

PHY407H1 Lab6

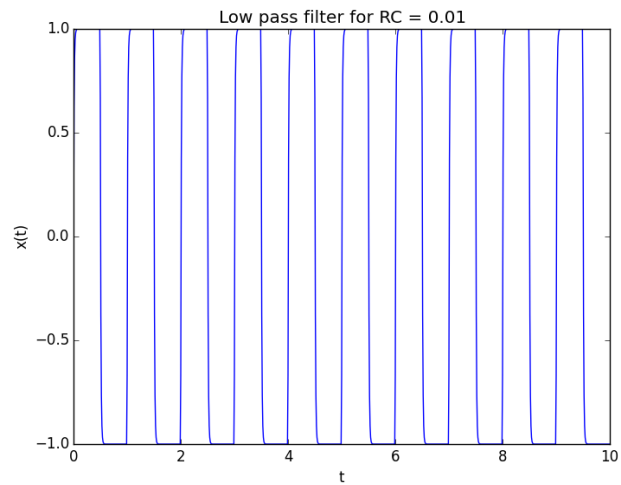
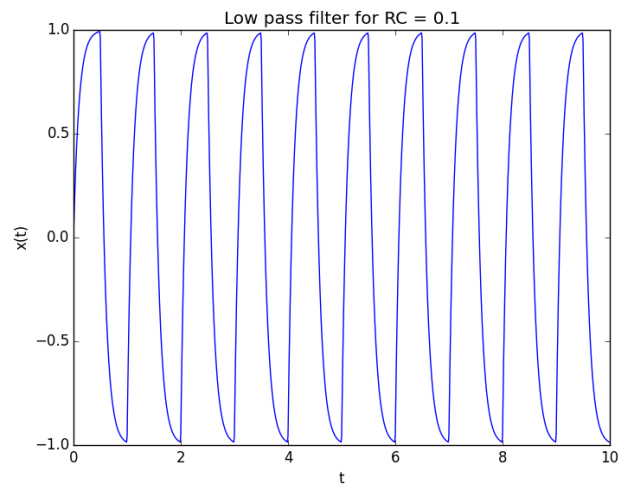
Eric Yeung*

Department of Physical Sciences, University of Toronto, Toronto M1C 1A4, Canada

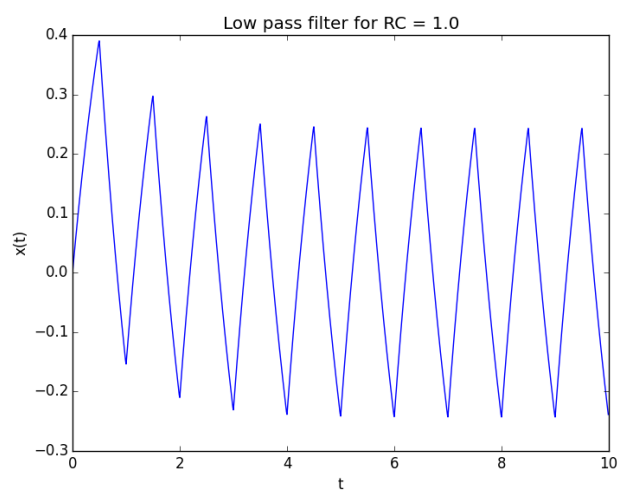
(Dated: October 24, 2015)

QUESTION 1

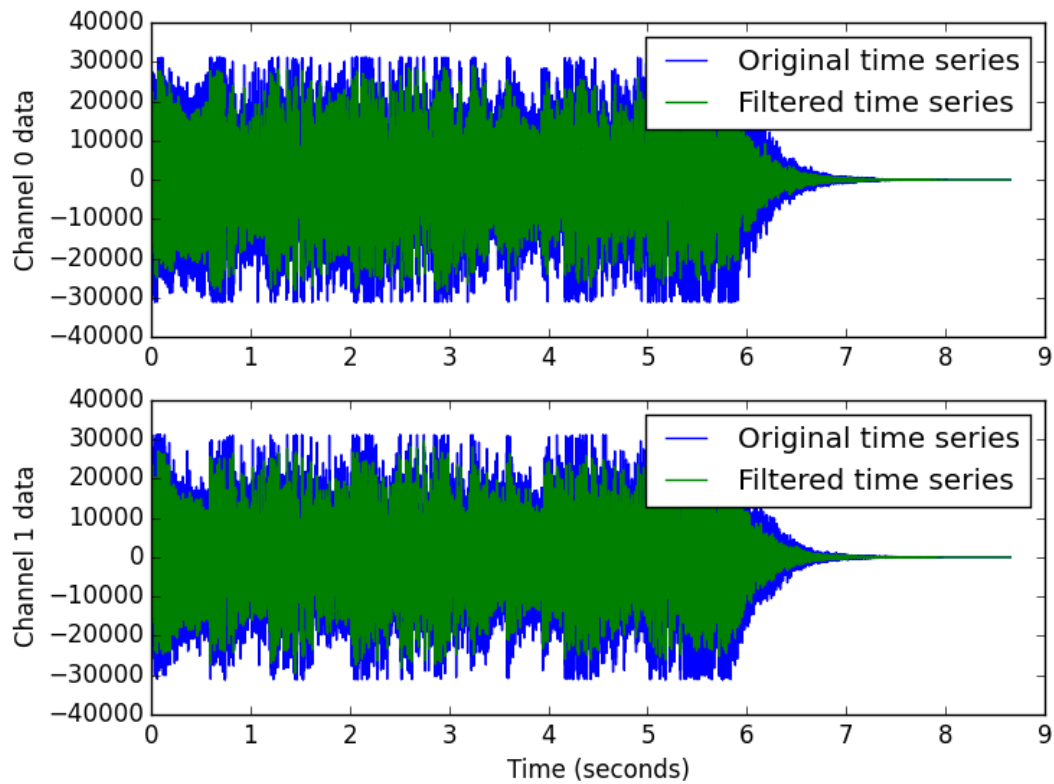
Plots from the textbook,



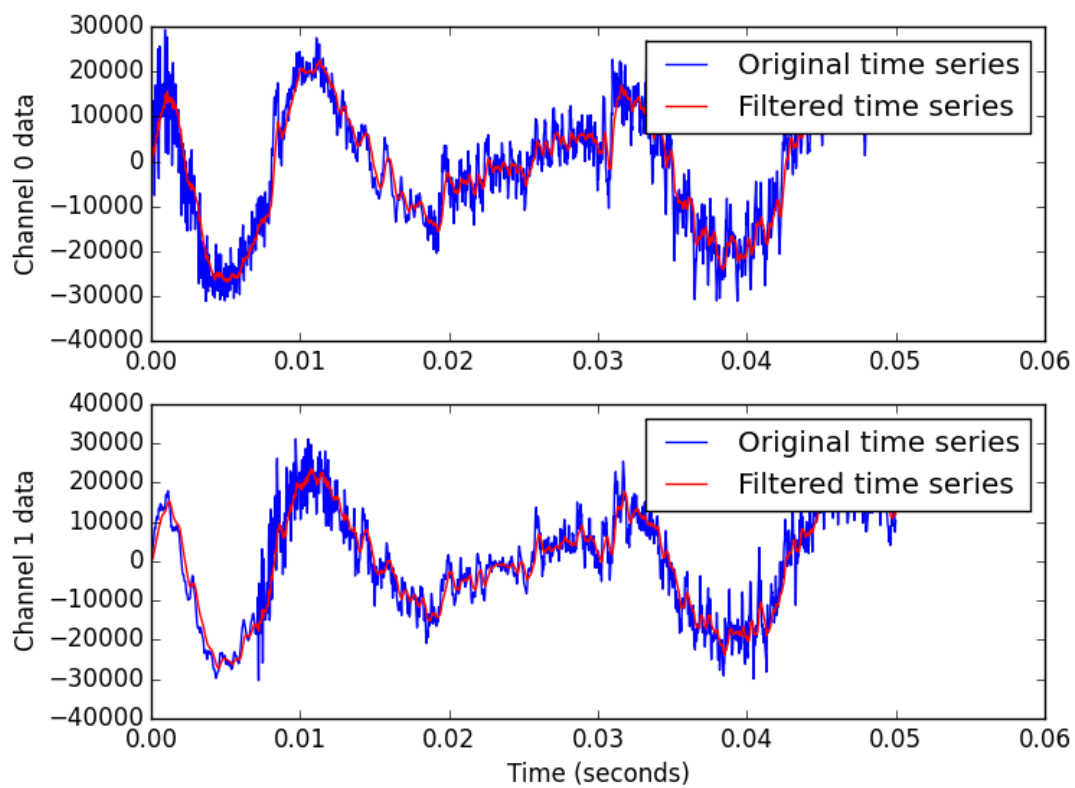
* eric.yeung@mail.utoronto.ca



a) See [Lab6_q1a.py](#) by Chi.

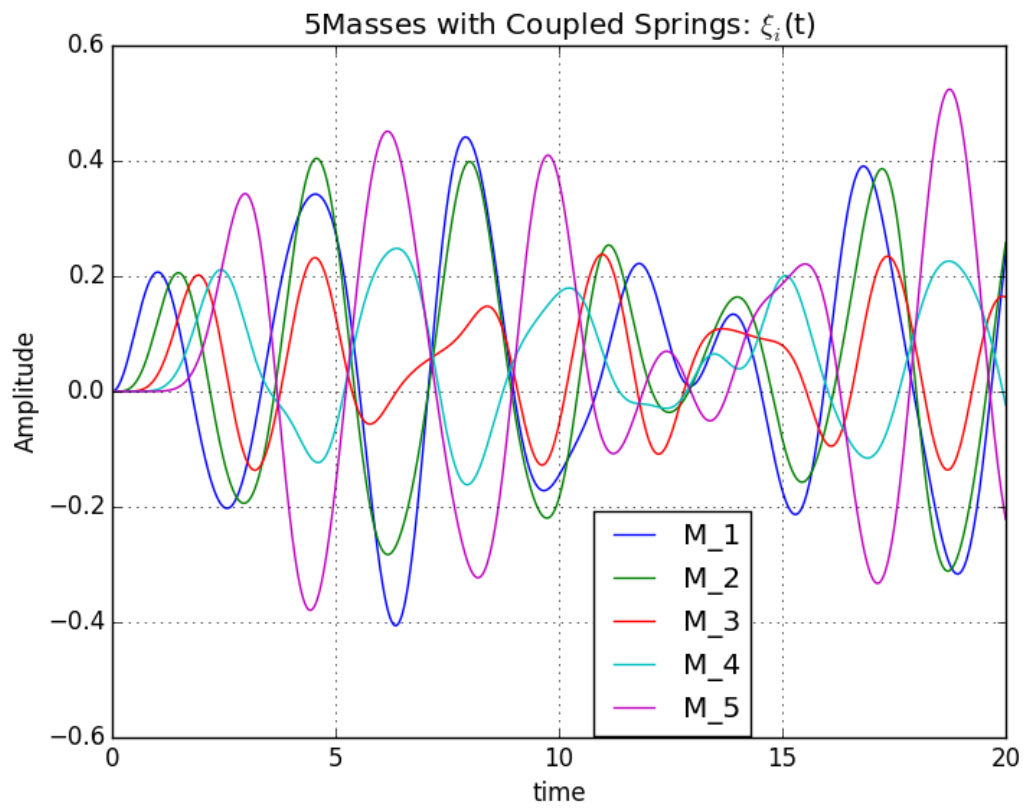


The graph below was for a shorter timescale (0, 0.05). The one above has a larger timescale. Advantages of this method is it is faster and it is relatively easier to code.



QUESTION 2

See [Lab6-q2a.py](#) by me. The animation was attempted but not submitted because I was told you do not need to hand in the animation.



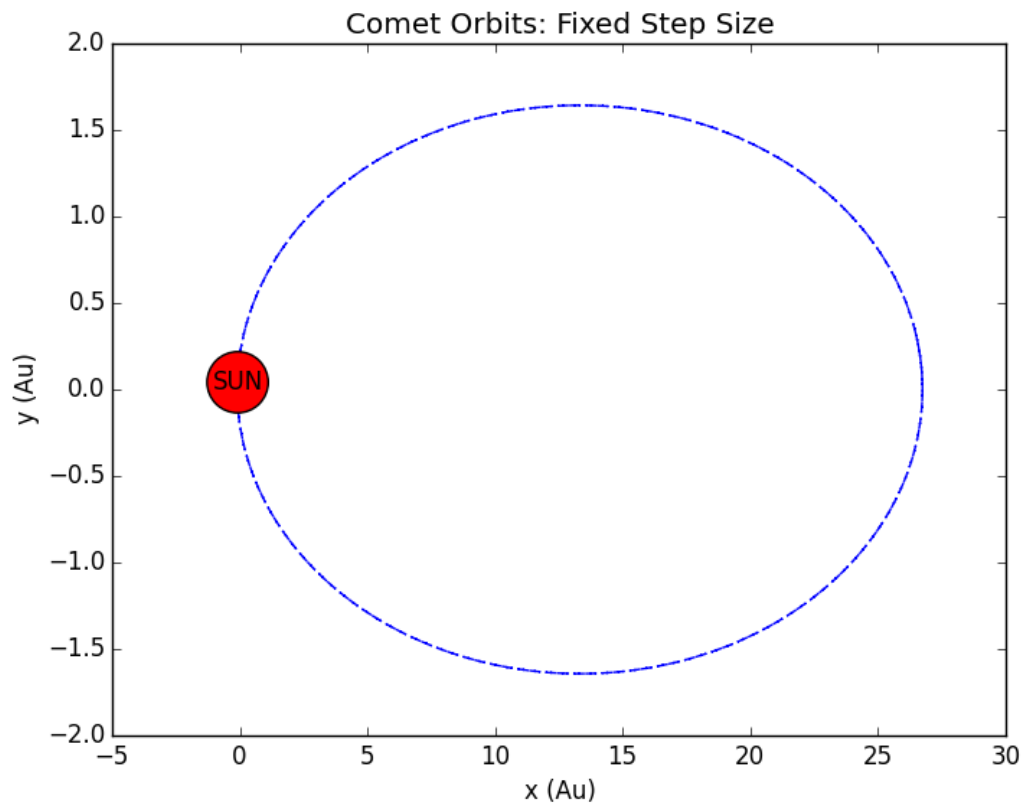
QUESTION 3

a) Reduced the order of the SODEs to ODEs.

$$\begin{aligned} dx/dt &= v_x \\ dy/dt &= v_y \\ dv_x/dt &= -G * M * x / radius ** 3 \\ dv_y/dt &= -G * M * y / radius ** 3 \end{aligned}$$

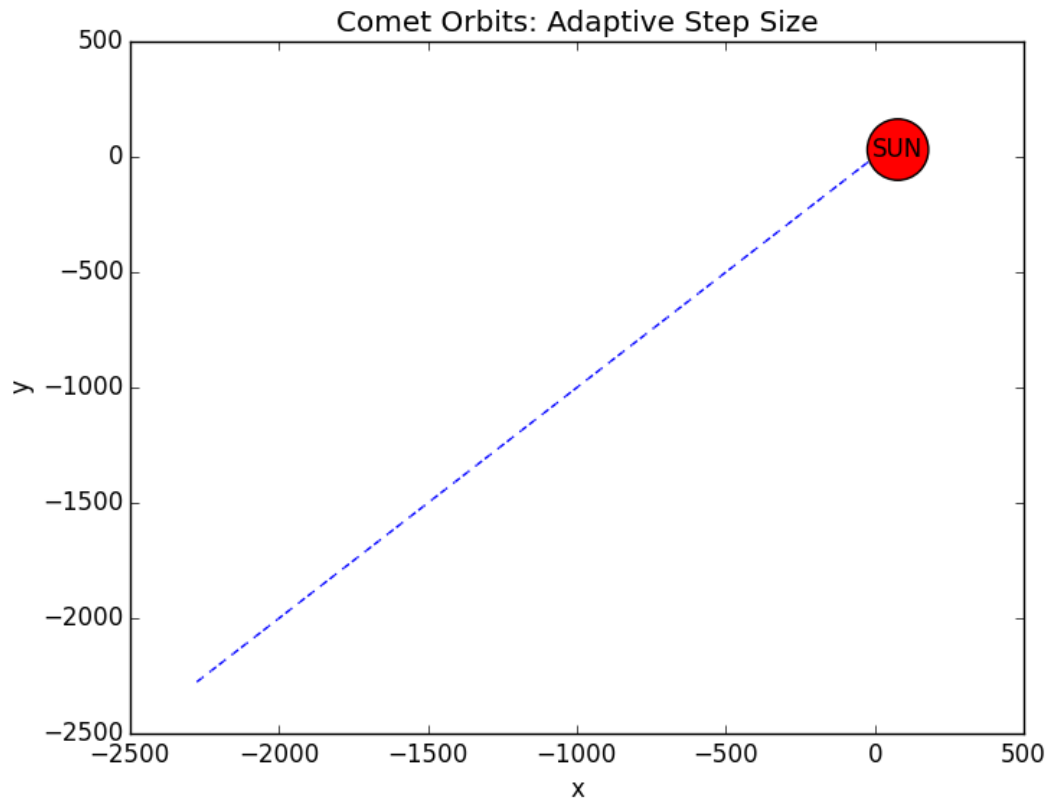
b) See [Lab6_q3b.py](#) by me.

I used $h = 5e-4$, and total time of 200 years. The program took 22.48 seconds to run, this was quite long.



c) See [Lab6_q3c.py](#) by me.

I couldn't get the adaptive code to work- because I ran out of time. After debugging, I found that ρ kept on returning not a number. And this was because my two estimates at different time steps h and $2 \times h$ were the same. But I would expect it to look like my first graph but with bigger spaces near the aphelion and smaller spaces near the the perihelion. Below is the obviously incorrect plot.



d) As one can imagine, the steps get closer when the comet is near the perihelion (closest to sun), and farther apart when the comet is near the aphelion (furthest point from the sun)