# Tracing Baryons in the Warm Hot Intergalactic Medium using Broad Lyman- $\alpha$ Absorbers

Mid-Term 2

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# Thesis Phase I: Recap

## Recap

- ▶ The missing baryon problem
- ▶ BLAs : Way to probe WHIM
- ▶ Absorber towards PG 0003+158
- ▶ BLA survey : 28 BLA candidates

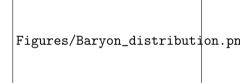


Figure 1: Baryon budget at  $z \sim 0$ . Shull et al. (2012)

# **Ionisation Modelling**

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- ►  $Z_0 = 0.1Z$ ⊙

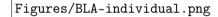


Figure 2: A BLA towards the LOS of quasar H 1821+643 ( $z_{em}=0.297$ ) Philipp Richter (2005)

#### How to detect WHIM?

- Quasars as backlight
  - O VI-VIII, Ne IX-X, N VII, etc.
  - BLAs

Figures/BLA.png

Ref. : Tepper-García et al. (2013) Figure 3: A BLA blended with other Lyα absorption lines towards the LOS of quasar PG1116+215 ( $z_{em}=0.176$ ). Philipp Richter (2020)

Savage et al. (2014)

## Data

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- $R \sim 17,000 \approx 17 \text{ km s}^{-1}$

## Phase I

Phase I

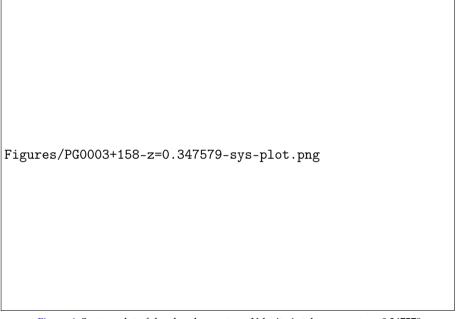
## Absorber system towards PG0003+158

## **Absorber system**

- Quasar at  $z_{em} = 0.45089$
- $z_{abs} \sim 0.347$
- ▶ 3 component system

Figures/Component\_structure2.png

# Voigt profile fitting



# Voigt profile fitting

- ▶ HI: 3 components
- ► O VI: 2 components
- ► CII, CIII, Si II, Si III: 1 component

Figures/param.png

#### **CLOUDY**

Figures/cloudy-transparent.png

Figure 5: Schematic diagram of CLOUDY simulations.

## **Ionization Modelling**

- ▶ Component I :-
- ► Component II : Hybrid Collisional + Photo-ionization
- Component III : Photo-ionization (PI)

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- Solution : Model that best matches the observed column densities

Ref.: Acharya and Khaire (2021)

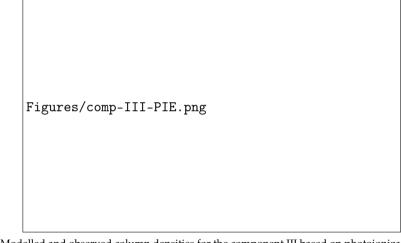


Figure 6: Modelled and observed column densities for the component III based on photoionization modelling

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- ▶ O VI and size as constraining factors

Figures/physical-params.png

# The Survey

$$b \geqslant 45 \,\mathrm{km \, s^{-1}}$$

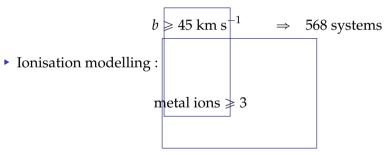
$$b \geqslant 45 \text{ km s}^{-1} \implies 568 \text{ systems}$$

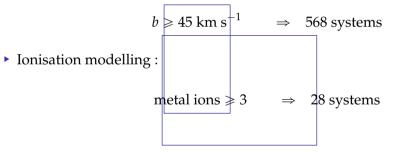
• Broad Ly $\alpha$ :

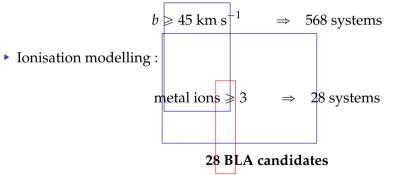
$$b \geqslant 45 \text{ km s}^{-1} \qquad \Rightarrow \quad 568 \text{ systems}$$

▶ Ionisation modelling :









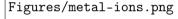


Figure 7: No. of different metal ions in all the 28 candidate BLAs

#### References

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