
Lindblad

Master-Projekt

Sander Stammbach
Prof. Patrick Plotts
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1 Introduction

One of the most important question in thermodynamics is, how to convert thermal energy into work. For such tasks it exist many classical engines, as example the steam-machine. To quantify heat-engines, its commen to look at the ergotropy. In my master-project I will quantify the a three level maser. The three-level maser is a Quantum heat engine (short: QHE). The work extraction from a classicle heat engine is a moving piston. But in this case it is a driving field. Experimentally we need two different reservoir. The high-temperature reservoir can be realized by a gas noise lamp and the filter by a wave guide cutting off the lower frequencies. (Quelle main Paper)

2 Theory

2.1 3-level-Maser

In a three level system you have the three energy levels E_1, E_2, E_3 . Pumping is from the lowest level to the highest level E_3 . The condition for the third level is that it falls to the middle level E_2 very quickly. On average, the system is almost not in the third state. The second system should then have a higher decay time, so that a population inversion can build up. This means that several particles are in the energetically higher state. From this state they come almost exclusively through stimulated emission into the deeper system E_1 . Stimulated emission is a necessary condition for coherent light. Coherent light Means they have the same phase and same frequency. (Quelle Wiki) In the case of a normal laser the population inversion is achieved through a pumping light. In the case of this calculation the higher level will be reached with a interaction of a warm bath.

2.2 phase-averaged coherent states (PHAV)

The output of a Laser is coherent light. The quantum description of coherent light is a coherent state. The photon number distribution of coherent light is a poisson distribution. The Pignerfunction from a coherent state itself is a Gaussian. But the Wignerfunction of a pahse-average-state has a non-Gaussian Wignerfunction. The mathematical description of a PHAV is the same as a normal coherent state but with a random phase. So we get a new term of $\exp(i\pi\phi)$ in it. The PHAV state could represented by following

formula (Quelle 5):

$$|\alpha\rangle = e^{-1/2|\alpha|^2} \sum_{n=0}^{\infty} \frac{|\alpha|^{2n} e^{in\phi}}{\sqrt{n!}} \quad (1)$$

2.3 Lindblad-Master-Equation

The Basic of our system is a three-level quantum system. This three level system is driven by a hot bath and a cold bath. Those have the temperature T_c and T_h . This quantum system lives in a cavity. In my calculation, the thermal bath is constant, so I can use the Lindblad-masterequation. In the following figure is shown a three-level system: The efficiency of the maser is given by the formula:

$$\eta_{maser} = \frac{\omega_f}{\omega_h} \quad (2)$$

The master equation is:

$$\dot{\rho}(t) = \frac{1}{i\hbar}[H, \rho] + \mathcal{L}\rho \quad (3)$$

The first part of formula (1) is the von Neuman-equation. The von Neuman-equation is the analog of the Schrödinger-equation but for density matrices. This part of the equation is unitary and therefore the process is reversible. The non-unitary part of the equation \mathcal{L} have three parts. \mathcal{L}_h describe the interaction with the hot bath. \mathcal{L}_c Describe the interaction with the cold bath. \mathcal{L}_{cav} Describe the Photons which leaves the cavity. so if we have a small κ means less photons will leave and stay in the cavity. we see that in the Fockplott.

The interaction Hamiltonian or Jaynes-Cummings Hamiltonian is:

$$H_{int} = \hbar g(\sigma_{12}a^\dagger + \sigma_{21}a) \quad (4)$$

The Hamiltonian of the photons is, which describes the phonons in the cavity:

$$H_{free} = \sum_{i=1}^3 \hbar\omega_i |i\rangle \langle i| + \hbar\omega_f a^\dagger a \quad (5)$$

The total Hamiltonian

$$H = H_{free} + H_{int} \quad (6)$$

The interaction with the various environmental heat baths is described by the Liouvillian:

$$\begin{aligned}\mathcal{L}\hat{\rho} = & \frac{\gamma_h}{2}\bar{n}(\omega_h, T_h) + 1 \cdot \mathcal{D}[\sigma_{13}] + \frac{\gamma_h}{2}\bar{n}(\omega_h, T_h) \cdot \mathcal{D}[\sigma_{31}] \\ & + \frac{\gamma_c}{2}\bar{n}(\omega_c, T_c) + 1 \cdot \mathcal{D}[\sigma_{23}] + \frac{\gamma_c}{2}\bar{n}(\omega_c, T_c) \cdot \mathcal{D}[\sigma_{32}] \\ & \kappa\bar{n}(\omega_f, T_f) + 1 \cdot \mathcal{D}[a] + \kappa\bar{n}(\omega_f, T_f) \cdot \mathcal{D}[a^\dagger]\end{aligned}\tag{7}$$

\mathcal{D} is defined with this formula:

$$\mathcal{D}[A] = (2A\rho A^\dagger - A^\dagger A\rho - \rho A^\dagger A)\tag{8}$$

The Bose-Einstein statistic is a probability distribution in quantum statistics . It describes the mean occupation number $\langle n(E) \rangle$ of a quantum state of energy E , in thermodynamic equilibrium at absolute temperature T for identical bosons as occupying particles. n depends on the temperature and the frequency. n is defined as:

$$n(\omega, T) = \frac{1}{\exp[\frac{\hbar\omega_i}{k_b T_i}] - 1}\tag{9}$$

The prefactor γ_i The Liouvillian have different constants. The coupling constants g for the Hamiltonian and the κ for the Liouvillian part.

2.4 Probability calculation

When we work with density matrices, its common to work with expectation values with $\langle A \rangle = Tr[A \cdot \rho]$. A is a operator and describe a measurement. With this we can calculate the probability To calculate the expected heat flow we can take the partial trace from

$$\langle J \rangle = Tr[\rho_{free} \cdot \mathcal{L}_h[\rho]] + Tr[\rho_{free} \cdot \mathcal{L}_c[\rho]] + Tr[\rho_{free} \cdot \mathcal{L}_{cav}[\rho]]\tag{10}$$

3 Calculation

3.1 Software

For the hole implementation of the tree-level-system in a cavity, I used qutip. Qutip is library in python, which allows to solve Masterequation pretty fast.

3.2 implementation of the tree-levelsystem in qutip

In our case only ω_f interact with the light. first i defined the frequencies ω_c , ω_h and ω_f . The constants \hbar and the bolzmanfactor k_b are 1. also defined as constants are the three different Boseinstein-distributions n_h , n_c and n_f . The transition-operators are made by following qutip implementation: "Trans13=tensor(vg*va.dag(),qutip.identity(nph))". nph is the maximum of the photonnumber in the cavity. so I get 90 x 90 matrices. similarly the projectors. with those its easy to construct the hamiltoniens, H_{free} and H_{int} as in formula 3 and 4. To calculate the the density matrices for steadystates we can also use a qutip function, call `steadystate()`. this function needs the total Hamiltonian and a list of the non-unitary operators as arguments. We can we can construct this list as a multiplication of our transition-operators and the tree different Boseinstein-distributions times the different γ -factors. as output of the function steady state we get the density-matrices for steady-states.

3.3 Out

First i made a Fockplot and a wigner function of the reduce density matrices ρ_{free} . Because $\rho = \rho_{atom} \otimes \rho_{free}$, I can make the partial trace of ρ with qutip, to trace out the reduced density matrices ρ_{free} . The Fock-plot and the Wigner-plot is also done with a qutip function. with the density matrix times the $L_p\rho$ i calculated on the heat flux by taking the trace of $H\mathcal{L} \cdot \rho$. and plot this for 200 different g 's so the goal of this work is to find Einstein-Bose-distributions which yield a RAHV state. as in the paper (Quelle2)

4 Results

5 Discussion

6 References

$$\begin{aligned} \mathcal{L}\hat{\rho} = & \frac{\gamma_h}{2} \left[\frac{1}{\exp[\frac{\hbar\omega_h}{k_b T_h}] - 1} + 1 \right] \cdot \left(2\sigma_{13} \cdot \rho \cdot \sigma_{13}^\dagger - \sigma_{13}^\dagger \sigma_{13} \rho - \rho \sigma_{13}^\dagger \sigma_{13} \right) \\ & + \frac{\gamma_h}{2} \left[\frac{1}{\exp[\frac{\hbar\omega_h}{k_b T_H}] - 1} \right] \cdot \left(2\sigma_{31} \cdot \rho \cdot \sigma_{31}^\dagger - \sigma_{31}^\dagger \sigma_{31} \rho - \rho \sigma_{31}^\dagger \sigma_{31} \right) \end{aligned}$$

$$\begin{aligned}
& + \frac{\gamma_c}{2} \left[\frac{1}{\exp[\frac{\hbar\omega_c}{k_b T_c}] - 1} + 1 \right] \cdot \left(2\sigma_{23} \cdot \rho \cdot \sigma_{23}^\dagger - \delta_{23}^\dagger \sigma_{23} \rho - \rho \sigma_{23}^\dagger \sigma_{23} \right) \\
& + \frac{\gamma_c}{2} \left[\frac{1}{\exp[\frac{\hbar\omega_c}{k_b T_c}] - 1} \right] \cdot \left(2\sigma_{32} \cdot \rho \cdot \sigma_{32}^\dagger - \sigma_{32}^\dagger \sigma_{32} \rho - \rho \sigma_{32}^\dagger \sigma_{32} \right) \\
& \kappa \left[\frac{1}{\exp[\frac{\hbar\omega_f}{k_b T_f}] - 1} + 1 \right] \cdot \left(2a\rho a^\dagger - a^\dagger a \rho - \rho a^\dagger a \right) \\
& \kappa \left[\frac{1}{\exp[\frac{\hbar\omega_f}{k_b T_f}] - 1} \right] \cdot \left(2a^\dagger \rho a - a a^\dagger \rho - \rho a a^\dagger \right)
\end{aligned} \tag{11}$$