# Project 2

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# 1 Problem Statement

Our purpose is to solve 1-D Euler Equations System:

$$\begin{aligned} \frac{\partial \rho}{\partial t} + \frac{\partial (\rho u)}{\partial x} &= 0\\ \frac{\partial (\rho u)}{\partial t} + \frac{\partial (\rho u^2 + p)}{\partial x} &= 0\\ \frac{\partial (\rho v)}{\partial t} + \frac{\partial (\rho u v)}{\partial x} &= 0\\ \frac{\partial e}{\partial t} + \frac{\partial (u(e + p))}{\partial x} &= 0 \end{aligned}$$

1 Solve Riemann Problem with Initial Condition:

$$p(x,0) = 1, u(x,0) = 0, v(x,0) = \begin{cases} -1, & x > 0 \\ 1, & x < 0 \end{cases}, \rho(x,0) = \begin{cases} 1 & x > 0 \\ 3 & x < 0 \end{cases}$$

2 Solve the Woodward-Collela blast wave Problem

$$\rho(x,0) = 1, u(x,0) = 0, \, , \, p(x,0) = \left\{ \begin{array}{ll} 1000, & \text{if } x \in (0,0.1) \\ 0.01, & \text{if } x \in (0.1,0.9) \\ 100, & \text{if } x \in (0.9,1) \end{array} \right., \, v(x,0) = \left\{ \begin{array}{ll} -10, & x \in (0,0.5) \\ 20, & x \in (0.5,1) \end{array} \right.$$

With reflective boundary conditions:

$$u(0,t) = u(1,t) = 0$$

# 2 Numerical Scheme

We use the HLLC method to solve the Riemann Problem. Before our method, we first introduce the HLL Method. The Godunov Scheme can be expressed as:

$$\mathbf{Q}_{i}^{n+1} = \mathbf{Q}_{i} - \frac{\Delta t}{\Delta x} (\mathbf{F}_{i+1/2} - \mathbf{F}_{i-1/2})$$

Where

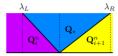
$$\mathbf{Q}_{i}^{n} = \frac{1}{\Delta x} \int_{x_{i-1/2}}^{x_{i+1/2}} \mathbf{q}(x, t_{n}) \, \mathrm{d}x$$

# 2.1 HLL Flux

Consider a Riemann Problem:

$$\begin{split} &\frac{\partial \mathbf{q}}{\partial t} + \frac{\partial f(\mathbf{q})}{\partial x} = 0 \\ &\mathbf{q}(x,0) = \left\{ \begin{array}{ll} \mathbf{Q}_i^n, & x < 0 \\ \mathbf{Q}_{i+1}^n, & x > 0 \end{array} \right. \end{split}$$

We assume that the region in the center and both sides are connected by waves with velocities: $\lambda_L$  and  $\lambda_R$ :



Where:

$$\lambda_L = \min \{\lambda(\mathbf{Q}_i^n), \lambda(\mathbf{Q}_{i+1}^n)\}, \lambda_R = \max \{\lambda(\mathbf{Q}_i^n), \lambda(\mathbf{Q}_{i+1}^n)\}$$

Obviously: $\lambda_L < 0, \lambda_R > 0$ . So we have  $\mathbf{F}_{i+1/2}^{HLL} = \mathbf{f}(\mathbf{Q}_{i+1}^n)$  and  $\mathbf{F}_{i+1/2}^{HLL} = \mathbf{f}(\mathbf{Q}_i^n)$ . So we only need to consider  $\lambda_L < 0 < \lambda_R$ . Consider the intergral form of the PDE:

$$\int_{x_L}^{x_R} \mathbf{q}(x,t) \, dx = \int_{x_L}^{x_R} \mathbf{q}(x,0) \, dx + \int_0^T \mathbf{F}(\mathbf{q}(x_L,t)) \, dt - \int_0^T \mathbf{F}(\mathbf{q}(x_R,t)) \, dt$$

By direct calculation:

$$\int_{x_{I}}^{x_{R}} \mathbf{q}(x,t) dx = x_{R} \mathbf{Q}_{i+1}^{n} - x_{R} \mathbf{Q}_{i}^{n} + T(\mathbf{f}(\mathbf{Q}_{i}^{n}) - \mathbf{f}(\mathbf{Q}_{i+1}^{n}))$$

On the other hand, we have:

$$\begin{split} \int_{x_L}^{x_R} \mathbf{q}(x,t) \, dx = & \int_{x_L}^{T\lambda_L} \mathbf{q}(x,t) \, dx + \int_{T\lambda_L}^{T\lambda_R} \mathbf{q}(x,t) \, dx + \int_{T\lambda_R}^{x_R} \mathbf{q}(x,t) \, dx \\ = & (T\lambda_L - x_L) \mathbf{Q}_i^n + \int_{T\lambda_L}^{T\lambda_R} \mathbf{q}(x,t) \, dx + (T\lambda_R - x_R) \mathbf{Q}_{i+1}^n \end{split}$$

Combining two equations, we get:

$$\int_{T\lambda L}^{T\lambda_R} \mathbf{q}(x,t) \, dx = \frac{\lambda_R \mathbf{Q}_{i+1}^n - \lambda_L \mathbf{Q}_i^n + \mathbf{f}(\mathbf{Q}_i^n) - \mathbf{f}(\mathbf{Q}_{i+1}^n)}{\lambda_R - \lambda_L}$$

As  $T \to 0^+$ , we get the approximated estimator  $\mathbf{Q}^*$ :

$$\mathbf{Q}^* = \frac{\lambda_R \mathbf{Q}_{i+1}^n - \lambda_L \mathbf{Q}_i^n + \mathbf{f}(\mathbf{Q}_i^n) - \mathbf{f}(\mathbf{Q}_{i+1}^n)}{\lambda_R - \lambda_L}$$

Then:

$$\tilde{\mathbf{Q}}(x,t) = \left\{ \begin{array}{ll} \mathbf{Q}_i^n, & x < \lambda_L t \\ \mathbf{Q}_i^*, & \lambda_L t < x < \lambda_R t \\ \mathbf{Q}_{i+1}^n, & x > \lambda_R t \end{array} \right.$$

Let **F**\* denotes the flux in star region, by Rankine-Hugoniot condition:

$$\mathbf{F}^* = \mathbf{f}(\mathbf{Q}_i^n) + \lambda_L(\mathbf{Q}^* - \mathbf{Q}_i^n)$$
  
$$\mathbf{F}^* = \mathbf{f}(\mathbf{Q}_{i+1}^n) + \lambda_R(\mathbf{Q}^* - \mathbf{Q}_{i+1}^n)$$

Then we solved:

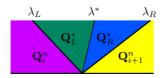
$$\mathbf{F}^* = \frac{\lambda_R \mathbf{f}(\mathbf{Q}_i^n) - \lambda_L \mathbf{f}(\mathbf{Q}_{i+1}^n) + \lambda_L \lambda_R (\mathbf{Q}_{i+1}^n - \mathbf{Q}_i^n)}{\lambda_R - \lambda_L}$$

Then the flux of HLL can be defined as:

$$\mathbf{F}_{i+1/2} = \begin{cases} \mathbf{f}(\mathbf{Q}_i^n), & \lambda_L \geqslant 0 \\ \mathbf{F}^*, & \lambda_L < 0 < \lambda_R \\ \mathbf{f}(\mathbf{Q}_{i+1}^n), & \lambda_R \leqslant 0 \end{cases}$$

# 2.2 HLLC Solver

Consider a three-wave approximate Riemann solution:



By the Rankine-Hugoniot condition:

$$\mathbf{F}_{L}^{*} - \mathbf{f}(\mathbf{Q}_{i}^{n}) = \lambda_{L}(\mathbf{Q}_{L}^{*} - \mathbf{Q}_{i}^{n})$$

$$\mathbf{F}_{R}^{*} - \mathbf{F}_{L}^{*} = \lambda^{*}(\mathbf{Q}_{R}^{*} - \mathbf{Q}_{L}^{*})$$

$$\mathbf{f}(\mathbf{Q}_{i+1}^{n}) - \mathbf{F}_{R}^{*} = \lambda_{R}(\mathbf{Q}_{i+1}^{n} - \mathbf{Q}_{R}^{*})$$

For Euler's equation, by the property of the equation itself, the velocity and pressure are actually continuous:

$$u_L^* = u_R^* = u^*, p_L^* = p_R^* = p^*$$

Meanwhile, at the discontinuity:

$$\lambda^* = u^*$$

Cancelling  $\mathbf{F}_{L}^{*}$  and  $\mathbf{F}_{R}^{*}$ , we get:

$$\begin{split} \mathbf{f}(\mathbf{Q}_{i+1}^{n}) - \mathbf{f}(\mathbf{Q}_{i}^{n}) - \lambda^{*}(\mathbf{Q}_{k}^{*} - \mathbf{Q}_{L}^{*}) = & \lambda_{L}(\mathbf{Q}_{L}^{*} - \mathbf{Q}_{i}^{n}) + \lambda_{R}(\mathbf{Q}_{i+1}^{n} - \mathbf{Q}_{R}^{*}) \\ (\lambda_{R} - \lambda^{*})\mathbf{Q}_{k}^{*} + (\lambda^{*} - \lambda_{L})\mathbf{Q}_{L}^{*} = & \lambda_{R}\mathbf{Q}_{i+1}^{n} - \lambda_{L}\mathbf{Q}_{i}^{n} - [\mathbf{f}(\mathbf{Q}_{i+1}^{n}) - \mathbf{f}(\mathbf{Q}_{i}^{n})] \end{split}$$

The first two equations are:

$$\begin{split} (\lambda_R - \lambda^*) \rho_R^* + (\lambda^* - \lambda_L) \rho_L^* = & \lambda_R \rho_{i+1}^n - \lambda_L \rho_i^n - (\rho_{i+1}^n u_{i+1}^n - \rho_i^n u_i^n) \\ (\lambda_R - \lambda^*) \rho_R^* \lambda^* + (\lambda^* - \lambda_L) \rho_L^* \lambda^* = & \lambda_R \rho_{i+1}^n u_{i+1}^n - \lambda_L \rho_i^n u_i^n - \\ & [\rho_{i+1}^n (u_{i+1}^n)^2 + p_{i+1}^n - \rho_i^n (u_i^n)^2 - p_i^n] \end{split}$$

We get:

$$\lambda^* \!=\! \! \frac{\lambda_R \rho_{i+1}^n u_{i+1}^n - \lambda_L \rho_{i}^n u_{i}^n - [\rho_{i+1}^n (u_{i+1}^n)^2 + p_{i+1}^n - \rho_{i}^n (u_{i}^n)^2 - p_{i}^n]}{\lambda_R \rho_{i+1}^n - \lambda_L \rho_{i}^n - (\rho_{i+1}^n u_{i+1}^n - \rho_{i}^n u_{i}^n)}$$

Then:

$$\mathbf{F}_{L}^{*} - \mathbf{f}(\mathbf{Q}_{i}^{n}) = \lambda_{L}(\mathbf{Q}_{L}^{*} - \mathbf{Q}_{i}^{n}) \Rightarrow \rho_{L}^{*} = \frac{\lambda_{L} - u_{i}^{n}}{\lambda_{L} - \lambda^{*}} \rho_{i}^{n}$$
$$\mathbf{f}(\mathbf{Q}_{i+1}^{n}) - \mathbf{F}_{R}^{*} = \lambda_{R}(\mathbf{Q}_{i+1}^{n} - \mathbf{Q}_{R}^{*}) \Rightarrow \rho_{L}^{*} = \frac{\lambda_{R} - u_{i+1}^{n}}{\lambda_{R} - \lambda^{*}} \rho_{i+1}^{n}$$

We solve  $p_L^*$  and  $p_R^*$ :

$$\begin{split} [\rho_L^*(\lambda^*)^2 + p_L^*] - [\rho_i^n(u_i^n)^2 + p_i^n] &= \lambda_L(\rho_L^*\lambda^* - \rho_i^n u_i^n) \\ \Downarrow \\ p_L^* &= \rho_i^n(u_i^n - \lambda^*)(u_i^n - \lambda_L) + p_i^n \\ p_R^* &= \rho_{i+1}^n(u_{i+1}^n - \lambda^*)(u_{i+1}^n - \lambda_L) + p_{i+1}^n \end{split}$$

Similar, we get also get  $e_L^*$  and  $e_R^*$ :

$$\begin{split} e_L^* &= \frac{\lambda_L - u_i^n}{\lambda_L - \lambda^*} e_i^n + \frac{\lambda^* - u_i^n}{\lambda_L - \lambda^*} p_i^n + \rho_i^n \frac{\lambda^* (\lambda^* - u_i^n) (\lambda_L - u_i^n)}{\lambda_L - \lambda^*} \\ e_R^* &= \frac{\lambda_R - u_{i+1}^n}{\lambda_R - \lambda^*} e_{i+1}^n + \frac{\lambda^* - u_{i+1}^n}{\lambda_R - \lambda^*} p_{i+1}^n + \rho_{i+1}^n \frac{\lambda^* (\lambda^* - u_{i+1}^n) (\lambda_R - u_{i+1}^n)}{\lambda_R - \lambda^*} \end{split}$$

Then the scheme can be expressed as follows:

$$\mathbf{Q}_{i}^{n+1} = \mathbf{Q}_{i} - \frac{\Delta t}{\Delta x} (\mathbf{F}_{i+1/2} - \mathbf{F}_{i-1/2})$$

The numerical flux is:

$$\mathbf{F}_{i+1/2} = \begin{cases} \mathbf{f}(\mathbf{Q}_i^n) & \lambda_L \geqslant 0 \\ \mathbf{f}(\mathbf{Q}_i^n) + \lambda_L(\mathbf{Q}_L^* - \mathbf{Q}_i^n) & \lambda_L < 0 < \lambda^* \\ \mathbf{f}(\mathbf{Q}_{i+1}^n) + \lambda_R(\mathbf{Q}_R^* - \mathbf{Q}_{i+1}^n) & \lambda^* \leqslant 0 < \lambda_R \\ \mathbf{f}(\mathbf{Q}_{i+1}^n) & \lambda_R \leqslant 0 \end{cases}$$

Where:

$$\begin{aligned} \mathbf{Q}_L^* = & \left( \begin{array}{ccc} \rho_L^* & \rho_L^* \lambda^* & \rho_L^* v_i^n & e_L^* \end{array} \right)^T \\ \mathbf{Q}_R^* = & \left( \begin{array}{ccc} \rho_R^* & \rho_R^* \lambda^* & \rho_R^* v_{i+1}^n & e_R^* \end{array} \right)^T \end{aligned}$$

# 2.3 Speed Estimate

In 2.2 we set:

$$\lambda_L = \min \{ \lambda(\mathbf{Q}_i^n), \lambda(\mathbf{Q}_{i+1}^n) \}, \lambda_R = \max \{ \lambda(\mathbf{Q}_i^n), \lambda(\mathbf{Q}_{i+1}^n) \}$$

Which is also equal to:

$$\lambda_L = \min \{u_i^n - c_i^n, u_{i+1}^n - c_{i+1}^n\}, \lambda_R = \max \{u_i^n + c_i^n, u_{i+1}^n + c_{i+1}^n\}$$

These estimates make use of data values only, are exceedingly simple but are not recommended for practical computations. There are some other ways to choose speed

#### 2.3.1 Average Eigenvalues

$$\lambda_L = \tilde{u} - \tilde{a}, \lambda_R = \tilde{u} + \tilde{a}$$

Where:

$$\tilde{u} = \frac{\sqrt{\rho_i^n} u_i^n + \sqrt{\rho_{i+1}^n} u_{i+1}^n}{\sqrt{\rho_i^n} + \sqrt{\rho_{i+1}^n}}, \tilde{a} = \sqrt{(\gamma - 1) \bigg(\tilde{H} - \frac{1}{2}\tilde{u}\bigg)}, \tilde{H} = \frac{\sqrt{\rho_i^n} H_i^n + \sqrt{\rho_{i+1}^n} H_{i+1}^n}{\sqrt{\rho_i^n} + \sqrt{\rho_{i+1}^n}}$$

Here H is defined as:

$$H = \frac{E + p}{\rho}$$

Motivated by the Roe eigenvalues, Einfeldt proposed the estimates:

$$\lambda_L = \bar{u} - \bar{d}, \lambda_R = \bar{u} + \bar{d}$$

where:

$$\begin{split} \bar{d} &= \frac{\sqrt{\rho_i^n a_i^n} + \sqrt{\rho_{i+1}^n} a_{i+1}^n}{\sqrt{\rho_i^n} + \sqrt{\rho_{i+1}^n}} + \eta_2 (u_{i+1}^n - u_i^n)^2 \\ \eta_2 &= \frac{1}{2} \frac{\sqrt{\rho_i^n} \sqrt{\rho_{i+1}^n}}{\left(\sqrt{\rho_i^n} + \sqrt{\rho_{i+1}^n}\right)^2} \end{split}$$

# 2.3.2 Pressure-Based Wave Speed Estimates

$$\lambda_L = u_i^n - a_i^n q_i^n, \lambda_R = u_{i+1}^n + a_{i+1}^n q_{i+1}^n$$

Where:

$$q_K = \left\{ \begin{array}{ll} 1, & p^* \leqslant p_K \\ \sqrt{1 + \frac{\gamma + 1}{2\gamma} \left(\frac{p^*}{p_K} - 1\right)}, & p^* > p_K \end{array} \right.$$

Where K = L,  $p_K = p_i^n$ ; K = R,  $p_K = p_{i+1}^n$ 

# 3 Numerical Result

# 3.1 For Riemann Problem

We present the result at t = 1, 2, 3:

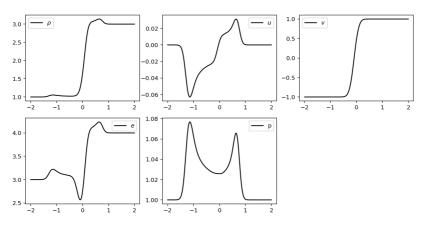


Figure 1. t=1

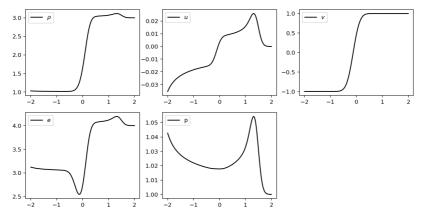
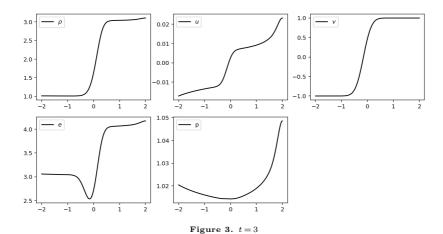


Figure 2. t=2



We test the result under different speed estimate, and find the speed estimate has litte

impact on final result (take t=1 as example:

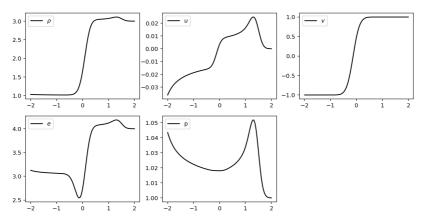


Figure 4. t=2 Under Average Eigenvalues Speed

# 3.2 For Woodward-Collela blast wave Problem

We present the numerical result at t = 0.01, 0.02, 0.03, 0.038

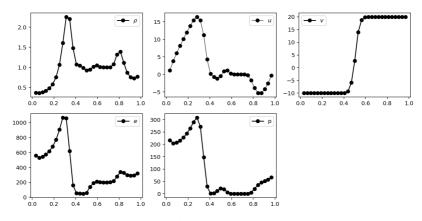


Figure 5. t = 0.01

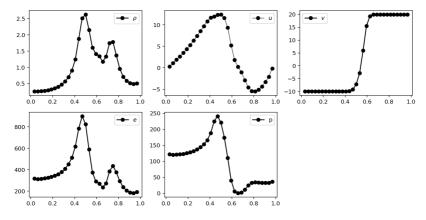


Figure 6. t = 0.02

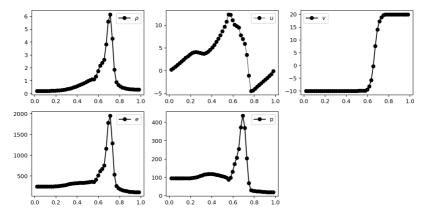
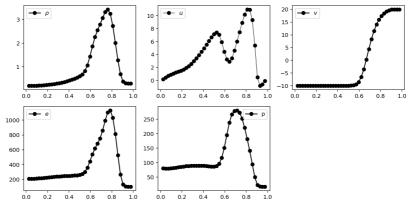


Figure 7. t = 0.03



**Figure 8.** t = 0.038

# **Bibliography**

[1] Toro(2009). Riemann Solvers and Numerical Methods for Fluid Dynamics, a Pratical Introduciton. Springer