Large-Scale Numerical Investigations into the Dynamics of Nonlinear Classical Systems

SIAM Conference on Applications of Dynamical Systems

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Acknowledgements

- I would like to thank A.I. Nicolin, and V. Băran for helping and motivating me.
- The author has been supported by the research project PN-III-P4-ID-PCE-2016-0792 funded by the Romanian Minisrty of Research and Inovation.
- All numerical simulations were performed on the computing cluster of Department of Computational Physics and Information Technologies, "Horia Hulubei" National Institute for Physics and Nuclear Engineering.

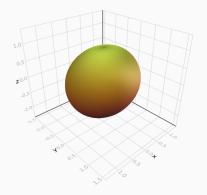
Introduction

- The physical system that we model [5] is the surface of heavy nuclei.
- We use a Hamiltonian that describes the constrained motion of the vibrational quadrupole degrees of freedom of the nuclear surface.

The Hamiltonian of the system

$$H = \frac{A}{2} \left(p_0^2 + p_2^2 \right) + \frac{A}{2} \left(q_0^2 + q_2^2 \right) + \frac{B}{\sqrt{2}} q_0 \left(3q_2^2 - q_0^2 \right) + \frac{D}{4} \left(q_0^2 + q_2^2 \right)^2$$

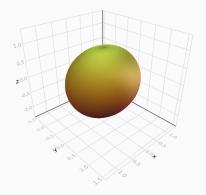
- · Harmonic oscillator part
- · Integrable part
- · Non-integrable term



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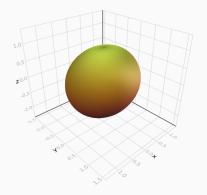
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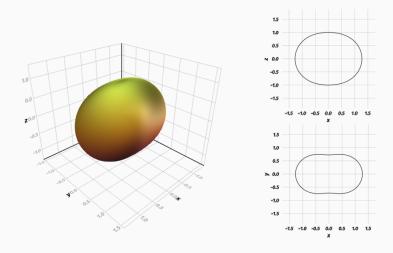


Figure 1: The nuclear surface and its sections

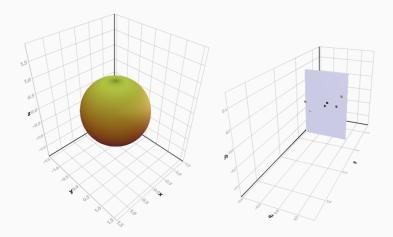


Figure 2: The nucleus and the corresponding trajectory in the phase space for a chaotic trajectory with B = 0.5, E = 0.3

Numerical simulations

Julia

- Numerical simulations and the visualizations of the results was done in Julia [2] (DifferentialEquations.jl
 [6] and DynamicalSystems.jl
 [3] for simulations and respectively Plots.jl
 and Makie.jl
 for visualizations).
- Having access to the implementations of a large number of integrators helps us taking an informed decision for the choiche of the integration algorithm.

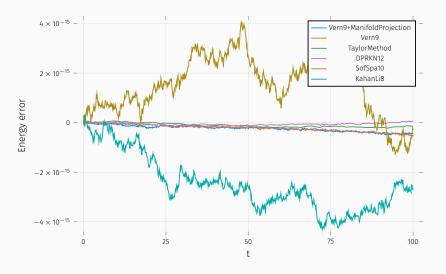


Figure 3: Energy error benchmark for short integration time

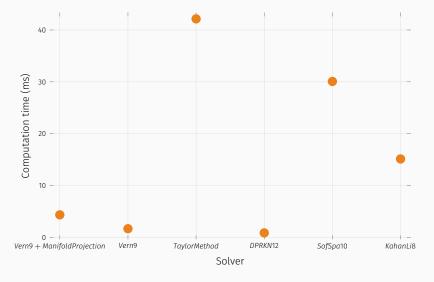


Figure 4: Computational time benchmark for short integration time

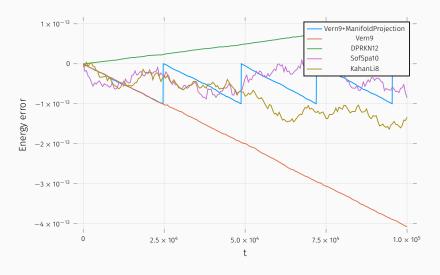


Figure 5: Energy error benchmark for long integration time

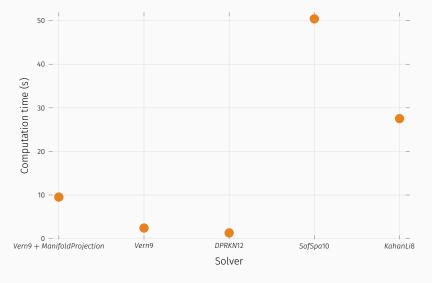


Figure 6: Computational time benchmark for long integration time

Quantifying chaoticity

• The maximal Lyapunov exponent is defined as

$$\lambda = \lim_{t \to \infty} \frac{1}{t} \ln \frac{\mathrm{d}x(t)}{\mathrm{d}x(0)}$$

- One of the signature characheristics of chaos is the sensitivity to the initial conditions ($\lambda > 0$).
- We can get some intuitive insight for the sensitivity to the initial conditions by following the evolution of two nearby initial conditions.

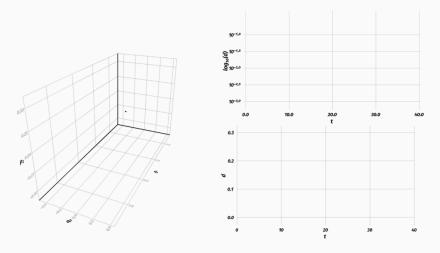


Figure 7: The distance between nearby trajectories for B = 0.5 and E = 0.3.

Interactive exploratinos

- For a given set of parameters (*B* and *E*) we have a set of compatible initial conditions.
- · Poincaré sections give us a global picture of the dynamics.
- For a better (visual) understanding of the dynamics we can show the Lyapunov exponents on the Poincaré map.

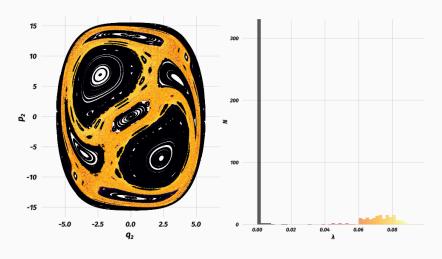


Figure 8: A Poincaré section at B = 0.5, E = 120

A note regarding the λ histogram

- Theoretically the Lyapunov exponent histogram should have two sharp peaks: one for the regular part and one for the chaotic one.
- The spread in the chaotic part is given by finite time effects.
- To better understand this we will take a look at how the integration time affects the results.

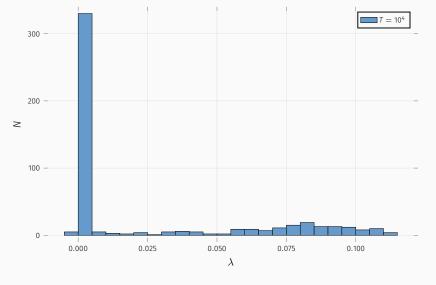


Figure 9: Maximal Lyapunov coefficient histogram for B=0.5 and E=120.

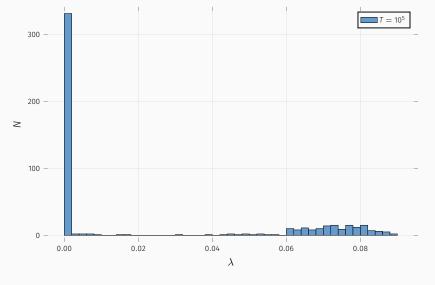


Figure 10: Maximal Lyapunov coefficient histogram for B=0.5 and E=120.

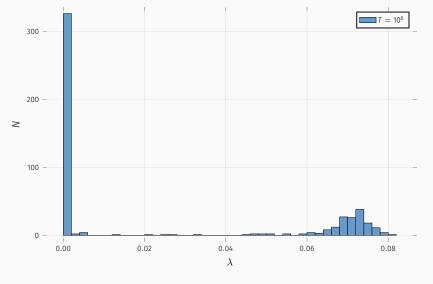


Figure 11: Maximal Lyapunov coefficient histogram for B=0.5 and E=120.

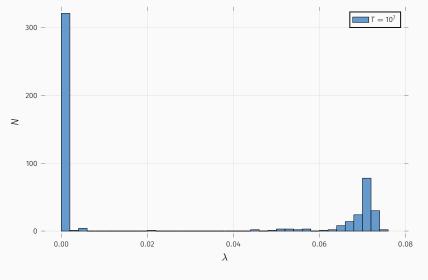


Figure 12: Maximal Lyapunov coefficient histogram for B=0.5 and E=120.

Averaging λ over the initial conditions

- We define the averaged Lyapunov coefficient as the mean of the maximal Lyapunov exponents in the chaotic region.
- We need a sufficiently robust method of selecting the λs in the chaotic region.
- We consider as chaotic everything after the first local maxima in the histogram.

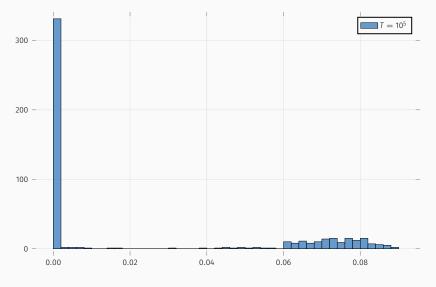


Figure 13: Selecting the chaotic trajectories for B = 0.5 and E = 120.

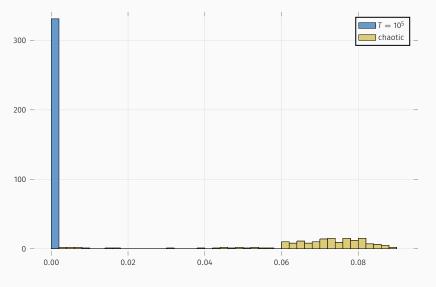


Figure 14: Selecting the chaotic trajectories for B = 0.5 and E = 120.

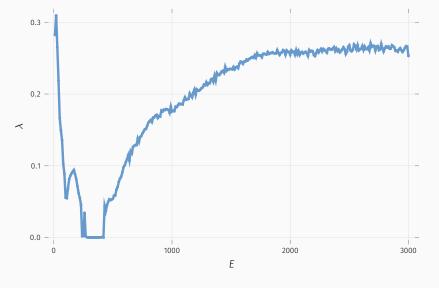


Figure 15: Averaged λ for B = 0.5 and $E \in (10, 3000)$.

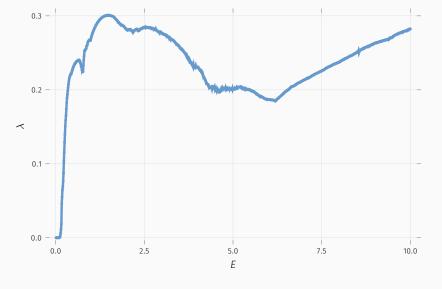


Figure 16: Averaged λ for B = 0.5 and $E \in (0.01, 10)$.

Other indicators

- We can look at the distance between two nearby trajectories in the limit of $T \to \infty$ in order to get an estimate of the phase space volume. We will denote this as d_{∞} .
- In a finite phase space volume we cannot have only an exponential divergence of trajectories, so there must be some something that folds the trajectories back after the initial divergence[1].
- \cdot We define Γ as a measure of folding

$$\Gamma = \frac{e^{\lambda} - 1}{d_{\infty}}$$

• We can apply similar averaging techniques as for λ .

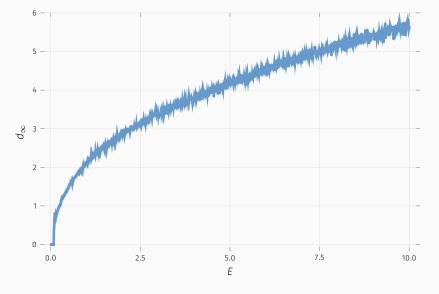


Figure 17: Averaged d_{∞} for B = 0.5 and $E \in (0.01, 10)$.

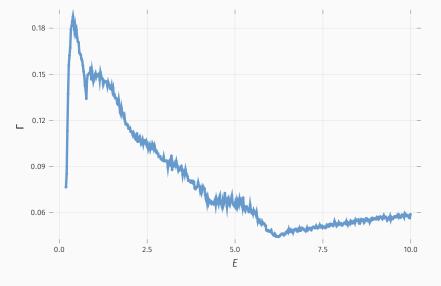


Figure 18: Averaged Γ for B = 0.5 and $E \in (0.01, 10)$.

Averaging λ over the energy

- We can get a global picture of the sistem by integrating the averaged Lyapunov coefficient over an energy interval.
- For a small energy interval in the low energy limit we get a monotonusly increasing dependence.
- For a large energy interval we have a non-trivial dependence of the averaged Lyapunov exponent with respect to the non-integrability parameter *B*.

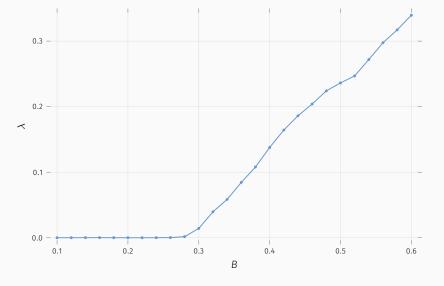


Figure 19: Averaged λ for $B = 0.5, E \in (10, 3000)$ and $B \in (0.1, 0.6)$.

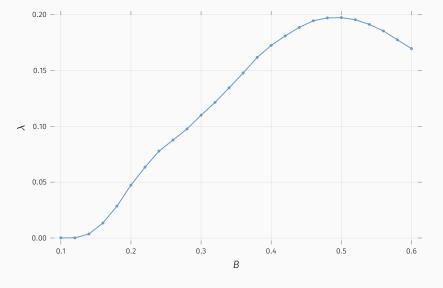


Figure 20: Averaged λ for $B = 0.5, E \in (0.01, 10)$ and $B \in (0.1, 0.6)$.

Conclusions

Conclusions

- The **Julia** ecosystem provides performant and flexible tools for exploring dynamical systems.
- We investigated in detail the classical dynamics of a non-integrable system and found a series of interesting phenomena with respect to its phase-space structure as function of energy and the non-integrability parameter.
- In order to propote open and reproducible[4, 7] science the presentation and the dataset needed for the visualizations are freely available online at https: //github.com/SebastianM-C/DS19Presentation.

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Quadrupole vibrations

· The radius of the nuclear surface

$$R(\theta, \varphi) = R_0 \left(1 + \sum_{\lambda, \mu} \alpha_{\lambda, \mu}^* Y_{\lambda, \mu}(\theta, \varphi) \right)$$

- For $\lambda=2$ we describe the surface in terms of quadrupole vibrations.
- $\alpha_{\lambda,\mu}$ become dynamical degrees of freedom, with $q_0=a_0$ and $q_2=a_2\sqrt{2}$.

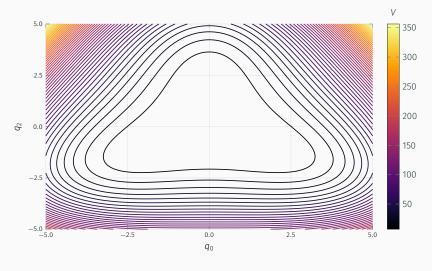


Figure 21: The equipotential lines for V at B=0.5.

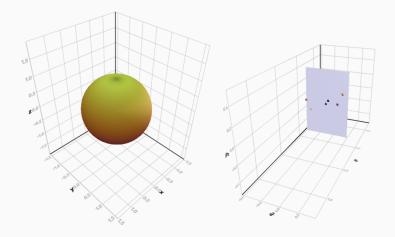


Figure 22: The nucleus and the corresponding trajectory in the phase space for a regular trajectory with B = 0.5, E = 0.3

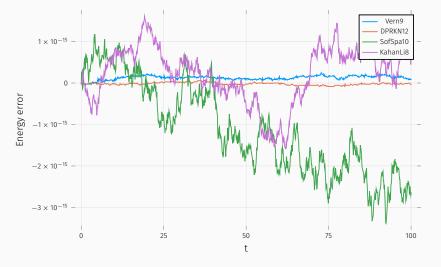


Figure 23: Energy error benchmark for short integration time with rescaling

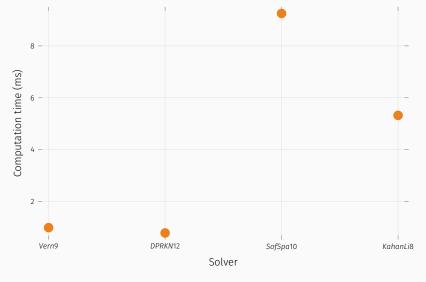


Figure 24: Computational time benchmark for short integration time with rescaling

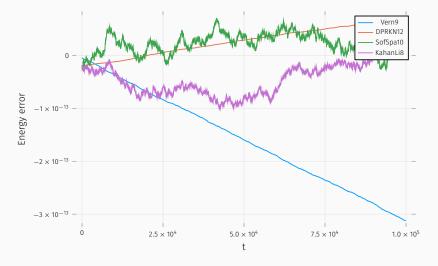


Figure 25: Energy error benchmark for long integration time with rescaling

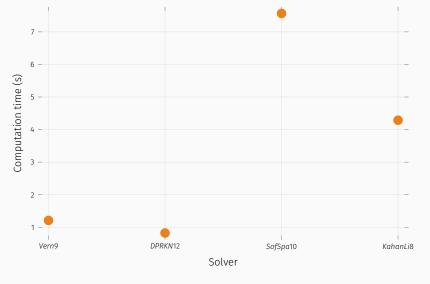


Figure 26: Computational time benchmark for long integration time with rescaling

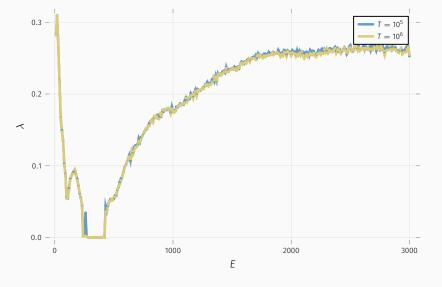


Figure 27: Averaged λ for B = 0.5 and $E \in (10, 3000)$.

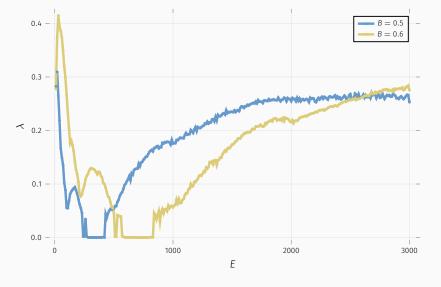


Figure 28: Averaged λ for $B \in 0.5, 0.6$ and $E \in (10, 3000)$.