

1. The Kerr-cat qubit

$$\text{Def: } |C_{\alpha}^{\pm}\rangle = N_{\alpha}^{\pm} (|+\alpha\rangle \pm |- \alpha\rangle)$$

$$N_{\alpha}^{\pm} = 1/\sqrt{2(1 \pm e^{-2|\alpha|^2})}$$

$| \pm \alpha \rangle$: same n , but opposite phase

$$| \pm X \rangle = (|C_{\alpha}^{\pm}\rangle \pm |C_{\bar{\alpha}}^{\pm}\rangle)/\sqrt{2}$$

$$| \pm Y \rangle = (|C_{\alpha}^{\pm}\rangle \pm i|C_{\bar{\alpha}}^{\pm}\rangle)/\sqrt{2}$$

$$| \pm Z \rangle = |C_{\alpha}^{\pm}\rangle$$

a) We define $|C_{\alpha}^{\pm}\rangle \hat{=} \begin{pmatrix} 1 \\ 0 \end{pmatrix}$

$$|C_{\bar{\alpha}}^{\pm}\rangle \hat{=} \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

$$11 = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \begin{pmatrix} 1 \\ 0 \end{pmatrix} + \begin{pmatrix} 0 \\ 1 \end{pmatrix} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \hat{=} |C_{\alpha}^+ X C_{\alpha}^+| + |C_{\bar{\alpha}}^+ X C_{\bar{\alpha}}^+|$$

$$X = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \begin{pmatrix} 0 & 1 \end{pmatrix} + \begin{pmatrix} 0 \\ 1 \end{pmatrix} \begin{pmatrix} 1 & 0 \end{pmatrix} \hat{=} |C_{\alpha}^+ X C_{\bar{\alpha}}^+| + |C_{\bar{\alpha}}^+ X C_{\alpha}^+|$$

$$Y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \hat{=} -i|C_{\alpha}^+ X C_{\bar{\alpha}}^+| + i|C_{\bar{\alpha}}^+ X C_{\alpha}^+|$$

$$Z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \hat{=} |C_{\alpha}^+ X C_{\alpha}^+| - |C_{\bar{\alpha}}^+ X C_{\bar{\alpha}}^+|$$

$$X| \pm X \rangle = X(|C_{\alpha}^{\pm}\rangle \pm |C_{\bar{\alpha}}^{\pm}\rangle)/\sqrt{2} = \frac{1}{\sqrt{2}}(|C_{\bar{\alpha}}^{\pm}\rangle \pm |C_{\alpha}^{\pm}\rangle) = \begin{cases} 1/\sqrt{2}(|C_{\alpha}^+ \rangle + |C_{\bar{\alpha}}^+ \rangle) \rightarrow EV=1 \text{ for } |+X\rangle \\ -1/\sqrt{2}(|C_{\alpha}^+ \rangle - |C_{\bar{\alpha}}^+ \rangle) \rightarrow EV=-1 \text{ for } |-X\rangle \end{cases}$$

$$Y| \pm Y \rangle = Y(|C_{\alpha}^{\pm}\rangle \pm i|C_{\bar{\alpha}}^{\pm}\rangle)/\sqrt{2} = \frac{1}{\sqrt{2}}(|C_{\bar{\alpha}}^{\pm}\rangle \pm |C_{\alpha}^{\pm}\rangle) = \begin{cases} 1/\sqrt{2}(|C_{\alpha}^+ \rangle + i|C_{\bar{\alpha}}^+ \rangle) \rightarrow & |+Y\rangle \\ -1/\sqrt{2}(|C_{\alpha}^+ \rangle - i|C_{\bar{\alpha}}^+ \rangle) \rightarrow & |-Y\rangle \end{cases}$$

$$Z| \pm Z \rangle = Z|C_{\alpha}^{\pm}\rangle = \pm |C_{\alpha}^{\pm}\rangle \quad \downarrow$$

b) Without Normalisation: $|C_{\alpha}^{\pm}\rangle \sim (|+\alpha\rangle \pm |- \alpha\rangle)$

We expand $|+\alpha\rangle = e^{1/2|\alpha|^2} \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} |n\rangle$ and $|- \alpha\rangle = e^{1/2|\alpha|^2} \sum_{n=0}^{\infty} \frac{(-\alpha)^n}{\sqrt{n!}} |n\rangle$

$$|C_{\alpha}^+\rangle = 2e^{1/2|\alpha|^2} \left(\frac{\alpha^0}{\sqrt{0!}} |0\rangle + 0 + \frac{\alpha^2}{\sqrt{2!}} |2\rangle + 0 + \dots \right) \rightarrow \text{odd terms cancel} \rightarrow \text{has even photon number.}$$

$$|C_{\bar{\alpha}}^+\rangle = 2e^{1/2|\alpha|^2} \left(0 + \frac{\alpha^1}{\sqrt{1!}} |1\rangle + 0 + \frac{\alpha^3}{\sqrt{3!}} |3\rangle + 0 + \dots \right) \rightarrow \text{even terms cancel} \rightarrow \text{has odd photon number.}$$

c) $\alpha=0$. We develop $e^{-2|\alpha|^2} = 1 - 2|\alpha|^2 + \dots$

$$e^{\pm 2|\alpha|^2} = 1 \pm 2|\alpha|^2 + \dots$$

$$|C_{\alpha}^+\rangle = 1/\sqrt{2(1+e^{-2|\alpha|^2})} \cdot 2e^{\pm 2|\alpha|^2} |0\rangle \approx 1/\sqrt{2(1+1-2|\alpha|^2)} \cdot 2(1-\frac{1}{2}|\alpha|^2) |0\rangle$$

$$\approx 1/\sqrt{1-1+2|\alpha|^2} \cdot (1-\frac{1}{2}|\alpha|^2) |0\rangle \approx |0\rangle$$

$$|C_{\bar{\alpha}}^+\rangle = 1/\sqrt{2(1-e^{-2|\alpha|^2})} \cdot 2e^{\pm 2|\alpha|^2} \cdot \alpha |1\rangle \approx \frac{1}{\sqrt{2(1-1+2|\alpha|^2)}} \cdot 2(1-\frac{1}{2}|\alpha|^2) \cdot \alpha |1\rangle = \frac{1}{2|\alpha|} \cdot 2(1-\frac{1}{2}|\alpha|^2) \alpha |1\rangle$$

$$\approx |1\rangle$$

Note: First order approximation would have been sufficient (probably, I didn't check explicitly)

d) limit $|x|^2 = \bar{n} \rightarrow \infty$

$$I \cdot P = N_{\alpha}^+ / N_{\alpha}^- = \frac{1/\sqrt{2(1+e^{2|x|^2})}}{1/\sqrt{2(1-e^{2|x|^2})}} \approx \frac{1/\sqrt{2(1+0)}}{1/\sqrt{2(1-0)}} = 1 \Rightarrow N_{\alpha}^+ \sim N_{\alpha}^- \sim \frac{1}{2}$$

$$II \quad |x\rangle = (|C_{\alpha}^+\rangle \pm |C_{\alpha}^-\rangle) / \sqrt{2} \approx \frac{1}{2} (1+\alpha) + 1-\alpha \pm (1+\alpha) - 1-\alpha = \frac{1}{2} 1 \pm \alpha = 1 \pm \alpha$$

$$III \quad |y\rangle = (|C_{\alpha}^+\rangle \pm i|C_{\alpha}^-\rangle) / \sqrt{2} \approx \frac{1}{2} (1+\alpha) + 1-\alpha \pm i(1+\alpha) - i(1-\alpha) = \frac{1}{2} (1 \pm i)(1+\alpha) + (1 \mp i)(1-\alpha)$$

Now we analyze $|y\rangle$ individually

$$|y\rangle = \frac{1}{2} ((1+i)(1+\alpha) + (1-i)(1-\alpha)) \Rightarrow \text{extract global phase}$$

$$\approx \frac{1}{2} (\sqrt{2}(1+\alpha) - i\sqrt{2}(1-\alpha)) = \frac{1}{\sqrt{2}} (1+\alpha - i(1-\alpha))$$

$$|y\rangle = \frac{1}{2} ((1-i)(1+\alpha) + (1+i)(1-\alpha)) \approx \frac{1}{\sqrt{2}} (1-\alpha + i(1+\alpha))$$

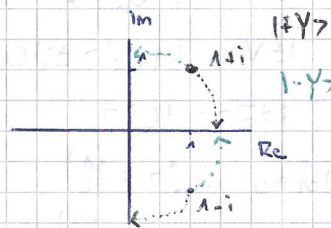
$$IV \quad |z\rangle = |C_{\alpha}^{\pm}\rangle = \frac{1}{\sqrt{2}} (1+\alpha \pm 1-\alpha)$$

$$V \quad 1-P = 1 - \frac{N_{\alpha}^+}{N_{\alpha}^-} \quad \text{for } \bar{n} = |x|^2 \in \{2, 4, 6\}$$

$$1-P = 1 - \frac{1/\sqrt{2(1+e^{2|x|^2})}}{1/\sqrt{2(1-e^{2|x|^2})}} \stackrel{\text{Wolfram-alpha}}{=} 1 - \sqrt{\tanh(|x|^2)} = 0.018 \quad \text{for } \bar{n}=2$$

$$= 0.00034 \quad \bar{n}=4$$

$$0.06061 \quad \bar{n}=6$$



1.2 Photon loss, error suppression and stabilization

Master equation $\dot{\rho} = -\frac{i}{\hbar} [H_{eff}, \rho] + \kappa a \rho a^\dagger$ with κ photon loss rate and $H_{eff} = H - \frac{i}{2} \kappa a^\dagger a$

e) Apply $\sqrt{\kappa} a$ to $\hat{P} = |C_{\alpha}^+\rangle \langle C_{\alpha}^+| + |C_{\alpha}^-\rangle \langle C_{\alpha}^-|$

$$\sqrt{\kappa} a \hat{P} = \sqrt{\kappa} a (|C_{\alpha}^+\rangle \langle C_{\alpha}^+| + |C_{\alpha}^-\rangle \langle C_{\alpha}^-|)$$

$$= \sqrt{\kappa} a (N_{\alpha}^{+2} (1+\alpha) + 1-\alpha) (1+\alpha) + N_{\alpha}^{+2} (1+\alpha) - 1-\alpha) (1-\alpha)$$

$$= \sqrt{\kappa} a (N_{\alpha}^{+2} (1+\alpha) - 1-\alpha) (1+\alpha) + N_{\alpha}^{+2} (1+\alpha) + 1-\alpha) (1-\alpha)$$

$$= \sqrt{\kappa} a \left(\frac{N_{\alpha}^+}{N_{\alpha}^-} |C_{\alpha}^+\rangle \langle C_{\alpha}^+| + \frac{N_{\alpha}^-}{N_{\alpha}^+} |C_{\alpha}^-\rangle \langle C_{\alpha}^-| \right) \stackrel{\text{in } C_{\alpha} \text{-Basis}}{=} \sqrt{\kappa} a \begin{pmatrix} 0 & N_{\alpha}^+ / N_{\alpha}^- \\ N_{\alpha}^- / N_{\alpha}^+ & 0 \end{pmatrix}$$

We can write this in terms of Paulis (here off-diagonal, i.e. only X, Y)

$$\Rightarrow \sqrt{\kappa} a \hat{P} = \frac{1}{2} \sqrt{\kappa} a \left(\left(\frac{N_{\alpha}^+}{N_{\alpha}^-} + \frac{N_{\alpha}^-}{N_{\alpha}^+} \right) X + i \left(\frac{N_{\alpha}^-}{N_{\alpha}^+} - \frac{N_{\alpha}^+}{N_{\alpha}^-} \right) Y \right) = \frac{1}{2} \sqrt{\kappa} a \left(\left(\frac{N_{\alpha}^{+2} + N_{\alpha}^{-2}}{N_{\alpha}^+ N_{\alpha}^-} \right) X + i \left(\frac{N_{\alpha}^{-2} - N_{\alpha}^{+2}}{N_{\alpha}^+ N_{\alpha}^-} \right) Y \right)$$

Wolfram-alpha

$$= \frac{1}{2} \sqrt{\kappa} a \left(\left(\frac{2}{1-e^{-2|x|^2}} \right) X + i \left(-\frac{2e^{2|x|^2}}{1-e^{-2|x|^2}} \right) Y \right) = \frac{\sqrt{\kappa} a}{\sqrt{1-e^{-2|x|^2}}} (X - i e^{2|x|^2} Y)$$

We can see, that the jump operator corresponds to a superposition of bit-flips and bit-phase-flips with different complex amplitudes.

f) For large \bar{n} , bit-phase flips (and therefore phase flips) are suppressed exponentially.

Physically, we can argue, that photon loss cannot cause a transition from $|x\rangle$, since both are eigenstates of \hat{a} .

However, for their superposition, it holds $\hat{a} |C_{\alpha}^{\pm}\rangle = \alpha |C_{\alpha}^{\pm}\rangle$, which corresponds to a bit-flip.



g) Limit $\bar{n} \rightarrow \infty$

$$p_j = \tau \text{Tr}[k a \hat{\rho} a^\dagger]$$

Without loss of generality $\hat{\rho} = |\beta\rangle\langle\beta|$ with $|\beta\rangle = a|\alpha\rangle + b|\alpha^*\rangle \Rightarrow \hat{\rho} = \begin{pmatrix} |\alpha|^2 & \alpha\bar{b} \\ \alpha^*b & |\alpha^*|^2 \end{pmatrix}$

$$\Rightarrow X\hat{\rho}X = \begin{pmatrix} |\alpha|^2 & \alpha\bar{b} \\ \alpha^*b & |\alpha^*|^2 \end{pmatrix}$$

$$\Rightarrow p_j = \tau \cdot k \cdot |\alpha|^2 \cdot \text{Tr}(X\hat{\rho}X) = \tau \cdot k \cdot |\alpha|^2 = \tau \cdot k \cdot \bar{n}$$

\Rightarrow Linear dependence on $\bar{n} \Rightarrow$ The higher the average photon number, the higher the chance for a jump
... but at least it is not exponential !!

h) Parametrically driven nonlinear Hamiltonian $\frac{\hat{H}}{\hbar} = -k a^{\dagger 2} + \epsilon_2 a^{\dagger 2} + \epsilon_2^* a^2$

$$\text{We factorize the Hamiltonian } \frac{\hat{H}}{\hbar} = -k(a^{\dagger 2} - \frac{\epsilon_2^*}{k})(a^2 - \frac{\epsilon_2}{k}) + \frac{|\epsilon_2|^2}{k}$$

With this form of the Hamiltonian we can guess the Eigenstates $| \pm \alpha \rangle = | \pm \sqrt{\epsilon_2/k} \rangle$

which are Energy degenerate with Energy $\frac{|\epsilon_2|^2}{k}$ (see paper: Engineering the quantum states of light...
Puri S. et al. 2017)

Eigenstates will stay the same under evolution with this Hamiltonian.

Since they are Energy-degenerate, the same holds for all of their superpositions.

i) Microwave drive $\hat{H}_{\text{drive}}/\hbar = \epsilon_x \hat{a} + \epsilon_x^* \hat{a}^\dagger$

$$a \approx \alpha X \rightarrow a^\dagger \approx \alpha^* X$$

$$\hat{H}_{\text{drive}} = \hbar (\epsilon_x \alpha X + \epsilon_x^* \alpha^* X) = \hbar (\epsilon_x \alpha + \epsilon_x^* \alpha^*) X$$

We already know the two Eigenstates of this Hamiltonian: $| \pm X \rangle$

Note, that the Rabi frequency is given by $\Omega = \frac{E_+ + E_-}{\hbar}$, where E_\pm are the Eigenenergies of $| \pm X \rangle$
(formula from physics 3)

$$E_+ = \hbar (\epsilon_x \alpha + \epsilon_x^* \alpha^*) \quad E_- = -E_+$$

$$\Rightarrow \Omega = 2(\epsilon_x \alpha + \epsilon_x^* \alpha^*) = 2\text{Re}(\epsilon_x \alpha)$$

For a transmission, the Rabi frequency would be $\Omega(t) = 2g\alpha(t)$ where g is the vacuum Rabi coupling, which corresponds to the Hamiltonian

$$\hat{H}^d(t) = \hbar(t) (\bar{e}^{i\omega t} a^\dagger + e^{i\omega t} a), \text{ which is very similar to the drive, we applied with } \hat{H}_{\text{drive}}$$