## PHY305A Exercise Set 5

- 1. Determine the cross section for scattering of a billiard ball from another billiard ball. Both have a radius equal to R.
- 2. Assume that a galaxy is moving away from us at a speed of 300 Km/s. Determine the observed Doppler shift of a spectral line with frequency  $\nu_0 = 5 \times 10^{14}$  Hz produced by this galaxy.
- 3. Consider an atomic state that has a lifetime of (i) 1 second and (ii) 1 nanosecond. Determine the width,  $\Delta E$ , of this state. A transition from this state produces a spectral line at visible frequency  $\nu_0 = 5 \times 10^{14}$  Hz. Determine the width of this line caused by natural broadening. Compare this with  $\nu_0$ .
- 4. Using the surface temperature of the Sun, make a rough estimate of the broadening of Balmer H-alpha line at  $\lambda = 6563$  Angstroms.
- 5. Show that the solution to the differential equation

$$\frac{\mathrm{d}\mathrm{I}_{\lambda}}{\mathrm{d}s} = -\rho \kappa_{\lambda} \mathrm{I}_{\lambda}$$

is given by

$$I_{\lambda}(s) = I_{\lambda}(0) \exp \left(-\int_{0}^{s} \kappa_{\lambda} \rho ds'\right)$$

- 6. The visual extinction coefficient  $A_V$  for a star at 1 Kpc is approximately 1.8 magnitude. Determine its blue extinction coefficient  $A_B$ .
- 7. The absolute magnitude  $M_V$  of a star is known to be -3.0. Its apparent visual magnitude V is observed to be 12.
  - (a) Determine its distance by ignoring  $A_V$ .
  - (b) If we include  $A_V$ , the distance cannot be computed directly. We can obtain a rough estimate as follows: Compute  $A_V$  for r=1,2,3,4 Kpc. Insert these values in the equation

 $V = M_V + 5 \log \frac{r}{10 \, \, \mathrm{pc}} + A_V \, ,$ 

and compute the distances. The value that is closest to the input distance provides a rough approximation of the true distance. This procedure can now be refined by computing  $A_V$  values over intervals of 100 pc and inserting in the above equation. Notice that (a) provides a very poor approximation of the distance.