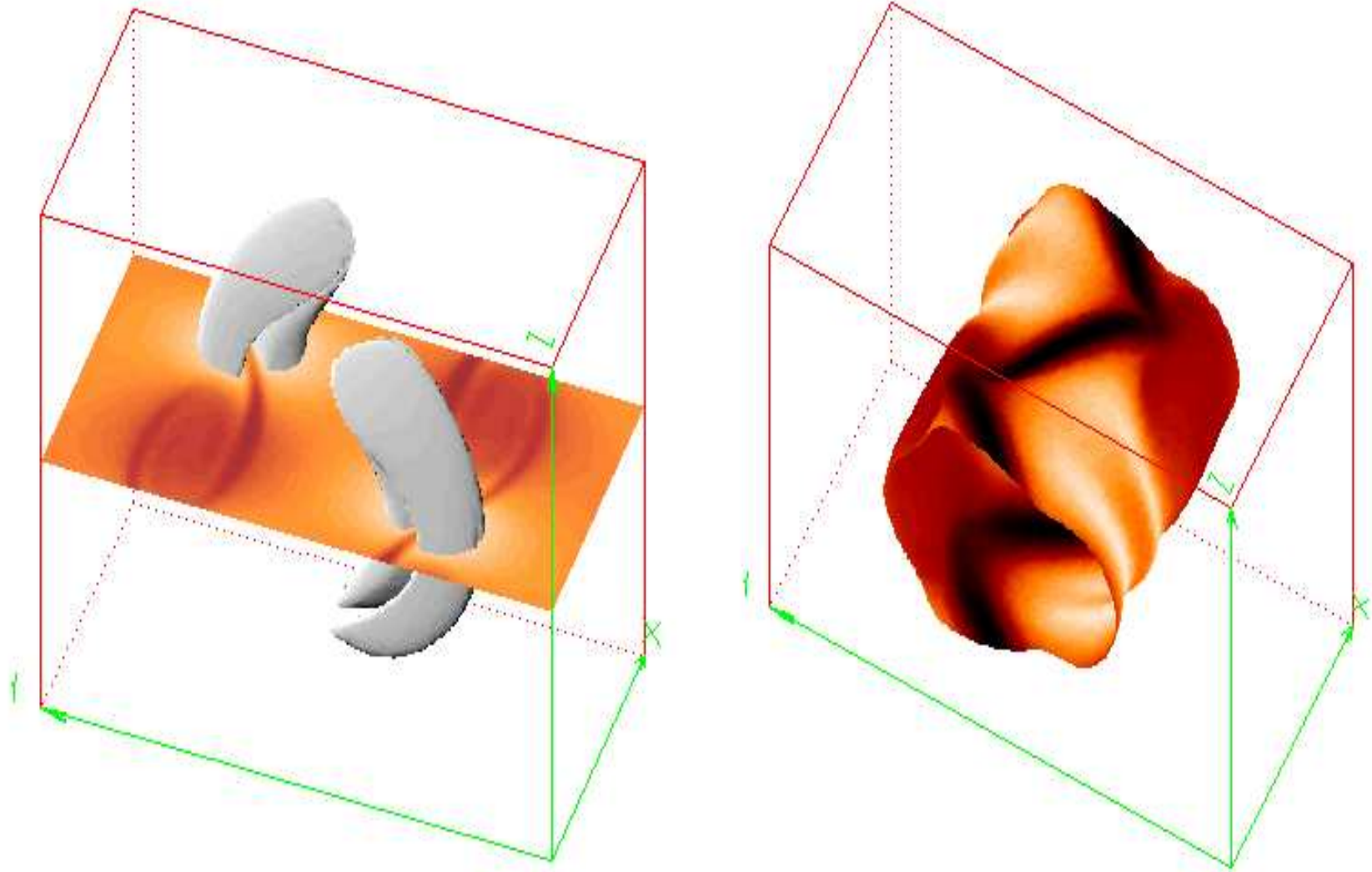
- 3D perturbation wavenumber n along jet, m about jet axis

⇒ sideways 'kink' perturbation $m = 1 = n$

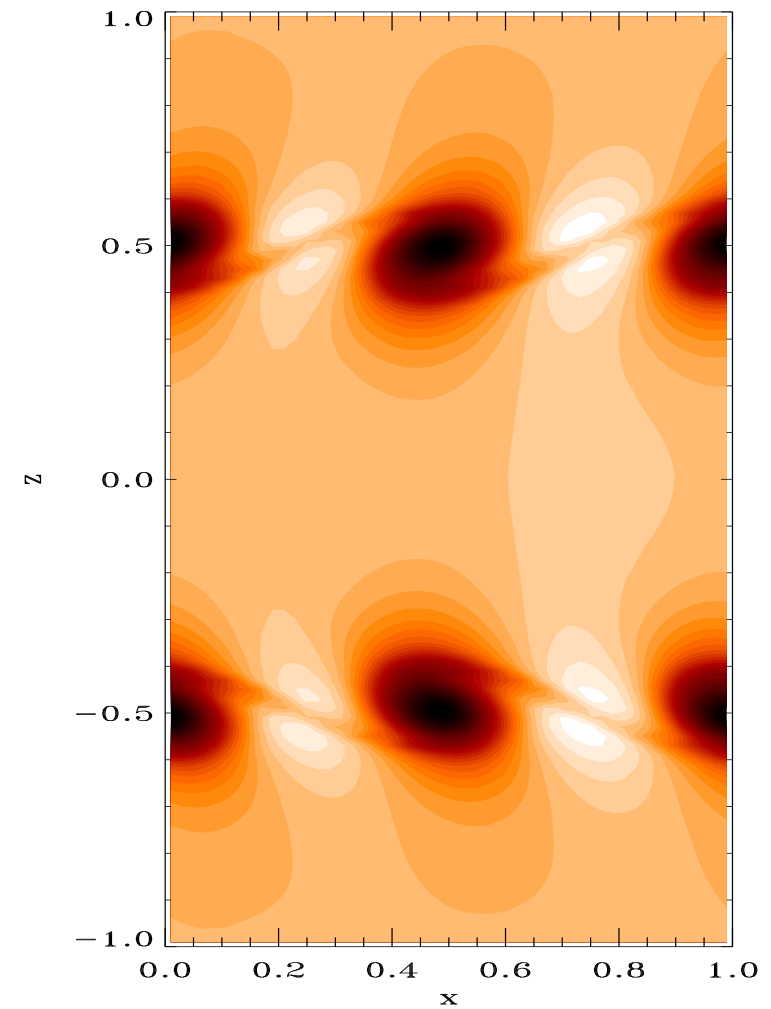
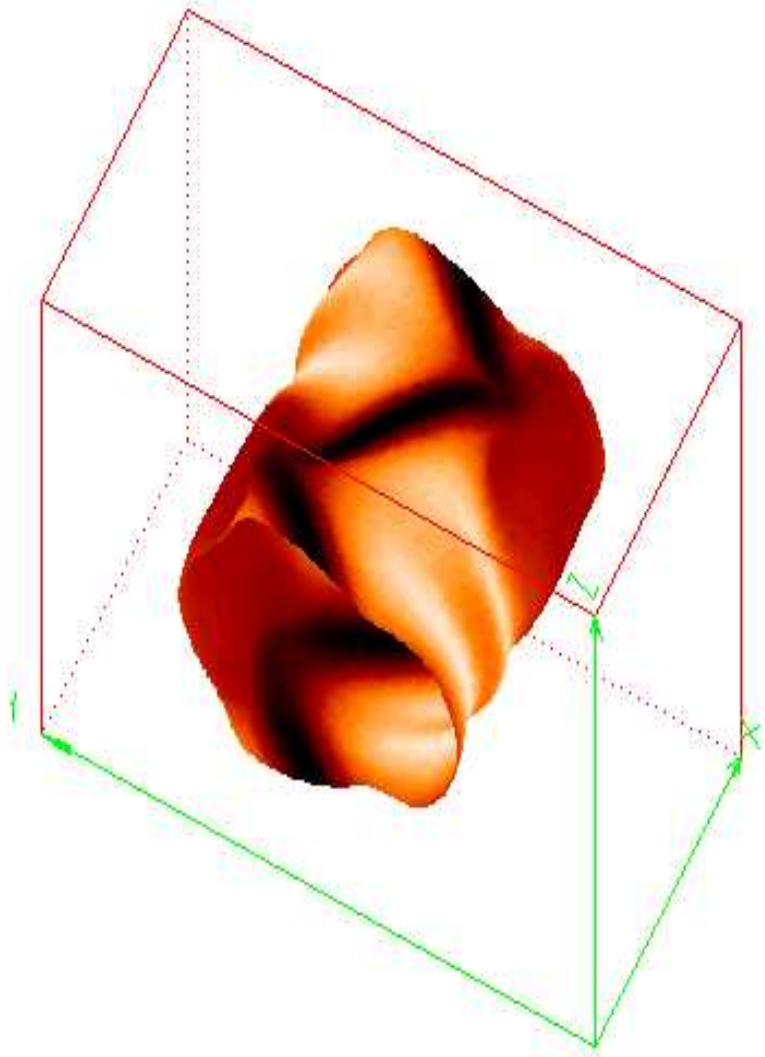
- in horizontal cross-cut: doubled 2D result: ρ at $t = 4$



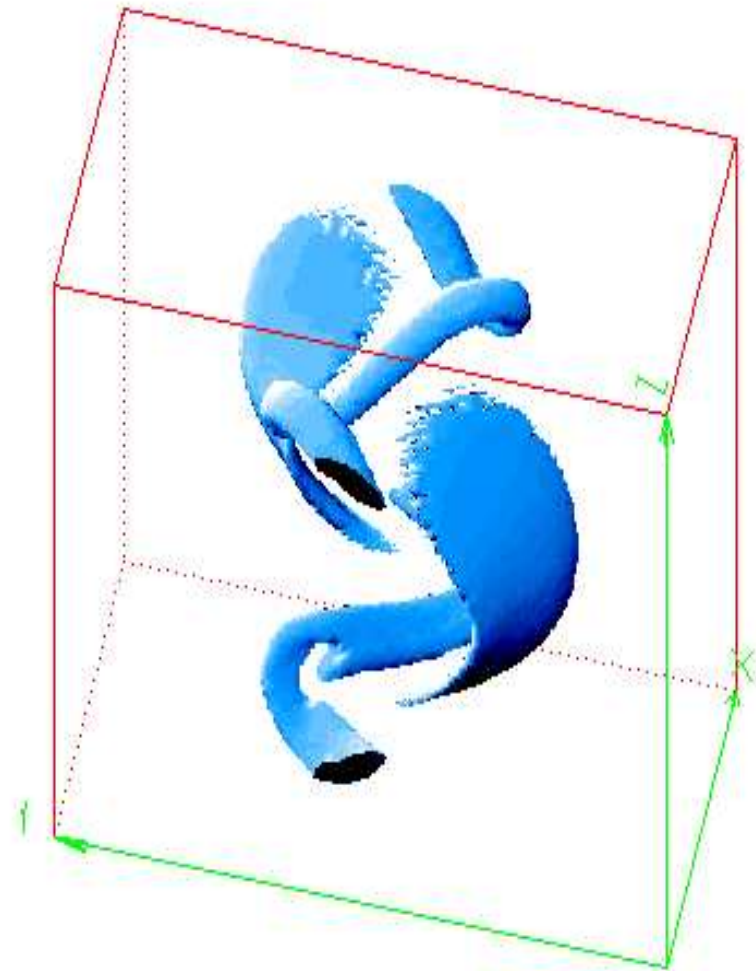
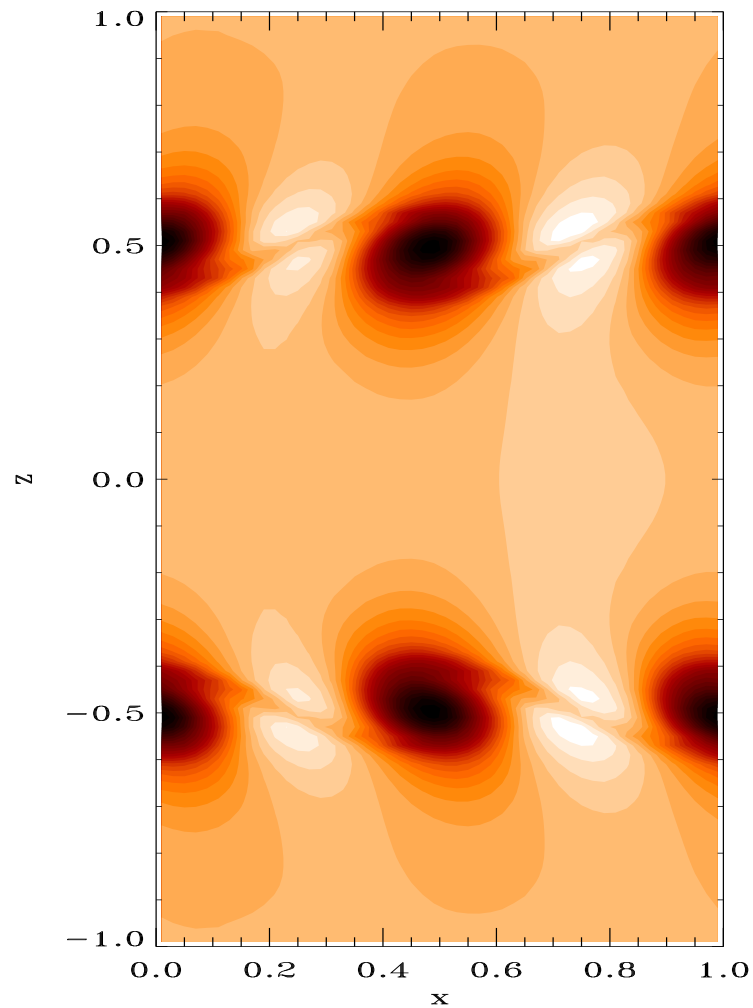
- high ρ isosurface and $v_x = 0$ jet surface colored by p_{th}



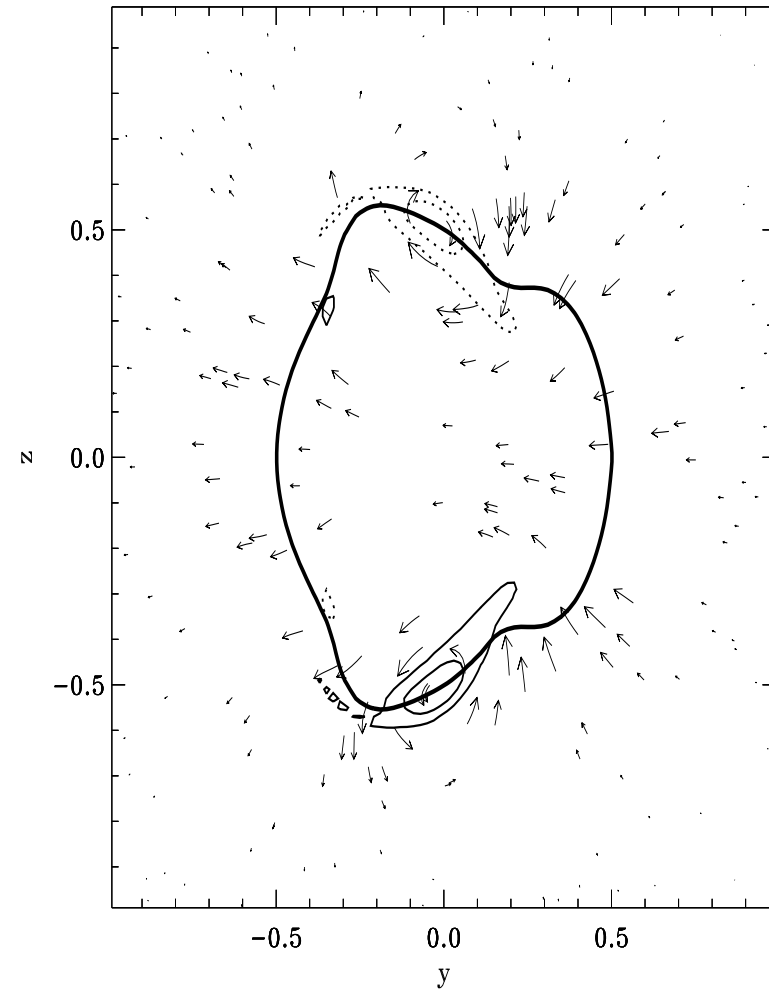
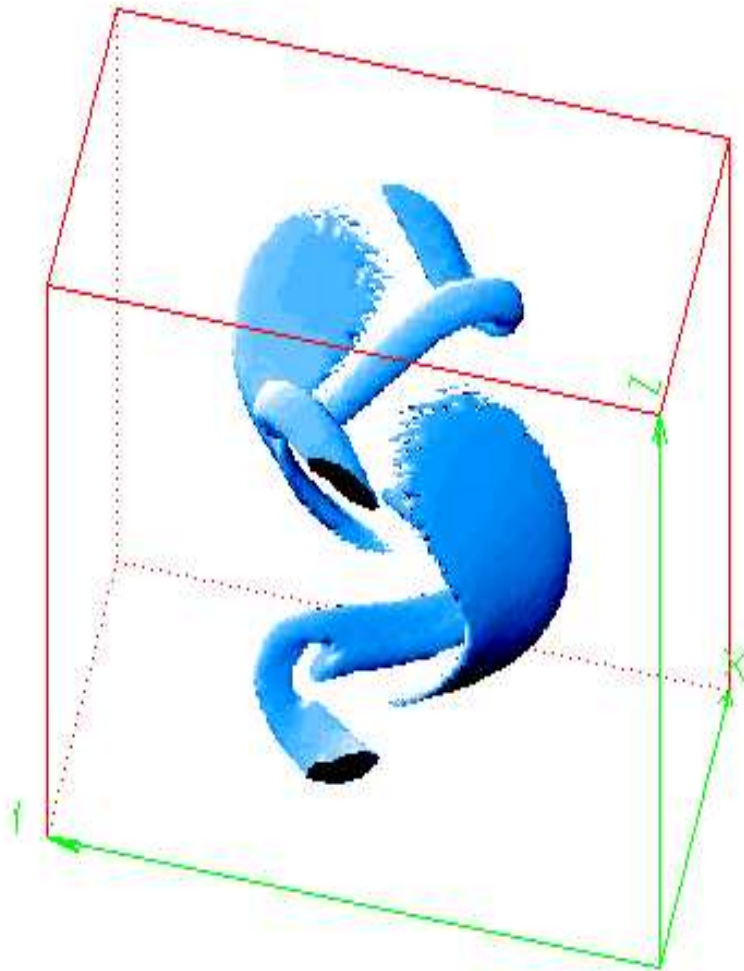
- p_{th} gradient induces wavenumber doubling on top/bottom



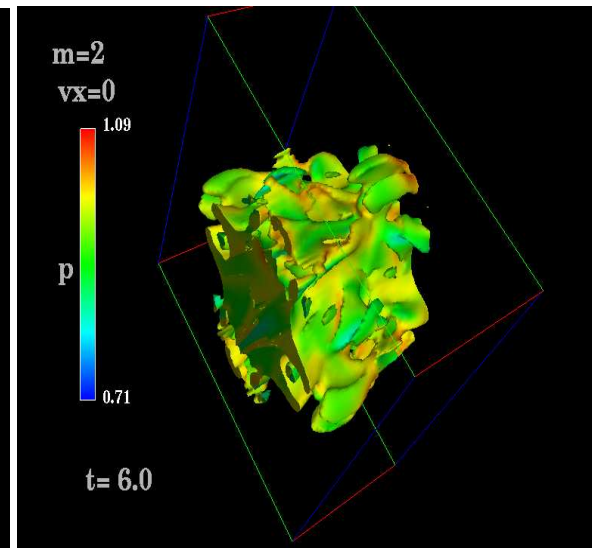
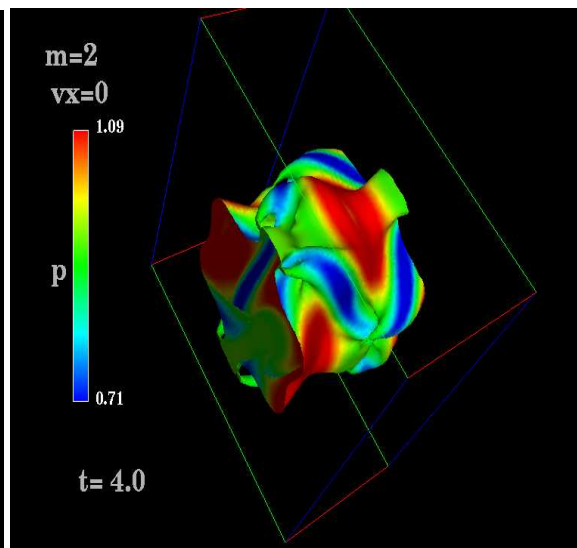
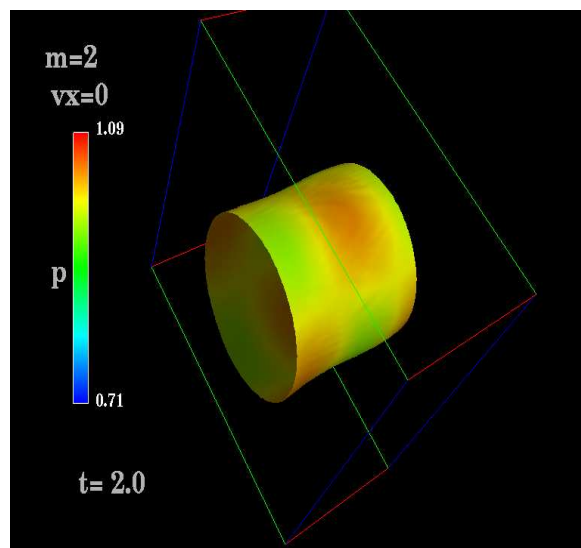
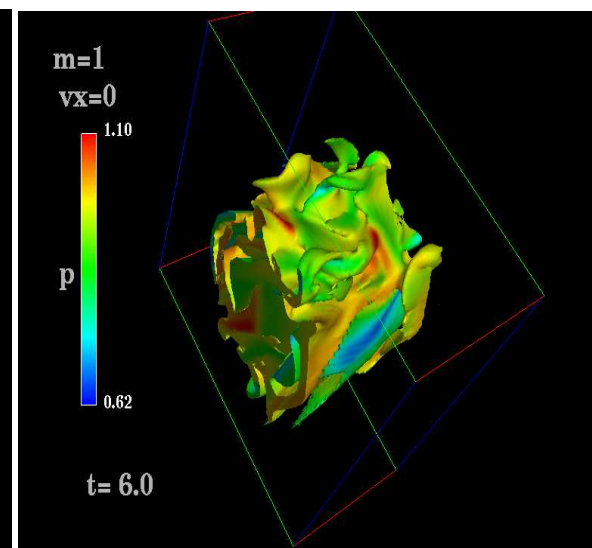
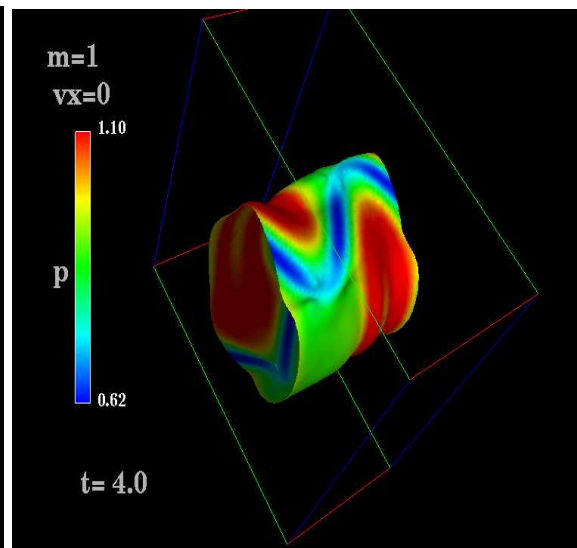
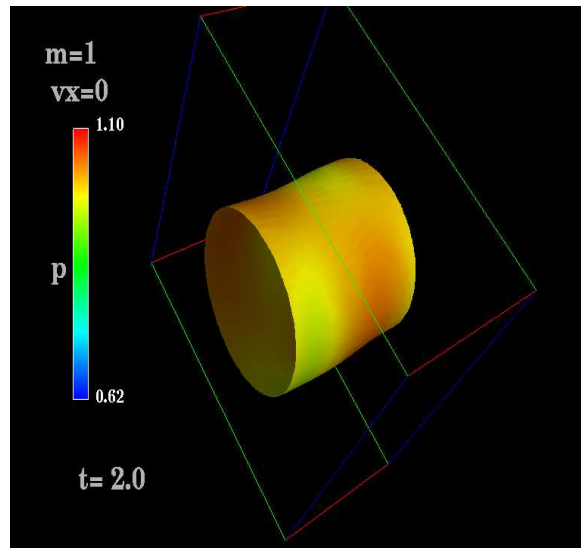
- low ρ lanes: 3D fibril and sheet structures: cospatial with high B_{pol}



- localized 3D high B_{pol} regions control jet deformation



Kelvin-Helmholtz unstable jet : $m=1$ vrs. $m=2$ breakup.



References

- R. Keppens et al., *J. Plasma Physics* **61**, 1 (1999)
- R. Keppens, G. Tóth, *Physics of Plasmas* **6**(5), 1461 (1999)

