# Tracing Baryons in the Warm Hot Intergalactic Medium using Broad Lyman- $\alpha$ Absorbers

Mid-Term 2

Sameer Patidar SC19B161

Dual Degree (Astronomy & Astrophysics)
Indian Institute of Space Science and Technology

Supervisors: Dr. Vikram Khaire and Dr. Anand Narayanan





### Thesis Phase I: Recap

- ▶ The missing baryon problem
- ▶ BLAs : Way to probe WHIM
- ▶ Absorber towards PG 0003+158
- ▶ BLA survey : 28 BLA candidates

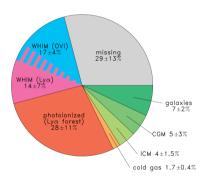


Figure 1: Baryon budget at  $z \sim 0$ . Shull et al. (2012)

- ▶ The missing baryon problem
- **▶** BLAs : Way to probe WHIM
- ▶ Absorber towards PG 0003+158
- ▶ BLA survey : 28 BLA candidates

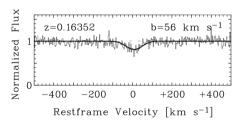


Figure 2: A BLA towards the LOS of quasar H 1821+643. Philipp Richter (2005)

- ▶ The missing baryon problem
- ▶ BLAs : Way to probe WHIM
- ► Absorber towards PG 0003+158
- ▶ BLA survey : 28 BLA candidates

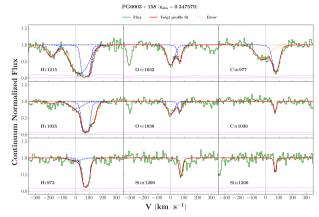


Figure 3: System plot of the absorber system towards PG 0003+158. Velocity is taken zero at z = 0.347579

- ▶ The missing baryon problem
- ▶ BLAs : Way to probe WHIM
- ▶ Absorber towards PG 0003+158
- ▶ BLA survey : 28 BLA candidates

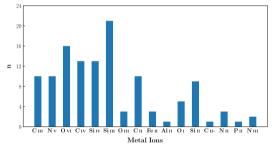


Figure 4: Distribution of metal ions in all 28 candidate BLAs

### The BLA Survey

# Survey so far...

### Survey so far...

▶ Voigt profile fitting : 16 (O VI) + 6

### Survey so far...

- ▶ Voigt profile fitting : 16 (O VI) + 6
- ► Ionisation Modelling : **16 (O VI)**

### **Insights**

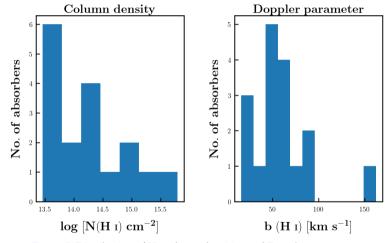


Figure 5: Distribution of H I column densities and Doppler parameters.

# **Insights**

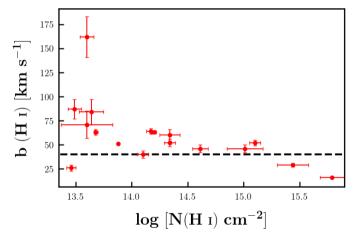


Figure 6: H I column density vs. Doppler parameter



# **Insights**

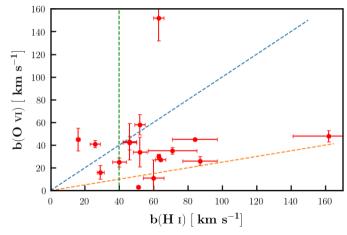


Figure 7: b(O VI) vs. b(H I)



# **Ionisation Modelling**



▶ Grid of PI CLOUDY models : Density and Metallicity

- ▶ Grid of PI CLOUDY models : Density and Metallicity
- ▶  $\log (n_H/cm^{-3})$ : -5 to 1 in steps of 0.02



- ▶ Grid of PI CLOUDY models : Density and Metallicity
- ▶  $\log (n_H/cm^{-3})$ : -5 to 1 in steps of 0.02
- ▶  $\log (Z/Z \odot)$ : -3 to 2 in steps of 0.05

- ▶ Grid of PI CLOUDY models : Density and Metallicity
- ▶  $\log (n_H/cm^{-3})$ : -5 to 1 in steps of 0.02
- ▶  $\log (Z/Z \odot)$ : -3 to 2 in steps of 0.05
- Solution : Model that best predicts the observed column densities





▶ 2d CLOUDY models : computationally expensive



- 2d CLOUDY models : computationally expensive
- Scale column density with metallcity



- 2d CLOUDY models : computationally expensive
- Scale column density with metallcity
- $N(n_H, Z) = N(n_H, Z_0) + \log(Z/Z_0)$

- 2d CLOUDY models : computationally expensive
- Scale column density with metallcity
- $N(n_H, Z) = N(n_H, Z_0) + \log(Z/Z_0)$
- ►  $Z_0 = 0.1Z$ ⊙

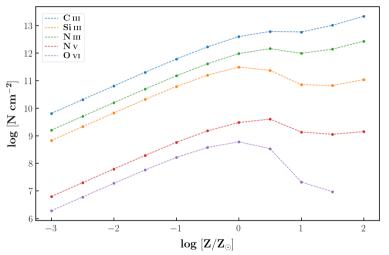


Figure 8: Column densities of various ions at different metallcity. N(H I)= $10^{14}$  cm $^{-2}$  and  $n_H=10^{-3}$  cm $^{-3}$ 

#### **Results**

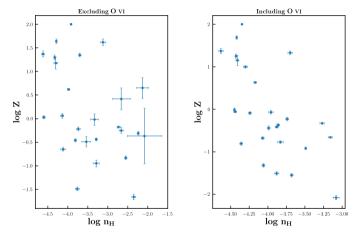


Figure 9: Ionisation modelling solutions for both excluding and including O VI cases.



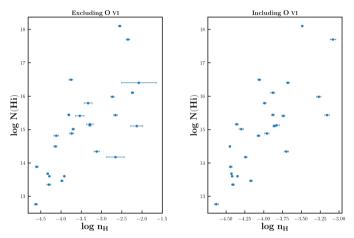


Figure 10:  $n_H$  vs. N(H I) for both excluding and including O VI cases.



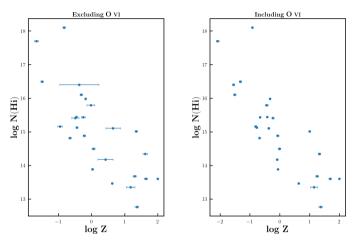


Figure 11: Z vs. N(H I) for both excluding and including O VI cases.



### Towards the end

Voigt profile fitting : 6



- ▶ Voigt profile fitting : 6
- ▶ Ionisation modelling : 12



- ▶ Voigt profile fitting : 6
- ▶ Ionisation modelling : 12
- Exploring the survey results



- ▶ Voigt profile fitting : 6
- ▶ Ionisation modelling : 12
- Exploring the survey results
- *Finally*, calculating  $\Omega_b$

#### References

Ionisation Modelling

Acharva A., Khaire V., 2021, MNRAS, 509, 5559

Danforth C. W., et al., 2016, ApJ, 817, 111

Lehner N., Savage B. D., Richter P., Sembach K. R., Tripp T. M., Wakker B. P., 2007, ApJ, 658, 680

Richter P., Savage B. D., Sembach K. R., Tripp T. M., 2006, A&A, 445, 827

Savage B. D., Kim T. S., Wakker B. P., Keeney B., Shull J. M., Stocke J. T., Green J. C., 2014, ApJS, 212, 8



References