# Tracing Baryons in the Warm Hot Intergalactic Medium using Broad Lyman- $\alpha$ Absorbers

Mid-Term 2

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### Thesis Phase I: Recap

- ▶ The missing baryon problem
- ▶ BLAs : Way to probe WHIM
- ▶ Absorber towards PG 0003+158
- ▶ BLA survey : 28 BLA candidates

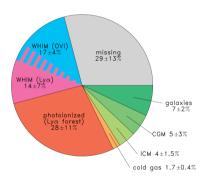


Figure 1: Baryon budget at  $z \sim 0$ . Shull et al. (2012)

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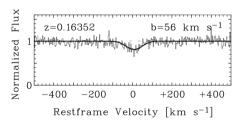


Figure 2: A BLA towards the LOS of quasar H 1821+643. Philipp Richter (2005)

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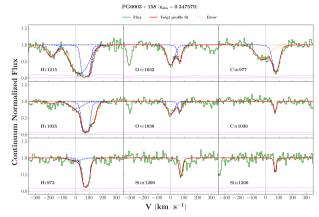


Figure 3: System plot of the absorber system towards PG 0003+158. Velocity is taken zero at z = 0.347579

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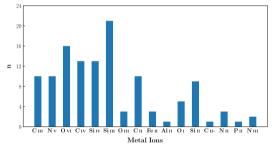


Figure 4: Distribution of metal ions in all 28 candidate BLAs

### The BLA Survey

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- ► Ionisation Modelling : **16 (O VI)**

### **Insights**

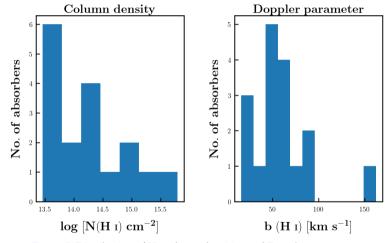


Figure 5: Distribution of H I column densities and Doppler parameters.

# **Insights**

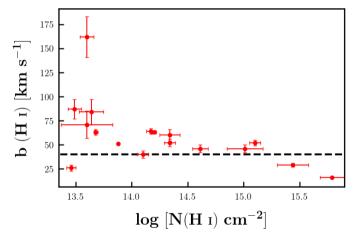


Figure 6: H I column density vs. Doppler parameter



# **Insights**

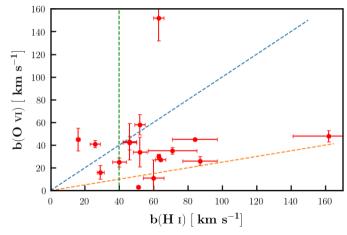


Figure 7: b(O VI) vs. b(H I)



# **Ionisation Modelling**



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- ▶  $\log (n_H/cm^{-3})$ : -5 to 1 in steps of 0.02
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- Solution : Model that best predicts the observed column densities





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- ►  $Z_0 = 0.1Z$ ⊙

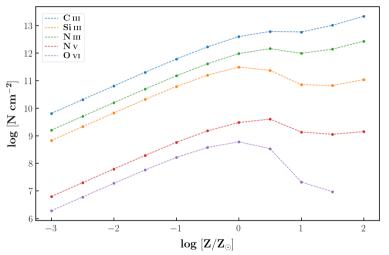


Figure 8: Column densities of various ions at different metallcity. N(H I)= $10^{14}$  cm $^{-2}$  and  $n_H=10^{-3}$  cm $^{-3}$ 

#### **Results**

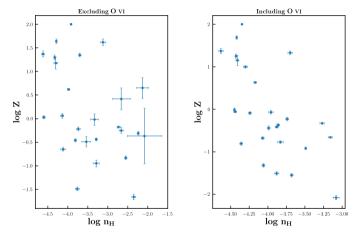


Figure 9: Ionisation modelling solutions for both excluding and including O VI cases.



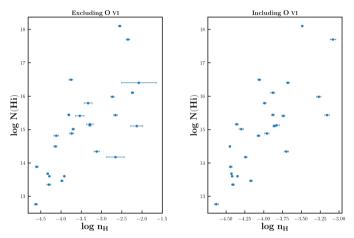


Figure 10:  $n_H$  vs. N(H I) for both excluding and including O VI cases.



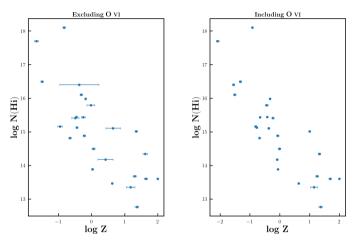


Figure 11: Z vs. N(H I) for both excluding and including O VI cases.



### Towards the end

Voigt profile fitting : 6



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- ▶ Ionisation modelling : 12



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- Exploring the survey results



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- ▶ Ionisation modelling : 12
- Exploring the survey results
- *Finally*, calculating  $\Omega_b$  (*BLA*)

#### References

**Ionisation Modelling** 

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