

Problem Set - 4

Compact Binary Evolution, Rates and Population Modelling

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Problem 1:

Gaussian - Gaussian Model: The noise and observations are both normally distributed. We consider the simplest case where observed signal is dependent only on a single parameter θ . Then, the distribution of parameter:

$$p(\theta|\vec{\lambda} \equiv \{\mu, \Sigma^2\}) = \frac{1}{\sqrt{2\pi\Sigma^2}} \exp\left[-\frac{(\theta - \mu)^2}{2\Sigma^2}\right] \quad (1)$$

where $\vec{\lambda} \equiv \{\mu, \Sigma^2\}$ are the population hyper-parameters (parameters characterising the shape of distribution of θ). The noise is also Gaussian with zero mean and variance σ^2 . We can write the observed data as

$$d = \theta + n \quad (2)$$

The likelihood of data d given parameter θ is given by

$$p(d|\theta) = \frac{1}{\sqrt{2\pi\sigma^2}} \exp\left[-\frac{(d - \theta)^2}{2\sigma^2}\right] \quad (3)$$

- (a) Compute the likelihood of data d given hyper-parameters $\vec{\lambda}$: $p(d|\vec{\lambda})$. *Hint: You would like to use the probability multiplicative rule here, i.e.*

$$p(d|\vec{\lambda}) = \int d\theta p(d|\theta)p(\theta|\vec{\lambda}) \quad (4)$$

and compute the integral analytically. We have ignored the selection effects here for the sake of simplicity. For N number of data points, we will have

$$p(D|\vec{\lambda}) = \prod_{i=1}^N \int d\theta p(d_i|\theta)p(\theta|\vec{\lambda}) \quad (5)$$

where $D = \{d_1, d_2, \dots, d_N\}$.

- (b) Assume true values of hyper-parameters as $\mu_{\text{true}} = 0.5, \Sigma_{\text{true}} = 1$. Choosing width of noise distribution as $\sigma = 0.1$, simulate the observations D for, say $N = 10^5$ samples and plot the histogram.
- (c) Assuming uniform priors on μ and Σ , find out the posteriors on the hyper-parameters $\vec{\lambda}$. Here, given the analytical form of the likelihood, you would like to use Emcee sampler to sample the likelihood over the hyper-parameters.
- (d) **[Bonus]** Put the detection threshold $d_{\text{th}} = 0$ and see how many of the above events are detected? Plot the histogram of the detected events, i.e. $D_{\text{det}} = \{d_i\}_{i=1}^{N_{\text{det}}}$ with $d_i \geq d_{\text{th}}$.
- (e) **[Bonus]** Calculate the selection function $\alpha(\vec{\lambda})$, which can be computed as

$$\alpha(\vec{\lambda}) = \int_{d_{\text{th}}}^{\infty} dd \int d\theta p(d|\theta)p(\theta|\vec{\lambda}) \quad (6)$$

Hint: A clever choice of the order of integration here might make your life simple to evaluate it analytically. It is worth mentioning here that most of the times this integral can not be computed analytically.

- (f) **[Bonus]** Calculate the posteriors on the hyper-parameters using the detected events, i.e. after including the selection effects, using Emcee sampler. *Hint: Notice that selection effects will modify the expression for likelihood for hyper-parameters as*

$$p(D_{\text{det}}|\vec{\lambda}) = \prod_{i=1}^{N_{\text{det}}} \frac{\int d\theta p(d_i|\theta)p(\theta|\vec{\lambda})}{\alpha(\vec{\lambda})} \quad (7)$$

Problem 2:

GWPopulation Tutorial: Please download the Jupyter-Notebook from *LIGO-Virgo Open Data Workshop-2022* on population inference of binary black hole (BBH) mergers using GWPopulation, a python package for doing population inference.

- (a) Make statements about the formation channels of the BBHs.

Problem 3:

Reweighting the posteriors with the population prior: Download posterior samples for GWs (low SNR and high SNR), and reweight with a population prior (the population prior would be fixed, as performing a hierarchical inference would be too much to do in the time). Plot the results.

- (a) For GW150914, reweight m_1 and q posteriors according to the mass models as used in Tutorial 2. Do the posteriors change?
- (b) Re-do the above in the next day using the power-law + peak mass model.