

Transient phenomena in foreshock, shock, and magnetosheath – Expectations from large separation campaign

Savvas Raptis^{1,2,3}

¹ APL/JHU, Laurel, MD, USA

² KTH, Stockholm, Sweden

³ ESA/ESTEC, Leiden, The Netherlands

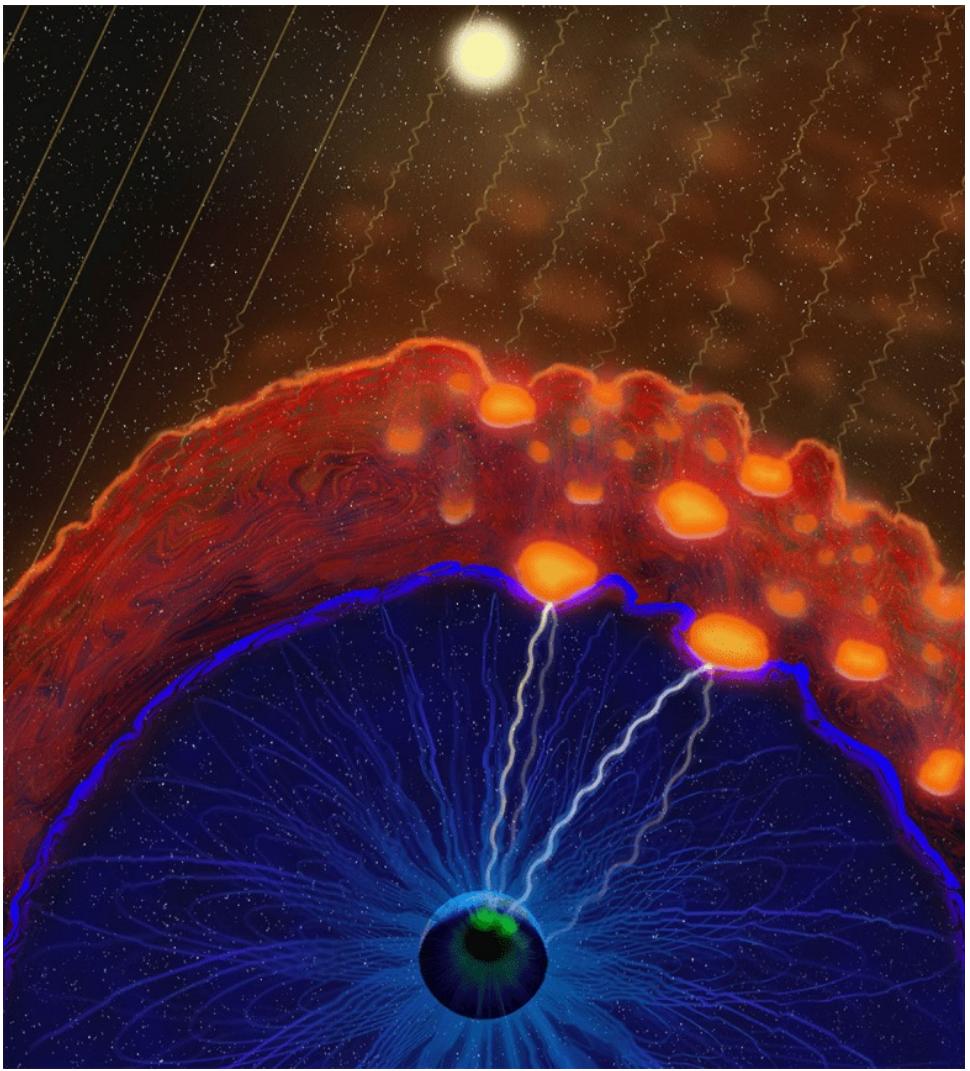
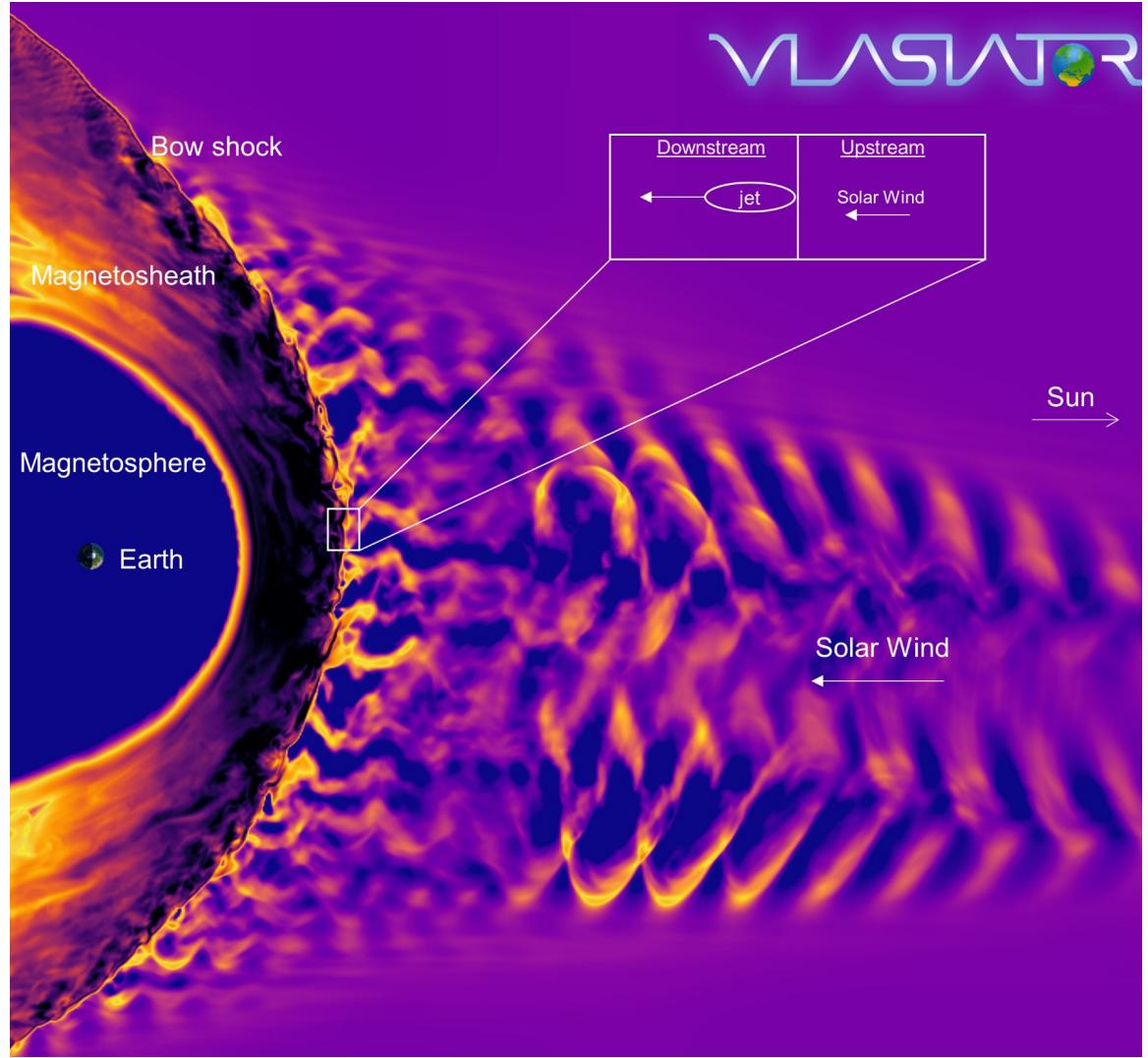
SR acknowledges the support by John Hopkins University Applied Physics Laboratory independent R&D fund

savvas.raptis@jhuapl.edu / <https://savvasraptis.github.io>



CENTER FOR
GEOSPACE STORMS

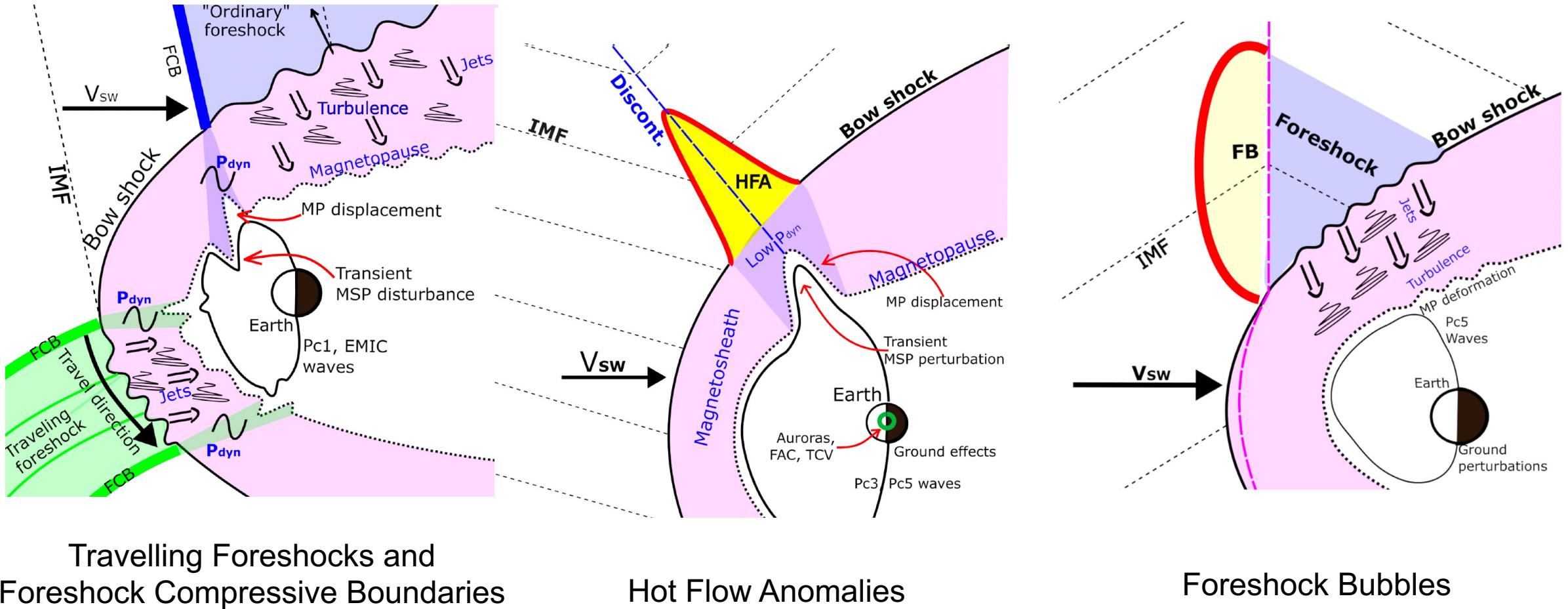
Transient Phenomena in Foreshock & Magnetosheath



Courtesy of M. Palmroth, U Helsinki

Credits: Vuorinen et al. (2022) <https://eos.org/features/space-raindrops-splashing-onearths-magnetic-umbrella>

Foreshock Transients

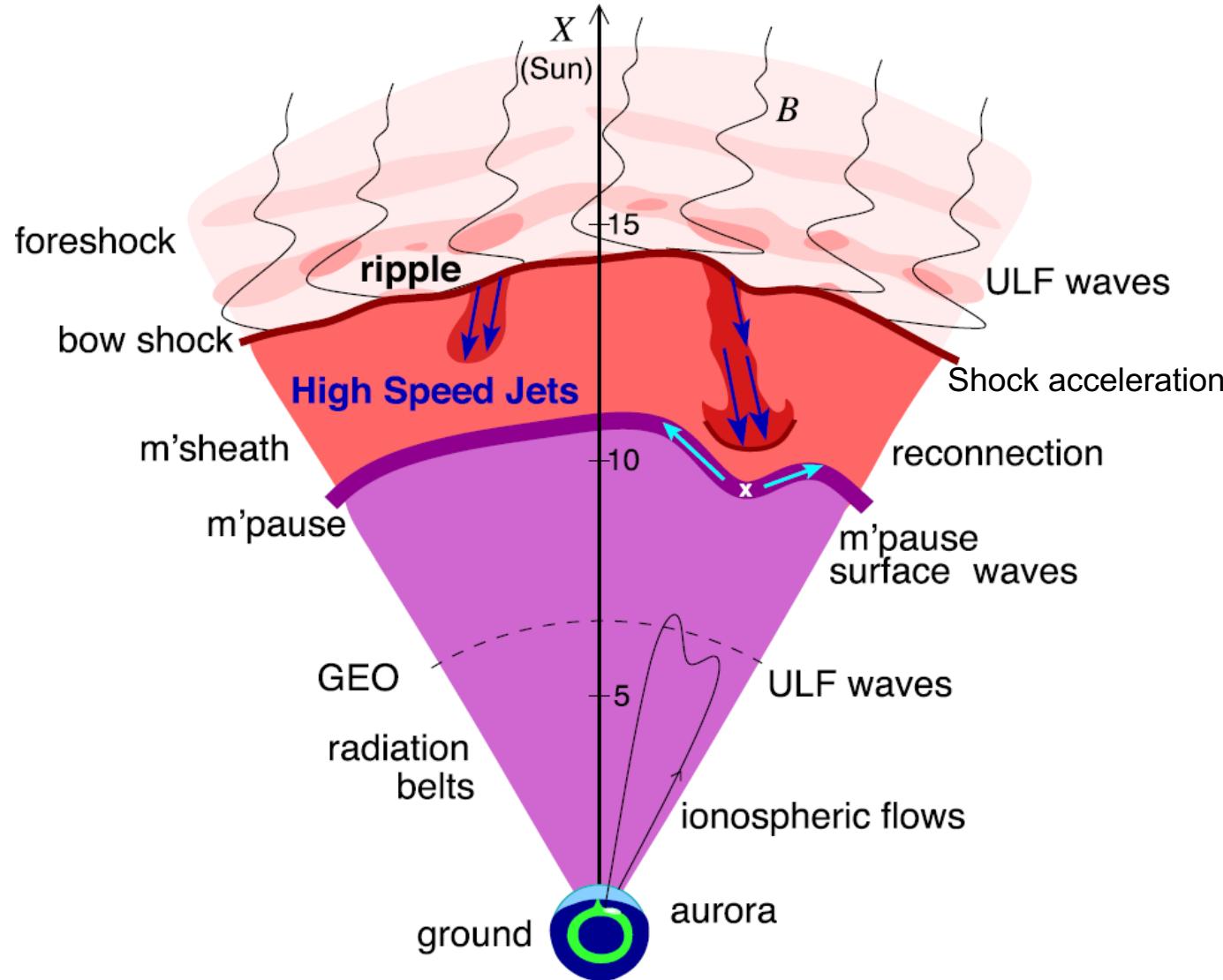


Travelling Foreshocks and
Foreshock Compressive Boundaries

Hot Flow Anomalies

Foreshock Bubbles

Magnetosheath jets



Definition

Magnetosheath jets are **transient localized enhancements** of **dynamic pressure** (density and/or velocity increase)

e.g., 200% dynamic pressure enhancement compared to background magnetosheath

Related phenomena

*Radiation belts
Throat aurora
Magnetopause reconnection
Magnetopause penetration
Shock acceleration
Magnetopause surface eigenmodes
ULF waves
Substorms
Ground magnetometer detection*

Jets – references update (>2019)

~110 citations
Jets Downstream of Collisionless Shocks

Plaschke et al. (2018)

<https://link.springer.com/article/10.1007/s1214-018-0516-3>

- **Excitation** of surface **eigenmodes** at magnetopause: [Archer et al. \(2019, 2021\)](#)
- **Mirror mode waves** and jets: [Bianco-Cano et al. \(2020\)](#)
- **Bursty magnetic reconnection** at the Earth's magnetopause: [Ng et al. \(2021\)](#)
- **Ground-based magnetometer** response: [Norenius et al. \(2021\)](#)
- Generation of **Pi2 pulsations**: [Katsavrias et al. \(2021\)](#)
- **B** in jets, **Bz variations near magnetopause**: [Vuorinen et al. \(2021\)](#)
- High-Speed Jets **Triggering Dayside Ground ULF**: [Wang et al. \(2022\)](#)
- **Waves** and **jets** using burst MMS data: [Krämer et al. \(2023\)](#)

Associated phenomena & effects

Modeling & formation

- **Velocity & magnetic field alignment** in jets: [Plaschke et al. \(2020\)](#)
- **Classification** of jets using MMS & Neural Networks: [Raptis et al. \(2020a,2020b\)](#)
- Comparison **MMS vs simulations**: [Palmroth et al. \(2021\)](#)
- **Solar wind effect** on jet formation: [LaMoury et al. \(2021\)](#)
- Magnetosheath Jets and **Plasmoids** - Hybrid Simulations: [Preisser et al. \(2020\)](#)
- **Formation** of jets in **Quasi-perpendicular magnetosheath**: [Primoz et al. \(2021\)](#)
- **Occurrence** in relation to **CMEs and SIRs**: [Koller et al. \(2022\)](#)
- **Shock reformation** and the **formation of high-speed jets**: [Raptis et al. \(2022a\)](#)
- **Electron acceleration** and **bow waves** in jets: [Vuorinen et al. \(2022\)](#)
- **Kinetic structure** of jets and **partial plasma moments**: [Raptis et al. \(2022b\)](#)

And more : [Liu et al. \(2020a,2020b\)](#), [Omelchenko et al \(2021\)](#), [Sibeck et al. \(2021\)](#), [Suni et al. \(2021\)](#), [Tinoco-Arenas et al. \(2022\)](#) ... etc.

Jets – references update (>2019)

~110 citations

Jets Downstream of Collisionless Shocks

Plaschke et al. (2018)



<https://link.springer.com/article/10.1007/s1214-018-0516-3>

Jets Downstream of Collisionless Shocks: The Last Five Years

Ongoing review (TBD)

Dayside transient community & expectations from MMS

1. ISSI Science International Team 555 “Impact of Upstream Mesoscale Transients on the Near-Earth Environment”
2. The Graz Jet Workshop (Group working on magnetosheath jets)
3. ISSI Science International Team 546 “Magnetohydrodynamic Surface Waves at Earth's Magnetosphere (and Beyond)”

Martin Archer (leader)	Imperial College London, UK
Katariina Nykyri (co-leader)	Embry-Riddle Aeronautical University, USA
Simone Di Matteo	Catholic University of America, USA
Tom Elsden	St Andrews University, UK
Megan Gillies	University of Alberta, USA
Michael Hartinger	Space Science Institute, USA
Anatoly Leonovich	Institute of Solar-Terrestrial Physics, Irkutsk, Russia
Bo Li	Shandong University Weihai, China
Valery Nakariakov	University of Warwick, UK
Vyacheslav Pilipenko	Space Research Institute, Moscow, Russia
Ferdinand Plaschke	Technische Universität Braunschweig, Germany
Xueling Shi	Virginia Tech, USA
Kareem Sorathia	Johns Hopkins University Applied Physics Laboratory, USA
Maria Walach	Lancaster University, UK
Zhonghua Yao	Chinese Academy of Sciences, Beijing, China
Kevin Blasl	Austrian Academy of Sciences, Graz, Austria
Rachel Rice	University of Maryland, USA
Frances Staples	University of California, Los Angeles, USA

3

Primož Kajdič	Geophysics Institute, UNAM, Mexico
Xóchitl Blanco-Cano	Geophysics Institute, UNAM, Mexico
Lucile Turc	University of Helsinki, Finland
Yann Pfau-Kempf	University of Helsinki, Finland
Terry Z. Liu	University of California, Los Angeles, USA
Hui Zhang	Geophysical Institute, University of Alaska Fairbanks, Fairbanks, USA
Boyi Wang	Harbin Institute of Technology (Shenzhen), China
Yufei Hao	KLPS, PMO, Chinese Academy of Sciences, China
Martin O. Archer	Imperial College London, UK
Nojan Omidi	Solana Scientific Inc., USA
Luis Preissler	Space Research Institute, Austrian Academy of Sciences, Graz, Austria
Yu Lin	Physics Department, Auburn University, USA
Savvas Raptis	ESA/ESTEC, Noordwijk, The Netherlands
Adrian LaMoury	Imperial College London, UK
Philippe Escoubet	ESA/ESTEC, Noordwijk, The Netherlands
Marcos Vinicius Dias Silveira	University of São Paulo, EEL, Brasil
Shutao Yao	Shandong University, China
Sun Hee Lee	CUA, NASA GSFC, USA

1

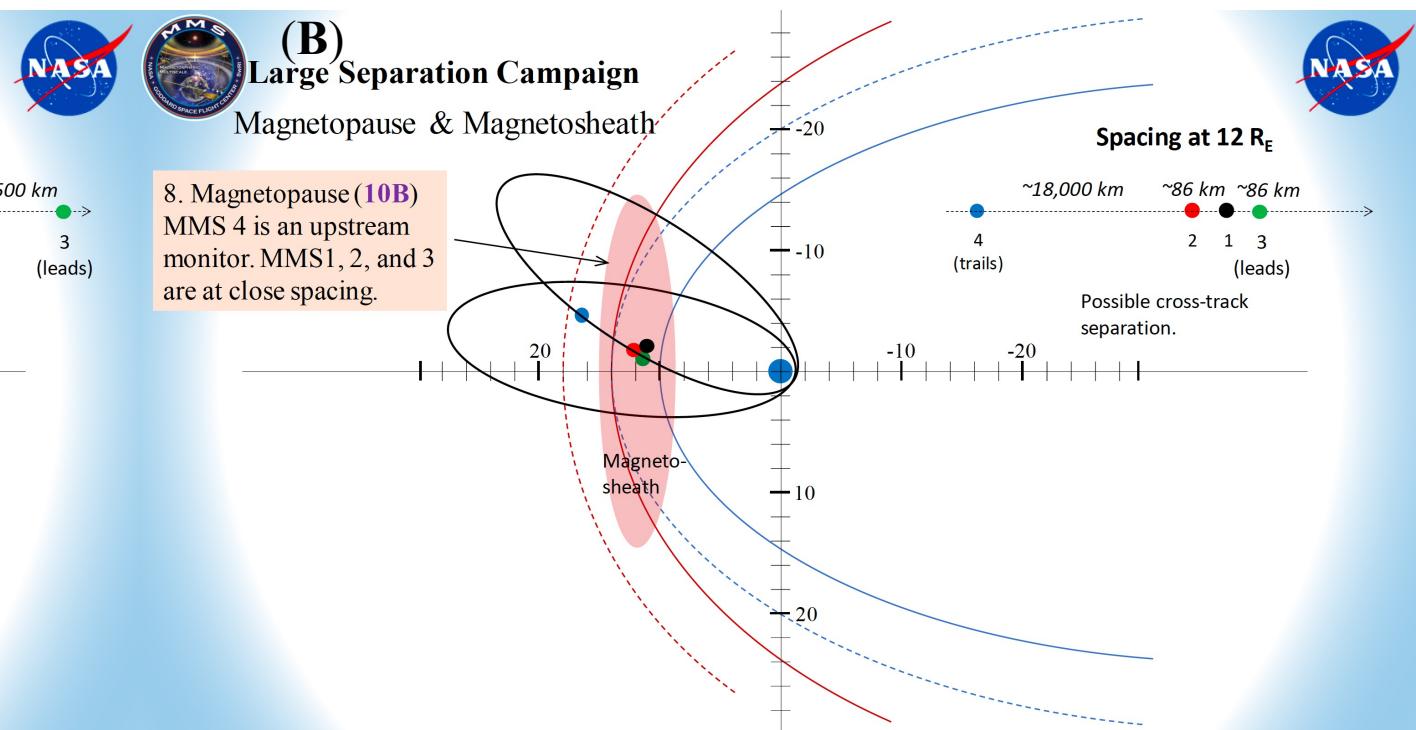
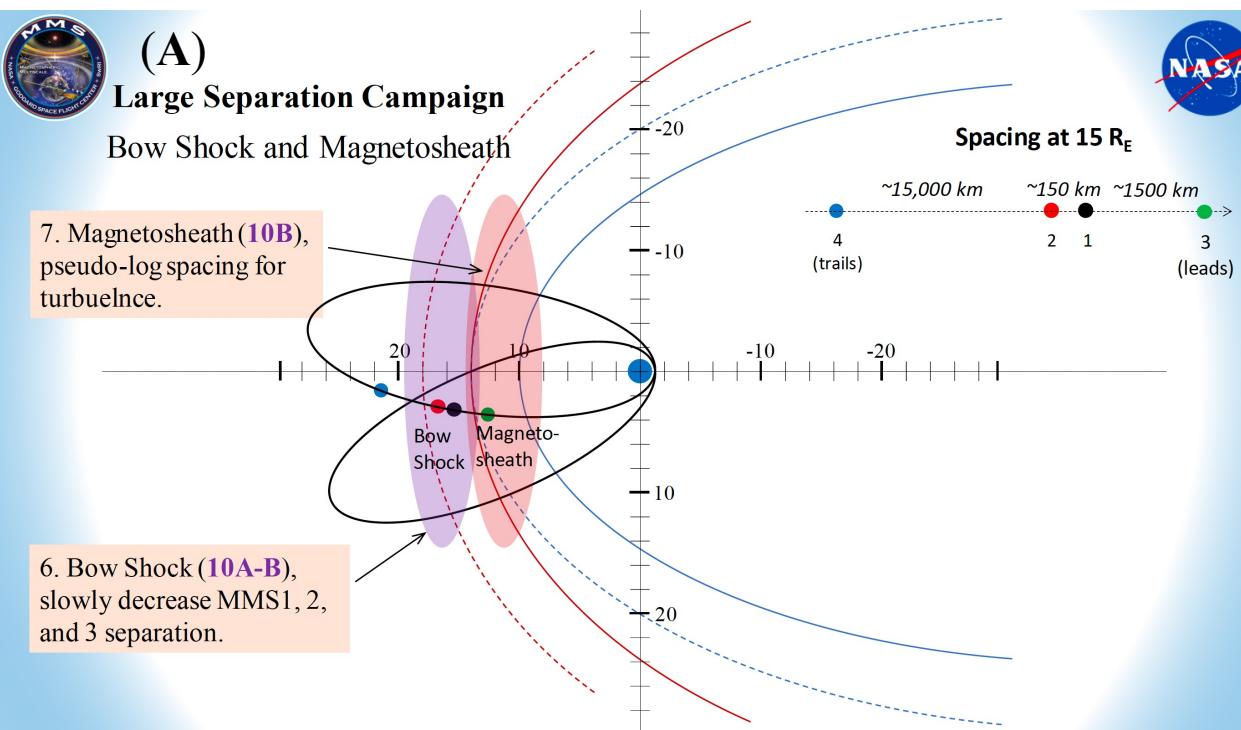
Savvas Raptis	ESA/ESTEC, Noordwijk, The Netherlands
Adrian LaMoury	Imperial College London, UK
Laura Vuorinen	University of Turku, Finland
Herbert Gunell	Umeå University, Sweden
Eva Krämer	Umeå University, Sweden
Niki Xirogiannopoulou	Charles University, Czech Republic
Jonas Suni	University of Helsinki, Finland
Adrian Pöppelwirth	TU Braunschweig, Germany
Cyril Simon Wedlund	Space Research Institute, Austrian Academy of Sciences, Graz, Austria
Manuela Temmer	Institute of Physics, University of Graz, Austria
Luis Preissler	Space Research Institute, Austrian Academy of Sciences, Graz, Austria
Tomas Karlsson	KTH Royal Institute of Technology, Sweden
Florian Koller	Institute of Physics, University of Graz, Austria
Xochitl Blanco-Cano	Instituto de Geofísica, UNAM, México
Heli Hietala	Queen Mary University of London

2

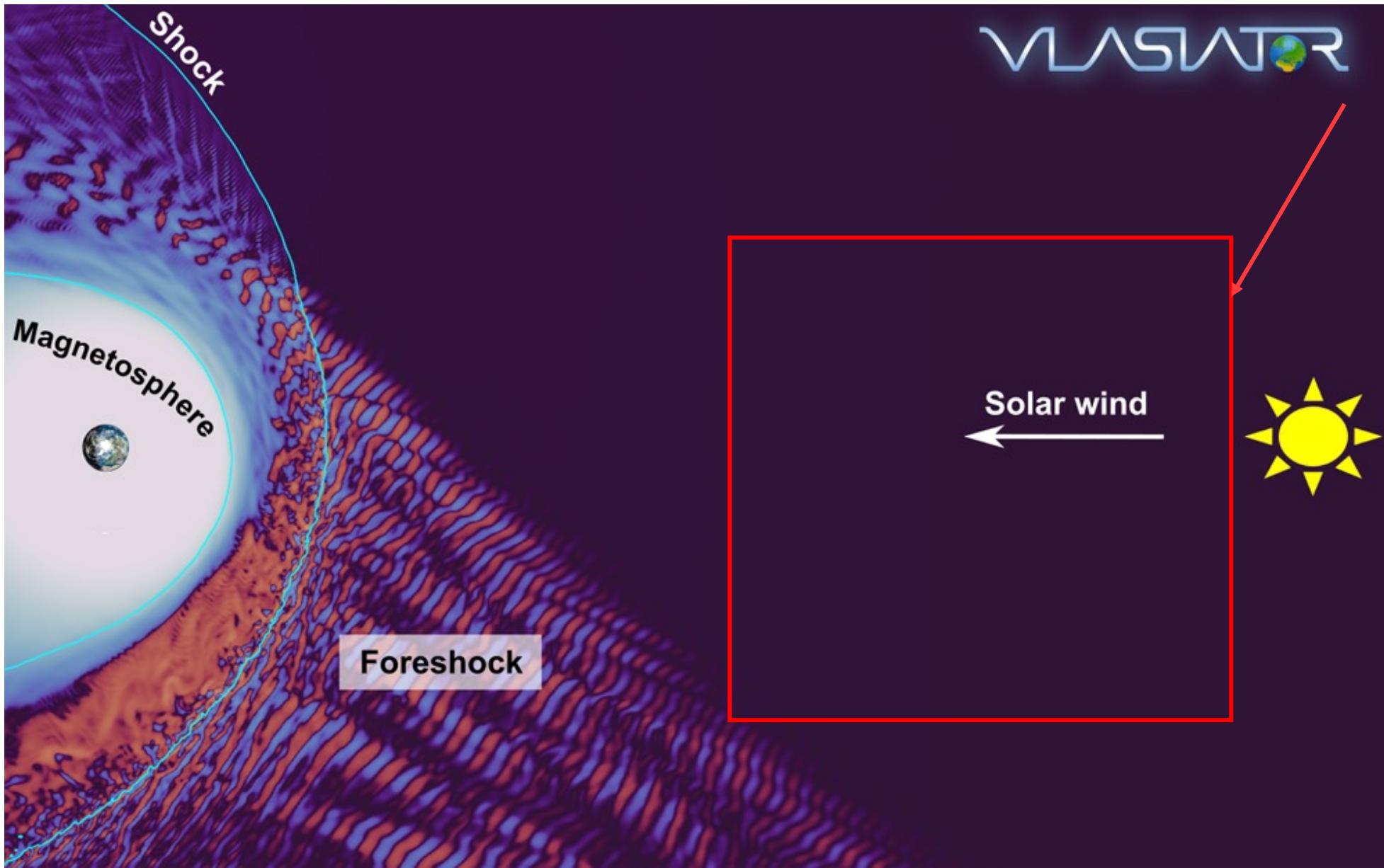
What we wished for?

String-of-pearls campaigns with variable separation:

- (a) 1000s kms to investigate the effect of dayside transients on the magnetosphere
- (b) 1-2 satellites upstream and 1-2 downstream to evaluate the transmission and potential effects for the shock
- (c) Magnetopause response (1000s km separation) while upstream SW/foreshock (several Re)



Why we need that ?
(i.e., test cases)

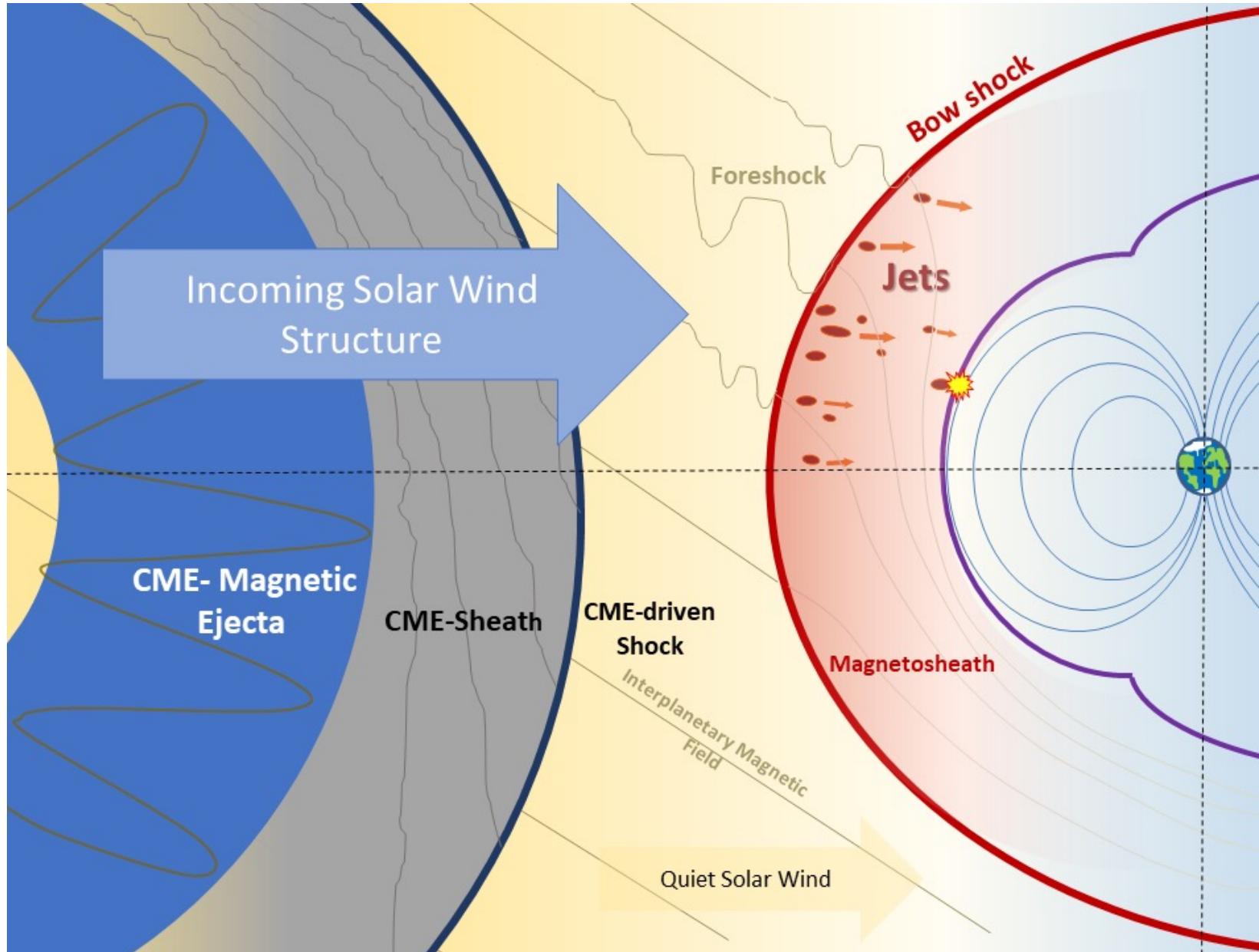


Credits: Lucille Turc / VlaSiator team

10

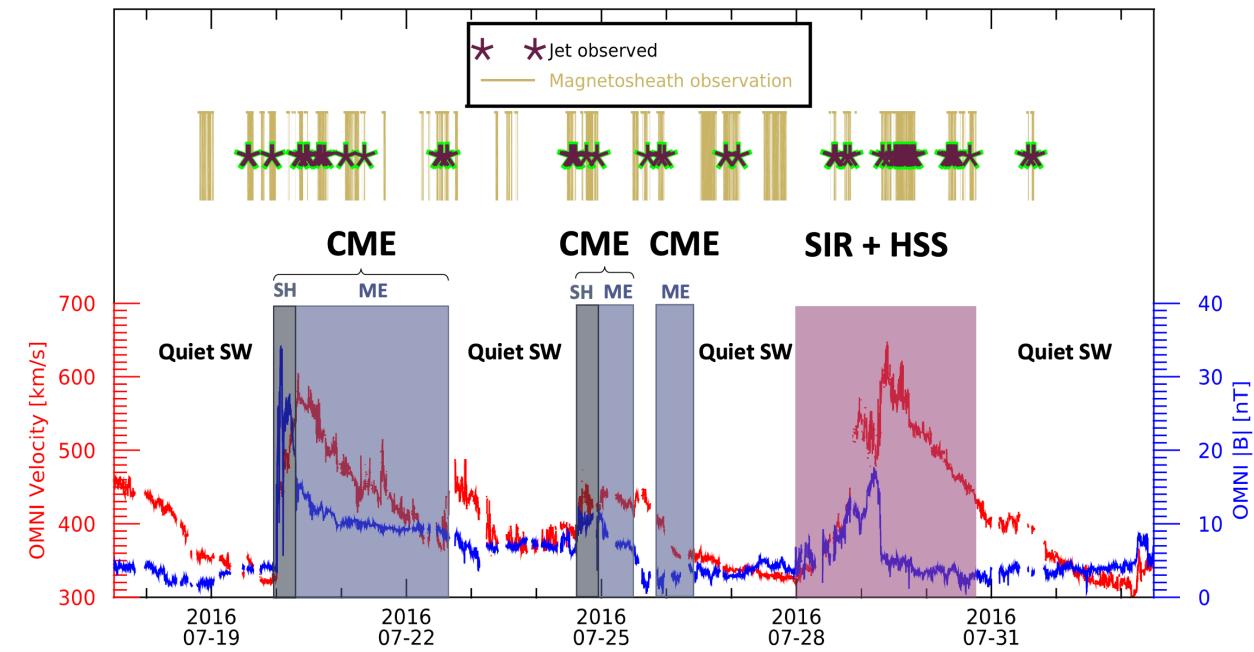
Savvas Raptis – Foreshock & Magnetosheath Transients

MMS SWT Meeting 2023 | 25 Oct 23

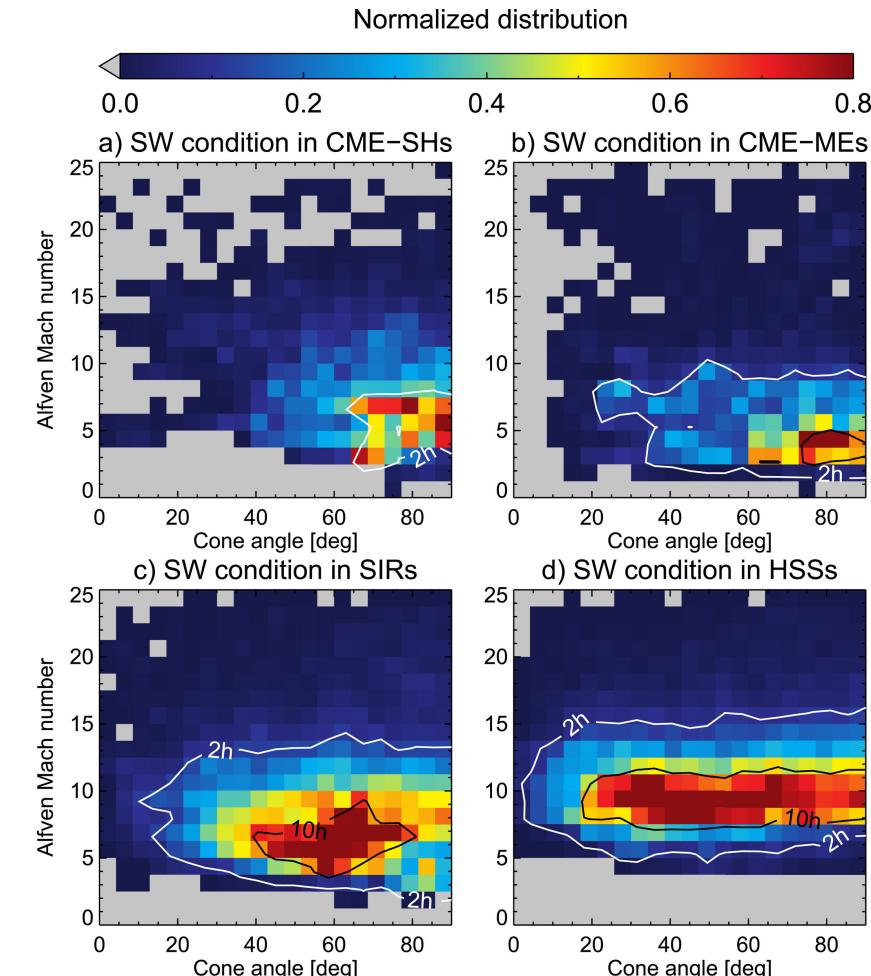


Jet formation – Solar Transients

Magnetic Ejecta ↓ #jets
SIR + HSS ↑ # jets

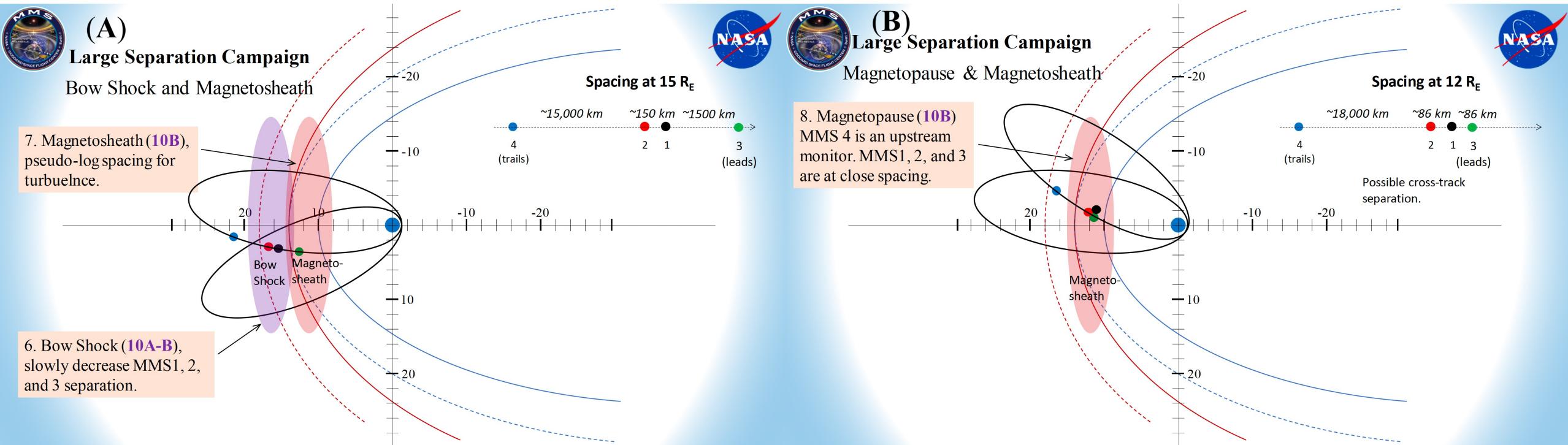


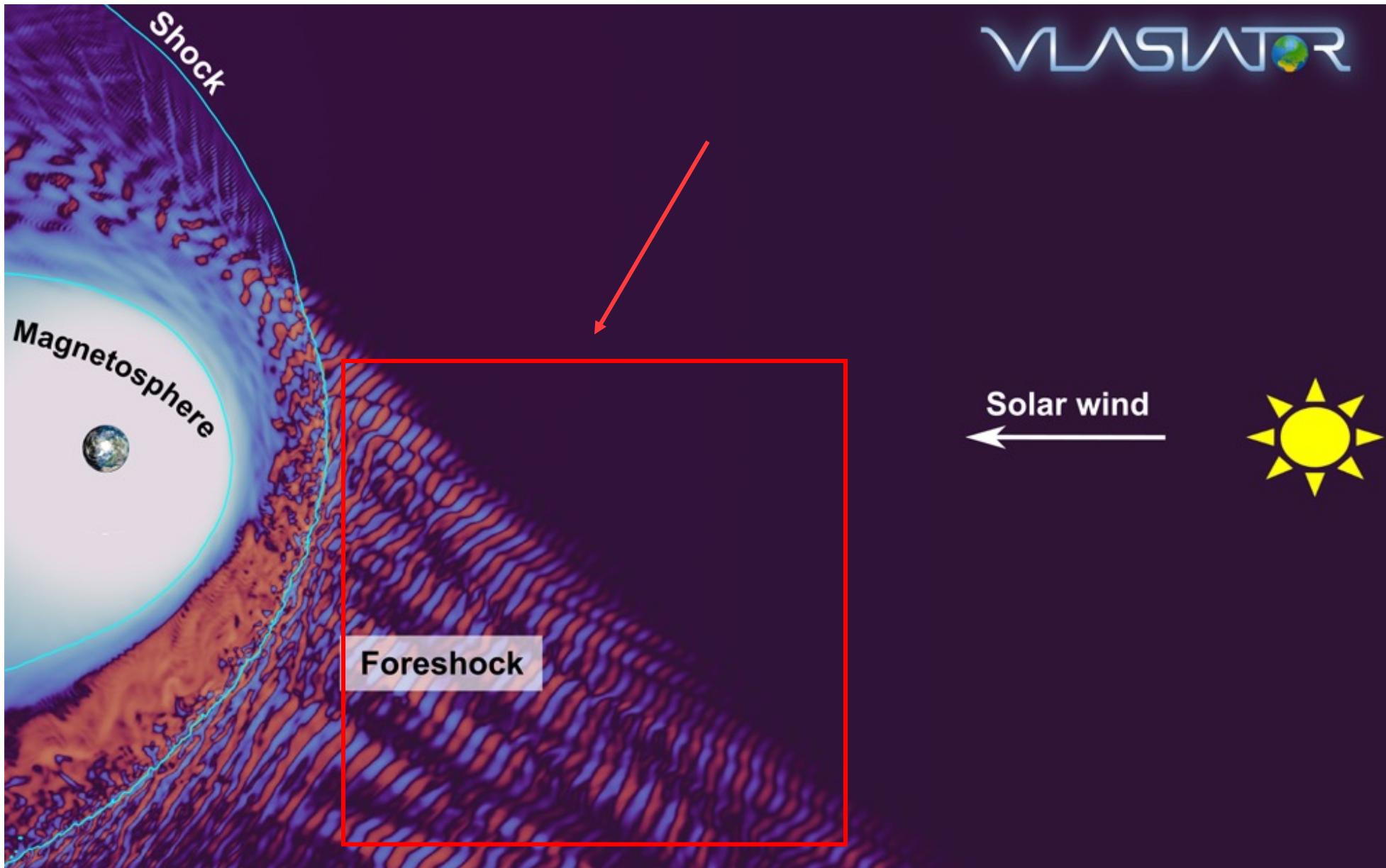
Variable or extreme parameters modify the occurrence and properties of transients in the foreshock / magnetosheath



Open questions on Solar transients

- 1) How do close to shock conditions vary with the presence of SW transient phenomena (magnetic clouds, CMEs, high speed streamers etc.) ? (A)
- 2) How does the formation of foreshock and magnetosheath transient change when there are SW transients? (A)
- 3) What are the exact effects of the solar transients to the magnetopause (B)





Credits: Lucille Turc / Vlasiator team

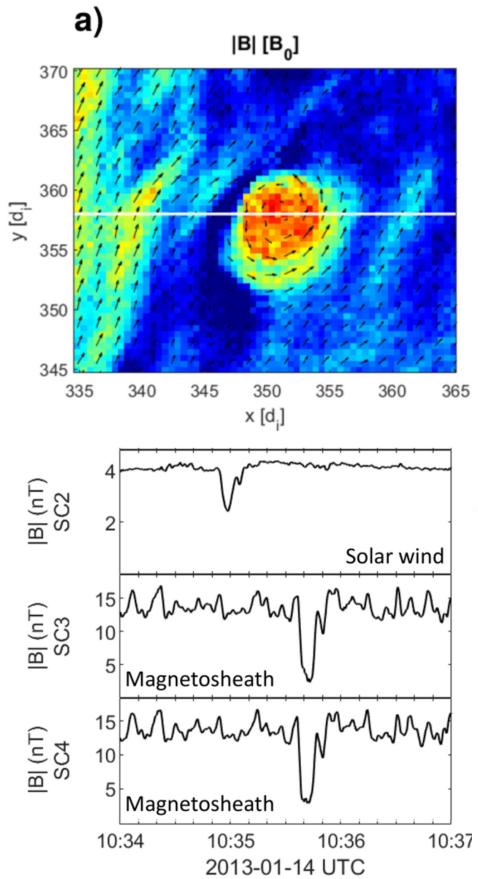
14

Savvas Raptis – Foreshock & Magnetosheath Transients

MMS SWT Meeting 2023 | 25 Oct 23

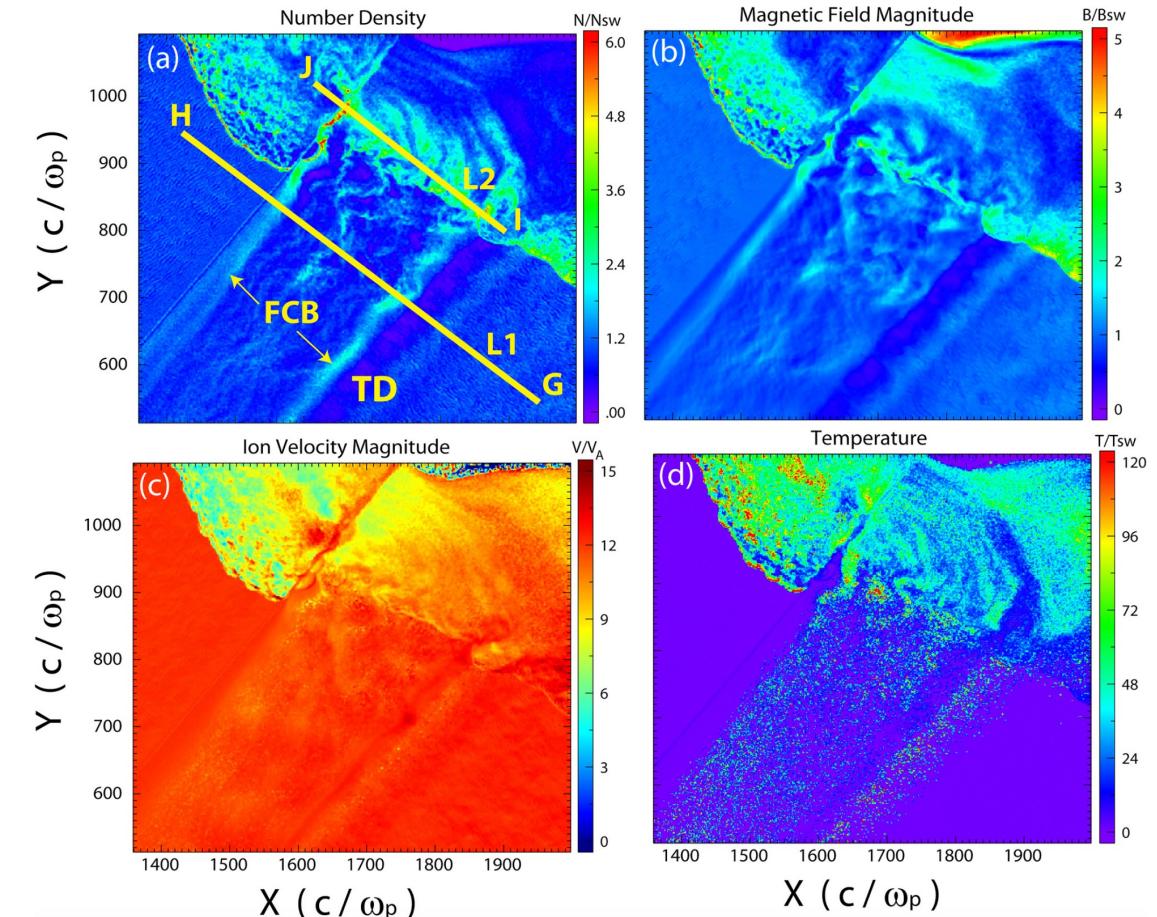
Transmission of Foreshock Transients

ULF waves are transmitted and so are the non-linear associated phenomena (Shocklets, SLAMS, etc.)



Karlsson T., et al. 2022 | ANGEO

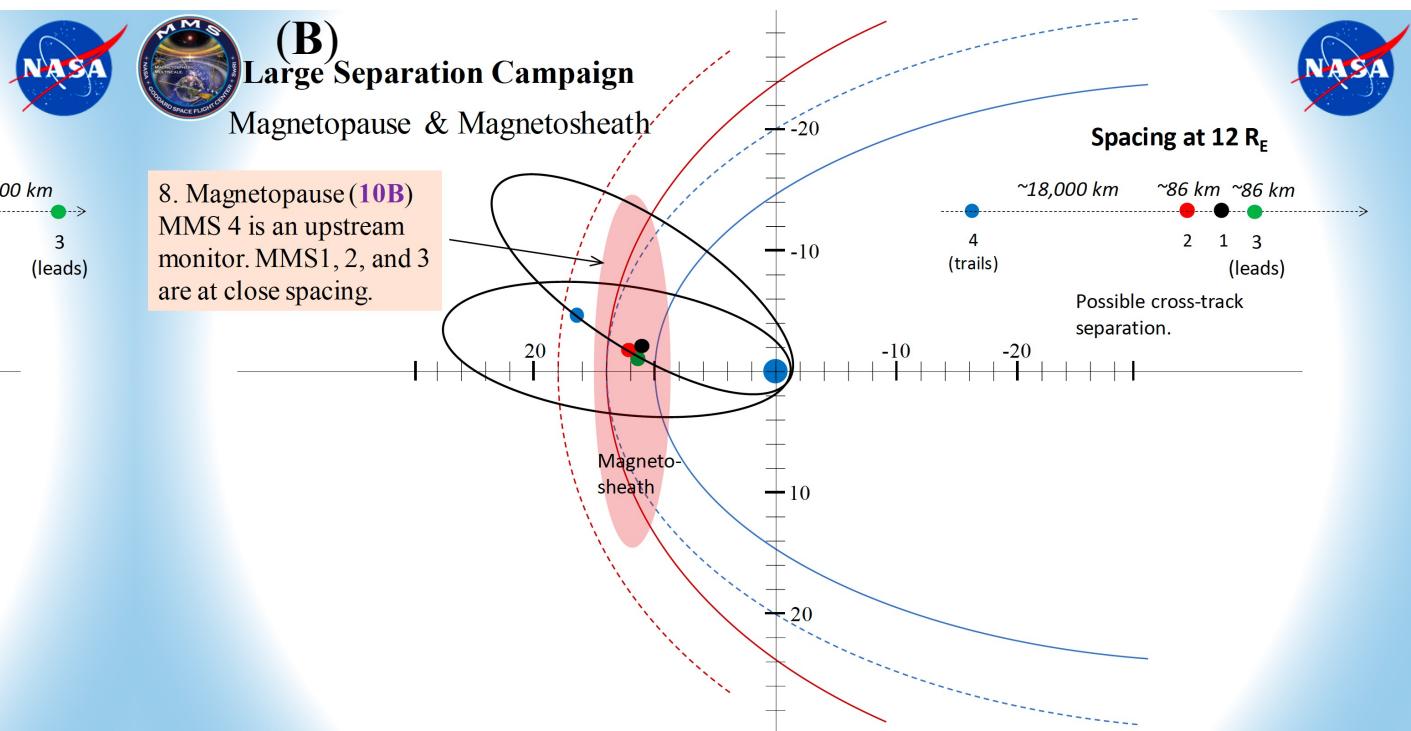
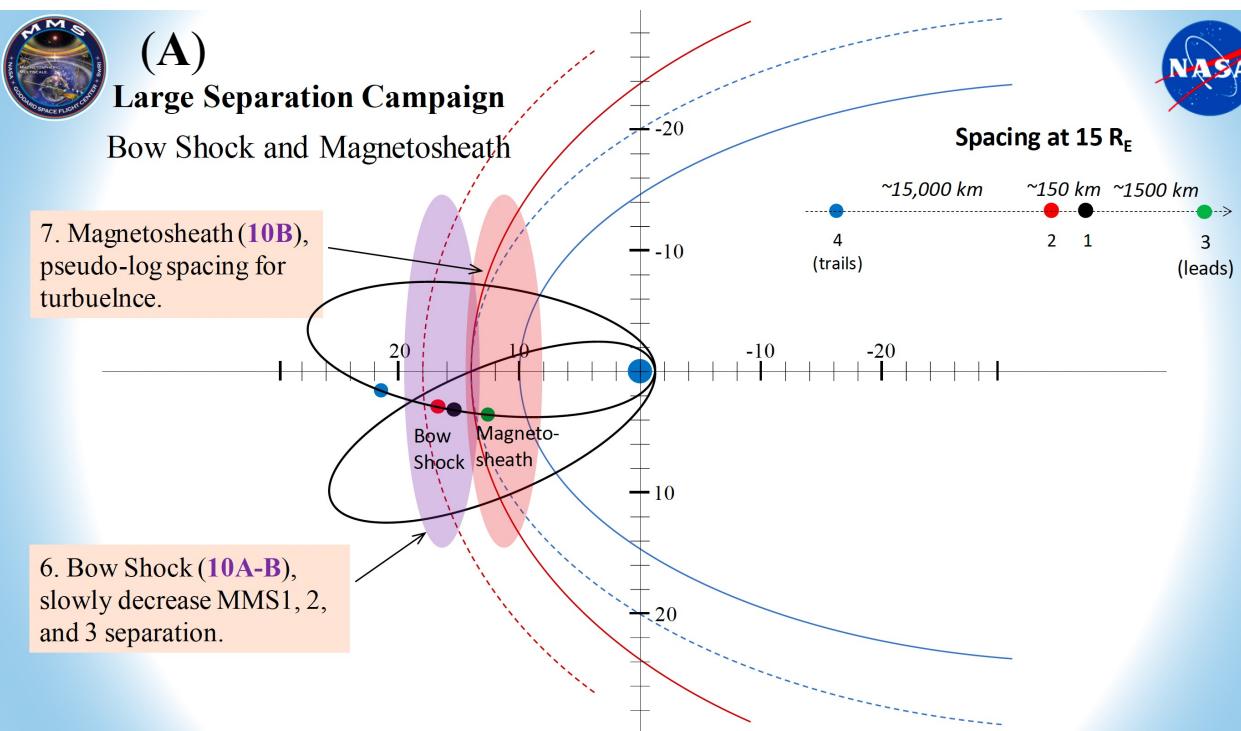
Transmission of FCB, FBs, HFAs, etc. has been shown in simulation and observations.

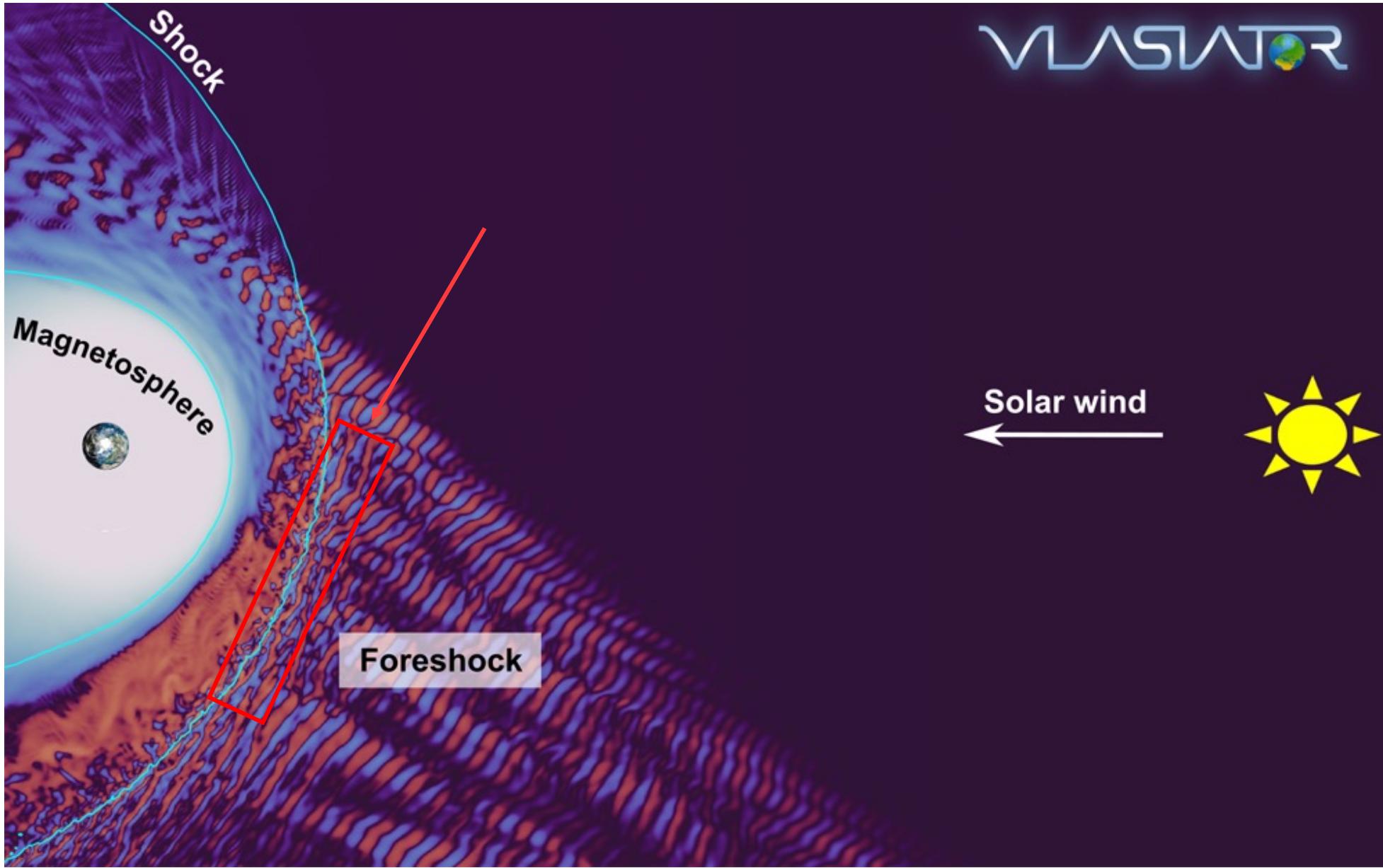


Sibeck D., et al. 2021 | JGR

Open questions on Foreshock transients

1. What are the observational signatures of ULF (and associated phenomena) transmission ? How are the properties changing ? (A)
2. What effects can be caused in the magnetosheath plasma and in the magnetopause through the interaction of these structures with the ambient plasma (A, B)

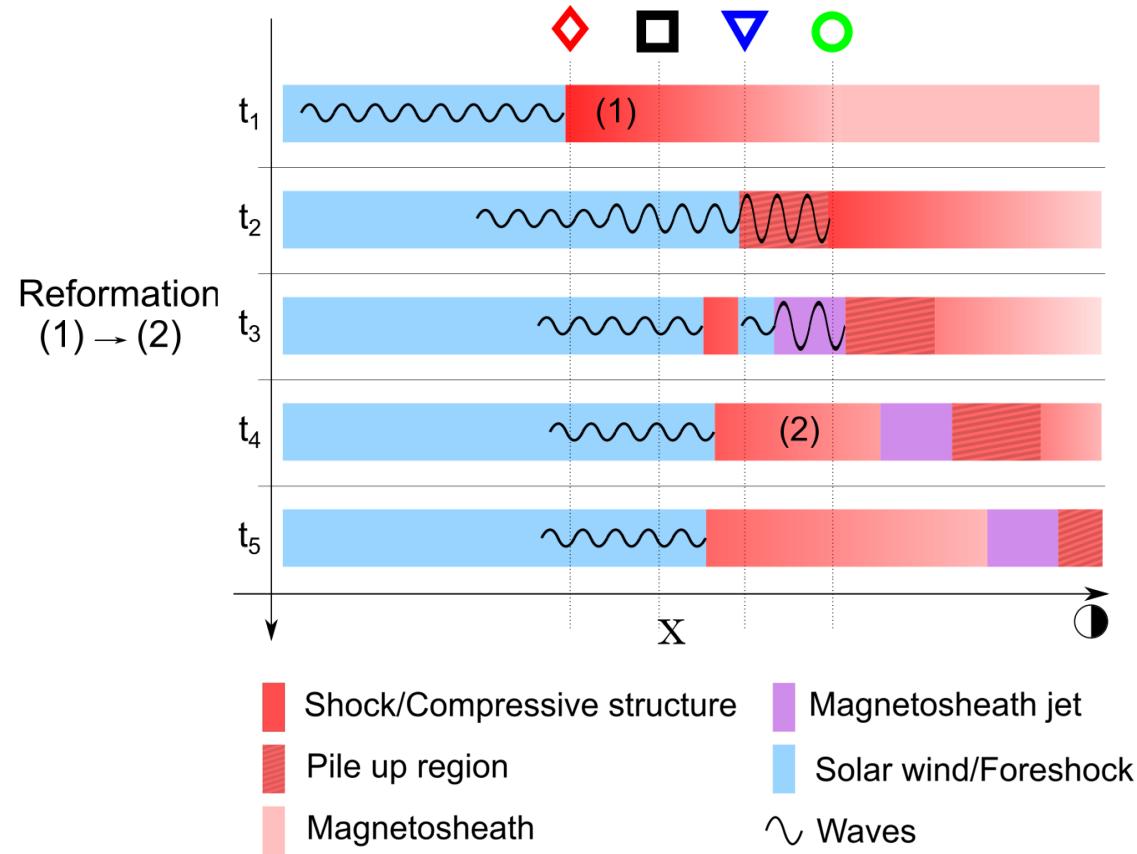




Vlasiator

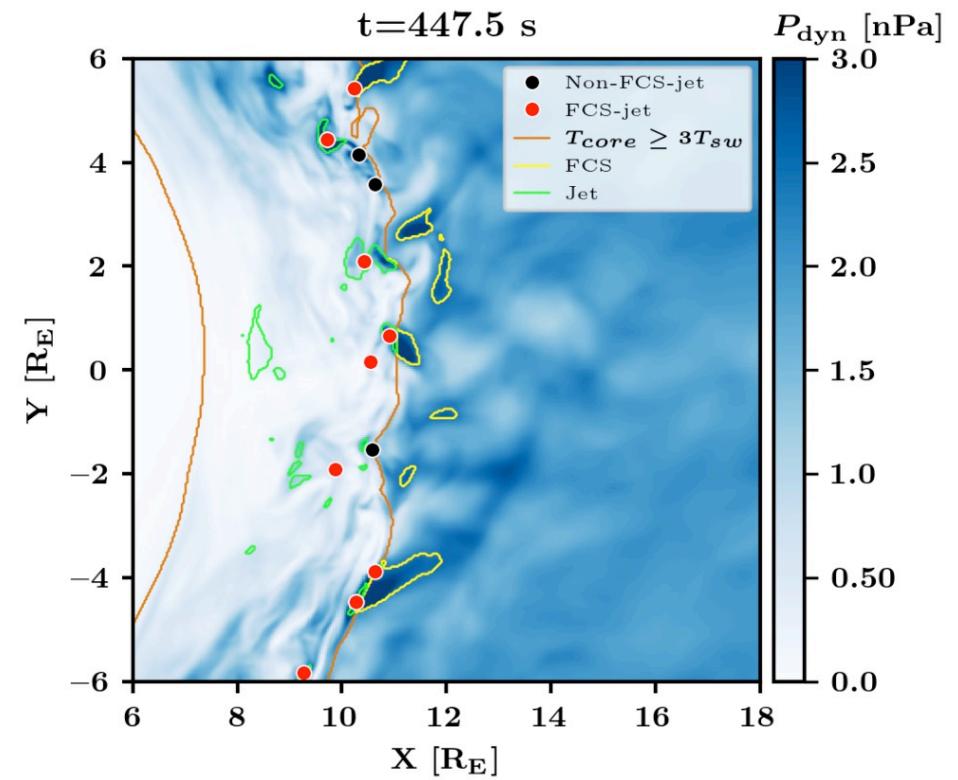
Jet formation

Jets forming from shock's fundamental non-stationarity



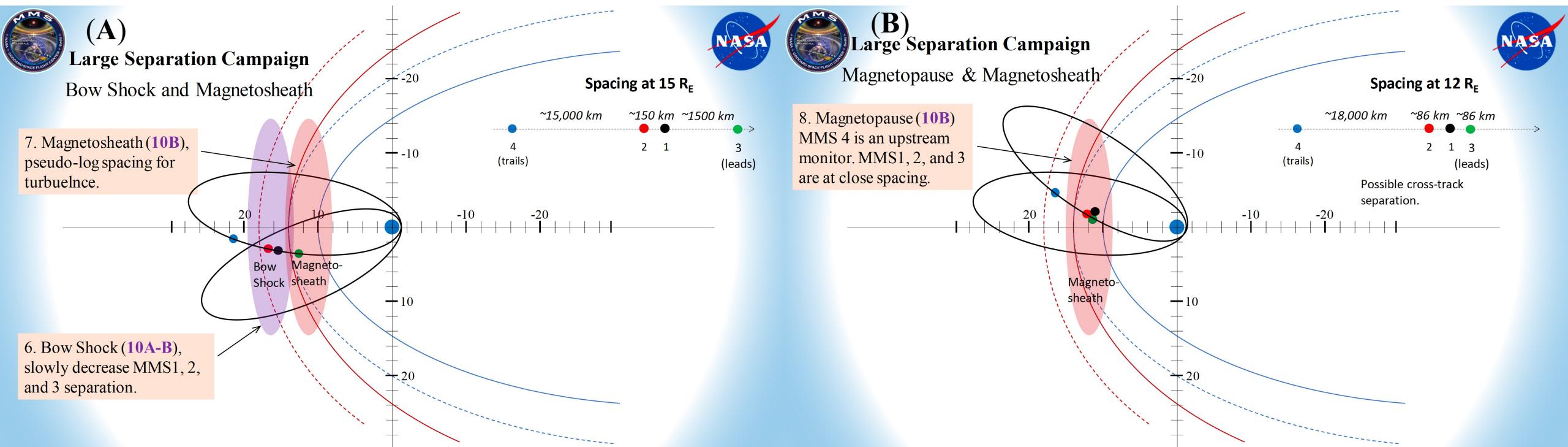
Turbulence Campaign 2019

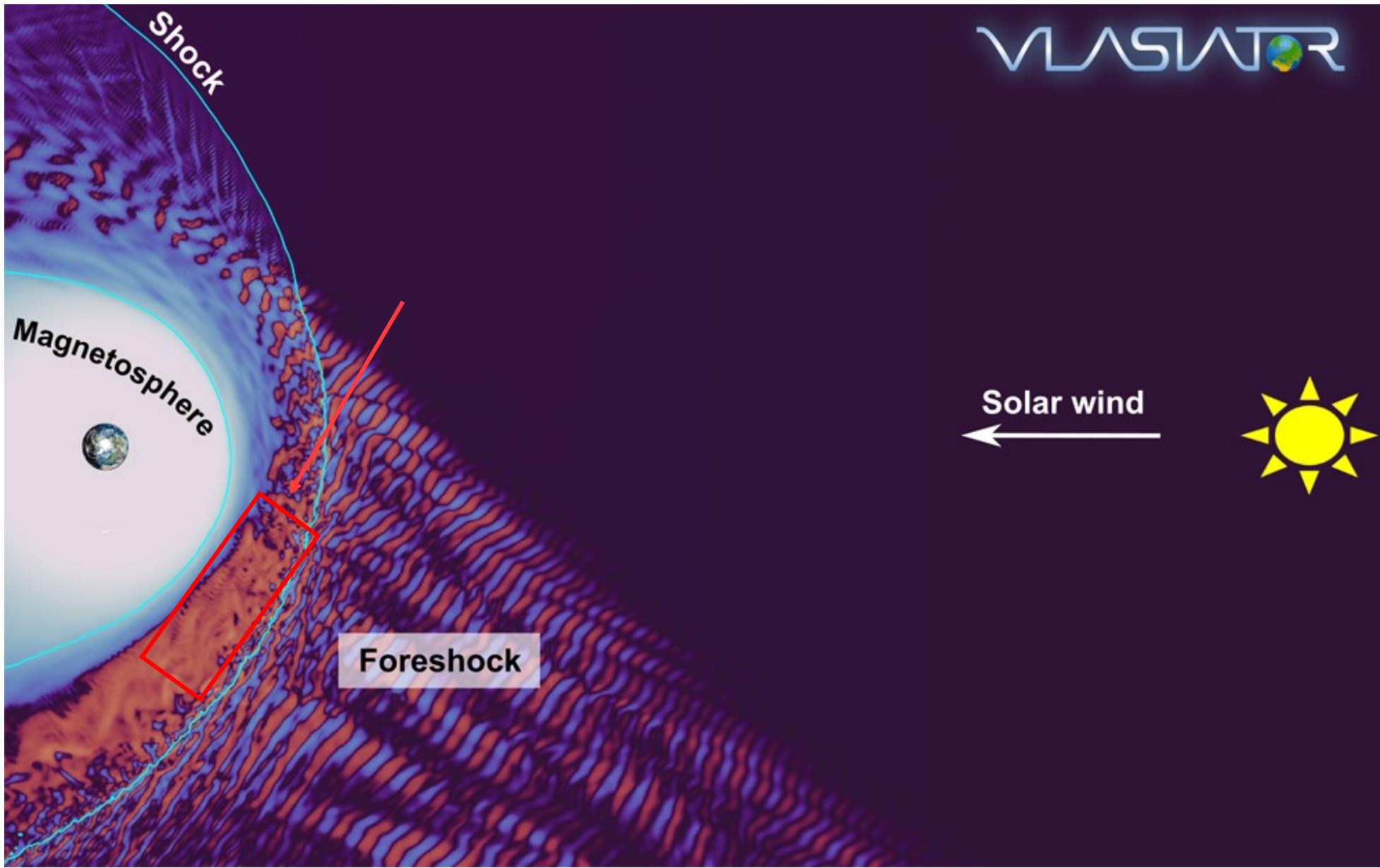
75% of jets caused by Foresight Compressive Structures



Open questions on formation

- 1) Is global shock reformation the answer to jets ? or maybe faster flows get dissipated and become part of the background population before reaching the MP ? (A)
- 2) How are close to the shock upstream conditions affecting the formation of downstream jets ? (A, B)





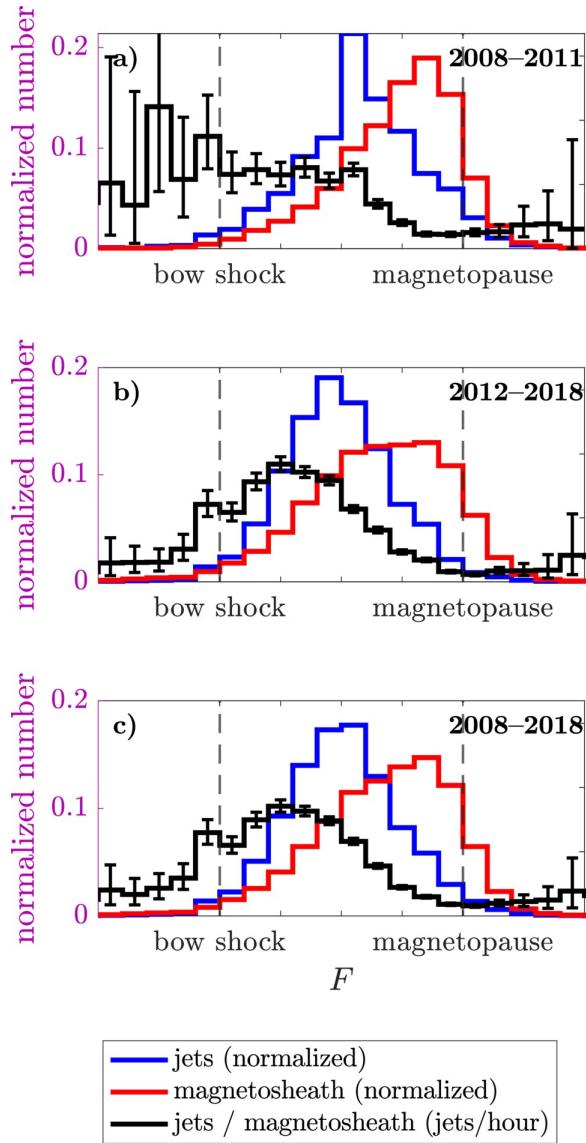
Credits: Lucille Turc / Vlasiator team

20

Savvas Raptis – Foreshock & Magnetosheath Transients

MMS SWT Meeting 2023 | 25 Oct 23

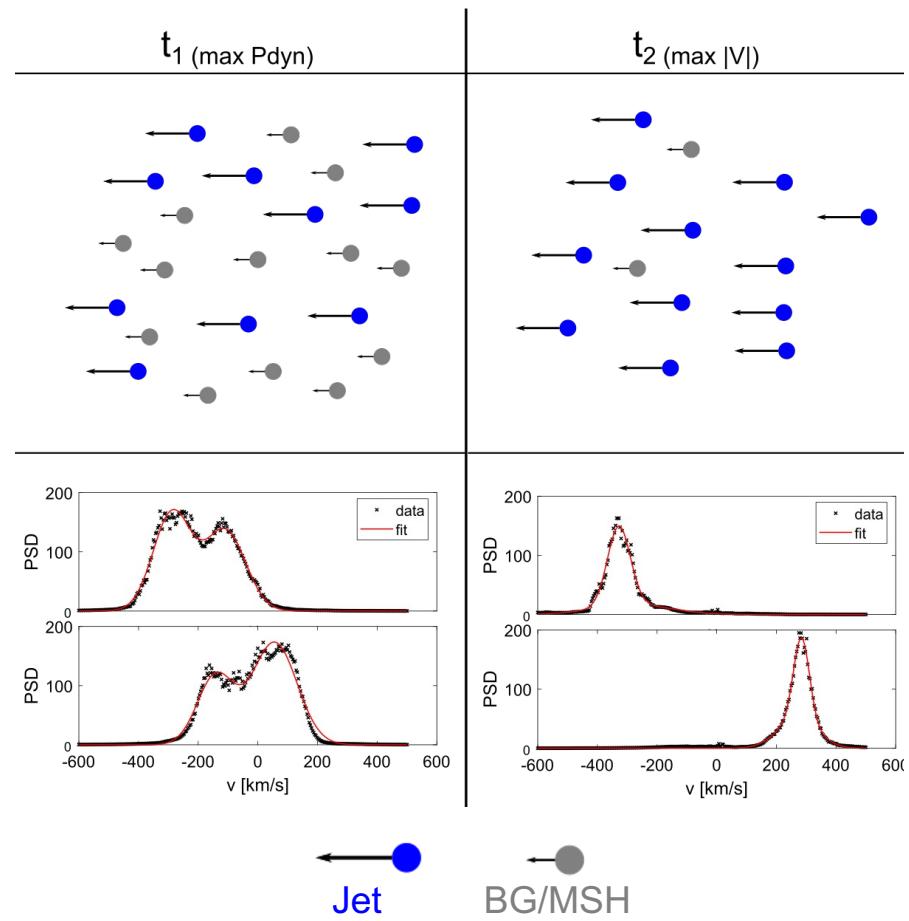
Jet evolution



When can a jet reach the MP ?

- 8 times more likely for high solar wind speed
- 17 more likely for low cone angles

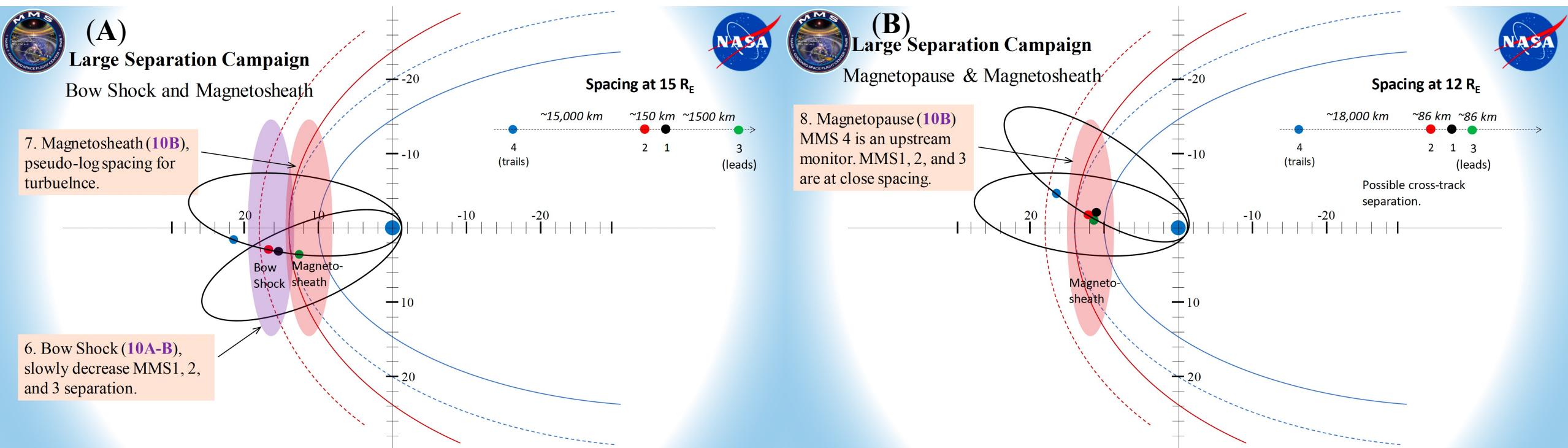
Jets exhibiting 2+ populations = partial moments needed
= similar properties to upstream SW



Jets **thermalize** towards the magnetopause, the shape of the jets **flatten** and they are able to **maintain their flow velocity** and direction within the magnetosheath flow

Open questions on Evolution

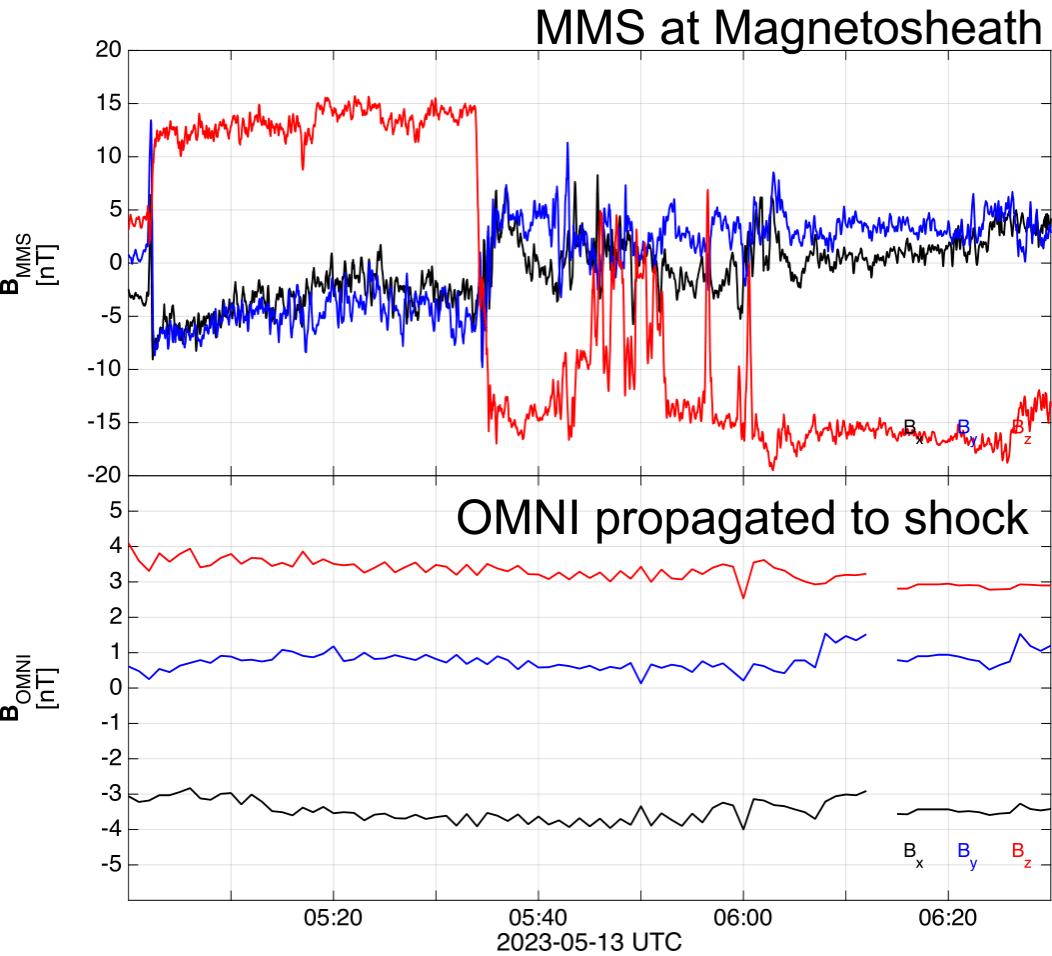
- 1) Do jets continue to show a clear double population? Does this change over time or closer to the MP ?
(A)
- 2) How are the properties and the shape vary over distance from the bow shock/magnetopause? (A)
- 3) How do jets contribute to the generation of wave activity, turbulence and particle energization? (A)



Discussion

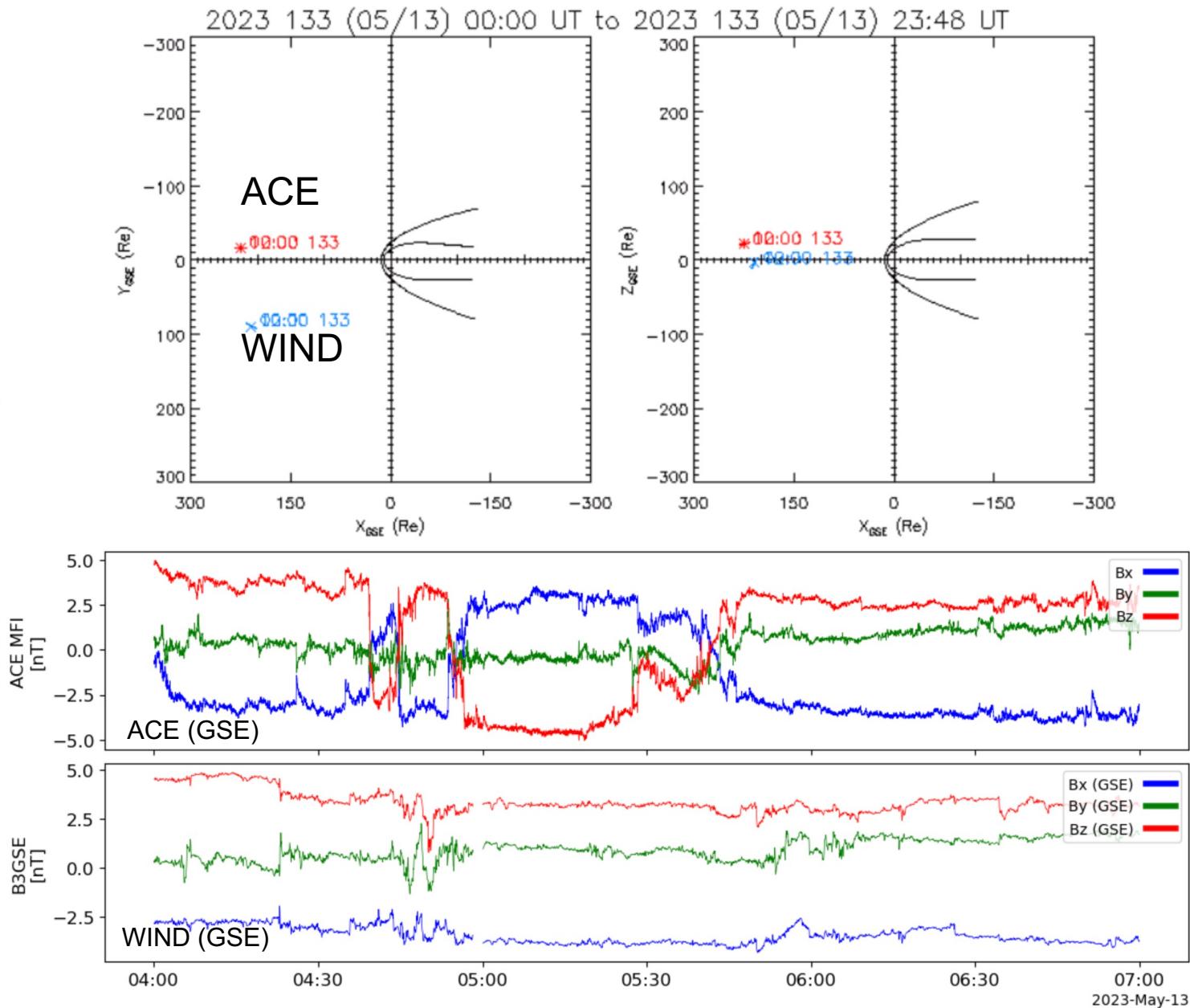
1. Foreshock & magnetosheath phenomena, in particular transient events = Interesting multi-scale physics, space weather effects, and overall unique environments to expand our understanding and test our codes.
2. MMS future campaigns: Opportunity to investigate different scales in high time-resolution, revealing previously unknown relationships between phenomena.
3. (Skipped): Effects on inner magnetosphere, excitation of ULF waves, etc.
4. (Skipped): Potential conjunctions with THEMIS/Cluster and other missions for investigating even larger scale evolution.

+1 bonus reason for the upstream monitor campaign



Key-point: Several minutes of Dt is typical,
but cases like above are not rare either.

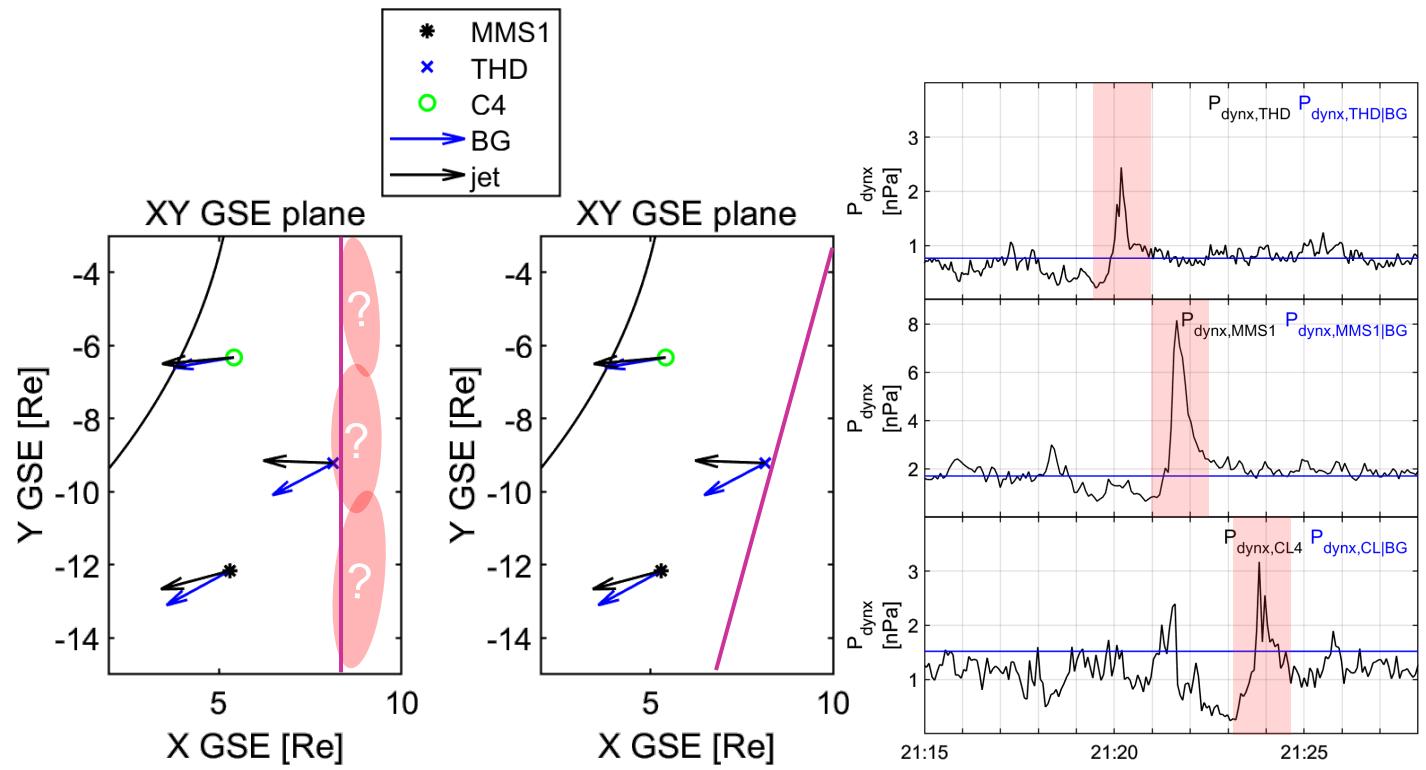
Be careful when using propagated values



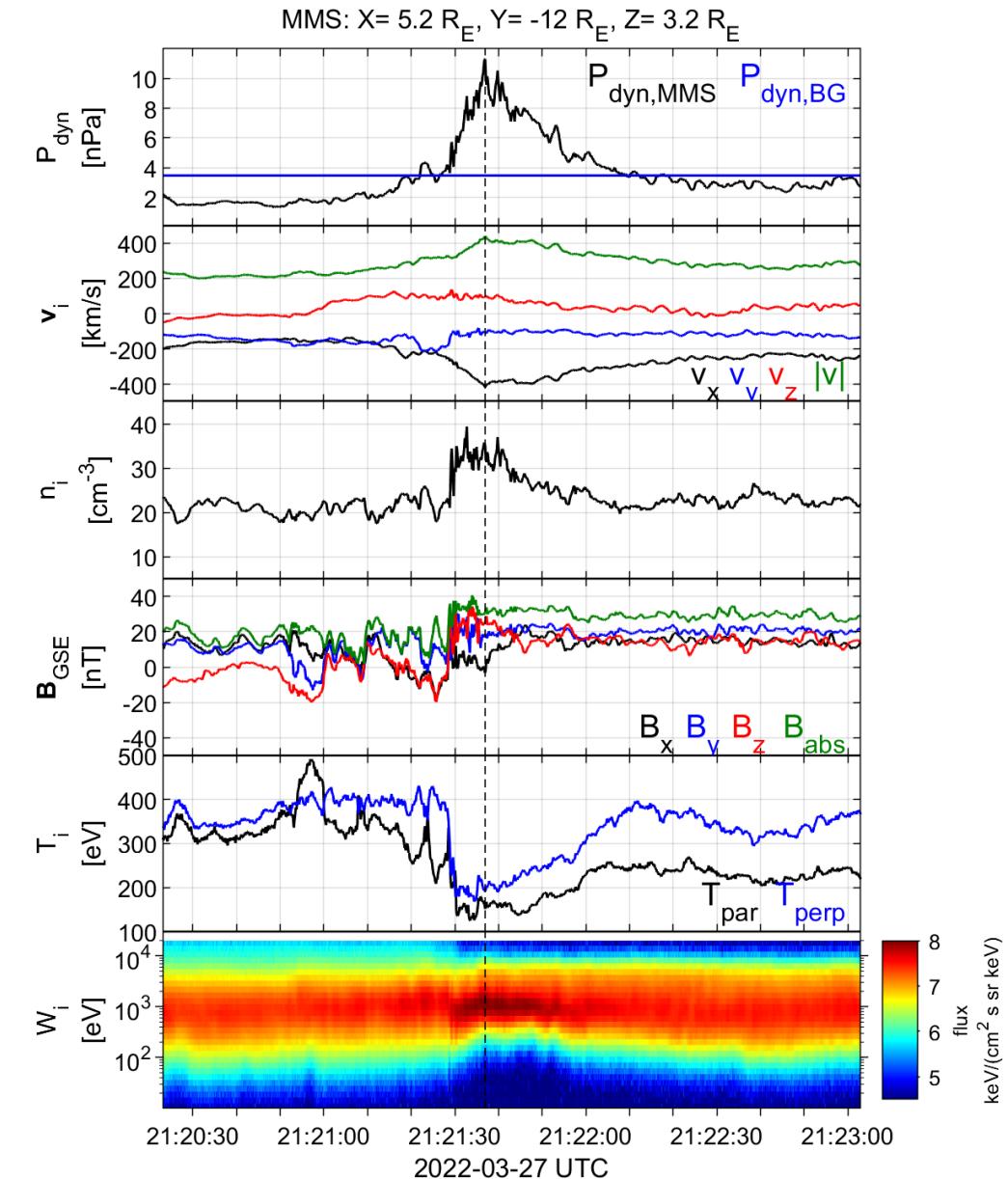
Extras

General

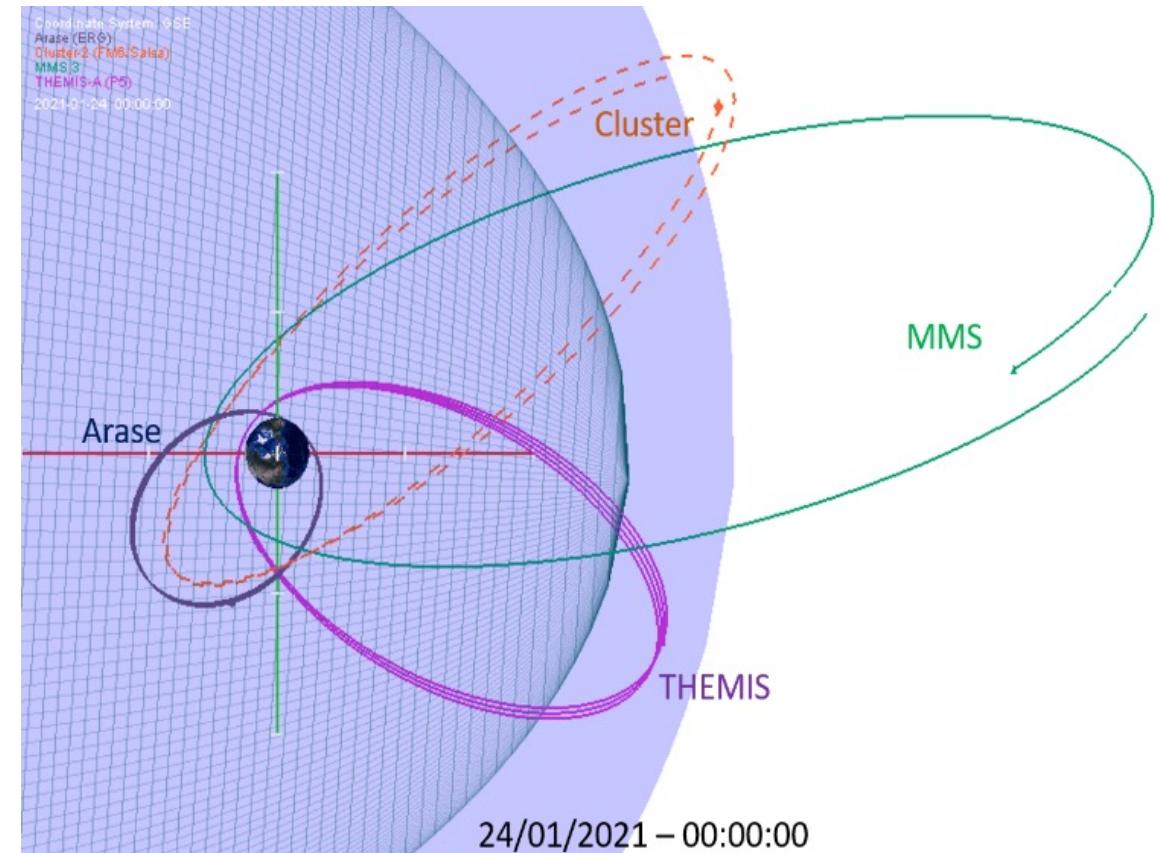
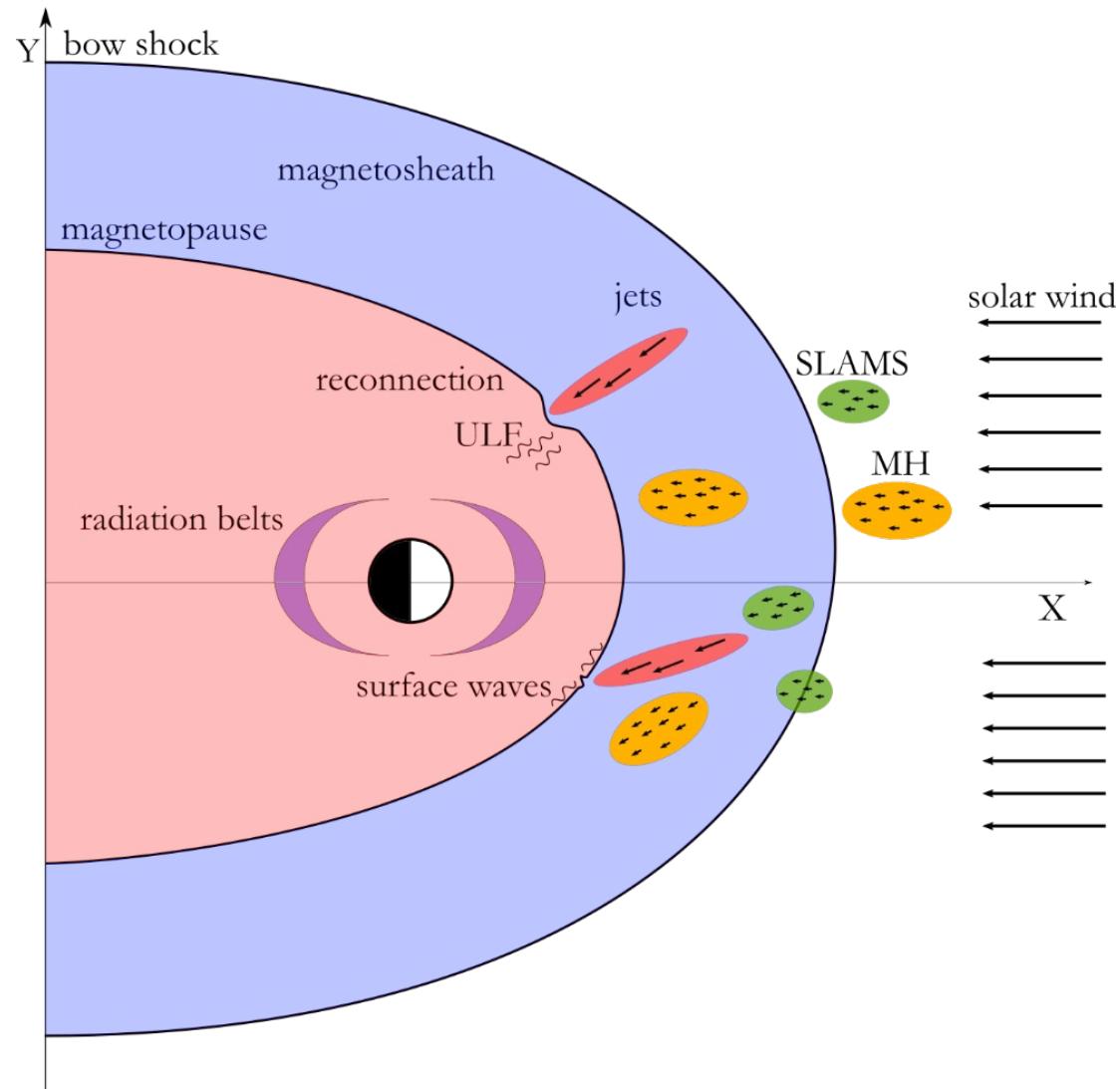
Example of a multi-SC Jet



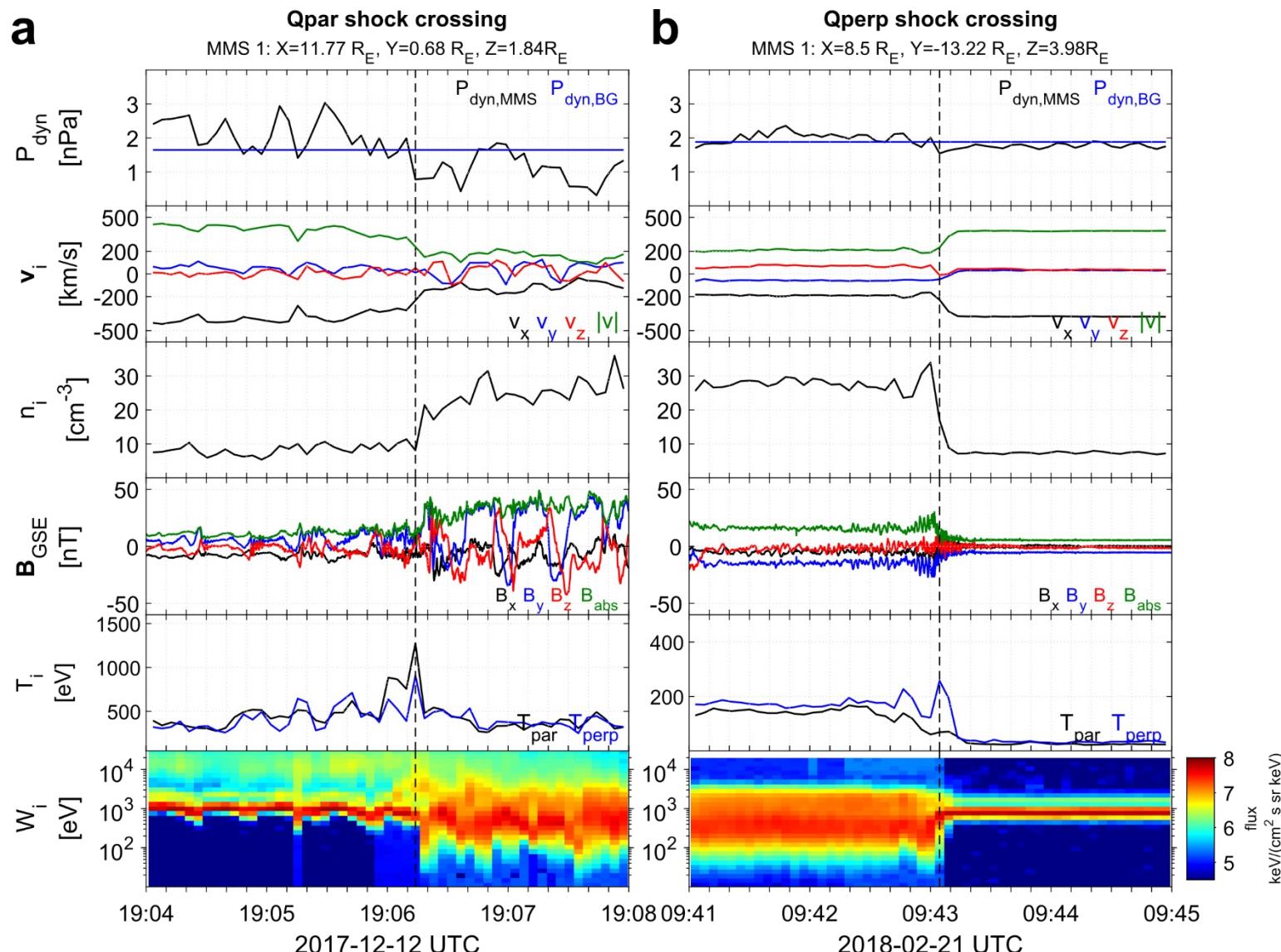
(Ongoing work – TBD)



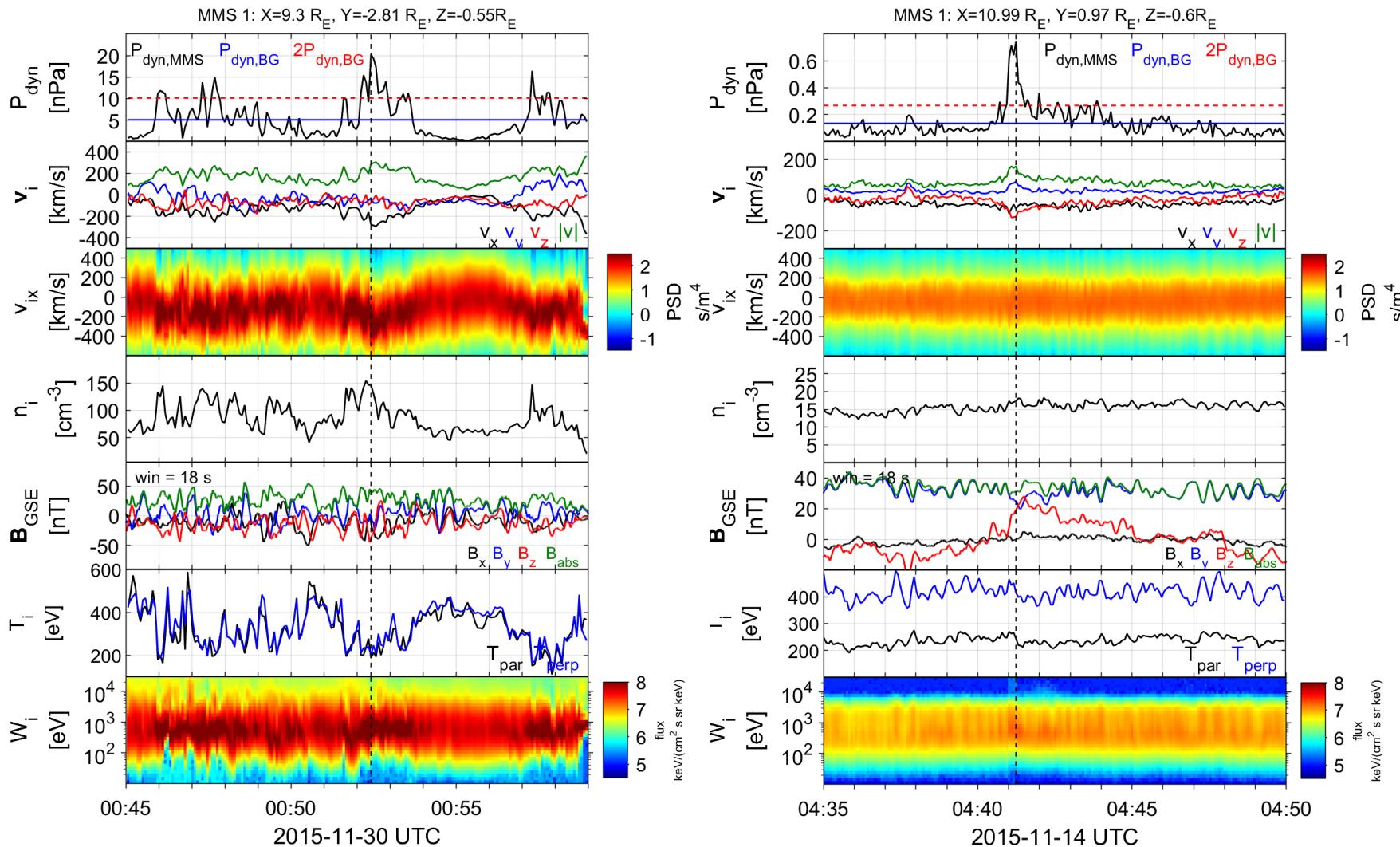
A lot of data are not fully used (conjunctions)



Qpar – Qperp crossings



Qpar – Qperp Jets



Why do we care? “big picture”

Dayside Transient Phenomena and Their Impact on the Magnetosphere and Ionosphere

Hui Zhang✉, Qiugang Zong✉, Hyunju Connor, Peter Delamere, Gábor Facskó, Desheng Han, Hiroshi Hasegawa, Esa Kallio, Árpád Kis, Guan Le, Bertrand Lembège, Yu Lin, Terry Liu, Kjellmar Oksavik, Nojan Omidi, Antonius Otto, Jie Ren, Quanqi Shi, David Sibeck & Shutao Yao

Space Science Reviews 218, Article number: 40 (2022) | [Cite this article](#)

Transmission of foreshock waves through Earth's bow shock

L. Turc✉, O. W. Roberts, D. Verscharen, A. P. Dimmock, P. Kajdič, M. Palmroth, Y. Pfau-Kempf, A. Johlander, M. Dubart, E. K. J. Kilpua, J. Soucek, K. Takahashi, N. Takahashi, M. Battarbee & U. Ganse

Nature Physics (2022) | [Cite this article](#)

Downstream high-speed plasma jet generation as a direct consequence of shock reformation

Savvas Raptis✉, Tomas Karlsson, Andris Vaivads, Craig Pollock, Ferdinand Plaschke, Andreas Johlander, Henriette Trollvik & Per-Arne Lindqvist

Nature Communications 13, Article number: 598 (2022) | [Cite this article](#)



Foreshock and magnetosheath transient phenomena and their effects on planetary magnetospheres.

Co-organized by PS2

Convener: Savvas Raptis✉ | Co-conveners: Heli Hietala, Ferdinand Plaschke, Tomas Karlsson, Christian Mazelle

Abstract submission

Geophysical Research Letters*

Research Letter | Free Access

Investigating the Role of Magnetosheath High-Speed Jets in Triggering Dayside Ground Magnetic Ultra-Low Frequency Waves

Boyi Wang✉, Yukitoshi Nishimura, Heli Hietala, Vassilis Angelopoulos

First published: 07 November 2022 | <https://doi.org/10.1029/2022GL099768>

Geophysical Research Letters*

Research Letter | Open Access | CC BY \$

Connection Between Foreshock Structures and the Generation of Magnetosheath Jets: Vlasiator Results

J. Suni✉, M. Palmroth, L. Turc, M. Battarbee, A. Johlander, V. Tarvus, M. Alho, M. Bussov, M. Dubart, U. Ganse, M. Grandin, K. Horaites, T. Manglayev, K. Papadakis, Y. Pfau-Kempf, H. Zhou



Impact of Upstream Mesoscale Transients on the Near-Earth Environment

ISSI team lead by Primož Kajdič & Xóchitl Blanco-Cano

Foreshocks Across The Heliosphere: System Specific Or Universal Physical Processes?

ISSI Team led by H. Hietala (UK) & F. Plaschke (AT)