# Advanced image synthesis

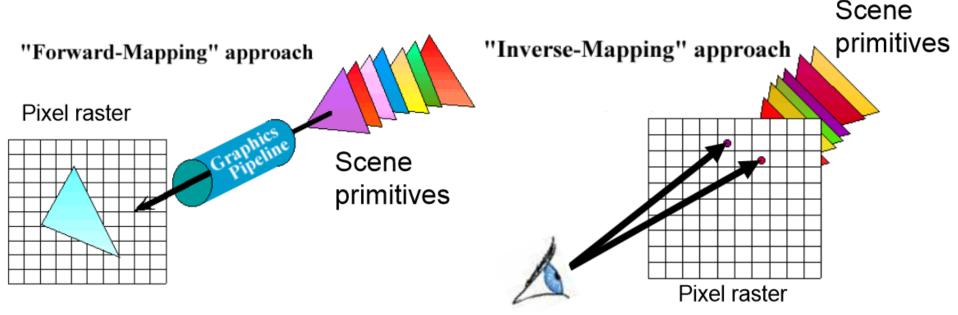
Romain Vergne – 2014/2015



### Rasterization VS ray-casting

- For each triangle
  - Project triangle to image plane
  - For each pixel
    - Check pixel in triangle
    - Resolve visibility with z-buffer

- For each pixel
  - Compute pixel ray
  - For each triangle
    - Check ray-triangle intersection
    - Get closest intersection





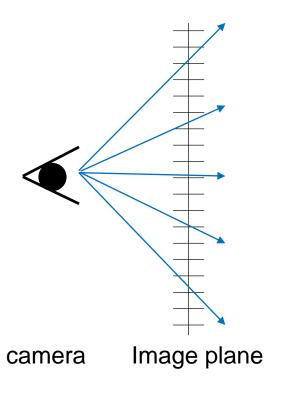
#### Eye ray and camera

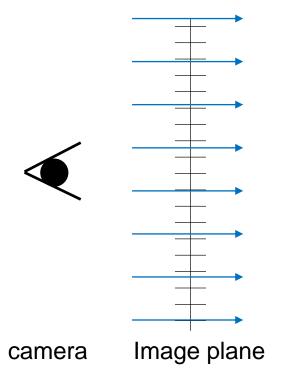
Persective

 $r = (x^*u,aspect^*y^*v,D^*w), normalized$  $P(t) = e + t^*r$  Orthographic

$$P(t) = o + t*w$$

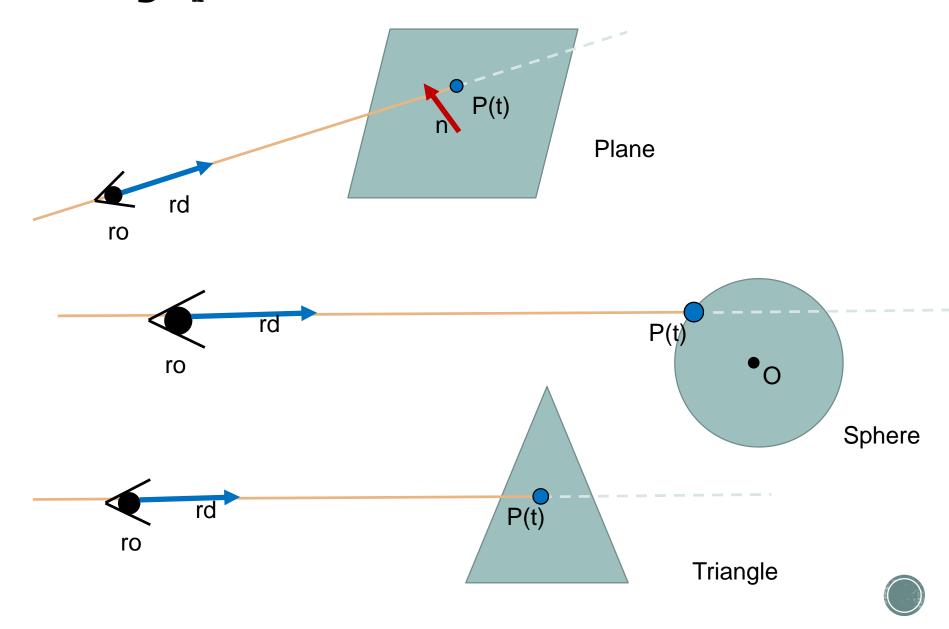
$$o = e + x*size*u + y*size*v$$







### Ray-plane intersection

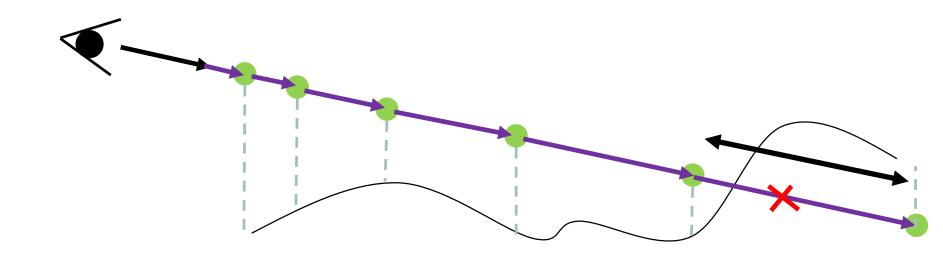


# Ray marching height-fields

$$P(t) = r_0 + r_d * t$$

#### **Optimizations:**

- Interpolate between the 2 last positions
- Increase deltaT with distance from eye
- See: <a href="http://www.iquilezles.org">http://www.iquilezles.org</a>



Height field f(x,z) = y

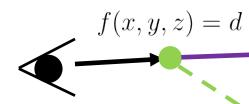




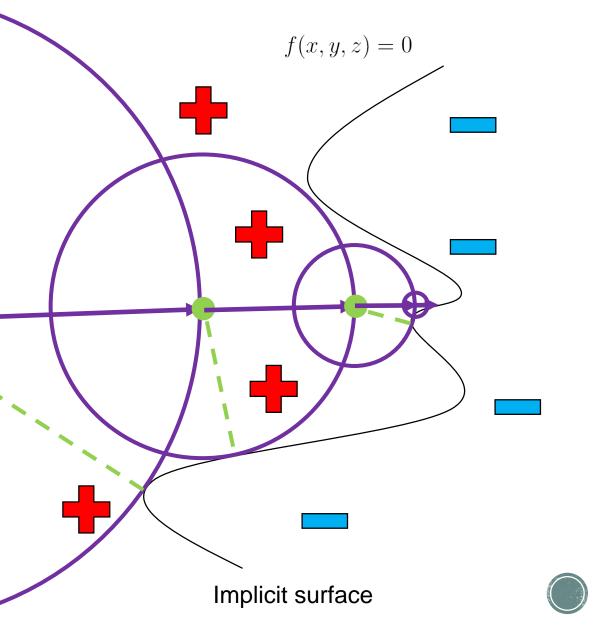
### Ray marching implicit surfaces



- 1: t = 0
- $/\!\!\!/ 2: d=0$ 
  - 3: while  $t < D \operatorname{do}$
  - 4: d = f(r(t))
  - 5: if  $d < \epsilon$  then return t
  - 6: t = t + d
  - 7: **return** 0



camera



#### Common implicit surfaces

#### Sphere - signed

```
float sdSphere( vec3 p, float s )
{
  return length(p)-s;
}
```

#### Plane - signed

```
float sdPlane( vec3 p, vec4 n )
{
   // n must be normalized
  return dot(p,n.xyz) + n.w;
}
```

#### Box - signed

#### Torus - signed

```
float sdTorus( vec3 p, vec2 t )
{
  vec2 q = vec2(length(p.xz)-t.x,p.y);
  return length(q)-t.y;
}
```

#### Round Box - unsigned

```
float udRoundBox( vec3 p, vec3 b, float r )
{
  return length(max(abs(p)-b,0.0))-r;
}
```

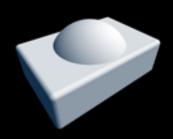
#### Cone - signed

```
float sdCone( vec3 p, vec2 c )
{
    // c must be normalized
    float q = length(p.xy);
    return dot(c,vec2(q,p.z));
}
```

#### Distance operations

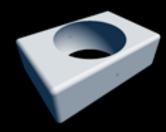
#### Union

```
float opU( float d1, float d2 )
{
    return min(d1,d2);
}
```



#### Substraction

```
float opS( float d1, float d2 )
{
    return max(-d1,d2);
}
```



#### Intersection

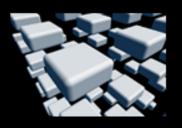
```
float opI( float d1, float d2 )
{
    return max(d1,d2);
}
```



#### Domain operations

#### Repetition

```
float opRep( vec3 p, vec3 c )
{
    vec3 q = mod(p,c)-0.5*c;
    return primitve( q );
}
```



#### Scale

```
float opScale( vec3 p, float s )
{
    return primitive(p/s)*s;
}
```



#### Rotation/Translation

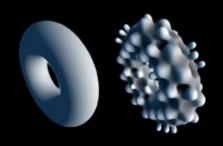
```
vec3 opTx( vec3 p, mat4 m )
{
    vec3 q = invert(m)*p;
    return primitive(q);
}
```



#### Distance deformations

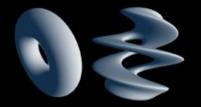
#### Displacement

```
float opDisplace( vec3 p )
{
    float d1 = primitive(p);
    float d2 = displacement(p);
    return d1+d2;
}
```



#### **Twist**

```
float opTwist( vec3 p )
{
    float c = cos(20.0*p.y);
    float s = sin(20.0*p.y);
    mat2 m = mat2(c,-s,s,c);
    vec3 q = vec3(m*p.xz,p.y);
    return primitive(q);
}
```



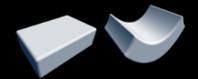
#### Blend

```
float opBlend( vec3 p )
{
    float d1 = primitiveA(p);
    float d2 = primitiveB(p);
    return smin( d1, d2 );
}
```

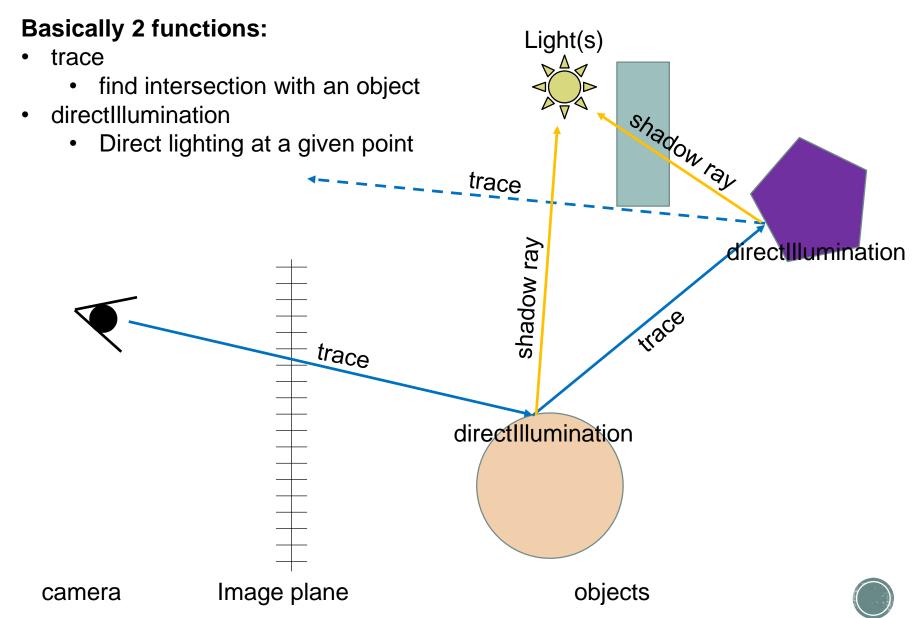


#### **Cheap Bend**

```
float opCheapBend( vec3 p )
{
    float c = cos(20.0*p.y);
    float s = sin(20.0*p.y);
    mat2 m = mat2(c,-s,s,c);
    vec3 q = vec3(m*p.xy,p.z);
    return primitive(q);
}
```



# Ray tracing



### Ray tracing

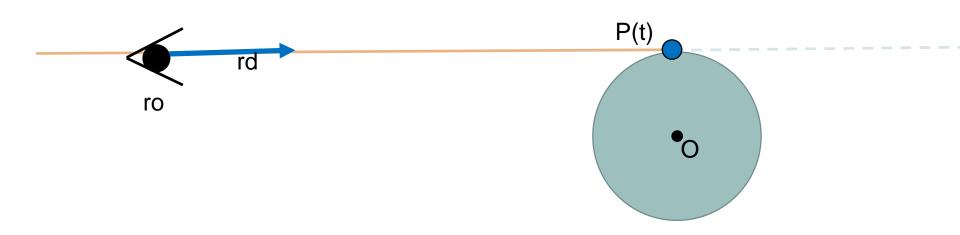
```
color trace(ray) {
  • hit = intersectScene(ray)
  • if(hit) {
    color = directIllumination(hit)
    if hit is reflective
       color += c_refl * trace(reflected ray)
    if hit is transmissive
       color += c_trans * trace(refracted ray)
  } else
    color = background_color
  return color
```

### Ray tracing

```
color directIllumination(hit) {
  - color = (0,0,0)
  for each light L {
    T = cast shadow ray to L
    if hit is not shadowed by L
      color += Ambient+diffuse+specular terms(L,hit)
  • }
  return color
```

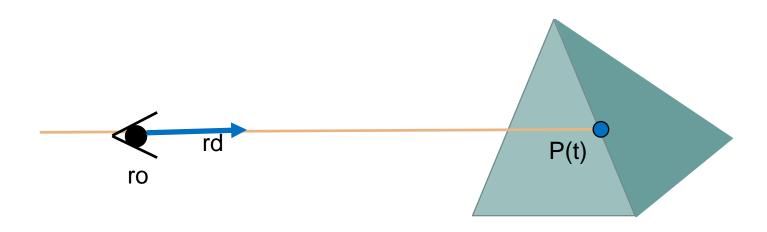
#### Precision

- Issues
  - Ray origin on an object surface
  - Grazing rays
- Floating point approximation



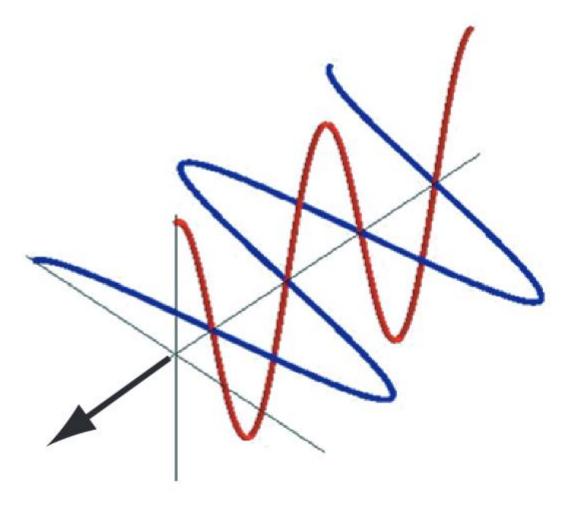
#### Precision

- Issues
  - Ray origin on an object surface
  - Grazing rays
- Floating point approximation
  - Must report intersection on triangles





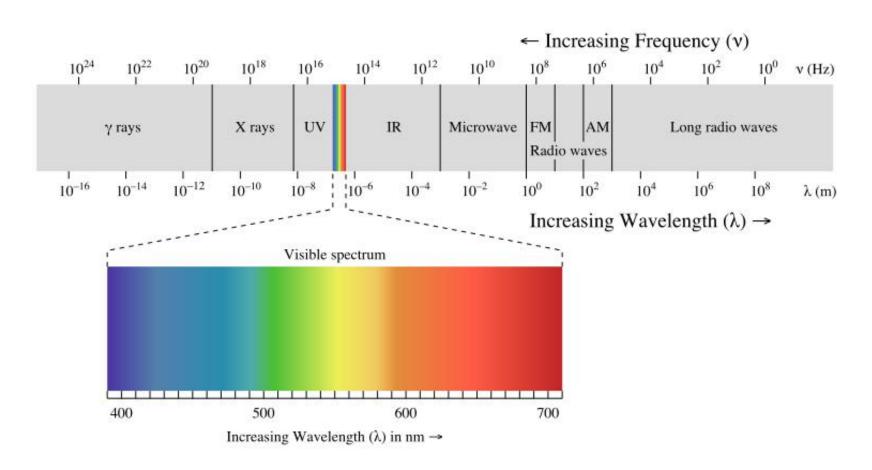
Light: electromagnetic transverse wave



See Siggraph 2014 course by Naty Hoffman – almost everything from there!! And images from « Real-time rendering » 3rd edition (A K Peters - 2008)

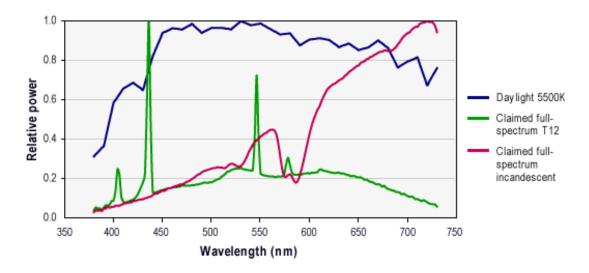


Visible light: between 400 and 700 nm



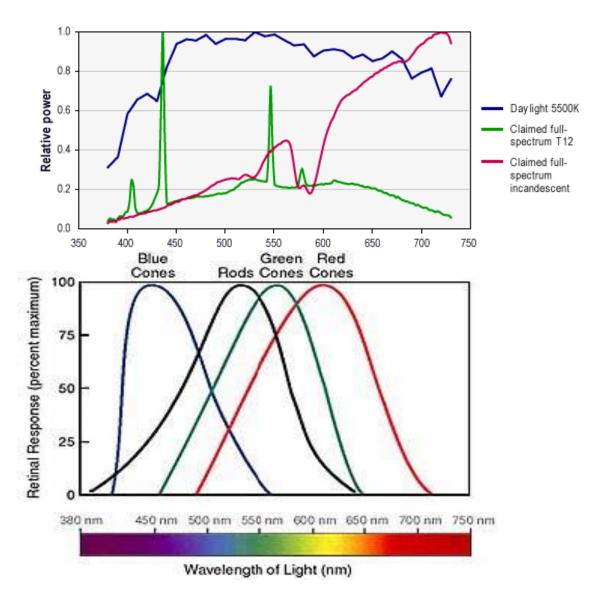


Visible light: between 400 and 700 nm





Visible light: between 400 and 700 nm





Light travels in straight line (homogeneous medium)

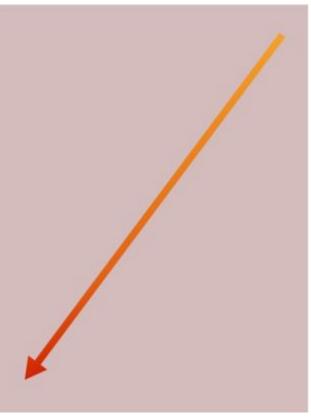






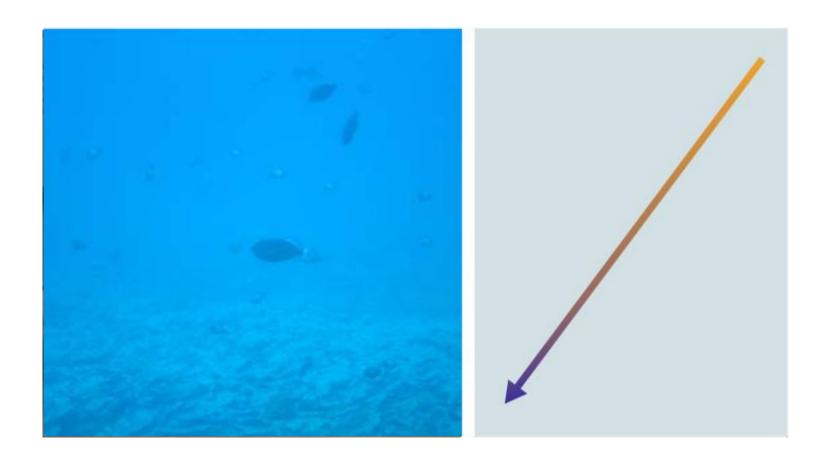
Absorption of parts of visible light





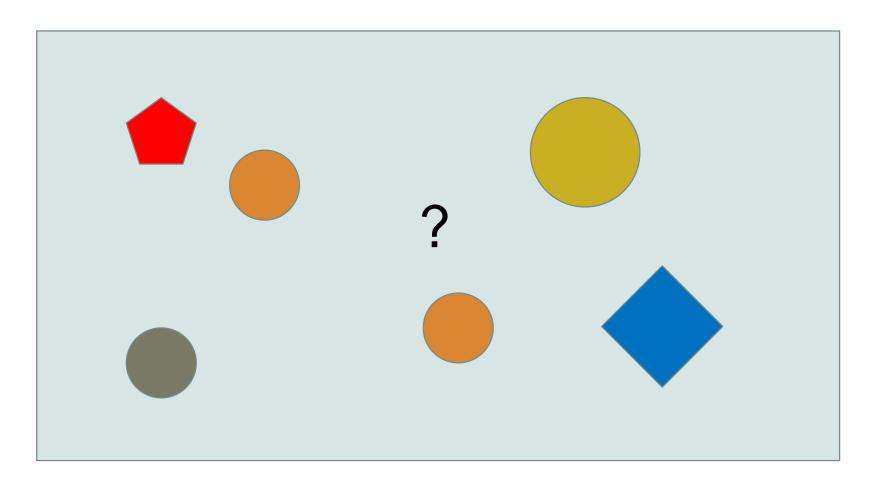


Absorption of parts of visible light



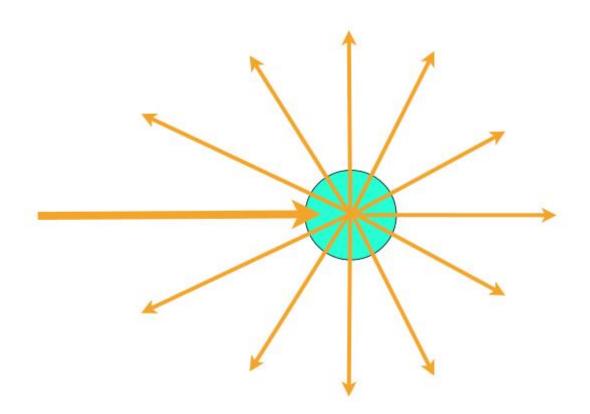


Non homogeneous media?





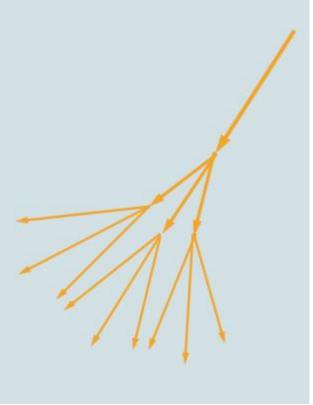
- Non homogeneous media?
  - Scattering due changes in the index of refraction





- Non homogeneous media?
  - Scattering due changes in the index of refraction

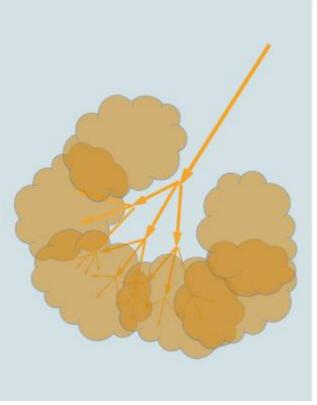






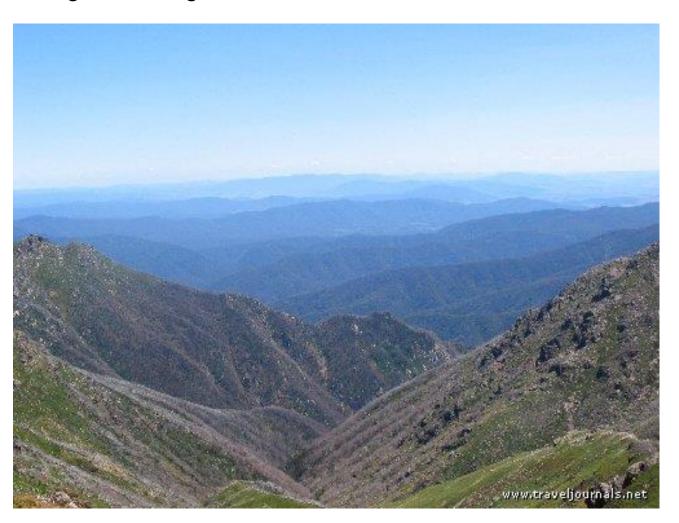
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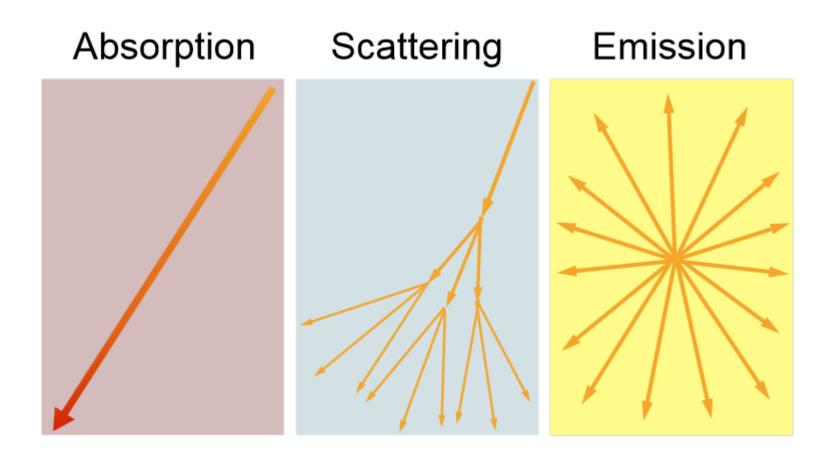




- Non homogeneous media?
  - Scattering due changes in the index of refraction



3 Modes of light / matter interaction





3 Modes of light / matter interaction

Absorption (color)



Scattering (cloudiness)



What about surfaces?











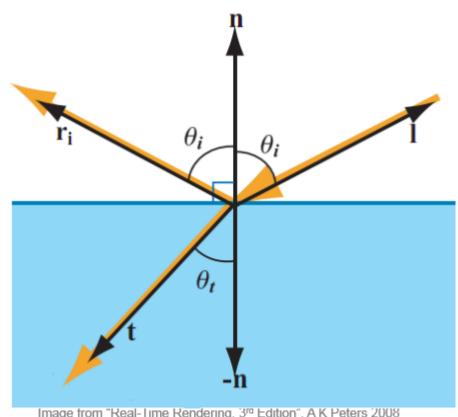








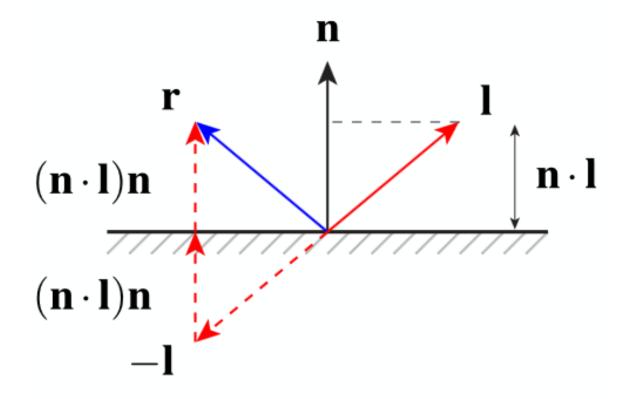
Case of (optically) flat surface: Snell Descartes laws







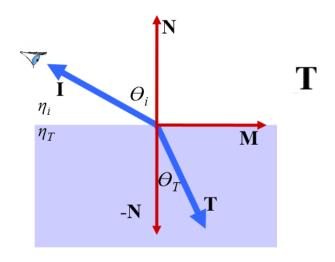
- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...



$$\mathbf{r} = 2(\mathbf{n} \cdot \mathbf{l})\mathbf{n} - \mathbf{l}$$



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted



$$n_i \sin \theta_i = n_T \sin \theta_T$$

$$\frac{\sin \theta_T}{\sin \theta_i} = \frac{n_i}{n_T} = n_r$$



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted

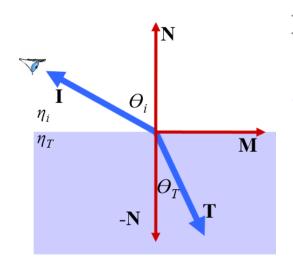
$$\mathbf{I} = \mathbf{N} \cos \theta_i - \mathbf{M} \sin \theta_i$$
$$\mathbf{M} = (\mathbf{N} \cos \theta_i - \mathbf{I}) / \sin \theta_i$$

$$n_i \sin \theta_i = n_T \sin \theta_T$$

$$\frac{\sin \theta_T}{\sin \theta_i} = \frac{n_i}{n_T} = n_r$$



- Case of (optically) flat surface: Snell Descartes laws
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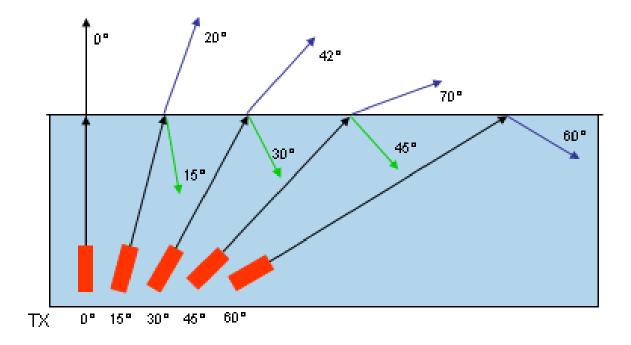
$$n_i \sin \theta_i = n_T \sin \theta_T$$

$$\frac{\sin \theta_T}{\sin \theta_i} = \frac{n_i}{n_T} = n_r$$

$$\begin{split} \mathbf{I} &= \mathbf{N} \cos \theta_{i} - \mathbf{M} \sin \theta_{i} \\ \mathbf{M} &= (\mathbf{N} \cos \theta_{i} - \mathbf{I}) / \sin \theta_{i} \\ \mathbf{T} &= -\mathbf{N} \cos \theta_{T} + \mathbf{M} \sin \theta_{T} \\ &= -\mathbf{N} \cos \theta_{T} + (\mathbf{N} \cos \theta_{i} - \mathbf{I}) \sin \theta_{T} / \sin \theta_{i} \quad Plug \, \mathbf{M} \\ &= -\mathbf{N} \cos \theta_{T} + (\mathbf{N} \cos \theta_{i} - \mathbf{I}) \, \eta_{r} \quad let's \, get \, rid \, of \\ &= [\, \eta_{r} \cos \theta_{i} - \cos \theta_{T}] \, \mathbf{N} - \eta_{r} \, \mathbf{I} \\ &= [\, \eta_{r} \cos \theta_{i} - \sqrt{1 - \sin^{2} \theta_{T}}] \, \mathbf{N} - \eta_{r} \, \mathbf{I} \\ &= [\, \eta_{r} \cos \theta_{i} - \sqrt{1 - \eta_{r}^{2} \sin^{2} \theta_{i}}] \, \mathbf{N} - \eta_{r} \, \mathbf{I} \\ &= [\, \eta_{r} \cos \theta_{i} - \sqrt{1 - \eta_{r}^{2} (1 - \cos^{2} \theta_{i})}] \, \mathbf{N} - \eta_{r} \, \mathbf{I} \\ &= [\, \eta_{r} (\mathbf{N} \cdot \mathbf{I}) - \sqrt{1 - \eta_{r}^{2} (1 - (\mathbf{N} \cdot \mathbf{I})^{2})}] \, \mathbf{N} - \eta_{r} \, \mathbf{I} \end{split}$$



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted



Total internal reflection



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted



Fig. 3.7A The optical manhole. From under water, the entire celestial hemisphere is compressed into a circle only 97.2° across. The dark boundary defining the edges of the manhole is not sharp due to surface waves. The rays are analogous to the crepuscular type seen in hazy air, Section 1.9. (Photo by D. Granger)

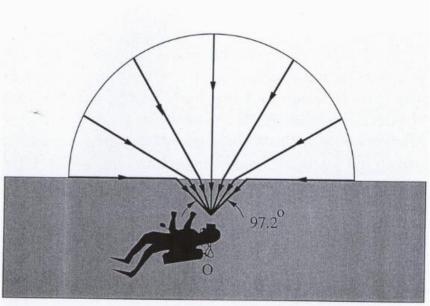
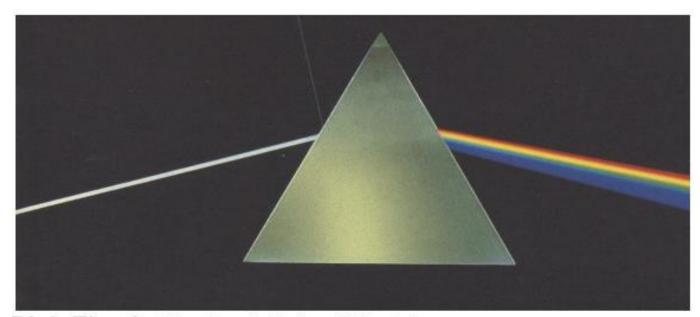


Fig. 3.7B The optical manhole. Light from the horizon (angle of incidence =  $90^{\circ}$ ) is refracted downward at an angle of  $48.6^{\circ}$ . This compresses the sky into a circle with a diameter of  $97.2^{\circ}$  instead of its usual  $180^{\circ}$ .



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted



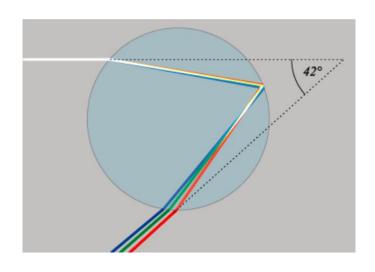
Pink Floyd, The Dark Side of the Moon

Refraction is wavelet dependent



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted





#### How rainbows work:

- Refraction + internal reflection + refraction
- Max for angle around 42 deg



- Case of (optically) flat surface: Snell Descartes laws
  - Incident ray is reflected...
  - ... and refracted
- The amount of reflection vs refraction.
  - Controlled with Fresnel law (electromagnetic wave)

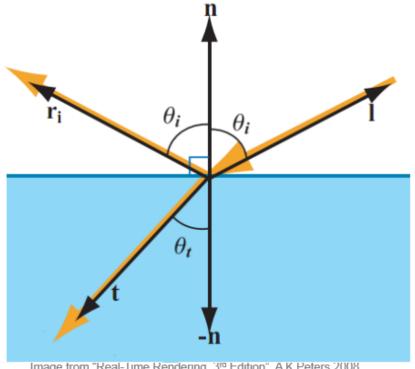




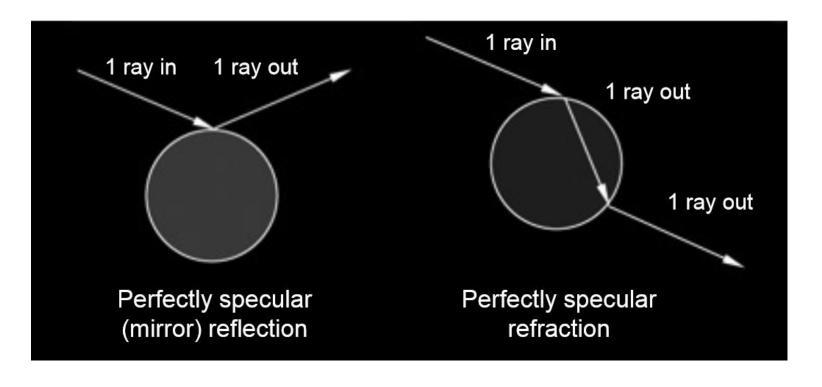
Image from "Real-Time Rendering, 3rd Edition", A K Peters 2008

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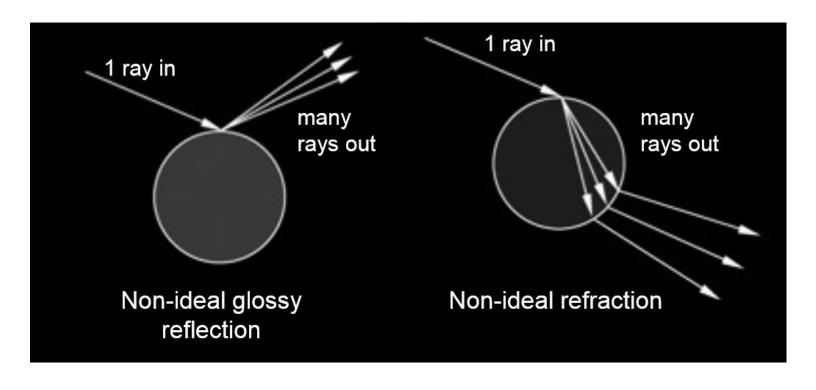


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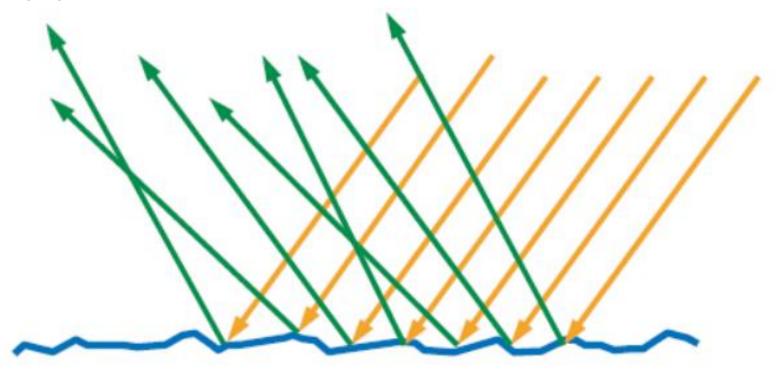
- Case of (optically) flat surface: Snell Descartes laws
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# Micro geometry

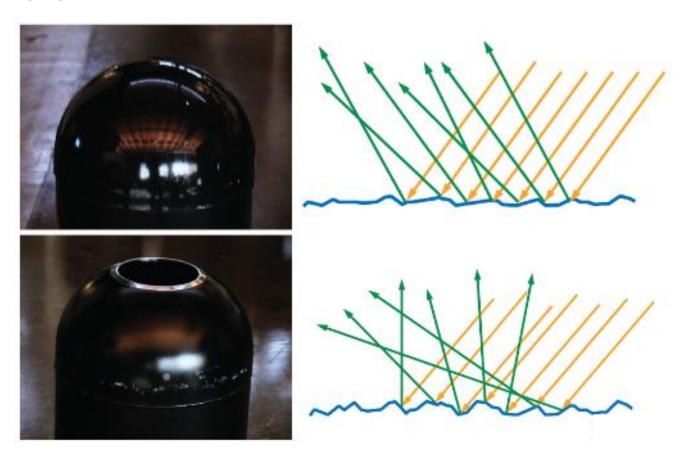
- Microgeometry bumps
  - Bigger than light wavelength
  - But too small to be visible!
  - Agregate of all response





## Micro geometry

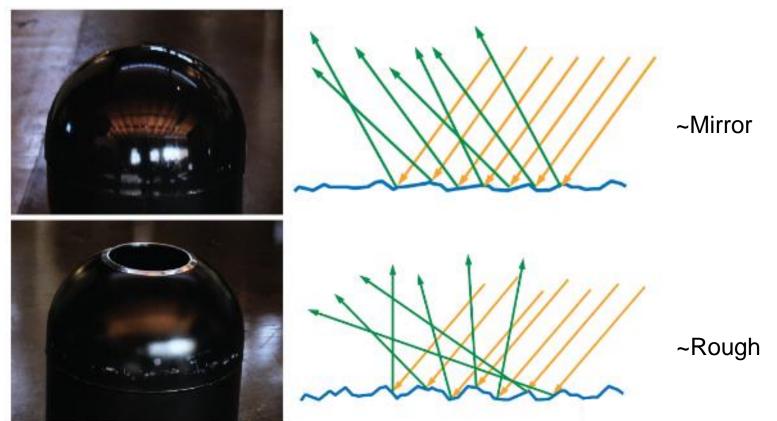
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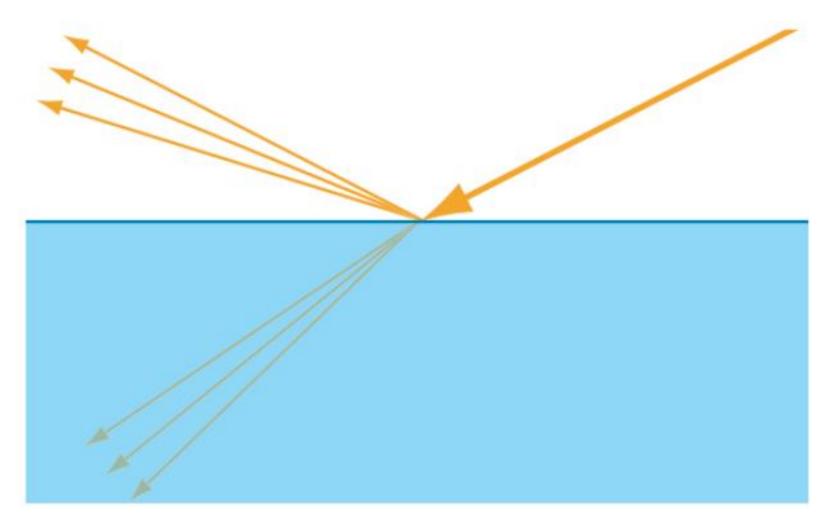
## Micro geometry

- Microgeometry bumps
  - Bigger than light wavelength
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  - Agregate of all response





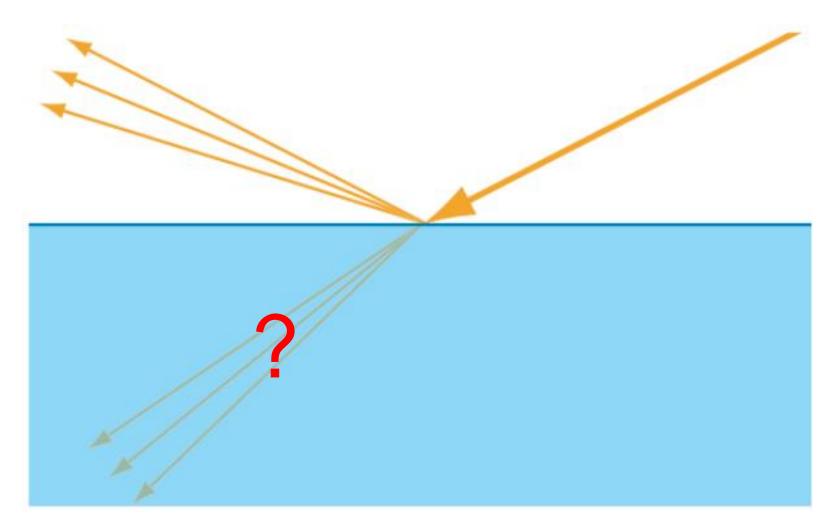
## Macroscopic view





## Macroscopic view

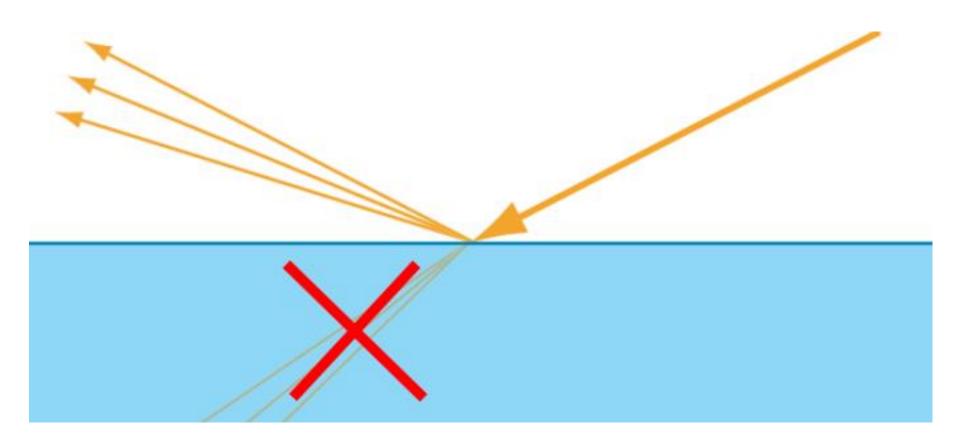
Refractions?





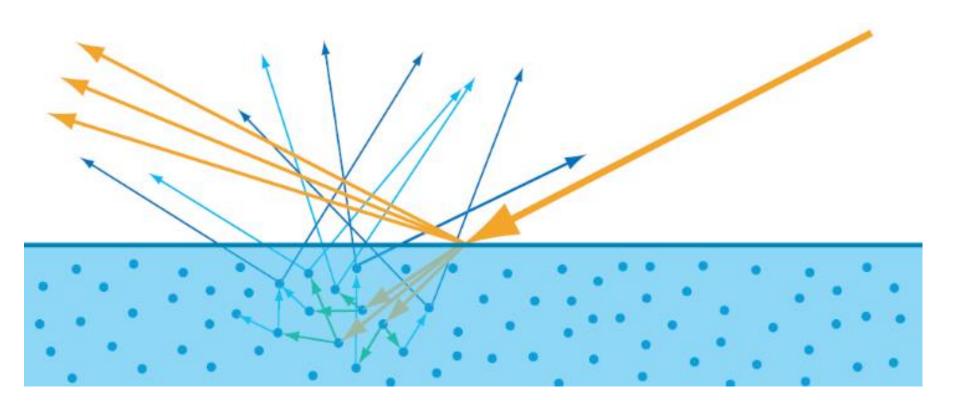
#### Metals

Refracted light immediatly absorbed by free electrons



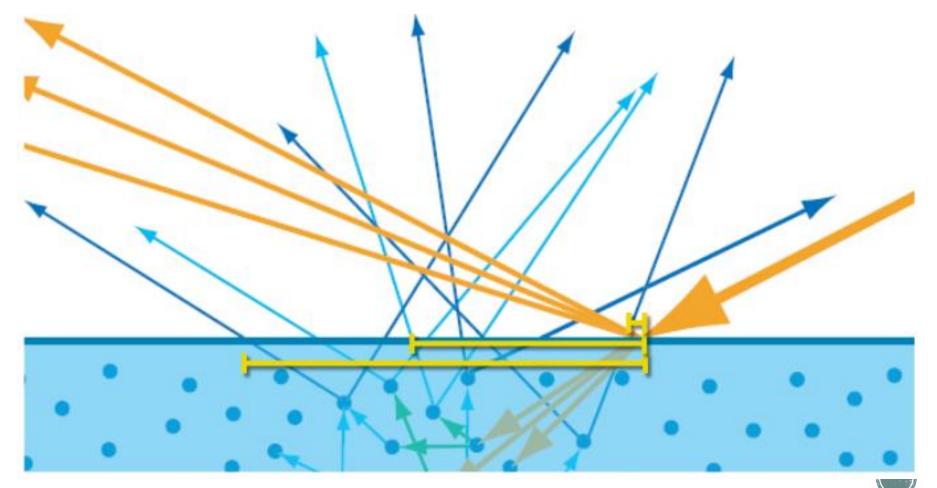


- Behave like regular participating media
  - Light is scattered (enough) are re-emited

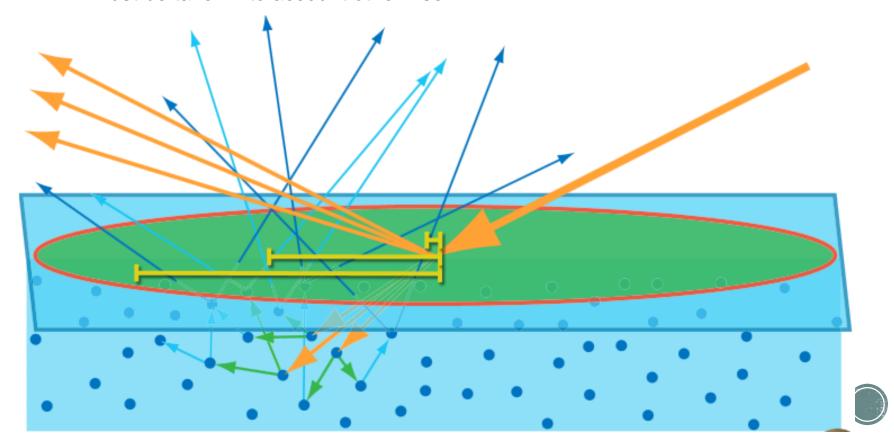




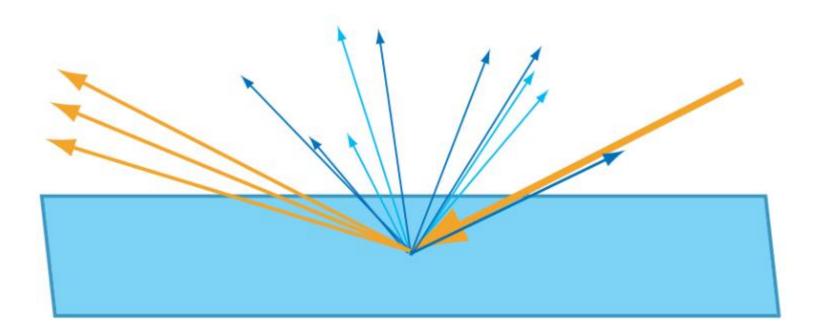
- Behave like regular participating media
  - Light is scattered (enough) are re-emited
  - Distance depends on particle densities



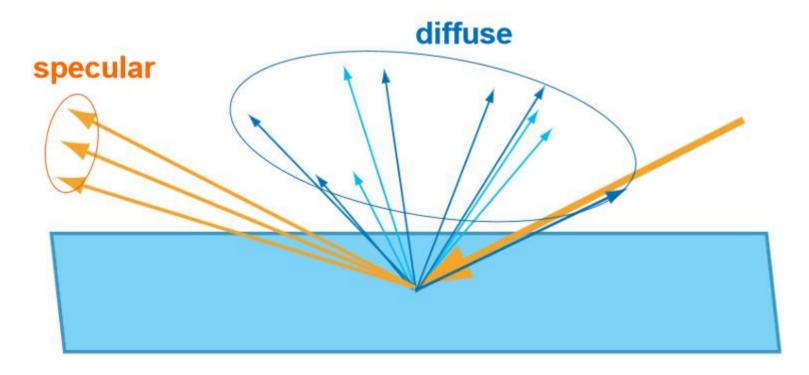
- Behave like regular participating media
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    - Assume = 0 if shading area (surface point) large enough
    - Must be taken into account otherwise



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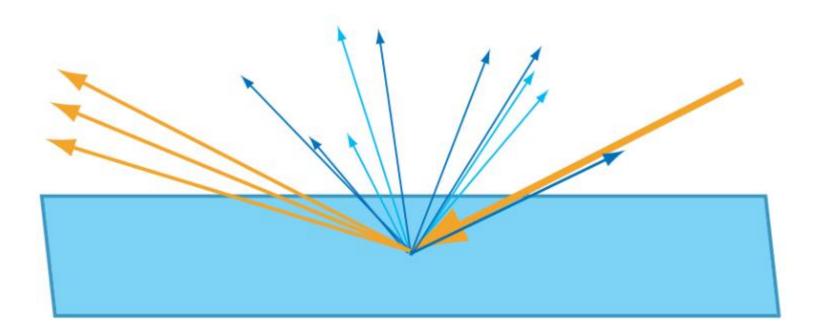


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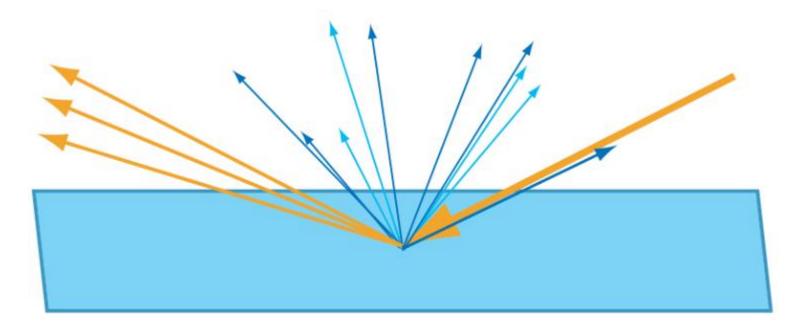


- Radiance
  - radiometric quantity used mesure the amount of light along a single ray
  - Spectral quantity (RGB in practice), Watt per steradian per square meter





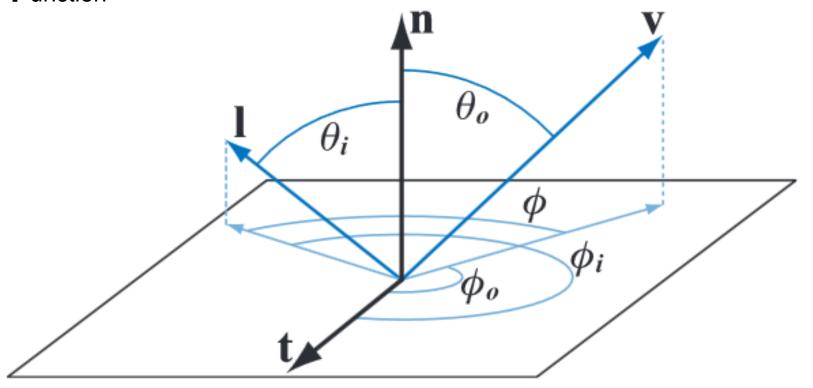
- Radiance
  - radiometric quantity used mesure the amount of light along a single ray
  - Spectral quantity (RGB in practice), Watt per steradian per square meter
- If shading can be handled locally, light response depends on
  - Light direction
  - View direction





- Bidirectionnal
- Reflectance
- Distribution
- Function

$$f(\mathbf{l}, \mathbf{v})$$





- Bidirectionnal
- Reflectance
- Distribution
- Function

$$f(\mathbf{l}, \mathbf{v})$$

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$



- Bidirectionnal
- Reflectance
- Distribution
- Function

Outgoing

radiance

$$f(\mathbf{l}, \mathbf{v})$$

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

Reflectance equation



- Bidirectionnal
- Reflectance
- Distribution
- Function

$$f(\mathbf{l},\mathbf{v})$$

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

Outgoing Ingoing radiance Ingoing radiance

Reflectance equation



- Bidirectionnal
- Reflectance
- Distribution
- Function

$$f(\mathbf{l}, \mathbf{v})$$

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

Outgoing radiance Surface orientation

Reflectance equation

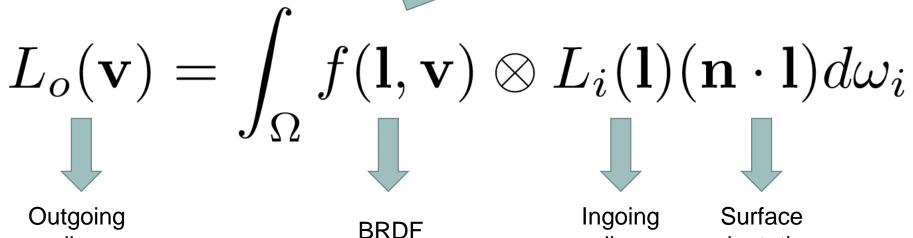


- Bidirectionnal
- Reflectance
- Distribution
- Function

radiance

$$f(\mathbf{l}, \mathbf{v})$$





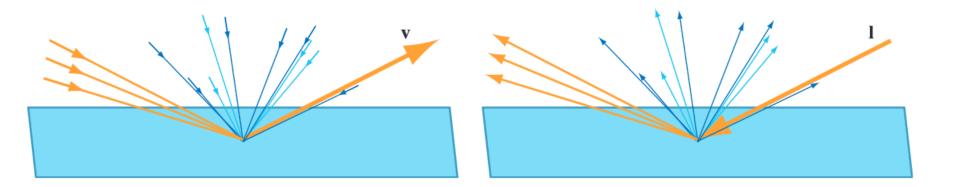
Reflectance equation

radiance

orientation

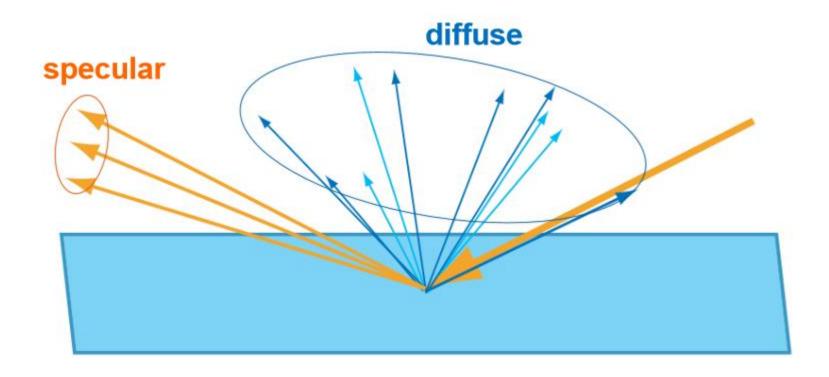


- 2 possible interpretations
  - Given outgoing view → relative contributions of incoming light
  - Given incoming light direction → distribution of outgoing light



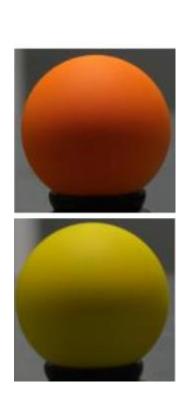


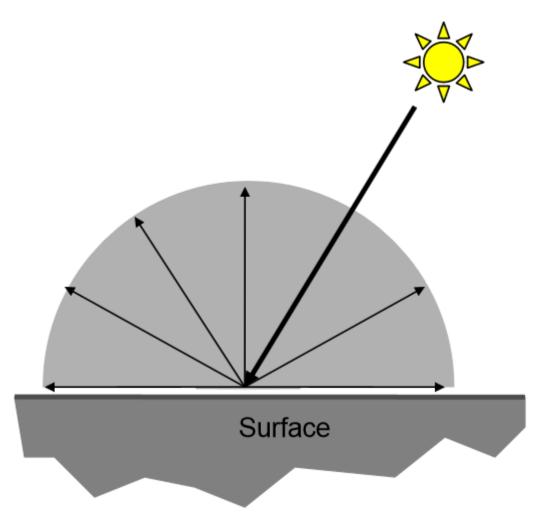
- Phenomena handled separatly:
  - Diffuse term
  - Specular term





- Ideal diffuse reflectance (matte materials)
  - Assume surface reflects equally in all directions



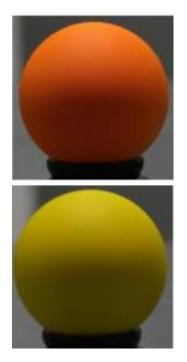


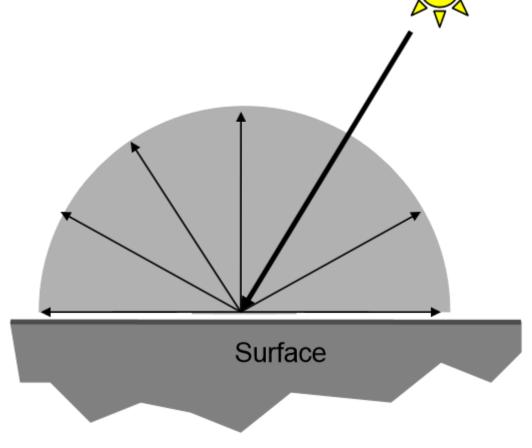


- Ideal diffuse reflectance (matte materials)
  - Assume surface reflects equally in all directions

Coefficient between 0 and 1 that says what fraction is reflected

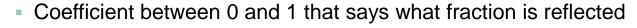
Usually called diffuse color







- Ideal diffuse reflectance (matte materials)
  - Assume surface reflects equally in all directions

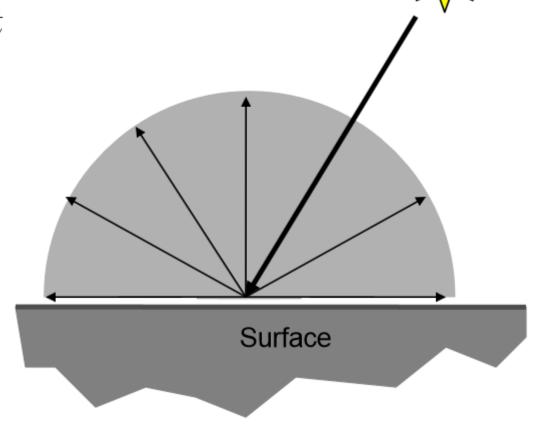


Usually called diffuse color

$$f(\mathbf{l}, \mathbf{v}) = const$$



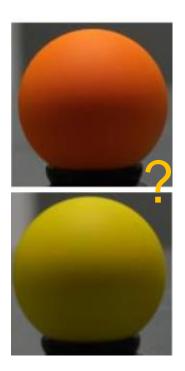






- Ideal diffuse reflectance (matte materials)
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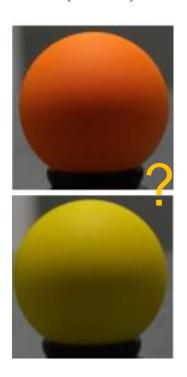


Why does color change?



- Ideal diffuse reflectance (matte materials)
  - Assume surface reflects equally in all directions
  - Coefficient between 0 and 1 that says what fraction is reflected
  - Usually called diffuse color

$$f(\mathbf{l}, \mathbf{v}) = const$$



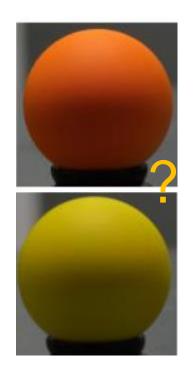
Why does color change?

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$



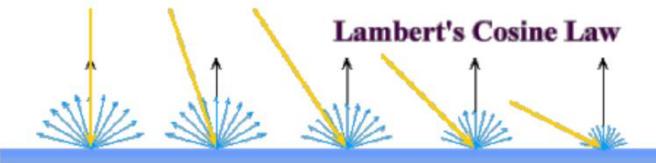
- Ideal diffuse reflectance (matte materials)
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  - Usually called diffuse color

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#### Why does color change?

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

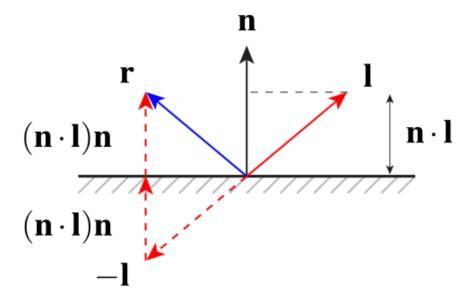




- Ideal specular reflectance (mirror materials)
  - Delta dirac in the reflected direction
  - Not usefull for point lights... better for reflections of other surfaces

$$f(\mathbf{l}, \mathbf{v}) = dirac$$

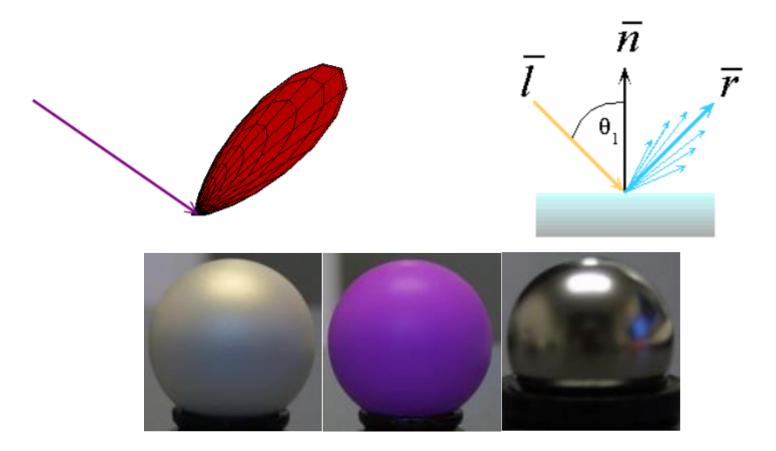




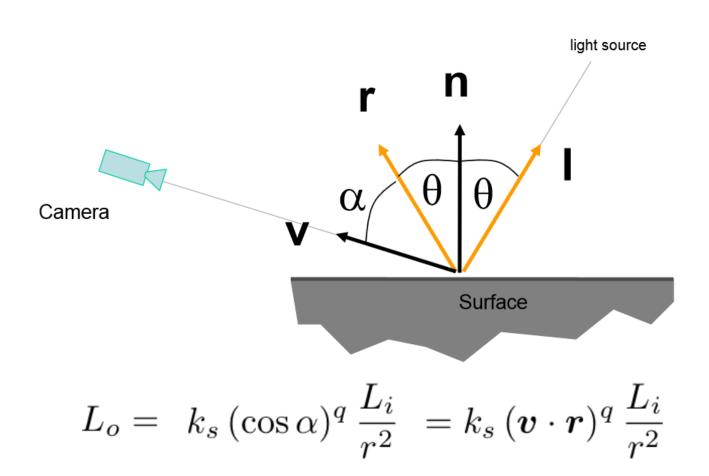
$$\mathbf{r} = 2(\mathbf{n} \cdot \mathbf{l})\mathbf{n} - \mathbf{l}$$



- Non ideal reflectors (glossy material)
  - Expect most of reflected light to travel in the direction of the ideal mirror ray
  - Some of the light should also be reflected slightly offsetted from the mirror ray
  - As we move farther and farther from the mirror ray, we expect to see less light reflected

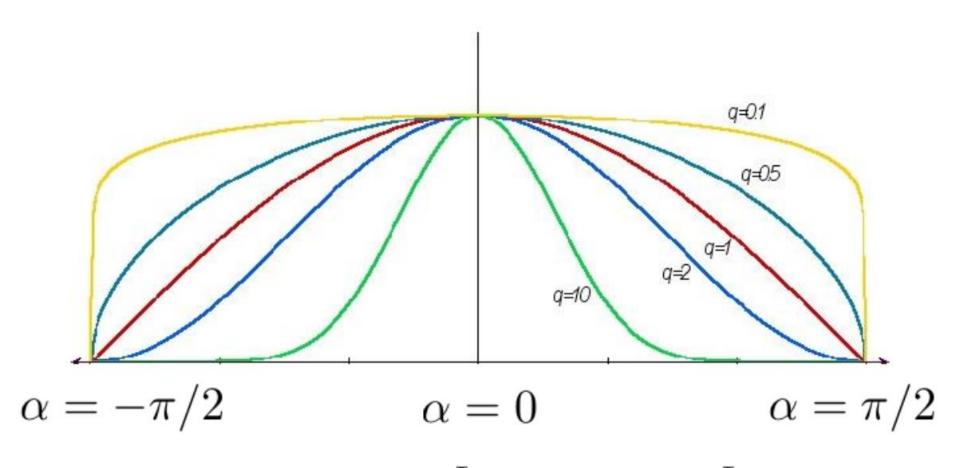


 Reflection depends on the angle between the ideal reflection and the view vectors





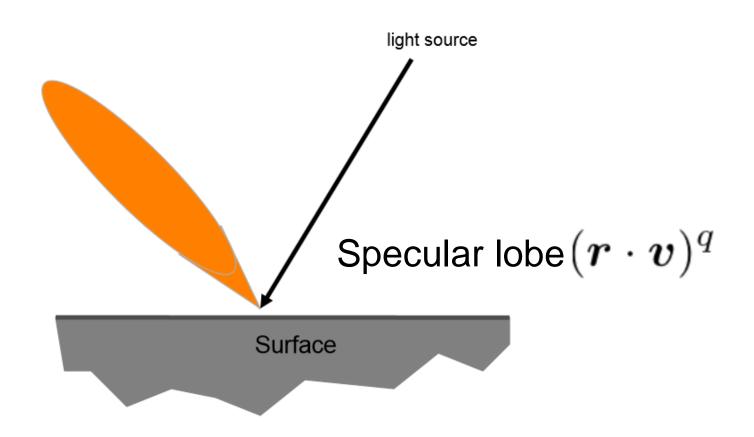
 Reflection depends on the angle between the ideal reflection and the view vectors



$$L_o = k_s (\cos \alpha)^q \frac{L_i}{r^2} = k_s (\boldsymbol{v} \cdot \boldsymbol{r})^q \frac{L_i}{r^2}$$



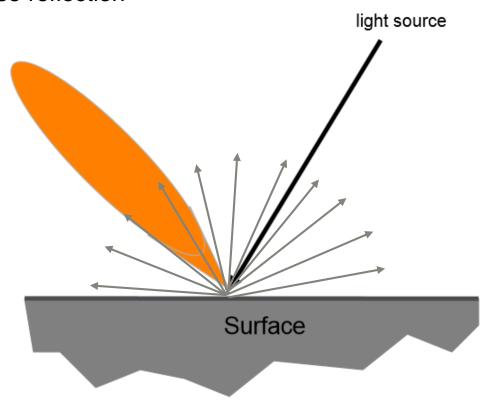
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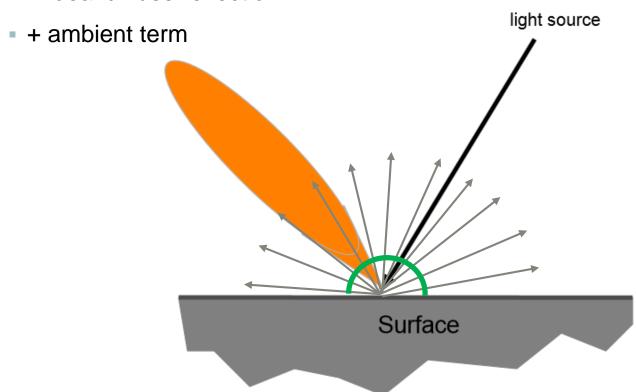
- Reflection depends on the angle between the ideal reflection and the view vectors
- + ideal diffuse reflection



$$L_o = \left[ k_d \left( \boldsymbol{n} \cdot \boldsymbol{l} \right) + k_s \left( \boldsymbol{v} \cdot \boldsymbol{r} \right)^q \right] \frac{L_i}{r^2}$$



- Reflection depends on the angle between the ideal reflection and the view vectors
- + ideal diffuse reflection



$$L_o = \left[ k_a + k_d \left( \boldsymbol{n} \cdot \boldsymbol{l} \right) + k_s \left( \boldsymbol{v} \cdot \boldsymbol{r} \right)^q \right] \frac{L_i}{r^2}$$

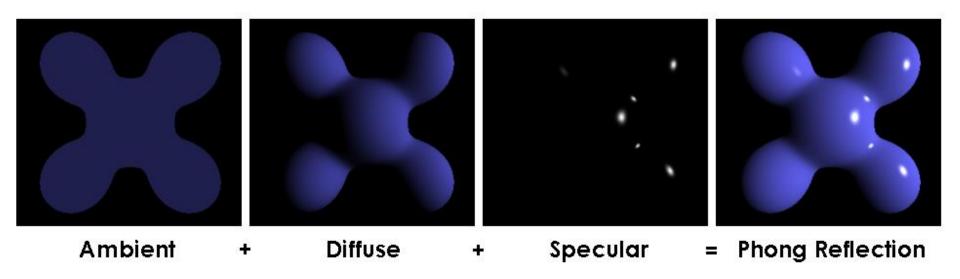


- Reflection depends on the angle between the ideal reflection and the view vectors
- + ideal diffuse reflection
- + ambient term

Phong	Pambient	P <sub>diffuse</sub>	Pspecular	P <sub>total</sub>
$\phi_i = 60^{\circ}$	•			
$\phi_i = 25^{\circ}$	•			
$\phi_i = 0^{\circ}$	•			

$$L_o = \left[ k_a + k_d \left( \boldsymbol{n} \cdot \boldsymbol{l} \right) + k_s \left( \boldsymbol{v} \cdot \boldsymbol{r} \right)^q \right] \frac{L_i}{r^2}$$

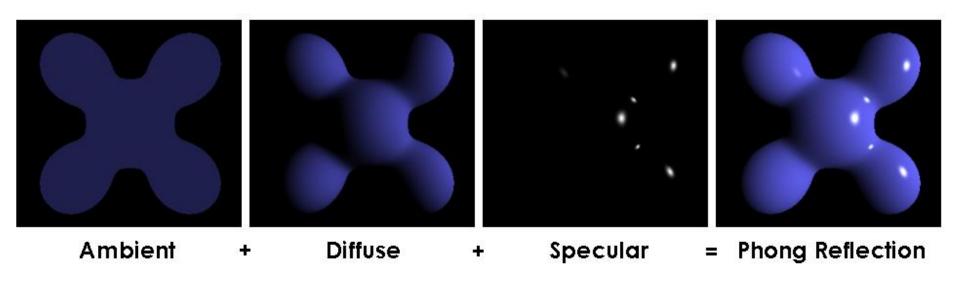
- Reflection depends on the angle between the ideal reflection and the view vectors
- + ideal diffuse reflection
- + ambient term



$$L_o = \left[ k_a + k_d \left( \boldsymbol{n} \cdot \boldsymbol{l} \right) + k_s \left( \boldsymbol{v} \cdot \boldsymbol{r} \right)^q \right] \frac{L_i}{r^2}$$



- Problems:
  - Does not conserve energy (may reflect more than it receives)
  - Not conform to BRDF model (cosine)
  - Ambient is a total hack



$$L_o = \left[ k_a + k_d \left( \boldsymbol{n} \cdot \boldsymbol{l} \right) + k_s \left( \boldsymbol{v} \cdot \boldsymbol{r} \right)^q \right] \frac{L_i}{r^2}$$



$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$



$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

• Positivity  $f(\mathbf{l}, \mathbf{v}) >= 0$ 



$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

• Positivity  $f(\mathbf{l}, \mathbf{v}) >= 0$ 

- Reciprocity  $f(\mathbf{l},\mathbf{v})=f(\mathbf{v},\mathbf{l})$ 

$$L_o(\mathbf{v}) = \int_{\Omega} f(\mathbf{l}, \mathbf{v}) \otimes L_i(\mathbf{l}) (\mathbf{n} \cdot \mathbf{l}) d\omega_i$$

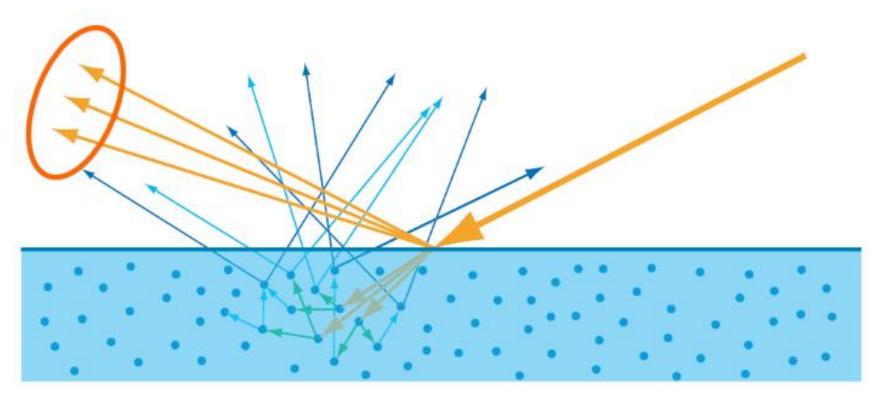
• Positivity  $f(\mathbf{l}, \mathbf{v}) >= 0$ 

- Reciprocity  $f(\mathbf{l},\mathbf{v})=f(\mathbf{v},\mathbf{l})$ 

Energy conservation  $\forall \mathbf{l}, \int_{\Omega} f(\mathbf{l}, \mathbf{v}) (\mathbf{n} \cdot \mathbf{v}) \, d\omega_o \leq 1$ 

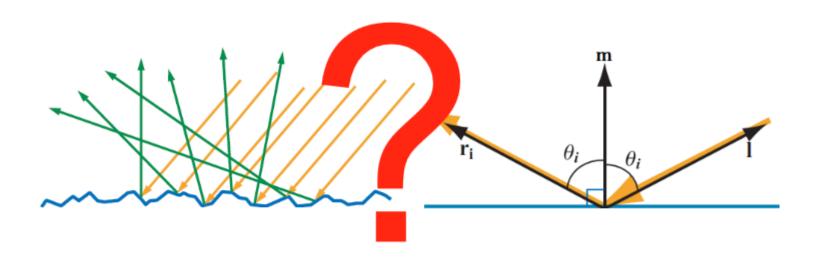


Surface reflection (specular term)



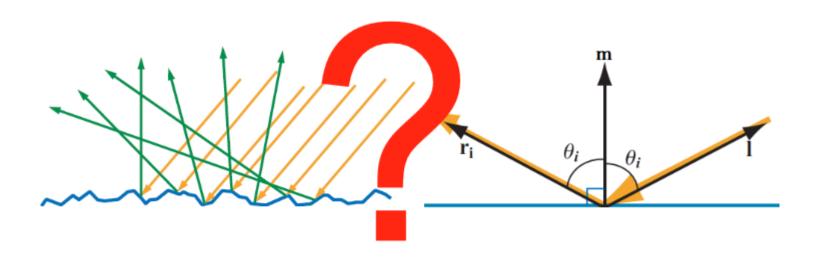


- Derive BRDF from non optically flat surfaces
  - Details too small to be visible
  - But large compared to light wavelength



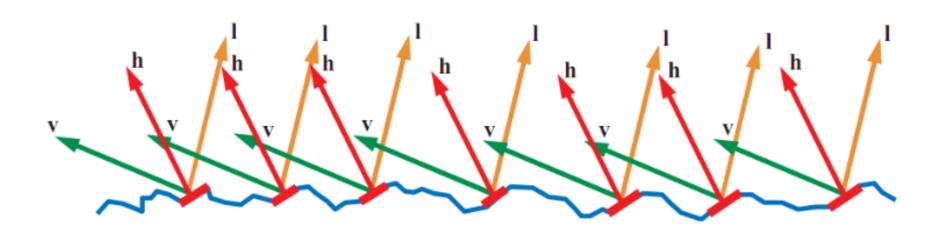


- Derive BRDF from non optically flat surfaces
  - Details too small to be visible
  - But large compared to light wavelength
- Each facet considered as a perfect mirror
  - Reflection depends on light direction and microfacet normal



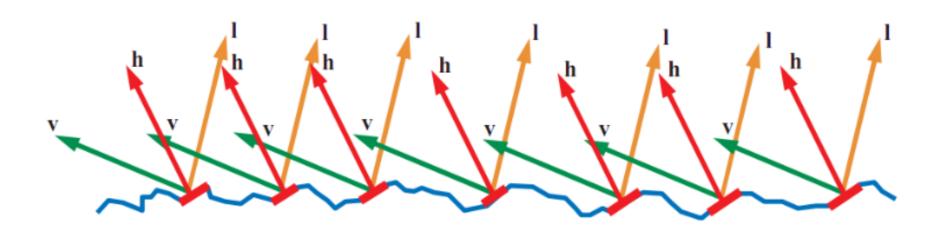


- Half vector
  - →microfact normal



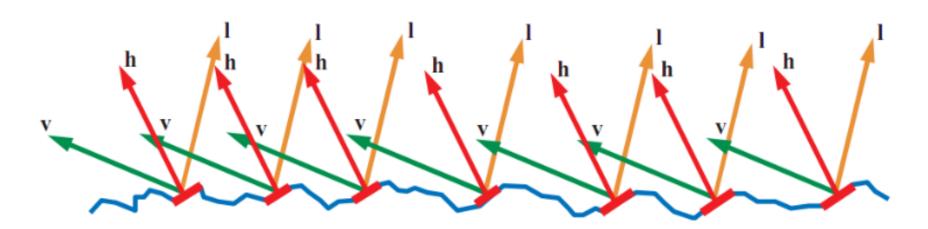


- Half vector
  - microfact normal
  - →Only microfacets having their normals halfway between the view and light direction will reflect something!



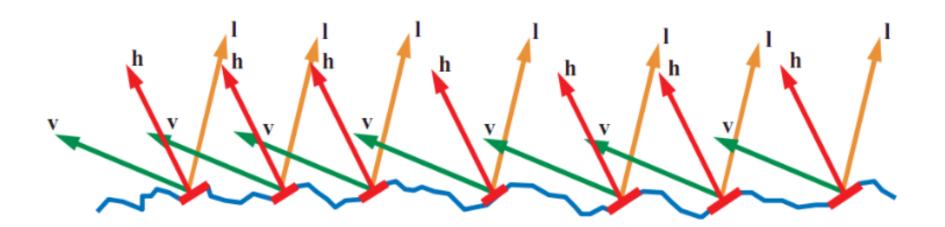


- Half vector
  - → microfact normal
  - →Only microfacets having their normals halfway between the view and light direction will reflect something!
  - →Parametrized by h: give me the percent number of facets having this orientation



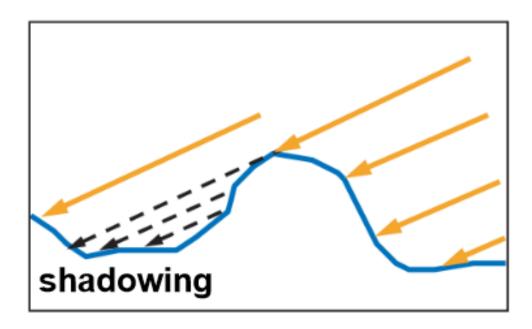


- Shadowing and masking
  - Not all microfacets oriented by a given h will contribute...
  - Some will be blocked by other microfacets from either



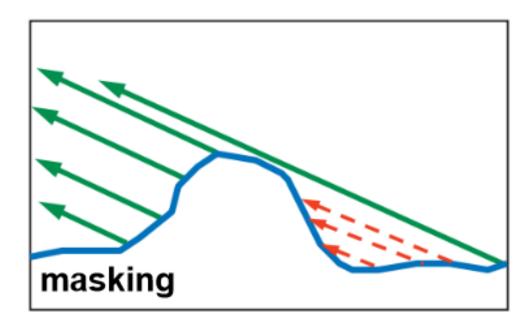


- Shadowing and masking
  - Not all microfacets oriented by a given h will contribute...
  - Some will be blocked by other microfacets from either
    - The light direction (shadowing)



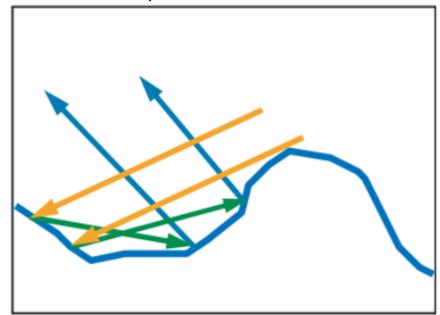


- Shadowing and masking
  - Not all microfacets oriented by a given h will contribute...
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    - The light direction (shadowing)
    - The view direction (masking)



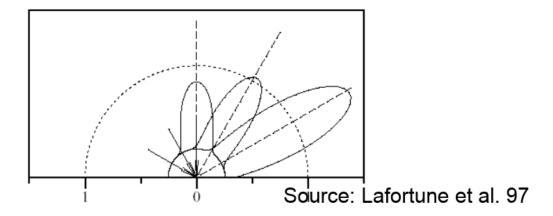


- Shadowing and masking
  - Not all microfacets oriented by a given h will contribute...
  - Some will be blocked by other microfacets from either
    - The light direction (shadowing)
    - The view direction (masking)
  - Not completely true (interreflections)
    - Microfacet limitation...





- Fresnel effect
  - Increase specularity near grazing angles









- Summary
  - Fresnel effect

$$f(\mathbf{l}, \mathbf{v}) = F(\mathbf{l}, \mathbf{h})$$

- Summary
  - Fresnel effect
  - Masking shadowing

$$f(\mathbf{l}, \mathbf{v}) = F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})$$

- Summary
  - Fresnel effect
  - Masking shadowing
  - Amount of microfacets at a particular orientation

$$f(\mathbf{l}, \mathbf{v}) = F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})$$



- Summary
  - Fresnel effect
  - Masking shadowing
  - Amount of microfacets at a particular orientation

$$f(\mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})}{4(\mathbf{n} \cdot \mathbf{l})(\mathbf{n} \cdot \mathbf{v})}$$

correction factor for quantities being transformed between the microgeometry local space and the overall macrosurface



- Summary (cook-terrance model)
  - Fresnel effect
  - Masking shadowing
  - Amount of microfacets at a particular orientation

$$f(\mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})}{4(\mathbf{n} \cdot \mathbf{l})(\mathbf{n} \cdot \mathbf{v})}$$

correction factor for quantities being transformed between the microgeometry local space and the overall macrosurface

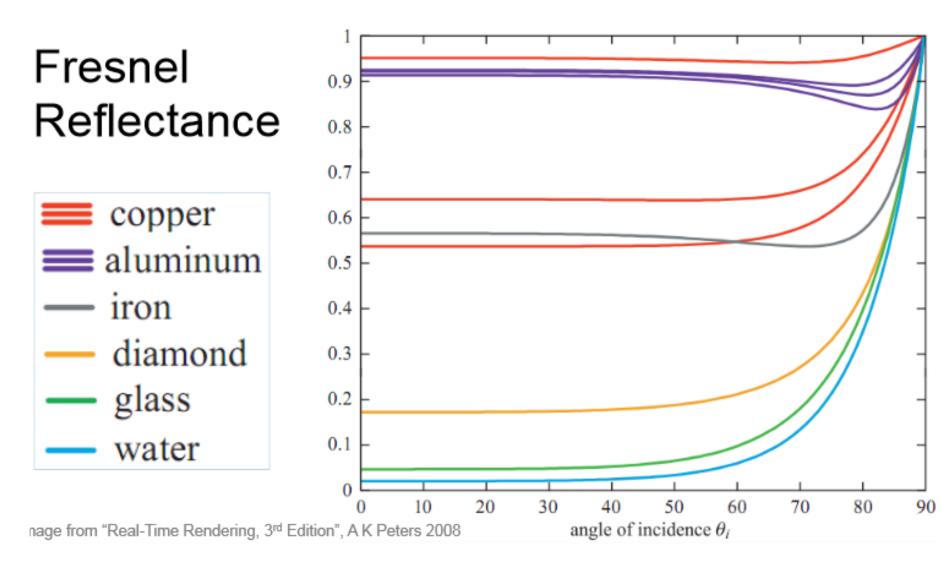


- Fraction of incoming light that is reflected
- In this case:
  - How much of the light hitting the relevant microfacets is reflected

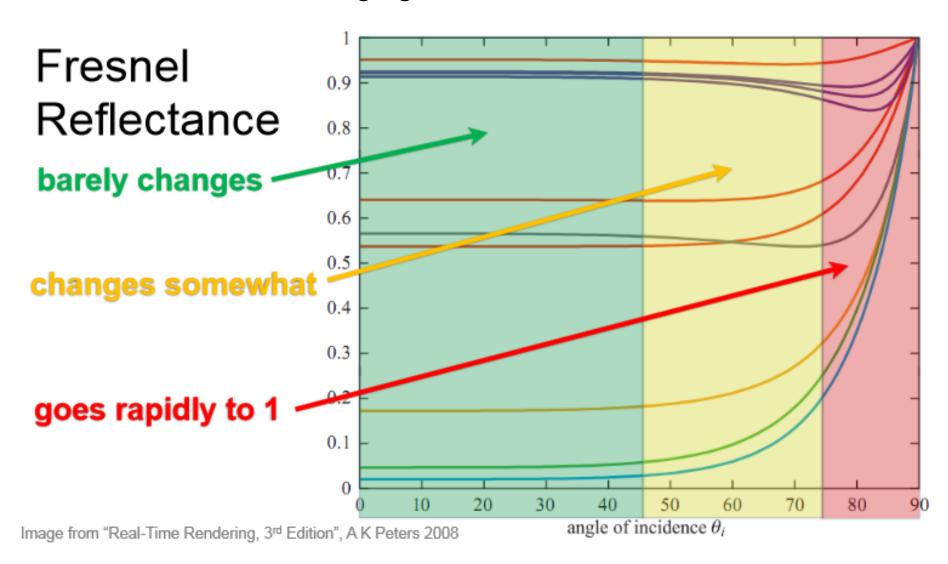
$$f(\mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})}{4(\mathbf{n} \cdot \mathbf{l})(\mathbf{n} \cdot \mathbf{v})}$$



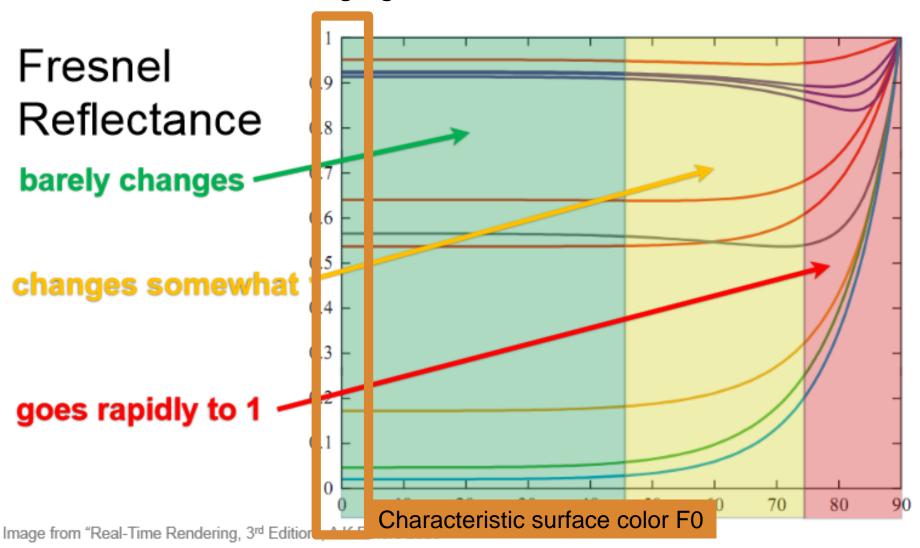
Fraction of incoming light that is reflected



Fraction of incoming light that is reflected



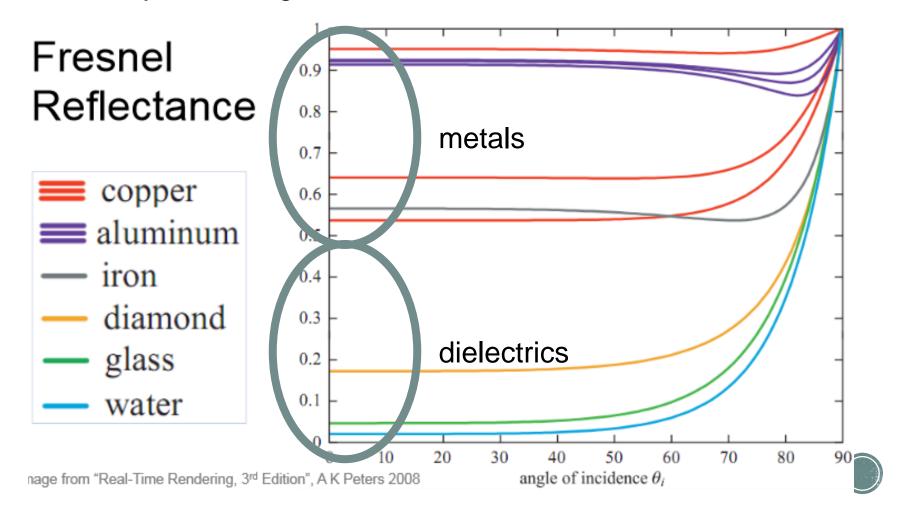
Fraction of incoming light that is reflected



- Fraction of incoming light that is reflected
  - Mainly affect edges



- Fraction of incoming light that is reflected
  - Mainly affect edges



- Fraction of incoming light that is reflected
  - Mainly affect edges
- Schlick approximation
  - Accurate, cheap and parametrized by F0

$$F_{\text{Schlick}}(F_0, \mathbf{l}, \mathbf{n}) = F_0 + (1 - F_0)(1 - (\mathbf{l} \cdot \mathbf{n}))^5$$

For microfacet models

$$F_{\text{Schlick}}(F_0, \mathbf{l}, \mathbf{h}) = F_0 + (1 - F_0)(1 - (\mathbf{l} \cdot \mathbf{h}))^5$$



#### Normal Distribution Function

- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight

$$f(\mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})}{4(\mathbf{n} \cdot \mathbf{l})(\mathbf{n} \cdot \mathbf{v})}$$



- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight

$$D_p(\mathbf{m}) = \frac{\alpha_p + 2}{2\pi} (\mathbf{n} \cdot \mathbf{m})^{\alpha_p}$$

$$D_{uabc}(\mathbf{m}) = \frac{1}{(1 + \alpha_{abc1} (1 - (\mathbf{n} \cdot \mathbf{m})))^{\alpha_{abc2}}}$$

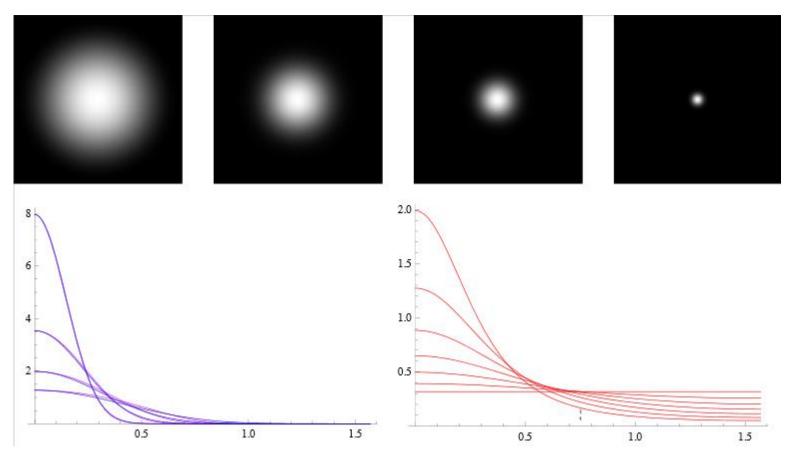
$$D_{tr}(\mathbf{m}) = \frac{\alpha_{tr}^2}{\pi ((\mathbf{n} \cdot \mathbf{m})^2 (\alpha_{tr}^2 - 1) + 1)^2}$$

$$D_b(\mathbf{m}) = \frac{1}{\pi \alpha_b^2 (\mathbf{n} \cdot \mathbf{m})^4} e^{-\left(\frac{1 - (\mathbf{n} \cdot \mathbf{m})^2}{\alpha_b^2 (\mathbf{n} \cdot \mathbf{m})^2}\right)}$$

$$D_{sgd}(\mathbf{m}) = \frac{p22 \left[\frac{1 - (\mathbf{n} \cdot \mathbf{m})^2}{(\mathbf{n} \cdot \mathbf{m})^2}\right]}{\pi (\mathbf{n} \cdot \mathbf{m})^4}$$

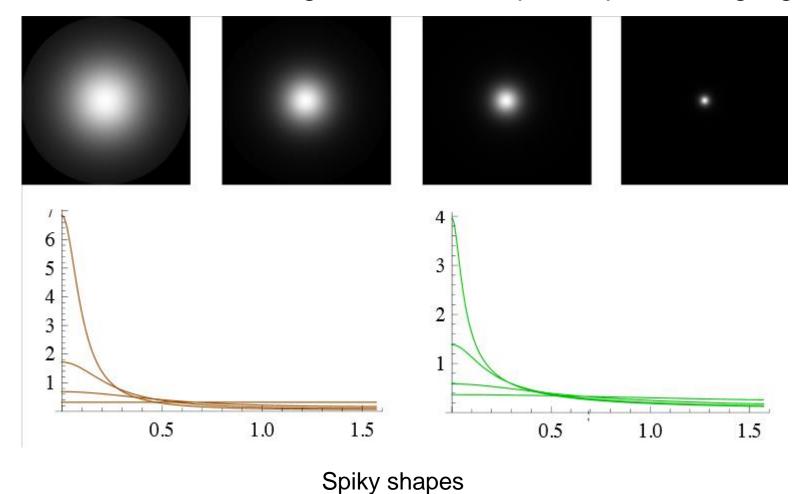


- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight





- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight





- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution

$$D_p(\mathbf{m}) = \frac{\alpha_p + 2}{2\pi} (\mathbf{n} \cdot \mathbf{m})^{\alpha_p}$$



- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution

$$D_p(\mathbf{m}) = \frac{\alpha_p + 2}{2\pi} (\mathbf{n} \cdot \mathbf{m})^{\alpha_p}$$

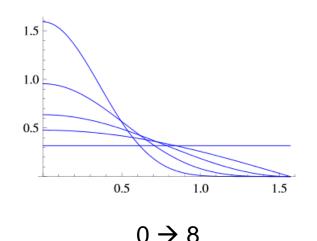


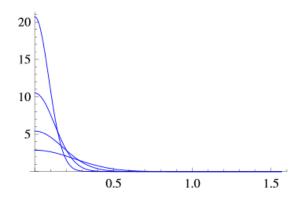
Normalization factor:  $(\mathbf{v} \cdot \mathbf{n}) = \int_{\Theta} D(\mathbf{m})(\mathbf{v} \cdot \mathbf{m}) \, d\omega_m$ 



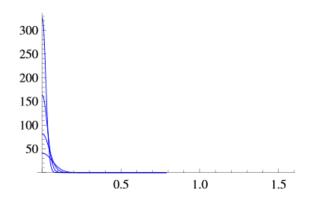
- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution

$$D_p(\mathbf{m}) = \frac{\alpha_p + 2}{2\pi} (\mathbf{n} \cdot \mathbf{m})^{\alpha_p}$$





 $16 \to 128$ 



 $256 \rightarrow 2048$ 

- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution
  - Beckmann distribution

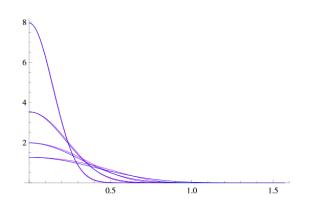
$$D_b(\mathbf{m}) = \frac{1}{\pi \alpha_b^2 (\mathbf{n} \cdot \mathbf{m})^4} e^{-\left(\frac{1 - (\mathbf{n} \cdot \mathbf{m})^2}{\alpha_b^2 (\mathbf{n} \cdot \mathbf{m})^2}\right)}$$

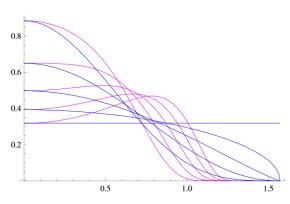


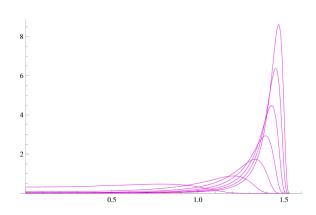
 $256 \rightarrow 2048$ 

- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution
  - Beckmann distribution

$$D_b(\mathbf{m}) = \frac{1}{\pi \alpha_b^2 (\mathbf{n} \cdot \mathbf{m})^4} e^{-\left(\frac{1 - (\mathbf{n} \cdot \mathbf{m})^2}{\alpha_b^2 (\mathbf{n} \cdot \mathbf{m})^2}\right)}$$







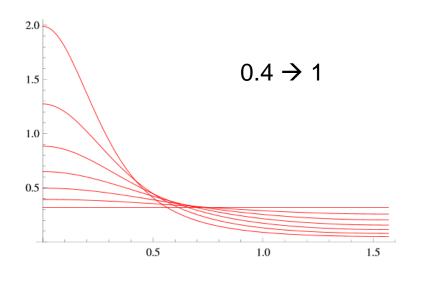
 $\alpha_p = 2\alpha_b^{-2} - 2$ 

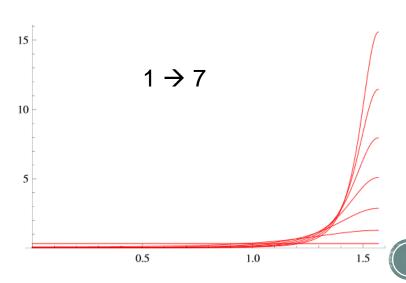
1→ 7
Super rough



- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution
  - Beckmann distribution
  - GGX distribution

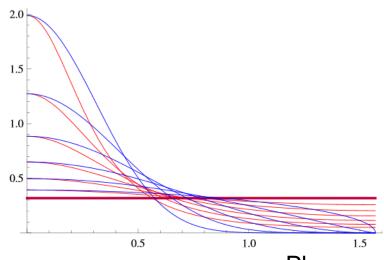
$$D_{\rm tr}(\mathbf{m}) = \frac{\alpha_{\rm tr}^2}{\pi \left( (\mathbf{n} \cdot \mathbf{m})^2 \left( \alpha_{\rm tr}^2 - 1 \right) + 1 \right)^2}$$

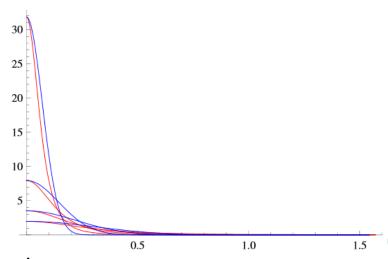




- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution
  - Beckmann distribution
  - GGX distribution

$$D_{\rm tr}(\mathbf{m}) = \frac{\alpha_{\rm tr}^2}{\pi \left( (\mathbf{n} \cdot \mathbf{m})^2 \left( \alpha_{\rm tr}^2 - 1 \right) + 1 \right)^2}$$





Phong comparison

- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution
  - Beckmann distribution
  - GGX distribution
  - And many others...



- Statistical distribution of orientation h
  - Determine size, brightness and shape of specular highlight
- Commonly used NDFs
  - Phong distribution
  - Beckmann distribution
  - GGX distribution
  - And many others...
- Choice of NDF?
  - Depends on evaluation cost (applications)
  - Material properties (rough, isotropic, etc)
  - Artistic controls



- Shadowing and masking
- Probability that points with given microfacet normal
  - is visible from light
  - And from view

$$f(\mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})}{4(\mathbf{n} \cdot \mathbf{l})(\mathbf{n} \cdot \mathbf{v})}$$



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  - No visibility  $G_{\text{implicit}}(\mathbf{l}, \mathbf{v}, \mathbf{m}) = (\mathbf{n} \cdot \mathbf{l_c})(\mathbf{n} \cdot \mathbf{v})$

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  - Smith  $G(\mathbf{l}, \mathbf{v}, \mathbf{h}) = G_1(\mathbf{l})G_1(\mathbf{v})$  depends on NDF



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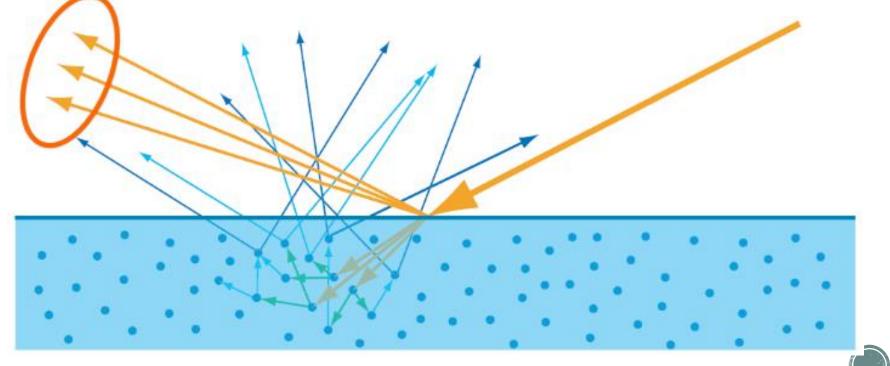
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- More about the masking shadowing function:
  - Understanding the Masking-Shadowing Function in Microfacet-Based BRDFs [Heitz - JCGT 2014]



#### Microfacet theory

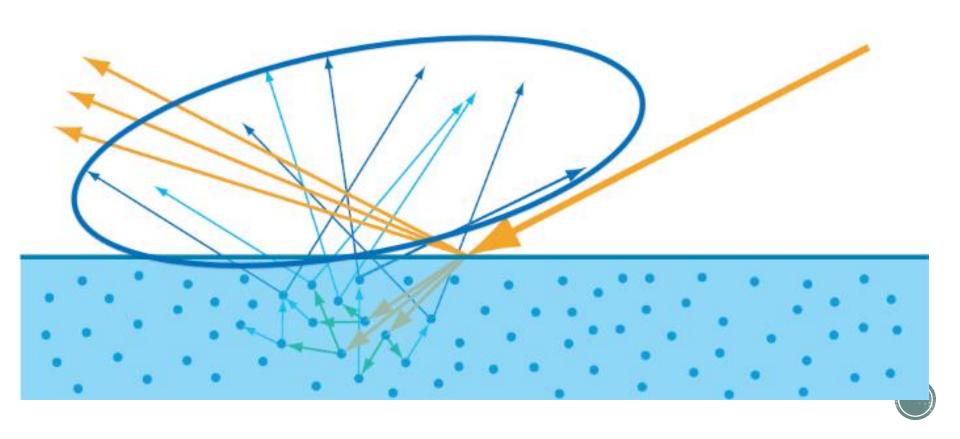
Surface reflection (specular term)

$$f(\mathbf{l}, \mathbf{v}) = \frac{F(\mathbf{l}, \mathbf{h})G(\mathbf{l}, \mathbf{v}, \mathbf{h})D(\mathbf{h})}{4(\mathbf{n} \cdot \mathbf{l})(\mathbf{n} \cdot \mathbf{v})}$$



## Microfacet theory

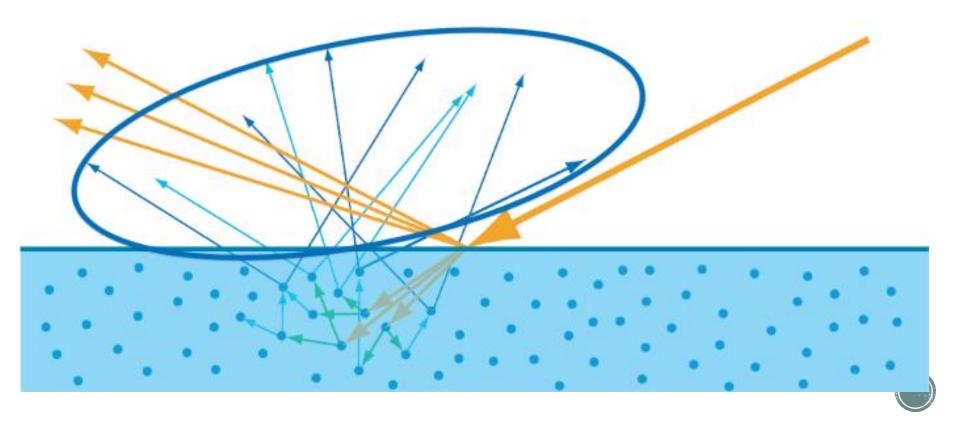
Subsurface reflection (diffuse term)

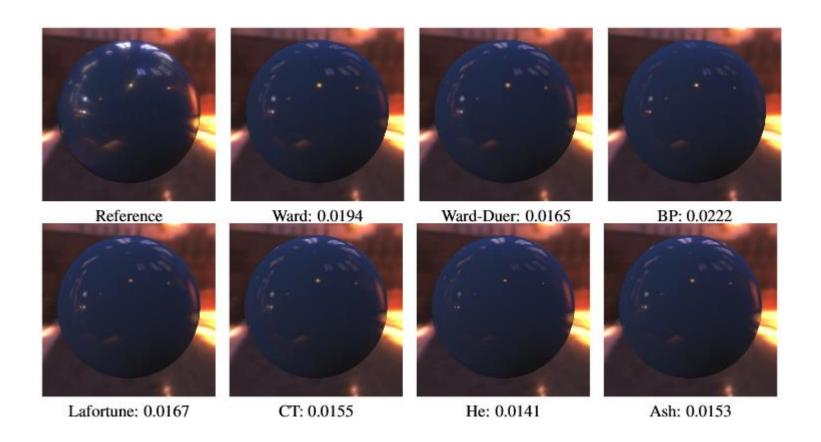


### Microfacet theory

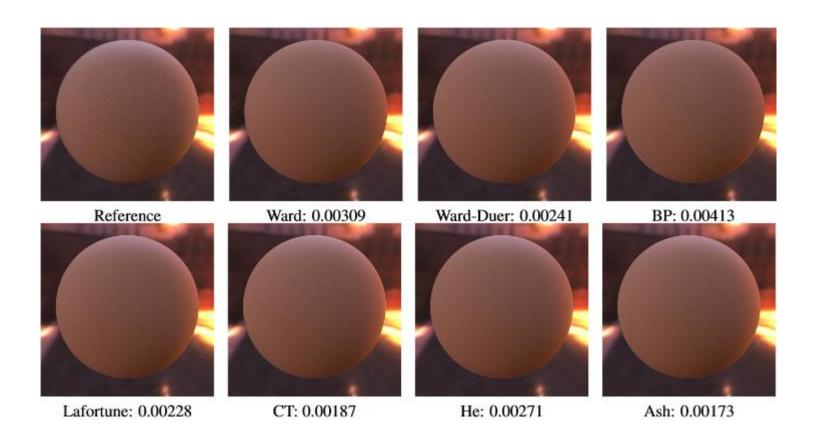
Subsurface reflection (diffuse term)

- Constant: 
$$f_{\mathrm{Lambert}}(\mathbf{l},\mathbf{v}) = rac{\mathbf{c}_{\mathrm{diff}}}{\pi}$$

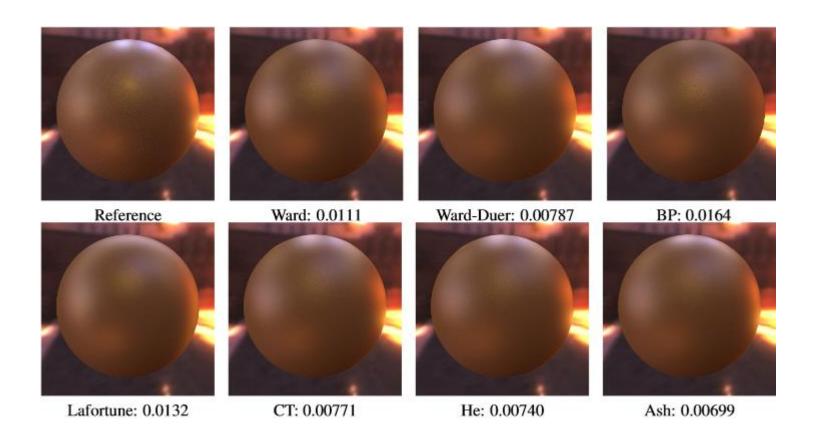




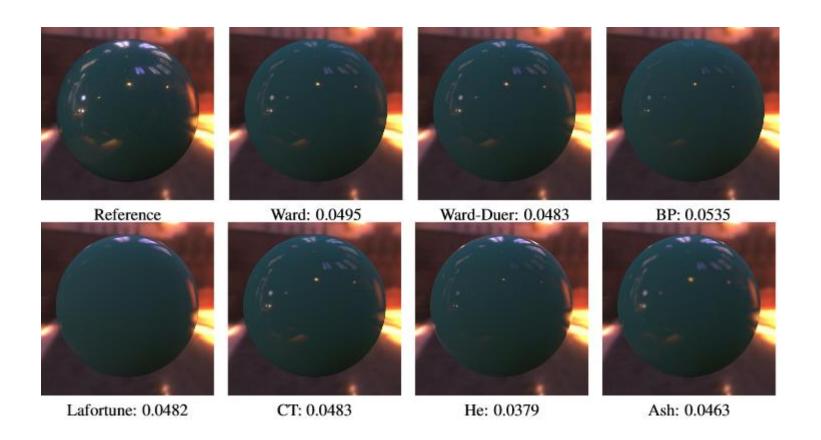




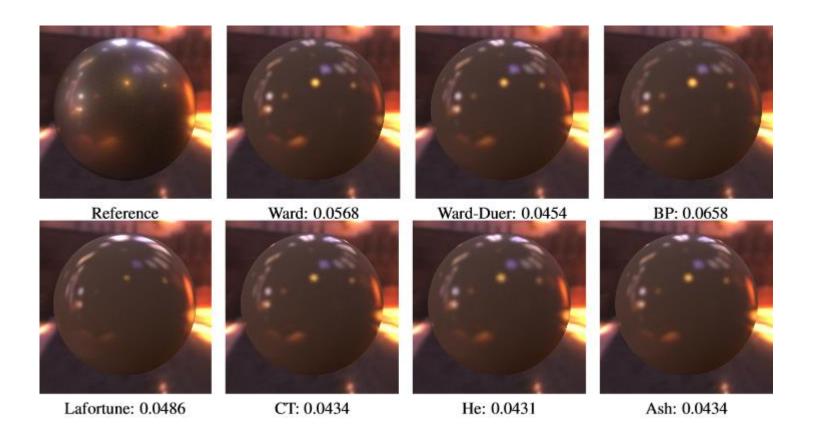














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- BSDF: Bidirectionnal Scattering Distribution Function
  - XD: General formulation  $f(\mathbf{l},\mathbf{x_l},\lambda_\mathbf{l},\mathbf{v},\mathbf{x_v},\lambda_\mathbf{v})$



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