

# Homework 2, Computational Physics

Seann Smallwood

October 13, 2020

## Abstract

This assignment we were asked to analyze, solve and model three physical systems. The solutions to which all involved differential equations. The three systems were as follows:

- 1) Double Pendulum
- 2) Sun, Earth, Satellite
- 3) Driven Mass-Spring

Lagrangian Mechanics, Newtonian Mechanics and O.D.E solvers in Fortran were utilized to model the systems.

## 1 Introduction

### 1.1 Double Pendulum

The Double Pendulum was to be set up in the following manner, defining  $\theta_1$  and  $\theta_2$  as shown in figure 1.

We had four main objectives for this problem.

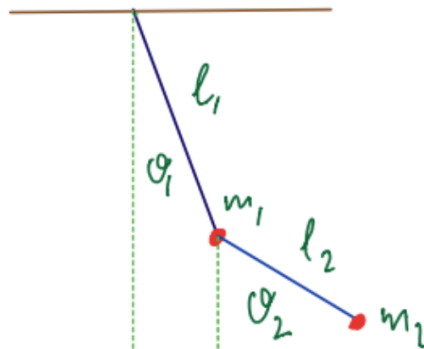


Figure 1: Structure of Double Pendulum system

- Derive the equation for  $\theta_1$ ,  $\theta_2$ ,  $\omega_1$ , and  $\omega_2$
- Derive the expression for total energy
- Write a program to solve the equation of motion for the case where  $l_1 = 0.5m$ ,  $l_2 = 0.4m$ ,  $m_1 = 3kg$ ,  $m_2 = 2kg$ . Start with initial conditions  $\theta_1 = \theta_2 = \pi/10, \pi/6, \pi/4, \pi/2$  and  $\omega_1 = \omega_2 = 0$ . Follow for 200 seconds.
- When  $\theta_1 = 0$  and  $\omega_1 > 0$  plot  $\omega_2$  as a function of  $\theta_2$

## 1.2 Sun, Earth, Satellite

The objective of this task was to model a system containing the Sun, Earth, and a satellite where  $M_{sat} \ll M_{earth} \ll M_{sun}$ . We were to find the Lagrange points at which the satellite would be in a stationary orbit relative to Earth.

## 1.3 Driven Mass-Spring

We were tasked with solving and plotting the motion of a spring mass system where the suspension point moves with a constant  $w = 7rad/s$ , at a radius of  $r = 0.5m$ .

Given the following parameters and initial conditions, we plotted the motion for 30 seconds. The equilibrium length was  $l_0 = 2m$ , mass = 1kg, spring constant  $k = 10m/N$ , initial angle  $\theta_0 = \pi/2$ , and the spring was stretched by  $\Delta l = 1m$ .

## 1.4 Setup and General Methods

Fortran was used to analyze each system. To implement common mathematical terms and define the precision to which values were calculated a file was created. This file, named numtype, is shown below.

```

≡ numtype.f90
1  module numtype
2
3      integer,parameter :: dp = selected_real_kind(15,307)
4      integer,parameter :: qp = selected_real_kind(33,4931)
5      real(dp), parameter :: pi = 4*atan(1._dp)
6      complex(dp), parameter :: iic = (0._dp,1._dp)
7
8  end module numtype

```

Figure 2: Numtype File

A Makefile, figure 3, was used to compile the fortran code and create and executable file.

The Runge-Kutta method for solving differential equations was used through a routine we named rk4, figure 4. The functions of which, lines 33 to 43, were

```

1 Makefile
2 OBJS1 = numtype.o double_pendulum.o rk4step.o
3
4 PROG1 = pend
5
6 F90 = gfortran
7
8 F90FLAGS = -O3 -funroll-loops -fexternal-blas
9
10 LIBS = -framework Accelerate
11
12 LDFLAGS = $(LIBS)
13
14 all: $(PROG1)
15
16 $(PROG1): $(OBJS1)
17     $(F90) $(LDFLAGS) -o $@ $(OBJS1)
18
19 clean:
20     rm -f $(PROG1) *.{o,mod} fort.*
21
22 .SUFFIXES: $(SUFFIXES) .f90
23
24 .f90.o:
25     $(F90) $(F90FLAGS) -c $<
26

```

Figure 3: Makefile

changed according to the system being solved. These functions input into rk4 were the equations of motion.

```

1 subroutine rk4step ( x, h, y )
2
3     use setup
4     implicit none
5     real(dp), intent(inout) :: x
6     real(dp), intent(in) :: h
7     real(dp), intent(inout), dimension(n_eq) :: y, f
8     real(dp), dimension(n_eq) :: k1, k2, k3, k4, dy
9
10
11     k1 = kv( x, h, y )
12     k2 = kv( x+h/2, h, y + k1/2 )
13     k3 = kv( x+h/2, h, y + k2/2 )
14     k4 = kv( x+h, h, y + k3 )
15     dy = ( k1 + 2*k2 + 2*k3 + k4 ) / 6
16
17     x = x + h
18     y = y + dy
19
20 contains
21
22     function kv(t, dt, y) result(k)
23
24         use setup
25         implicit none
26         real(dp), intent(in) :: t, dt
27         real(dp), dimension(n_eq), intent(in) :: y
28         real(dp), dimension(n_eq) :: f, k
29
30         real(dp) :: r_e, r_e_hat(1:3), r_s, r_c_hat(1:3), r_c_s, r_c_e_hat(1:3)
31
32         !-----
33         r_e = sqrt( y(1)**2 + y(2)**2 + y(3)**2 )      !magnitude of earths position
34         r_e_hat(1:3) = y(1:3) / r_e                  !unit vector in direction of earth
35
36         f(1:3) = y(4:6)
37         f(4:6) = - gravity * mass_sun / r_e**2 * r_e_hat
38
39         !-----
40
41         !-----
42         k = dt * f
43
44         !-----
45     end function kv
46
47 end subroutine rk4step
48

```

Figure 4: Runge-Kutte ODE Solver

## 2 Solutions

### 2.1 Double Pendulum

The main task was to derive the equations of motion for this system. From Figure 1 we see

$$x_1 = l_1 \sin(\theta_1)$$

$$y_1 = -l_1 \cos(\theta_1)$$

$$x_2 = l_1 \sin(\theta_1) + l_2 \sin(\theta_2)$$

$$y_2 = -l_1 \cos(\theta_1) - l_2 \cos(\theta_2)$$

The total energy is  $E = T + V$  where  $T$  is kinetic energy and  $V$  is potential. Taking the time derivative of the Cartesian equations and substituting into the kinetic and potential energy equations

$$T = 1/2 m_1 v_1^2 + 1/2 m_2 v_2^2$$

$$V = m_1 g y_1 + m_2 g y_2$$

We get both total energy  $E$  and the langrangian  $L$

$$E = 1/2 m_1 l_1^2 \dot{\theta}_1^2 + 1/2 m_2 [l_1^2 \dot{\theta}_1^2 + l_2^2 \dot{\theta}_2^2 + 2 l_1 l_2 \dot{\theta}_1 \dot{\theta}_2 \cos(\theta_1 - \theta_2)] - [(m_1 + m_2) g l_1 \cos(\theta_1) + m_2 g l_2 \cos(\theta_2)]$$

$$L = 1/2 m_1 l_1^2 \dot{\theta}_1^2 + 1/2 m_2 [l_1^2 \dot{\theta}_1^2 + l_2^2 \dot{\theta}_2^2 + 2 l_1 l_2 \dot{\theta}_1 \dot{\theta}_2 \cos(\theta_1 - \theta_2)] + [(m_1 + m_2) g l_1 \cos(\theta_1) + m_2 g l_2 \cos(\theta_2)]$$

From this Lagrange equation we then derived the Euler-Lagrange for each  $\theta$  which was input into the rk4 routine.

For  $\theta_1$

$$(m_1 + m_2) l_1 \ddot{\theta}_1 + m_2 l_2 \ddot{\theta}_2 \cos(\theta_1 - \theta_2) + m_2 l_2 \dot{\theta}_2^2 \sin(\theta_1 - \theta_2) + g(m_1 + m_2) \sin(\theta_1) = 0$$

For  $\theta_2$

$$m_2 l_2 \ddot{\theta}_2 + m_2 l_1 \ddot{\theta}_1 \cos(\theta_1 - \theta_2) - m_2 l_1 \dot{\theta}_1^2 \sin(\theta_1 - \theta_2) + m_2 g \sin(\theta_2) = 0$$

```

1 module setup
2
3   use ntype
4   implicit none
5
6   integer, parameter :: n_eq = 4
7   real(dp), parameter :: gravity = 9.8_dp, l_1 = 0.5_dp, &
8     mass_1 = 3.0_dp, l_2 = 0.4_dp, mass_2 = 2, &
9     alpha_1 = pi/4, alpha_2 = pi/4
10
11 end module setup
12
13
14
15 program double_pendulum
16
17   use setup
18   implicit none
19
20   real(dp) :: t, dt, tmax, eps, delt
21   real(dp), dimension(n_eq) :: y
22
23   eps = .01
24   delt = .01
25   t = 0
26   dt = .01
27   tmax = 200
28
29   y(1:2) = (/ alpha_1, 0._dp /)
30   y(3:4) = (/ alpha_1, 0._dp /)
31
32   do while ( t < tmax )
33
34     write(30,*) l_1 * sin(y(1)), -l_1 * cos(y(1))
35     write(40,*) l_1 * sin(y(1)) + l_2 * sin(y(3)), -l_1 * cos(y(1)) - l_2 * cos(y(3))
36     write(71,*) t, dt
37
38     if ((y(1) - eps) < 0 .AND. y(2) > 0) then
39       write(41,*) y(3), y(4)
40     endif
41
42     call rk4step ( t, dt, y )
43
44   end do
45
46 end program double_pendulum

```

Figure 5: Code File for Double Pendulum

```

function kv(t, dt, y) result(k)
  use setup
  implicit none
  real(dp), intent(in) :: t, dt
  real(dp), dimension(n_eq), intent(in) :: y
  real(dp), dimension(n_eq) :: r, k

  f(1) = (-gravity*(2*mass_1 + mass_2)*sin(y(2))-mass_2*gravity*sin(y(2)-2*y(4))-2*sin(y(2) &
    -y(4))*mass_2*((y(3)+2*l_2 + (y(1)+2*l_1*cos(y(2) - y(4))))/ &
    (l_1+2*mass_1 + mass_2 - mass_2*cos(2 * y(2) - 2 * y(4))))
  f(2) = y(1)

  f(3) = (2*sin(y(2)-y(4))*((y(1)+2*l_1*cos(y(2))+ &
    (y(3)+2*l_2*cos(y(2)-y(4))))/(l_2*(2*mass_1+mass_2-mass_2*cos(2*y(2)-2*y(4))))
  f(4) = y(3)

  k = dt * f
end function kv

```

Figure 6: Adjusted section of rk4 for Pendulum

## 2.2 Sun, Earth, Satellite

A ten year observational period with weekly time steps were used, and the rk4 force vectors were adjusted to incorporate the satellite with the gravitational effects of both the Sun and Earth. The effect of the satellite on both Earth and Sun were neglected as well as the Earth's pull on the Sun. The Lagrange points were calculated and input as initial conditions to be modeled.

- L1 was calculated using 0.99AU
- L2 was calculated using 1AU + 1.5e9m
- L4 used 1AU forming and equilateral triangle with Sun and Earth

The velocity was then calculated by  $v_{earth}/r_{earth} = v_{sat}/r_{sat}$  as a stationary

orbit is defined as one at which the body orbits with the same period as Earth. An L5 position exists just opposite L4, lagging behind earth rather than in front. L3 was not calculated.

```

module setup
  use namespace
  implicit none
  integer, parameter :: n_eq = 12
  real(dp), parameter :: gravity = 6.673e-11_dp, &
    mass_sun = 1.989e30_dp, mass_earth = 6.36e24_dp, mass_comet = 5e14_dp
end module setup

program planet
  use setup
  implicit none
  real(dp) :: t, dt, time
  real(dp), dimension(n_eq) :: y

  t = 0
  dt = 7 * 24 * 3600      ! hourly time steps
  time = 10 * 4000000000000 ! 10 year observational period

  y(1:3) = (/ 1.496e11_dp, 0._dp, 0._dp /) ! initial (x,y,z) of earth
  y(4:6) = (/ 0._dp, 29.78e3_dp, 0._dp /) ! initial component velocities of earth

  ! addition of satellite
  y(7:9) = (/ 1.496e11_dp * cos(pi/3), 1.496e11_dp * sin(pi/3), 0._dp /) ! initial (x,y,z) of satellite
  y(10:12) = (/ -29.78e3 * sin(pi/3), 29.78e3 * cos(pi/3), 0._dp /) ! initial component velocities of satellite

  do while ( t < time )
    write(10,*) y(1), y(9)
    write(40,*) y(7), y(10)
    write(70,*) t, dt
    call rk4step ( t, dt, y )
  end do
end program planet

```

Figure 7: Main Earth, Sun Satellite code file

```

35      r_e = sqrt( y(1)**2 + y(2)**2 + y(3)**2 ) ! magnitude of earth's position
36      r_e_hat(1:3) = y(1:3) / r_e
37
38      f(4:6) = - gravity * mass_sun / r_e**2 * r_e_hat
39
40      r_s = sqrt( y(7)**2 + y(8)**2 + y(9)**2 ) ! comet position relative to sun
41      r_s_hat(1:3) = y(7:9) / r_s
42
43      r_c_e = sqrt( (y(7)-y(1))**2 + (y(8)-y(2))**2 + (y(9)-y(3))**2 ) ! comet position relative to earth
44      r_c_e_hat(1:3) = ((y(7)-y(1))/r_c_e, (y(8)-y(2))/r_c_e, (y(9)-y(3))/r_c_e)
45
46
47
48      f(7:9) = y(10:12)
49      f(10:12) = (- gravity * mass_sun / r_s**2 * r_s_hat &
50        + (- gravity * mass_earth / r_c_e**2 * r_c_e_hat)
51
52
53      k = dt * f
54
55      end function kv
56

```

Figure 8: Adjusted section of rk4 file for satellite

## 2.3 Driven Mass-Spring

This system was solved numerically using the momentum of mass and the spring force as shown in Figure 5. The motion of the suspension point was parameterized with respect to time. The spring force was then calculated in each time step using the relative position of the mass to the suspension point.

```

2 module setup
3
4 use nunttype
5 implicit none
6
7 real(dp), parameter :: mass = 1.0_dp, g = 9.8_dp, &
8   spring_k = 10_dp, length_0 = 2.0_dp, tmax = 30_dp, &
9   r_sus = 0.5_dp
10
11
12 end module setup
13
14
15 program mass_spring
16
17 use setup
18 implicit none
19
20 real(dp) :: t, dt, dx, dy, mag_l, stretch
21 real(dp), dimension(3) :: bob_pos, bob_vel, bob_mom, F_spring, F_g, &
22   F, delta, l, l_hat
23
24 t = 0
25 dt = .01
26
27 bob_pos = (/3.5_dp, 0._dp, 0._dp/) !initial position of bob
28 bob_vel = (/0._dp, 0._dp, 0._dp/) !initial velocity of bob
29 bob_mom = mass * bob_vel
30
31 F_g = (/0._dp, -g*mass, 0._dp/)
32
33
34 do while (t < tmax)
35
36   dx = r_sus * cos(7*t) !time parameterization
37   dy = r_sus * sin(7*t) !of suspension point
38
39   delta = (/dx, dy, 0._dp/) !position of sus point at time t
40
41   l = bob_pos - delta
42   mag_l = sqrt(l(1)**2 + l(2)**2 + l(3)**2)
43   l_hat = l/mag_l
44   stretch = mag_l - length_0
45
46   F_spring = -spring_k * stretch * l_hat
47   F = F_spring + F_g
48
49   bob_mom = bob_mom + F * dt
50   bob_vel = bob_mom/mass
51   bob_pos = bob_pos + bob_vel * dt
52
53   write(13,*) bob_pos(1), bob_pos(2)
54   write(14,*) t
55   t = t + dt
56 end do
57
58 end program mass_spring

```

Figure 9: Numerical Solution to Driven Spring System

## 3 Results

### 3.1 Double Pendulum

At small angles the behavior was simple. The chaotic state increased with increased  $\theta_2$ . An IF statement was used to plot  $\omega_2$  as a function of  $\theta_2$  when  $\theta_1 = 0$  and  $\omega_1 > 0$ .

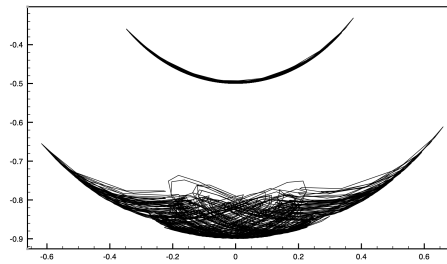


Figure 10: Position of both masses  $\text{Pi}/4$  I.C.

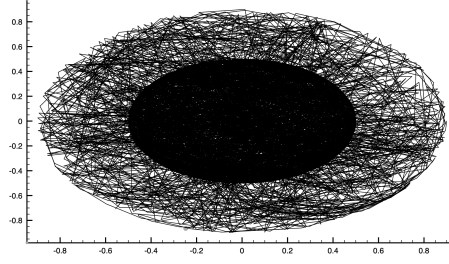


Figure 11: Position of both masses  $\text{Pi}/2$  I.C.

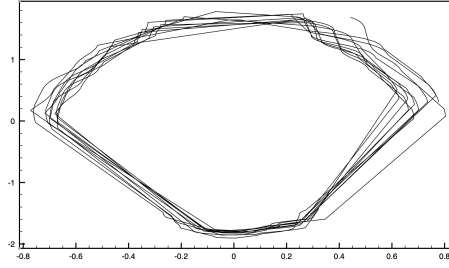


Figure 12:  $\theta_2$  vs.  $\omega_2$

### 3.2 Sun, Earth, Satellite

The first Satellite orbit generated was heavily elliptical.

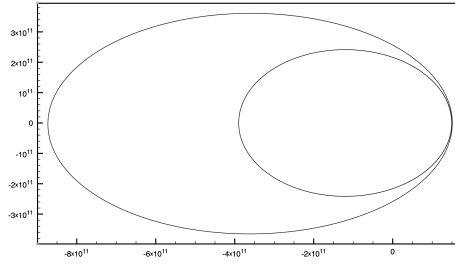


Figure 13: Satellite at arbitrary distance and speed

L1 and L2 had a narrow parameter for which they behaved as a stationary orbit.

The L4 and L5 positions matched Earth's orbit.



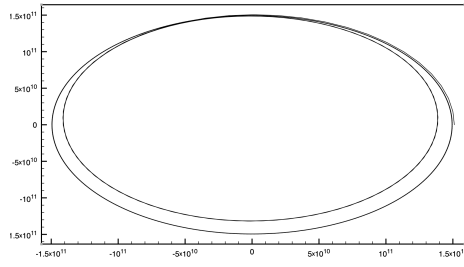


Figure 14: L2 Lagrange point orbit

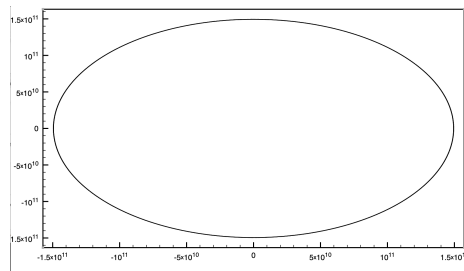


Figure 15: L4 orbit overlapping Earth's

### 3.3 Driven Mass-Spring

The resulting position of the mass was plotted in the x-y plane for a total of 30 seconds.

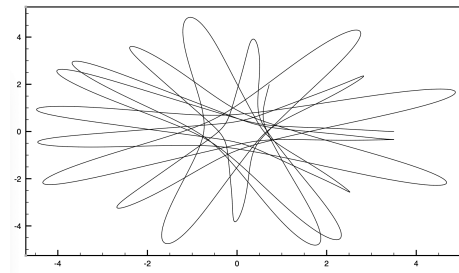


Figure 16: Position of Mass in x-y plane