BT5420: COMPUTER SIMULATIONS OF BIOMOLECULAR SYSTEMS. TAKE HOME EXAM: Submitted by Sahana G (BE178038)

I understand that it is a take home exam, not an assignment, and I'm aware that disciplinary action would be taken if I'm found to be involved in any form of malpractice

Question 1.

The three site TIP3P model (in which point charges are centred on each of the three atoms) is commonly used in AMBER simulations. TIP3P reproduces the key features of bulk water at 25°C, 1 atm (ie, dursity of 0.9979/cm3 & AHVap=10.53kcal/mol) but it underestimates the height of the second (tetrahedral) peak in the 0-0 radial distribution function of and overestimates the diffusion constant.

Both the four site TIPAP and five site TIPSP models gives better fit for experimental 0-0 radial distribution functions and is also more accurate than the TIPSP model in different conditions. TIPYP has a dumny atom near O along the bisector of H-O-H angle, with a negative charge. The TIPSP model places regative charges on dummy atoms (representing the lone pair of oxyger atom) and therefore has a tetrahedral-like geometry, thus better suproducing the experimental radial distribution functions from neutron diffraction and the temperature of maximal density of water.

However, as we move from TIPSP to TIPSP to TIPSP try adding additional extra point charges, increases the computational cost of higher versions. TIPAP is NIStimes and TIPSP is ~ 2.5 times as complex (computationally) & expensive as TIP3P is terms of simulation time

Variants of there water models:

- > TIP3P-PME and TIP3P-PME/LRC are generated using by changing a few parameters in the original TIP3P model. They show improved performance is suproducing bulk properties of liquid water (dielectric constant, diffusion coefficient, etc.) relative to TIP3P using short non-bonded cutoffs.
- > TIP4P- Ew model is similar to TIP4P, but used with Ewald Summation method.
- > TIPAP/Ia is used for simulation of solid water ice.
- > TIP4P/2005 is obtained by a general parameterization for simulating the entire phase diagram of condensed water.
- > TIPSP-E model is a reparameterization of TIPSP for use of Ewalds Surmation Method

The top file is a postable Topology file. There can be different specifications for the * it file The position restraint file (posse itp) contains the entries for all the heavy atoms Chence posing no restraints on the protons in the system) from the original PDB file and specifies the force field applied to it.

While specifying topology, the xitp file only lists the information of one particular molecule and allows you to resure the [moleculatype] directure in multiple systems with sursking pdb2gmx command/manually copy-pasting. Unlike *.tpr files, atom10s and residue tos in X. Hp files are indexed from I. Therefore a TPR file created from an ITP file will have different numbring during MD analysis, than while using ITP file. The wea. itp file consists of the following directives - molecule types, atoms, bonds,

diludrals, Examples [mdecularype] ; name nrexul residue number mesidue [atoms. amber C type 0.880229 12.01000 URE amber o -0.61335916.00000 0 URE 14.01000 amber -6.923545 NI URE 1.00800 amber 4 0.395055 H11 URE type 0.395055 1.06800 URE H12 amber H type 6 -0.923545 URE 14.01000 N2 amber N type URE N 21 0.395055 1.00800 amber H type N22 1.00800 URE 0.395055 amber H type [bonds] - meaning that atom 1 is connected to atoms 2,3,5. dihedrals al function definition -> details about all possible dihedral angles will be present

Difference with * top file-An *. itp file corresponds to a single molecule type, while a * top describes an entire System. Hence we would expect details for all the ruidues/small molecules with the * top file. If we have +: it file present for individual ousidues/molecules, we should use # include statement in the * top file. This allows for a modular supresentation of stesidues in the 4. top file #include "wea.itp"

Question 3

and Ux 181-821 mas: mi, m2. positions: 81, 82. velocities: 71,82 Imagnitude since Vis a scalar (1) Lagrangian of the system: quantity)

$$\mathcal{L}(r_1, r_2, \dot{r_1}, \dot{r_2}) = KE(\dot{r_1}, \dot{r_2}) - FE(\dot{r_1}, \dot{r_2})$$

$$\mathcal{L} = \frac{1}{2} m_1 \dot{s_1}^2 + \frac{1}{2} m_2 \dot{r_2}^2 - V(|\dot{r_1} - \dot{r_2}|) - O$$

(ii)
$$R = m_1 v_1 + m_2 v_2$$
, $M = m_1 + m_2$, $v = v_1 - v_2$, $\mu = \frac{m_1 m_2}{M}$

$$M \longrightarrow m_1 v_1 + m_2 v_2 = MR - 2$$

multiply both sides with $m_2 \Rightarrow m_2 v_1 - m_2 v_2 - m_2 v_3$

$$L = m_1 v_1 + m_2 v_2 = MR - 2$$

> Adding @ and 3 => (m1+m2) r1 = m2x+MR

$$\Rightarrow s_1 = \frac{m_2 r + RM}{M} - \Phi$$

> Substituting (1) in (3) => 82=81-8 = mx+RM-mix-mx => 72 = RM-mix

> Substituting @ & B is O, we get:

Lestituting (4) 8 (5) 0, we get
$$\mathcal{L} = \lim_{M \to \infty} \left[\frac{m_2 \dot{r}}{M} \right]^2 + \lim_{M \to \infty} \left[\frac{\dot{R}M}{M} \right]^2 + \lim_{M \to \infty} \left[\frac{\dot{R}M}{M} \right]^2 + \lim_{M \to \infty} \left[\frac{\dot{R}M}{M} \right]^2 - V(M)$$

$$2 = \frac{1}{2}MR^{2} + \frac{1}{2}\frac{m_{1}m_{2}\dot{v}^{2} - V(N)}{M}$$

dii) The second set of waitables with the Centre of mars town, oreduced mars and orelative motion is best suited for this problem. The cartesian variables have more complicated math and isolving is more computationally extensive. Also with the extensive second set of variables (R, M, µ, x), a potential (diatomic) two-body problem is reduced to a single-body problem. Here, we also decouple the mechanics as centre-of-mass related term and the relative motion term.

Question 4

Proof:
$$\Xi = \sum_{N=0}^{\infty} G^N g(N, V, T) \leftarrow Grand cononical partition function is terms of fugacity and cononical partition function.$$

Let us first durine conomical partition function:

Splitting the Kinetic and potential energy parts suparately, we get:

$$K = \frac{1}{h^{3N}} \int_{\infty}^{\infty} d\rho_{1} \dots \int_{\infty}^{\infty} d\rho_{N} e^{-\beta P_{2M}^{2}} = \frac{1}{\lambda^{3N}} - 2$$

In ideal gas, the particles don't interest and there is no attraction oupelsion. Therefore there is no overall change in their position.

une is no everall change is their formation
$$Z = \frac{1}{N!} \int_{-\infty}^{\infty} dq_1 \cdots \int_{-\infty}^{\infty} dq_N e^{-\frac{B[q]}{KT}}$$
 Here: $V\{q\} = 0. \implies e^{-\frac{B[q]}{KT}}$

Also, $\int dq_i \Rightarrow position of particle from <math>-\infty$ to ∞ . Since We can find particle anywhere within the volume $V \Rightarrow \int dq_i = V \Rightarrow Z = LV^N - 3$

$$\Rightarrow Q = \frac{V^{N}}{N!} \frac{1}{A^{2N}} = \left(\frac{V}{A^{3}}\right)^{N} \frac{1}{N!} - \Phi \rightarrow \text{Canonical partition function}.$$

Substitute @ is O, we get:

Also,
$$y = e^{\mu\beta N} \Rightarrow e^{\mu\nu/\kappa\tau}$$

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Vising
$$\sum_{n=0}^{\infty} \frac{x^n}{n!} = e^x$$
, we get

$$E = e^{\frac{x^n}{A^3}} \longrightarrow \text{Taking In on both sides} \Rightarrow \ln E = \frac{\text{Vehkt}}{A^3} - 6$$

We already know that: $PV = \text{KTIn} \exists \Rightarrow PV = \text{KT} \left[\frac{V}{A^3} e^{\frac{W}{A^3}} \right] \leftarrow 0$

To get the ideal gas equation / 9n the ideal gas kenario:

$$\langle N \rangle = \left[\frac{\partial (PV)}{\partial \mu} \right]_{T,V} = \frac{\text{KT}}{A^3} \left[\frac{\partial \left[\frac{Ve^{\frac{W}{A^3}}}{A^3} \right]}{\partial \mu} \right]_{T,V} \leftarrow 0$$

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$$\langle N \rangle = \left[\frac{\partial (PV)}{\partial \mu} \right]_{T,V} = \frac{\text{KT}}{A^3} \left[\frac{\partial \left[\frac{Ve^{\frac{W}{A^3}}}{A^3} \right]}{\partial \mu} \right]_{T,V} \leftarrow 0$$

$$\Rightarrow \langle N \rangle = \left[\frac{\langle KTV \rangle}{A^3} \cdot \left[\frac{1}{KT} \right] \cdot e^{\frac{W}{KT}} = \frac{Ve^{\frac{W}{KT}}}{A^3} = \ln \left[\frac{1}{A^3} \left\{ \text{Rem eq} (9) \right\} \right]_{T,V} \leftarrow 0$$

Comparing (6) (4), (8) we get: $\frac{PV}{KT} = \ln \left[\frac{1}{A^3} \left\{ \text{Rem eq} (9) \right\} \right]_{T,V} \leftarrow 0$

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Question 5

(i) Site site radial distribution function

This is a measure of the probability of firting a particle at a distance of away from a given reference particle, relative to that for an ideal gas

In a molecular system, tracking only the distance between two molecules does not suffice Two molecules can be at the same distance away from the oreference molecule, but they can be at very different orientations. To avoid this confusion, site-site radial distribution function, is hence used to track the distance between especific atoms (sites). Hence if we are looking at radial distribution function for water molecules, the different site-site interactions that can be tracked are goo (commonly measured), got (correspondence) -ds to the distribution function of all H atoms with O as the central atom) and 9 no Therefore, gas (Tab) is a set of site-site distribution function, can be calculated in the same way as the academic atomic g(r) for each type of site between molecules Han d(2) = 1/2 (\(\frac{1}{2} \) \(\frac{1}{2} \) \(\lambda \) \(\lambda \)

(i) Hydrogen bond life time analysis based on autocorrelation function.

The time autoconvelation function is calculated for all the selected hydrogen bonds that are passed to it. This set of hydrogen bonds, starting from t=to are measured based on geometric virturia and their lifetime is monitored over time

Multiple passes through the same trajectory is recorded to obtain the average Inhariows. Here the subscript & denotes whether we're defining the lifetime of the hydrogenbonds in continuum time or intornitant time. The continuum definition measures the time that a hydrogen bond remains continuously attached, whereas the intermittent definition also includes the same H bond if it gots deaved and re-attached during the measurement. The relevant lifetime can be measured as follows:

$$T_{x} = \int_{0}^{\infty} C_{x}(t)dt'$$

while performing MD simulations, this can be measured using the command:

· MD Analysis. aralysis. Hoonds. Hoond_autocorrel

with input as hydragen bond donors and acceptors

output as -> results (row result of autocorr dation function)

-> turns (time axis is ps)

→ fit - continuous lifetime: (A, A, E1, Te) and informittent liletime: (A, A, E1, Te)

intermittent lifetime: (A1, A2, T1, T2, T3) -> estimate

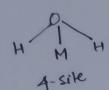
Cx(t) = A1 exp(-t/z) + A2 exp(-t/z) [+ A3 exp(-t/z3)]

Hore A Z Ai = 1 -> An = 1 - Z Ai

Supplementary information for question 1.

Water models:

3-sike



Proporties compared across TIP3P, TIPAP, TIPSIP.

Inoporties compared activity				
		TIPAP	TIPSP	Experimental
Property	TIP3P		0.9572	0.9584
distance (OH) A°	0.9572	0.9572		
	104.52	104.52	104.52	104.450
L HOH, digree		600.0	544,5	-
A, 103 Kcal A012/mol	582.0		590.3	-
B, Kcal A06/mol	595.6	610.0	X:L; a=-0.241	_
charge: 9(R)	X:0, 9=-0.834	X: M, Q=-1.04		_
	+0.417	+0.52	+0.241	
drange: qCH)		২. \৪	2.29	2.95
dipde moment	2.35		81.5	78.4
didectric constant	8 2	53		
self diffusion, 10 cm/s	5.19	3.29	2.62	2.30
Sey affactor, 10	-91	-25	+4	+ 3.984
density max @7, %		4.4	6.3	2.53
expansion coefficient (104°C-1)	9.2	4.4		

Using the Leonard Jones potential:

Eab =
$$\sum_{i} \sum_{j} \frac{1}{k_{i} + k_{i}} \frac{A}{8_{00}} + \frac{A}{8_{00}} = \frac{B}{8_{00}}$$

- -> Kc: dechostatic constant = 332.1 A kcal/mol e2
- 91,9; are the partial charges relative to charge on electron
 - -> mj : distance b/w charged atom
 - -> A,B: LJ parametous