Homework 6

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Problem 1

$$r(\phi) = \frac{a(1 - e^2)}{1 + e \cos \phi}$$

$$x = r \cos \phi$$

$$= \frac{a(1 - e^2) \cos \phi}{1 + e \cos \phi}$$

$$y = r \sin \phi$$

$$= \frac{a(1 - e^2) \sin \phi}{1 + e \cos \phi}$$

$$b = a\sqrt{1 - e^2}$$

$$T^{00} = M\delta(z) \left[\delta(x - x_1)\delta(y - y_1) + \delta(x - x_2)\delta(y - y_2)\right]$$

$$I_{xx} = \int d^3x \ x^2 T^{00}$$

$$= M(x_1^2 + x_2^2)$$

$$I_{yy} = M(y_1^2 + y_2^2)$$

$$I_{xy} = M(x_1y_1 + x_2y_2)$$

Problem 2

(a) First, contracting with the metric we can see that

$$0 = (G_{\mu\nu} + \Lambda g_{\mu\nu})$$

$$\implies 0 = g^{\mu\nu} \left(R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} + \Lambda g_{\mu\nu} \right)$$

$$= R - \frac{1}{2} 4R + 4\Lambda$$

$$\implies R = 4\Lambda.$$

Then,

$$0 = R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu}$$
$$= R_{\mu\nu} - \Lambda g_{\mu\nu}$$

Assuming the general form for the line element (as in Carrol)

$$ds^{2} = -e^{2\alpha(r)}dt^{2} + e^{2\beta(r)}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}).$$

the non-zero components of the Ricci tensor are

$$R_{tt} = e^{2(\alpha - \beta)} \left(\frac{2}{r} \alpha' + \alpha'^2 + \alpha'' \right)$$

$$R_{rr} = \frac{2}{r} \beta' + \alpha' \beta' - \alpha'^2 - \alpha''$$

$$R_{\theta\theta} = e^{-2\beta} \left(r\beta' - r\alpha' \right) + 1$$

$$R_{\phi\phi} = \sin^2 \theta \left(e^{-2\beta} \left((r\beta' - r\alpha') - 1 \right) + 1 + r^2 \right).$$

The Einstein equation then gives

$$0 = e^{2(\alpha - \beta)} \left(\frac{2}{r} \alpha' + \alpha'^2 + \alpha'' \right) - \Lambda e^{2\alpha}$$

$$0 = \frac{2}{r} \beta' + \alpha' \beta' - \alpha'^2 - \alpha'' + \Lambda e^{2\beta}$$

$$0 = e^{-2\beta} \left((r\beta' - r\alpha') - 1 \right) + 1 + \Lambda r^2$$

$$0 = \sin^2 \theta \left(e^{-2\beta} \left((r\beta' - r\alpha') - 1 \right) + 1 + \Lambda r^2 \right).$$

Multiplying the first equation by $e^{-2\alpha}$, the second by $e^{-2\beta}$, then adding gives

$$0 = e^{-2\beta} \left(\frac{2}{r} \alpha' + \alpha'^2 + \alpha'' \right) - \Lambda + e^{-2\beta} \left(\frac{2}{r} \beta' + \alpha' \beta' - \alpha'^2 - \alpha'' \right) + \Lambda$$

$$= \frac{2}{r} (\alpha' + \beta')$$

$$\implies \quad \alpha' = -\beta'$$

$$\implies \quad \alpha(r) = -\beta(r) + c.$$

However, we can set c=0 by taking $t\to e^{-c}t$. The third of the Einstein equations is now

$$0 = e^{-2\beta} \left((r\beta' - r\alpha') - 1 \right) + 1 + \Lambda r^2$$

$$= e^{2\alpha} \left(-2r\alpha' - 1 \right) + 1 + \Lambda r^2$$

$$= -\partial_r \left(re^{2\alpha} \right) + 1 + \Lambda r^2$$

$$\implies \partial_r \left(re^{2\alpha} \right) = 1 + \Lambda r^2$$

$$\implies re^{2\alpha} = r + \frac{1}{3}\Lambda r^3 + \alpha_0$$

$$\implies e^{2\alpha} = 1 + \frac{1}{3}\Lambda r^2 + \frac{\alpha_0}{r},$$

where α_0 is a yet to be determined constant. Plugging this back into the line element, we have

$$ds^{2} = -\left(1 + \frac{\alpha_{0}}{r} + \frac{1}{3}\Lambda r^{2}\right)dt^{2} + \left(1 + \frac{\alpha_{0}}{r} + \frac{1}{3}\Lambda r^{2}\right)^{-1}dr^{2} + r^{2}\left(d\theta^{2} + \sin^{2}\theta d\phi^{2}\right).$$

Clearly this reduces to the Schwarzschild metric when $\Lambda \to 0$ if we choose $\alpha_0 = -2GM$.

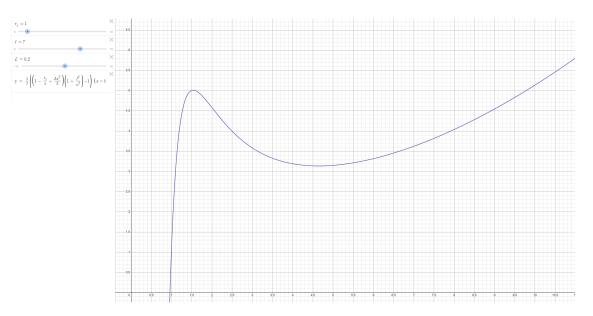


Figure 1: $V_{\text{eff}}(r)$ with $r_s=1,\,l=7,\,\Lambda=0.2.$

(b) We have the conserved quantities

$$e := -\xi_t^{\mu} \frac{\mathrm{d}x^{\mu}}{\mathrm{d}\tau} = \left(1 - \frac{r_s}{r} + \frac{1}{3}\Lambda r^2\right) \frac{\mathrm{d}t}{\mathrm{d}\tau}$$
$$l := -\xi_{\phi}^{\mu} \frac{\mathrm{d}x^{\mu}}{\mathrm{d}\tau} = -r^2 \sin^2 \theta \frac{\mathrm{d}\phi}{\mathrm{d}\tau}$$

Normalization of the four-velocity gives

$$\begin{split} \frac{\mathrm{d}x^{\mu}}{\mathrm{d}\tau} \frac{\mathrm{d}x_{\mu}}{\mathrm{d}\tau} &= -1 \\ &= -\left(1 + \frac{r_s}{r} + \frac{1}{3}\Lambda r^2\right) \left(\frac{\mathrm{d}t}{\mathrm{d}\tau}\right)^2 + \left(1 + \frac{\alpha_0}{r} + \frac{1}{3}\Lambda r^2\right)^{-1} \left(\frac{\mathrm{d}r}{\mathrm{d}\tau}\right)^2 + r^2 \left(\frac{\mathrm{d}\phi}{\mathrm{d}\tau}\right)^2 \\ &= -e^2 \left(1 + \frac{r_s}{r} + \frac{1}{3}\Lambda r^2\right)^{-1} + \left(1 + \frac{\alpha_0}{r} + \frac{1}{3}\Lambda r^2\right)^{-1} \left(\frac{\mathrm{d}r}{\mathrm{d}\tau}\right)^2 + \frac{l^2}{r^2} \\ \Longrightarrow \qquad e^2 &= 1 + \frac{r_s}{r} + \frac{1}{3}\Lambda r^2 + \left(\frac{\mathrm{d}r}{\mathrm{d}\tau}\right)^2 + \frac{l^2}{r^2} \left(1 + \frac{r_s}{r} + \frac{1}{3}\Lambda r^2\right) \\ &= \left(\frac{\mathrm{d}r}{\mathrm{d}\tau}\right)^2 + \left(1 + \frac{r_s}{r} + \frac{1}{3}\Lambda r^2\right) \left(1 + \frac{l^2}{r^2}\right) \\ \Longrightarrow \qquad \mathcal{E} := \frac{1}{2}(e^2 - 1) = \frac{1}{2}\left(\frac{\mathrm{d}r}{\mathrm{d}\tau}\right)^2 + V_{\mathrm{eff}}(r) \end{split}$$

where

$$\begin{split} V_{\text{eff}}(r) &= \frac{1}{2} \left(\left(1 - \frac{r_s}{r} + \frac{1}{3} \Lambda r^2 \right) \left(1 + \frac{l^2}{r^2} \right) - 1 \right) \\ &= \frac{1}{6} \Lambda (r^2 + l^2) - \frac{r_s}{2r} + \frac{l^2}{2r^2} - \frac{r_s l^2}{2r^3} \end{split}$$

Problem 3

- (a) Calculation done in Mathematica, see Appendix.
- (b) The transformation is given by

$$h_{\mu\nu} \to h_{\mu\nu} + \partial_{\mu}\xi_{\nu} + \partial_{n}u\xi_{\mu}$$

To kill off the cs, we need

$$2\partial_0 \xi_0 = -c \sin(k(x-t)),$$

$$2\partial_1 \xi_1 = c \sin(k(x-t)),$$

$$\partial_0 \xi_i = -\partial_i \xi_0 = 0,$$

$$\partial_i \xi_i = -\partial_i \xi_i = 0 \quad (i, j = y, z).$$

Integrating, we find that

$$\xi_{\mu} = \begin{pmatrix} -1 \\ -1 \\ 0 \\ 0 \end{pmatrix} \frac{c}{2k} \cos(k(x-t)) + b_{\mu}$$

where b_{μ} is any constant vector. Under this transformation a is unchanged.

Appendix

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$$\begin{pmatrix} c \sin(k (x-t)) & 0 & 0 & 0 \\ 0 & -c \sin(k (x-t)) & 0 & 0 \\ 0 & 0 & a \sin(k (x-t)) & 0 \\ 0 & 0 & 0 & -a \sin(k (x-t)) \end{pmatrix}$$

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