AE339 PROJECT REPORT

GAS TABLE CALCULTOR
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ABSTRACT:

 The aim of this project is to use the concepts of Isentropic Flows, Normal Shocks and Oblique Shocks learnt in class and apply them in order to develop a working Gas Table Calculator.

OBJECTIVES:

- To develop equations to find various parameters with a single given parameter in isentropic flow, normal shock and oblique shock
- To write a code for the above equations and make a working calculator

ISENTROPIC FLOW EQUATIONS:

- Mach Number is denoted as M and specific heat ratio is y
- Mach Angle (μ):

$$\mu = \sin^{-1}\left(rac{1}{M}
ight)$$

• Prandtl-Meyer Angle (ν):

$$u(M) = \sqrt{rac{\gamma+1}{\gamma-1}} an^{-1} \left(\sqrt{rac{\gamma-1}{\gamma+1}} \left(M^2-1
ight)
ight) - an^{-1} \left(\sqrt{M^2-1}
ight)$$

• Static pressure to Stagnation pressure (p/p0):

$$rac{p}{p_0} = \left(1 + rac{\gamma-1}{2}M^2
ight)^{-rac{\gamma}{\gamma-1}}$$

Static density to Stagnation density (ρ/ρ0):

$$rac{
ho}{
ho_0}=\left(1+rac{\gamma-1}{2}M^2
ight)^{-rac{1}{\gamma-1}}$$

• Static temperature to Stagnation temperature (T/T0):

$$rac{T}{T_0}=\left(1+rac{\gamma-1}{2}M^2
ight)^{-1}$$

Pressure Ratio at Critical Mach Number (p/p*):

$$rac{p}{p^*} = \left(rac{rac{\gamma+1}{2}}{1+rac{\gamma-1}{2}M^2}
ight)^{rac{\gamma}{\gamma-1}}$$

Density Ratio at Critical Mach Number (ρ/ρ*):

$$rac{
ho}{
ho^*}=\left(rac{rac{\gamma+1}{2}}{1+rac{\gamma-1}{2}M^2}
ight)^{rac{1}{\gamma-1}}$$

• Temperature Ratio at Critical Mach Number (p/p*):

$$rac{T}{T^*}=\left(rac{1+rac{\gamma-1}{2}M^2}{rac{\gamma+1}{2}}
ight)^{-1}$$

Area Ratio at Critical Mach Number (A/A*):

$$rac{A}{A^*} = rac{1}{M} \left(rac{2}{\gamma+1} \left(1+rac{\gamma-1}{2}M^2
ight)
ight)^{rac{\gamma+1}{2(\gamma-1)}}$$

NORMAL SHOCK EQUATIONS:

- Upstream Mach Number is denoted as M1 and specific heat ratio is y
- Downstream Mach Number (M2):

$$M_2 = \sqrt{rac{(\gamma-1)M_1^2+2}{2\gamma M_1^2-(\gamma-1)}}$$

Total Pressure Ratio (p02/p01):

$$rac{p_{02}}{p_{01}} = \left(rac{(\gamma+1)M^2}{2+(\gamma-1)M^2}
ight)^{rac{\gamma}{\gamma-1}} \left(rac{2\gamma M^2-(\gamma-1)}{\gamma+1}
ight)^{rac{1}{\gamma-1}}$$

• Static Pressure Ratio (p1/p02):

$$rac{p_1}{p_{02}} = rac{\left(1 + rac{2\gamma}{\gamma+1}(M^2 - 1)
ight)^{-1}}{\left(1 + 0.5(\gamma - 1)\left(\sqrt{rac{2 + (\gamma - 1)M^2}{2\gamma M^2 - (\gamma - 1)}}
ight)^2
ight)^{rac{\gamma}{\gamma - 1}}}$$

Static Pressure Ratio (p2/p1):

$$rac{p_2}{p_1}=rac{2\gamma M_1^2}{\gamma+1}-rac{\gamma-1}{\gamma+1}$$

Density Ratio (ρ2/ρ1):

$$rac{
ho_2}{
ho_1} = rac{p_2}{p_1} \cdot rac{T_1}{T_2} = rac{(\gamma+1)M_1^2}{2+(\gamma-1)M_1^2}$$

• Temperature Ratio (T2/T1):

$$rac{T_2}{T_1} = rac{(2\gamma M^2 - (\gamma - 1))(2 + (\gamma - 1)M^2)}{(\gamma + 1)^2 M^2}$$

OBLIQUE SHOCK EQUATIONS:

- Upstream Mach Number is denoted as M1, specific heat ratio is y, Turn Angle is θ and Wave Angle is β .
- Turn Angle (θ) with respect to wave angle (β) :

$$an heta = rac{2\coteta\left(M_1^2\sin^2eta-1
ight)}{M_1^2(\gamma+\cos2eta)+2}$$

Normal Mach Number Upstream (Mn1):

$$M_{n1} = M \cdot \sin(\beta)$$

Downstream Mach Number (M2):

$$M_2 = \sqrt{rac{2 + (\gamma - 1) M_{n1}^2}{2 \gamma M_{n1}^2 - (\gamma - 1)}} \cdot rac{1}{\sin(eta - heta)}$$

Normal Mach Number Downstream (Mn2):

$$M_{n2} = M_2 \cdot \sin(\beta - \theta)$$

Static Pressure Ratio (p2/p1):

$$rac{p_2}{p_1} = 1 + rac{2\gamma}{\gamma+1} \cdot (M_{n1}^2 - 1)$$

Density Ratio (ρ2/ρ1):

$$rac{
ho_2}{
ho_1} = rac{M_{n1}^2(\gamma+1)}{2+(\gamma-1)M_{n1}^2}$$

• Temperature Ratio (T2/T1):

$$rac{T_2}{T_1} = rac{1 + rac{2\gamma}{\gamma + 1}(M_{n1}^2 - 1)}{rac{M_{n1}^2(\gamma + 1)}{2 + (\gamma - 1)M_{n1}^2}}$$

Stagnation Pressure Ratio (p02/p01):

$$rac{p_{02}}{p_{01}} = rac{p_2}{p_1} \cdot \left(rac{1 + rac{\gamma - 1}{2} M_{n2}^2}{1 + rac{\gamma - 1}{2} M_{n1}^2}
ight)^{rac{\gamma}{\gamma - 1}}$$

LIBRARIES USED:

- math library: The math library in Python provides access to a collection of basic mathematical functions and constants.
- In the gas table, we sometimes have different parameters instead of the normal ones and we do not hav direct equations for them.
- Hence, we have to use the function fsolve from scipy.optimize. It is used to calculate the roots of non-linear equations. It uses the hybrid Powell method, which combines Newton-Raphson and other techniques for robust performance.
- We also use the function bisect from scipy.optimize. It uses numerical root-finding method that uses the bisection method.