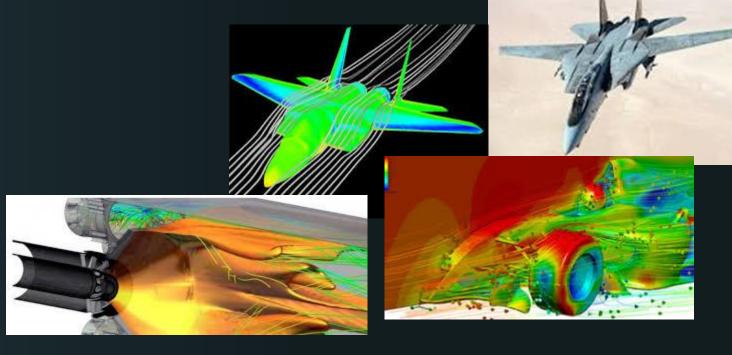
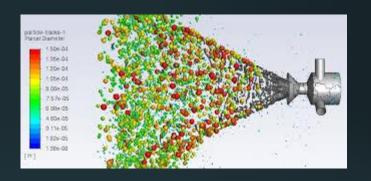
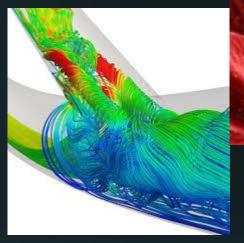
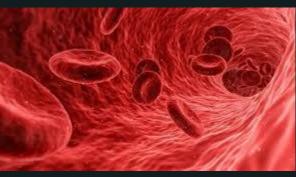
# Fluid Dynamics Lecture II – Inviscid ID equations

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#### **Outline**

Today we focus on the 1D inviscid (Euler) equations.

To understand it we start with simplified model problems that capture the physics of the full model.

- -The advection equation and numerical schemes
- Stability analysis Von Nuemann
- The Burgers equations, shocks and numerical schemes
- Diagonalization of systems of time dependent equations in oe space dimension
- Characteristics for the 1D Euler Equations
- Schemes for 1D Euler equations

### The Convection Equation

1D convection equation

$$\frac{\partial u}{\partial t} + a \frac{\partial u}{\partial x} = 0$$

Consider the problem t>0, and x in the whole space. Let the initial condition be g(x). Then the solution for all times is g(x - at). Check it!!

Note that this is a moving wave whose shape is g(x) and its speed is a (in the positive direction).

We want to develop numerical schemes to solve this equation as a preparation for more complex problems involving vectors valued functions u. We will start with a scheme that makes sense physically.

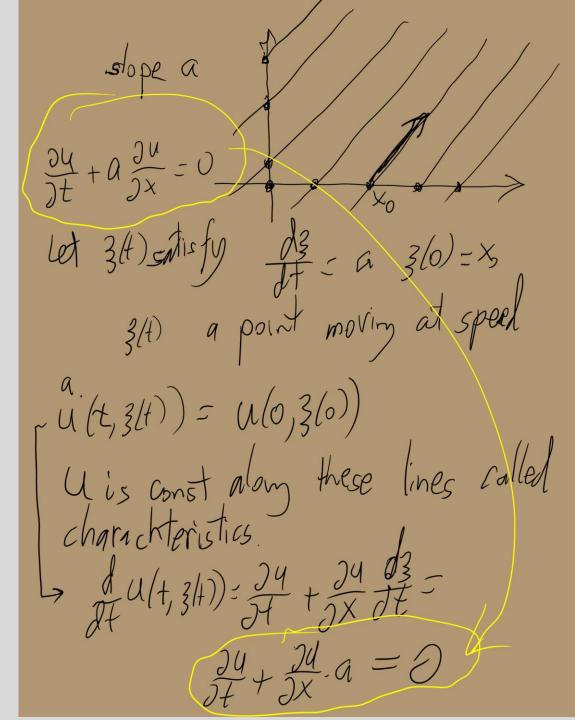
The wave is coming from the left so we will bias the scheme to the left.

How to build numerical schemes?

## The Convection Eq - cont

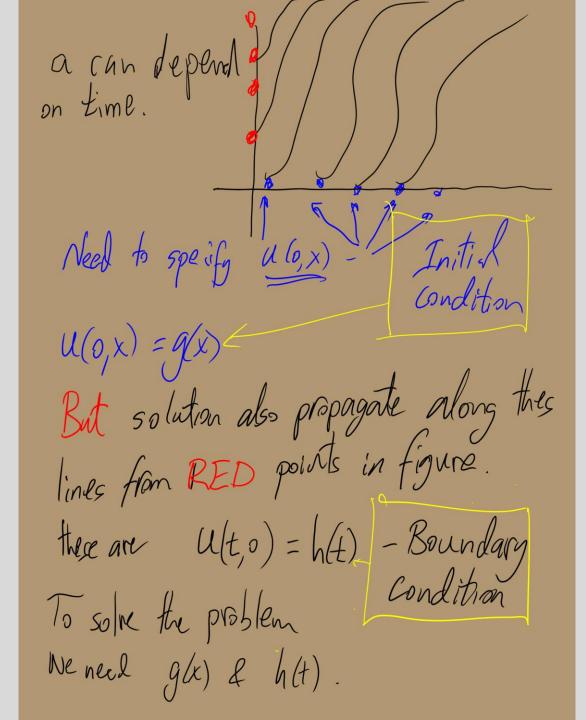
The solution propagates along lines, called characteristics.

See analysis on right.



### The Convection Eq -

Initial Conditions (IC) Boundary Condition (BC)



### The Convection Eq –

Initial Conditions (IC) Boundary Condition (BC)

Concept of inflow and outflow

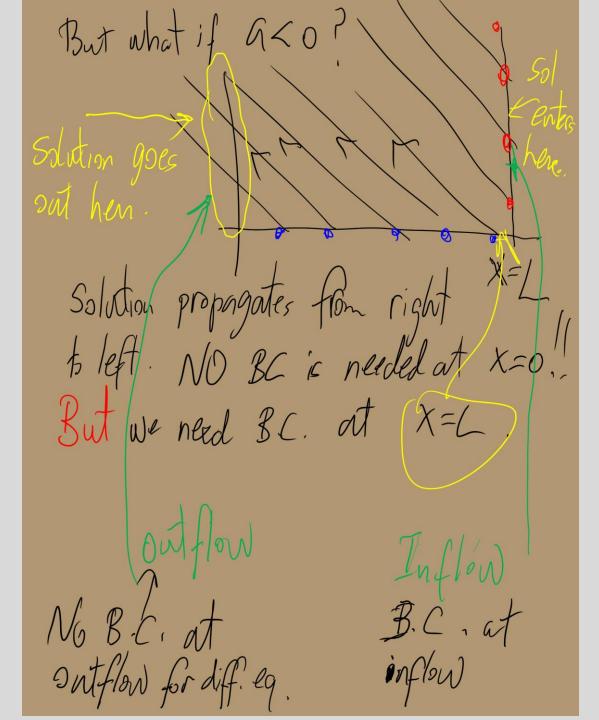
Based on U.N

- -U velocity vector
- N outward normal

U.N < 0 Inflow

U.N > 0 Outlflow

U.N = 0 Wall

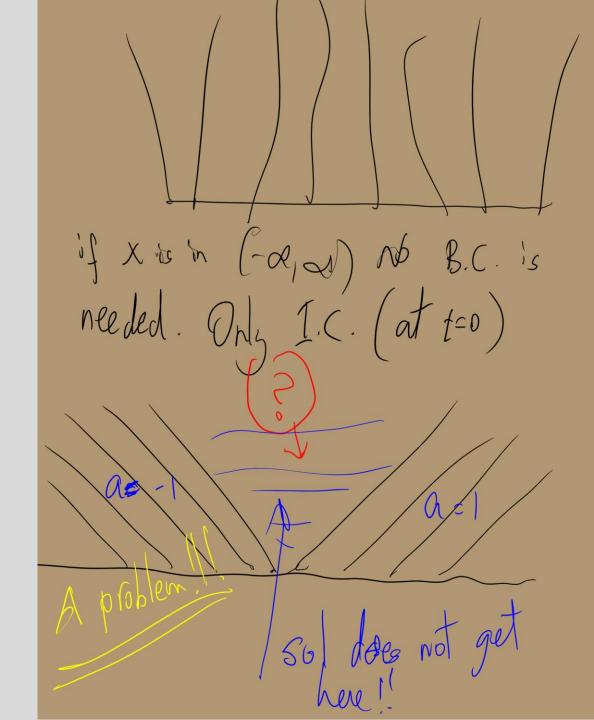


# The Convection Eq – cont

If the problem is in the whole space in x, no need to specify BC

What if characteristic do not cover the whole space?

- There is a region in which there is ambiguity
of the solution

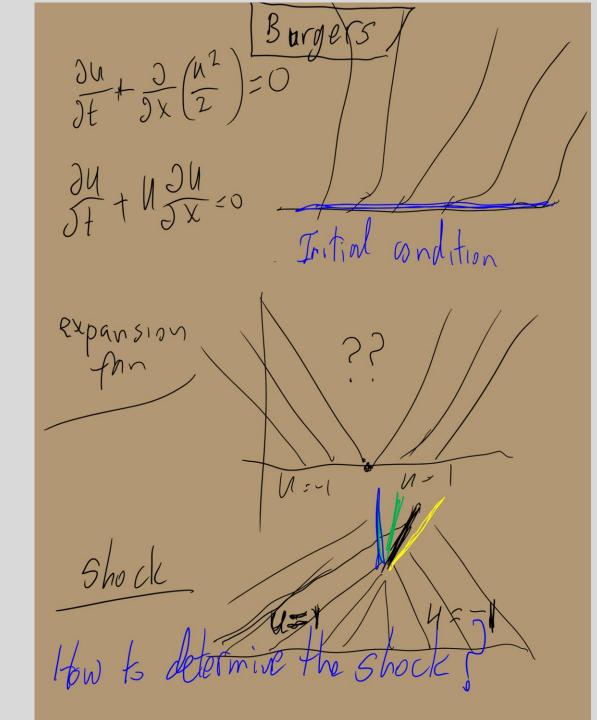


# **Burgers Equation**

Here the speed of the wave depends on the solution as it evolves over time. This model allows us to understand Important properties of the 1D Euler equations.

New things to consider.

- What if the characteristics do not fill up space?
   Expansion fan
- What if characteristics meet? Non-uniqueness Shocks



# **Burgers Equation - shocks**

Calculation of the shock speed done by integrating the equation across a shock on an infinitesimal region containing it.

The equation must be in conservation form in this analysis

We consider a general 1D scalar conservation law. u is the unknown function f(u) is called the flux

We integrate the equation in x, in an infinitesimal region around the shock. See figure on right

#### Shocks - cont.

Let [z] denotes the jump in quantity z across the shock

The jump condition for general scalar equation is shown to satisfy

$$c[u] = [f(u)]$$

Where c is the shock speed, [u] is the jump in u, and [f(u)] is the jump in f(u)

Note that for Burgers eq we get c = (u1+u2)/2

What is we have a system of equation, like the Euler 1D?

$$\int_{A}^{x_{s}(H)} \int_{A}^{x_{s}(H)} dx + \int_{A}^{x_{s}(H)} \int_{A}^{x_{s}(H)} dx = \int_{A}^{x_{s}(H)} \int_{A}^{x_{s}(H)} dx = \int_{A}^{x_{s}(H)} \int_{A}^{x_{s}(H)} \int_{A}^{x_{s}(H)} dx = \int_{A}^{x_{s}(H)} \int_{A}^{x_{s}(H)}$$

# Shocks in Systems

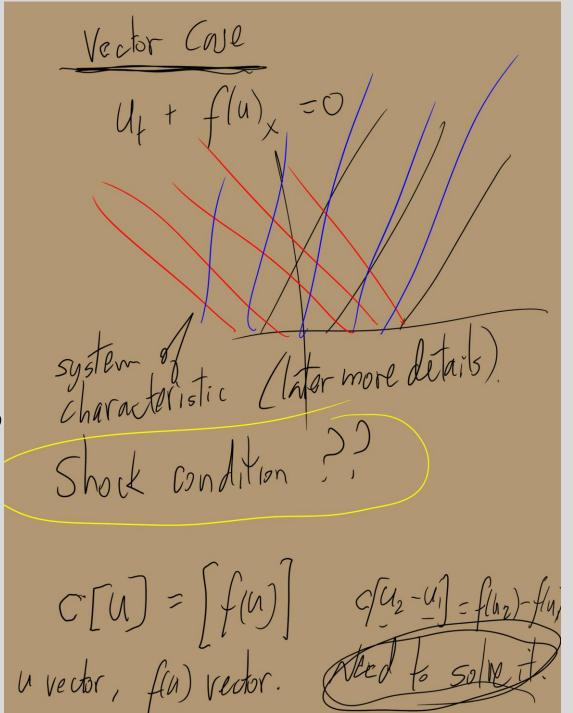
The jump condition for general 1D equation is shown to satisfy

$$c[u] = [f(u)]$$

Where c is the shock speed, [u] is the jump in u, and [f(u)] is the jump in f(u)

But to find c we need to solve a set of equations setting u to have two vector values across the shock and correspondingly f(u) is nonlinear vector valued expression in terms of these values.

Note that the analysis of shocks was done in conservation Form of the equations!!!



### Upwind Scheme – convection in 1D

We will solve the problem in t>0, x>0. This means that we also have to give a condition at x=0 (a boundary condition) in addition to the one at t=0 (initial condition).

The numerical scheme should be consistent with this physical information – the direction from which the wave is coming.

```
import matplotlib.pyplot as plt
import numpy as np
dt = 0.01
dx = 0.01
T = 0.9
a = 0.1
L = 128
u_prev = np.zeros(L)
u = np.zeros(L)
u_prev[0:int(L/4)] = 1.0
t = 0.0
plt.plot(u_prev)
fig = plt.figure()
```

```
k = 0
while t < T:
    for i in range(L-1):
        u[i+1] = u_prev[i+1] - dt*a * (u_prev[i+1]- u_prev[i])/dx
    if k % 20 == 0:
        plt.plot(u)
    for j in range(L):
        u_prev[j] = u[j]
    t = t + dt
    k = k + 1
plt.show()</pre>
```

## Instability – violating the physics

If we violate the physical information that the wave moves to th right. Say we change a to -1. **Solution** develops oscillations and blows up.

```
import matplotlib.pyplot as plt
import numpy as np
dt = 0.01
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T = 0.9
a = -0.1
L = 128
u_prev = np.zeros(L)
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t = 0.0
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```

```
k = 0
while t < T:
  for i in range(L-1):
       u[i+1] = u_prev[i+1] - dt*a * (u_prev[i+1] - u_prev[i])/dx
  if k \% 20 == 0:
     plt.plot(u)
  for j in range(L):
                           200000
     u_prev[j] = u[j]
  t = t + dt
                           100000
  k = k + 1
plt.show()
                          -100000
                          -200000
```

20

60

80

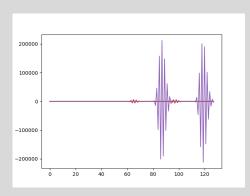
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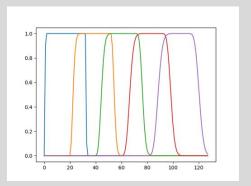
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# **Understanding Stability**

#### Von Neumann Analysis

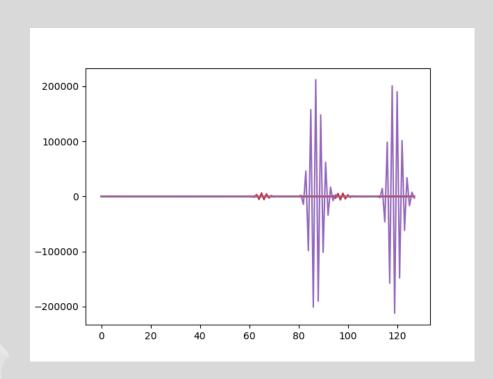
A simple analysis for constant coefficient problems that gives us information about stability or blowup





# Von Neumann Analysis

A simple analysis explaining the blowup we observed numerically for a > 0.



$$\frac{\sum x \alpha m \rho l e}{\int U + \alpha \int x} = 0$$

$$U_j^{n_1} - U_j^{n_2} + \alpha U_j^{n_3} - U_j^{n_4} = 0$$

$$U_j^{n_4} - U_j^{n_4} - \frac{\alpha \Delta t}{\Delta x} (U_j^{n_4} + U_j^{n_4})$$

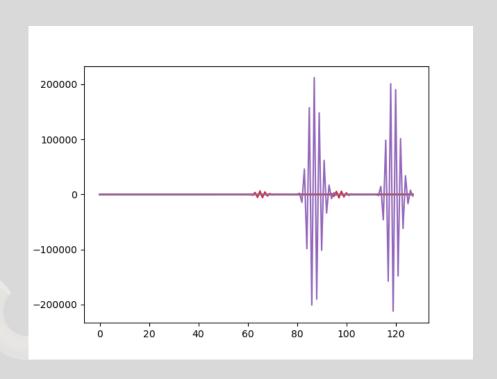
$$U_j^{n_4} = l^{n_4} - \frac{\alpha \Delta t}{\Delta x} (e^{-1}) e^{-1}$$

$$A = \left(1 + \frac{\alpha \Delta t}{\Delta x} - \frac{\alpha \Delta t}{\Delta x} e^{-1}\right)$$

$$Set \lambda = \frac{\alpha \Delta t}{\Delta x} \Delta x$$

# Von Neumann Analysis - cont

A simple analysis explaining the blowup we observed numerically if a > 0. For a<0, it is OK.



$$A = (14 A - A \cos \theta - i A \sin \theta)$$

$$= (14 A \cos \theta) = i A \sin \theta$$

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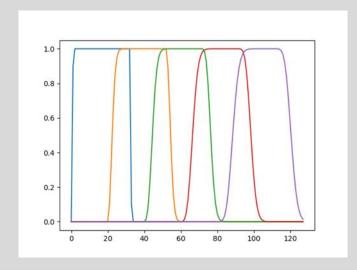
$$A = (14 A - A \cos \theta) - i A \cos \theta$$

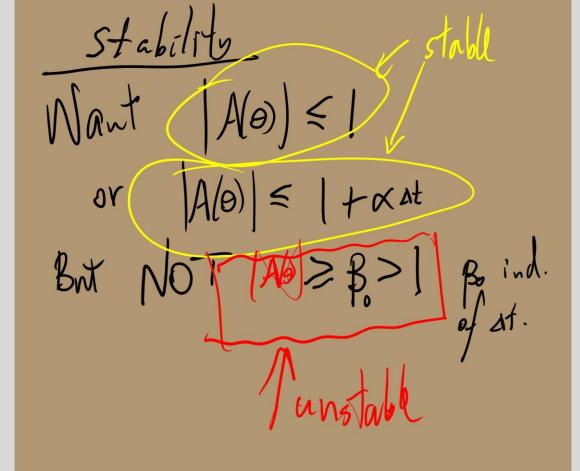
$$A = (14 A - A \cos \theta) - i A \cos \theta$$

$$A = ($$

# Stability Analysis

A simple analysis explaining good behavior of our upwind scheme





### Physical vs Numerical B.C.

In solving the advection equation, with a>0, in t>0, x>0, we need to specify a **physical B.C**. at x=0. When developing different numerical schemes, sometimes we also need to supply a **numerical boundary condition**, at the outflow boundary

The reason for this is that the equation used to defined values of u at the new time cannot be applied at the outlfow boundary since it uses points outside the domain.

The idea is then to supply a different condition for points at the outlow boundary, that are consistent with the equation. One choice is to extrapolate the solution from the inside. Another is to use an approximation of the equation but in a way that does not involve points outside the domain – e.g. a purely upwind scheme.

Nemerical schemes produce many time **spurious behavior** in the solution and stability of the numerical boundary condition also has to be done. This is beyond the scope of this short course. There is an adaptation of the Von-Neumann analysis to understand what happens at boundaries, especially reflection of waves that may have large amplitude and they are spurious.

#### The vectorial case

We consider the case where u(x,t) is a vector and A is a constant matrix with distinct real eigenvalues.

$$\frac{\partial u}{\partial t} + A \frac{\partial u}{\partial x} = 0$$

To understand the solution of such systems we introduce the diagonalization of A. Let S be a matrix whose columns are the eigenvectors of A, then we have

$$AS=S\Gamma$$
 or equivalently  $S^{-1}AS=\Gamma$ 

Where  $\Gamma$  is a diagonal matrix with the eigenvalues of A on the diagonal. Multiply the time dependent eq by  $S^{-1}$  we get

$$S^{-1}\frac{\partial u}{\partial t} + S^{-1}A\frac{\partial u}{\partial x} = 0 \qquad \text{or} \qquad S^{-1}\frac{\partial u}{\partial t} + S^{-1}ASS^{-1}\frac{\partial u}{\partial x} = 0$$

Introducing  $\delta w = S^{-1} \delta u \qquad \text{gives} \qquad \frac{\partial w}{\partial t} + \Gamma \frac{\partial w}{\partial x} = 0 \qquad \text{which are 3 independent equations} \qquad \frac{\partial w}{\partial w_i} + \frac{\partial w}{\partial x} = 0$ 

$$\frac{\partial w_j}{\partial t} + \lambda_j \frac{\partial w_j}{\partial x} = 0$$

The w's are called characteristic variables.

### Numerical Schemes for Systems

From the analysis of the vectorial case, we see that each of the w's might need a different scheme.

- Positive eigenvalue: wave goes from left to right
- Negative eigenvalue: wave goes from right to left.

The wave propagation direction also tells us where we need to give physical B.C.

- Positive eigenvalue: B.C. on the left boundary
- Negative eigenvalue: B.C. on the right boundary

Another issue is the proper treatment of shocks.

- Shocks are correctly described only in conservation form
- But waves are properly described in **characteristic form**
- How do we put these two together? More on this later.

### 1D Euler Equations – conservative form

The conservative unknowns are

$$U = \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} = \begin{pmatrix} \rho \\ m \\ \epsilon \end{pmatrix}$$

And the corresponding flux vector is

$$F = \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho u H \end{pmatrix} = \begin{pmatrix} m \\ m^2/\rho + p \\ m(\epsilon + p) \end{pmatrix}$$

And the equation is written in conservation form as

$$\frac{\partial U}{\partial t} + \frac{\partial F}{\partial x} = 0$$

$$\frac{\partial U}{\partial t} + A \frac{\partial U}{\partial x} = 0$$

Or as 
$$\frac{\partial U}{\partial t} + A \frac{\partial U}{\partial x} = 0$$
 where, 
$$A = \begin{pmatrix} 0 & 1 & 0 \\ -(3-\gamma)u^2/2 & (3-\gamma) & \gamma-1 \\ (\gamma-1)u^3 - \gamma uE & \gamma E - 3\frac{\gamma-1}{2}u^2 & \gamma u \end{pmatrix}$$

This equation uses the conservative variables.

## 1D Euler Equations – non-conservative form

The primitive variables are

$$U = \begin{pmatrix} \rho \\ u \\ p \end{pmatrix} = \begin{pmatrix} \rho \\ m/\rho \\ (\gamma - 1)(\epsilon - \frac{m^2}{2\rho}) \end{pmatrix}$$

And the corresponding equations are

$$\frac{\partial \rho}{\partial t} + u\rho_x + \rho u_x = 0$$
$$\frac{\partial u}{\partial t} + uu_x + \frac{1}{\rho}p_x = 0$$
$$\frac{\partial p}{\partial t} + up_x + \rho c^2 u_x = 0$$

Or

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ u \\ p \end{pmatrix} + \begin{pmatrix} u & \rho & 0 \\ 0 & u & \frac{1}{\rho} \\ 0 & \rho c^2 & u \end{pmatrix} \frac{\partial}{\partial x} \begin{pmatrix} \rho \\ u \\ p \end{pmatrix} = 0$$

In short 
$$\frac{\partial U}{\partial t} + A \frac{\partial U}{\partial x} = 0$$

where the matrix A is given by

$$A = \begin{pmatrix} u & \rho & 0 \\ 0 & u & \frac{1}{\rho} \\ 0 & \rho c^2 & u \end{pmatrix}$$

### 1D Euler Equations – characteristic variables

We will perform now a diagonalization procedure to discover the characteristic variables. We need the eigenvalues and eigenvectors of the matrix A. This is done as follows

$$\det \begin{pmatrix} u - \lambda & \rho & 0 \\ 0 & u - \lambda & \frac{1}{\rho} \\ 0 & \rho c^2 & u - \lambda \end{pmatrix} = 0 \quad \text{giving} \qquad \det = (u - \lambda)[(u - \lambda)^2 - c^2] = 0$$

There are 3 roots (eigenvalues)

$$\lambda_1 = u$$
 and eigenvector  $\lambda_2 = u + c$   $\lambda_3 = u - c$ 

 $\lambda_1 = u$  and eigenvectors

rs
$$l_1 = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix} \quad l_2 = \begin{pmatrix} \frac{\rho}{2c} \\ \frac{1}{2} \\ \frac{\rho c}{2} \end{pmatrix} \quad l_3 = \begin{pmatrix} -\frac{\rho}{2c} \\ \frac{1}{2} \\ -\frac{\rho c}{2} \end{pmatrix}$$

The characteristic variables are obtained form the inverse matrix and gives,

$$\delta w_1 = \delta \rho - \frac{1}{c^2} \delta p$$

$$\delta w_2 = \delta u + \frac{1}{\rho c} \delta p$$

$$\delta w_3 = \delta u - \frac{1}{\rho c} \delta p$$

and the equations for characteristic variables become

$$\frac{\partial}{\partial t} \begin{pmatrix} w_1 \\ w_2 \\ w_3 \end{pmatrix} + \begin{pmatrix} u & 0 & 0 \\ 0 & u+c & 0 \\ 0 & 0 & u-c \end{pmatrix} \frac{\partial}{\partial x} \begin{pmatrix} w_1 \\ w_2 \\ w_3 \end{pmatrix} = 0$$

#### Nozzle flow

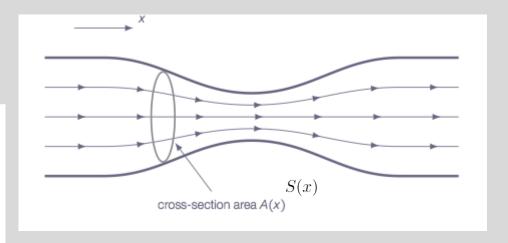
Cross section of nozzle is S(x)

The equations are

$$\frac{\partial(\rho S)}{\partial t} + \frac{\partial(\rho uS)}{\partial x} = 0$$

$$\frac{\partial(\rho uS)}{\partial t} + \frac{\partial(\rho u^2 + p)S}{\partial x} = p\frac{dS}{dx}$$

$$\frac{\partial(\rho ES)}{\partial t} + \frac{\partial(\rho uHS)}{\partial x} = 0$$



In primitive variables

$$\frac{\partial \rho}{\partial t} + u \frac{\partial \rho}{\partial x} + \rho \frac{\partial u}{\partial x} = -\frac{\rho u}{S} \frac{dS}{dx}$$
$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + \frac{1}{\rho} \frac{\partial p}{\partial x} = 0$$
$$\frac{\partial p}{\partial t} + u \frac{\partial p}{\partial x} + \rho c^2 \frac{\partial u}{\partial x} = -\frac{\rho u c^2}{S} \frac{dS}{dx}$$

RHS

$$\bar{Q} = \begin{pmatrix} -\rho u \\ 0 \\ -\rho c^2 u \end{pmatrix} \frac{1}{S} \frac{dS}{dx}$$

Note that these are the 1D Euler equations with a source term depending on the nozzle shape.

#### Nozzle flow – characteristic variables

The equations are

$$\begin{pmatrix} w_1 \\ w_2 \\ w_3 \end{pmatrix}_t + \begin{pmatrix} u & 0 & 0 \\ 0 & u+c & 0 \\ 0 & 0 & u-c \end{pmatrix} \begin{pmatrix} w_1 \\ w_2 \\ w_3 \end{pmatrix}_x = \begin{pmatrix} 0 \\ -uc \\ uc \end{pmatrix} \frac{1}{S} \frac{dS}{dx}$$

Using the expressions for the characteristic variables

$$\delta w_1 = \delta \rho - \frac{1}{c^2} \delta p$$
$$\delta w_2 = \delta u + \frac{1}{\rho c} \delta p$$
$$\delta w_3 = \delta u - \frac{1}{\rho c} \delta p$$

we get

$$d^{0}\rho - \frac{1}{c^{2}}d^{0}p = 0 \qquad d^{0} = \partial_{t} + u\partial_{x}$$

$$d^{+}u + \frac{1}{\rho c}d^{+}p = -\frac{uc}{S}\frac{dS}{dx} \qquad d^{+} = \partial_{t} + (u+c)\partial_{x}$$

$$d^{-}u - \frac{1}{\rho c}d^{-}p = \frac{uc}{S}\frac{dS}{dx} \qquad d^{-} = \partial_{t} + (u-c)\partial_{x}$$

#### Schemes

We start with the simplest equation –the advection eq.

We try a natural scheme (1) to use and analyze its stabile It turns out to be unstable

We introduce a modification (2). It is stable. We interpret the new scheme as (1) plus a special term. We call this term artificial viscosity.

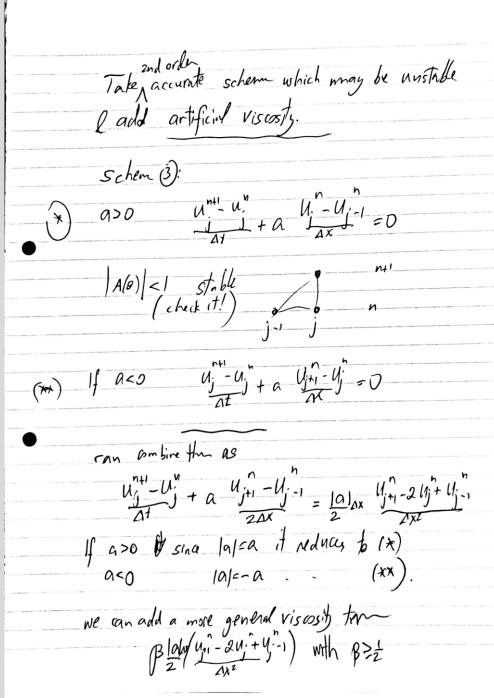
We continue with this approach. Central differencing plus artificial viscosity

$$(2) \quad (2) \quad (2)$$

#### Schemes - cont

Upwind schemes can be viewed also as central schemes plus Artificial viscosity.

We find the general formula, for both a>0 and a<0.



# Schemes – Burgers Equation

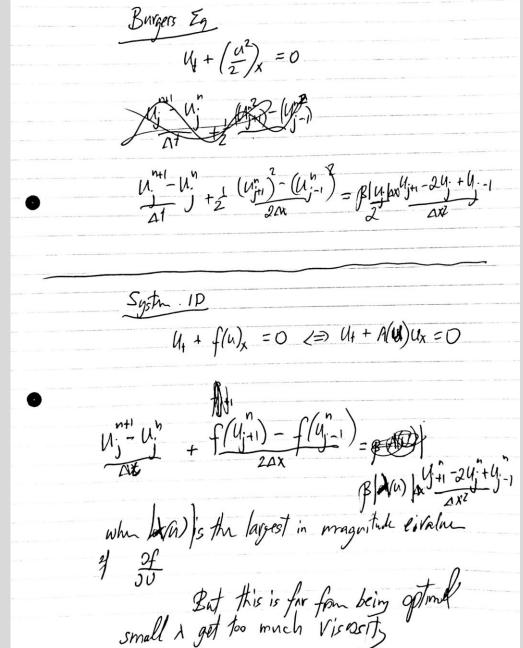
We follow what we have found for the advection equation.

We use central differencing for the conservation law and Artificial viscosity term inspired by the advection eq.

However, the scheme is not in conservation form. The artificial Viscosity breaks it down.

We will fix it later.

We then consider 1D systems of conservation laws. We add A viscosity which is based on the largest eigenvalue in absolute value. Also here we have lost the conservation and we fix it later.



# Schemes – systems cont

We can add artificial viscosity based on the speeds present in the problem. We do it using the diagonalization of A(U).

Also here we need to modify it to have a conservative scheme something properly.

AS= ST $S^{T}AS = \Gamma$ Let $ \Gamma $ be the diagonal matrix with $ \lambda_{j} $ on diagonal. Then defin $ A  = S^{-1} \Gamma S$ .  As a virtue of $ A  =  A  =$
B   A(v)   (Uj+1-2Uj+Uj-1)  B≥1
0

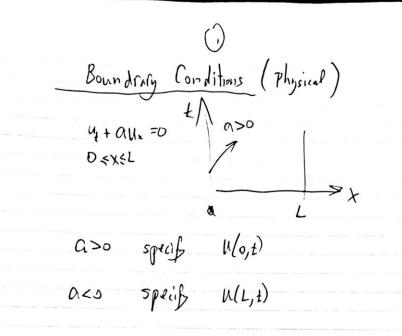
### **Boundary Conditions**

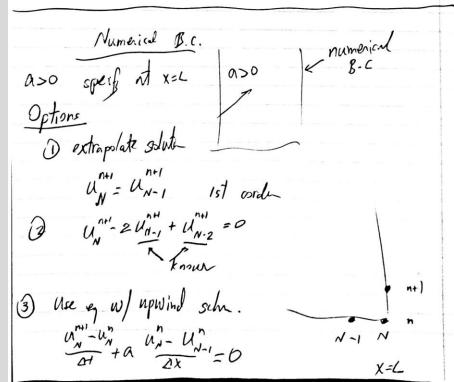
Physical and numerical

The first thing to understand is where to impose physical B.C. This depends on the waves.

At inflow we need to specify a physical BC At outflow we need to specify numerical BC.

Physical BC are done by specifying the function at the boundary. Numerical BC can be done in several ways. See (1)-(3) as examples. See if you can come up with more.



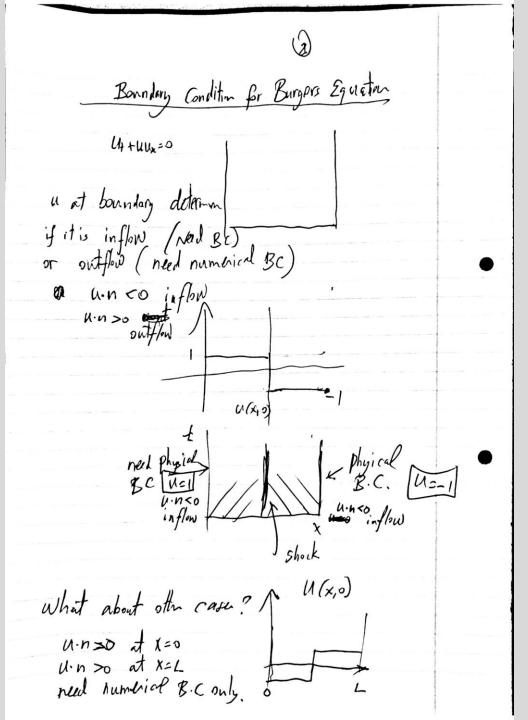


# B.C – Burgers equation

Physical and numerical

Depending on the sign of u, we may need to give physical BC Or not.

If not, then we need to specify numerical BC See figure.



#### B.C - cont.

Physical and numerical

For systems of equations, the **number** of physical BC needed depends on the eigenvalues.

Dextrapolate 1st only

2) 2 nd order

3) Note eq (same as prev.)
(age

Ny+aux=0

Sister of Greavation Laws  $U_4 + f(u)_x = 0$ 

X=/

evalue of  $A \rightarrow \lambda_1, \lambda_2, \dots$ 

 $\frac{\partial u_j}{\partial t} + \lambda_j \frac{\partial u_j}{\partial x} = 0$ 

of x=0 & numerial for 1:<0.

at X=L for each j>0 need numerical BC
for each 1, <0 need physical B-C.

#### B.C – 1D Euler

Physical and numerical

The eigenvalues for the 1D Euler are u, u+c, u-c. (we assume u> 0)

For subsonic flows, u < c and there are 2 positive eigenvalues and one negative. In this case we need 2 physical BC at inflow. one physical BC at outflow.

For supersonic flow u > c, all eigenvalues are positive and we need 3 physical BC at inflow and none at outflow.

In places where we do not specify physical BC, we specify numerical BC, in a way that ateach boundary we have three conditions, some may be physical and the other numerical depending on the situation

D Enler. P. Values are 11, 11+C, 11-C 4 M/cc subsonic, u+c,u>o Need 2 conditions at inflow At orthe need I condition need 3 conditions at inflow (X=0)

A physical conditions at outflow

example subsorie u-c<0 at x=0.

Sw3 = Su-pc Sp

Simplest BC. extrapolate W3 from insigh.

(W3)=(W3), (W3)-(W3)=0

#### B.C – 1D Euler

Physical and numerical

Implementing numerical BC is best done using characteristic Variables as shown here.



10-11, + postpo-P1)=0

I we spect fip at inflow fo po know
10=? we solve the above for No

(2) Using the 4, for W3 = 0

 $(W_3)^{hH} - W_0 + (U-C_1) \frac{(W_3)^2_2 - (W_3)^2_1}{4x} = 0$ 

 $\frac{1}{\sqrt{4}} \left[ u_{0}^{n+1} - u_{0}^{n} + \varphi_{0} \sqrt{p_{0}^{n+1} - p_{0}^{n}} \right] + (u_{0} \cdot c_{0}) \frac{u_{0}^{n} - u_{0}^{n} + \varphi_{0} \sqrt{p_{0}^{n} - p_{0}^{n}}}{4x} = 0$ 

Solu for vo if P.P are speifed at inflat.

Similarly for other cases

Key pt: Sw; = ... in terms of primitive vals