

# Large-Scale Numerical Investigations into the Dynamics of Nonlinear Classical Systems

SIAM Conference on Applications of Dynamical Systems

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Introduction

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Using the **Julia** ecosystem

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# Acknowledgements

- I would like to thank A.I. Nicolin, and V. Băran for helping and motivating me.
- The author has been supported by the research project PN-III-P4-ID-PCE-2016-0792 funded by the Romanian Ministry of Research and Innovation.
- All numerical simulations were performed on the computing cluster of Department of Computational Physics and Information Technologies, “Horia Hulubei” National Institute for Physics and Nuclear Engineering.

# Introduction

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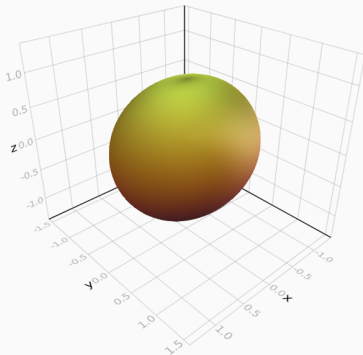
- The physical system that we model is the surface of heavy nuclei.
- We use a Hamiltonian that describes the constrained motion of the vibrational quadrupole degrees of freedom of the nuclear surface.

# The model

The Hamiltonian of the system

$$H = \frac{A}{2} (p_0^2 + p_2^2) + \frac{A}{2} (q_0^2 + q_2^2) + \frac{B}{\sqrt{2}} q_0 (3q_2^2 - q_0^2) + \frac{D}{4} (q_0^2 + q_2^2)^2$$

- Harmonic oscillator part
- Integrable part
- Non-integrable term

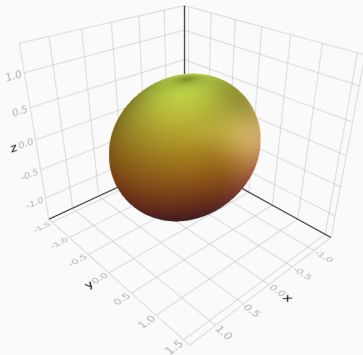


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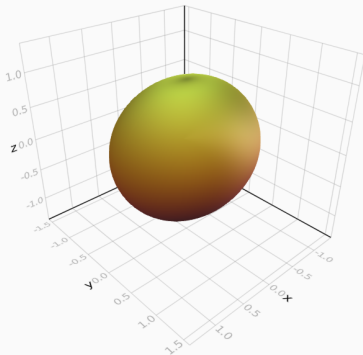


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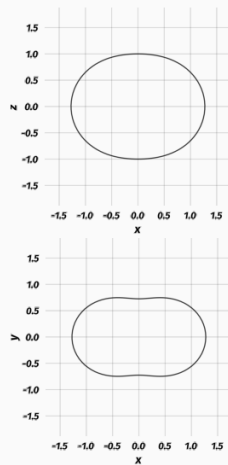
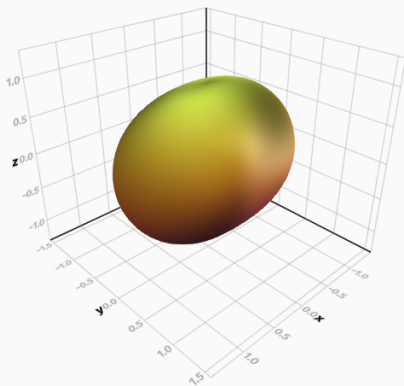


Figure 1: The nuclear surface and its sections

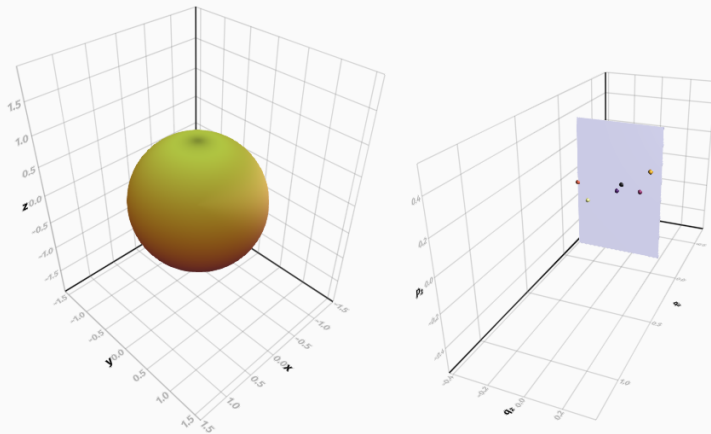


Figure 2: The nucleus and the corresponding trajectory in the phase space for a chaotic trajectory with  $B = 0.5, E = 0.3$

# Numerical simulations

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- Numerical simulations and the visualizations of the results was done in `Julia` [1] (`DifferentialEquations.jl` [4] and `DynamicalSystems.jl` [2] for simulations and respectively `Plots.jl` and `Makie.jl` for visualizations).
- Having access to the implementations of a large number of integrators helps us taking an informed decision for the choice of the integration algorithm.

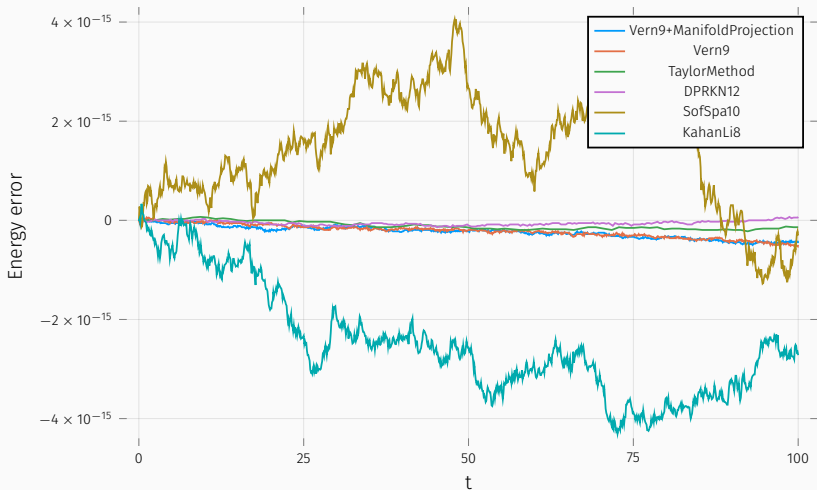


Figure 3: Energy error benchmark for short integration time

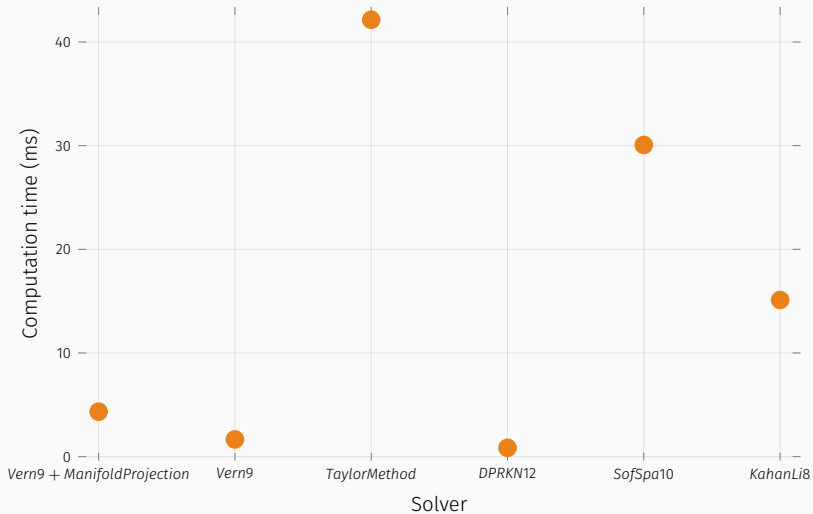


Figure 4: Computational time benchmark for short integration time



Figure 5: Energy error benchmark for long integration time

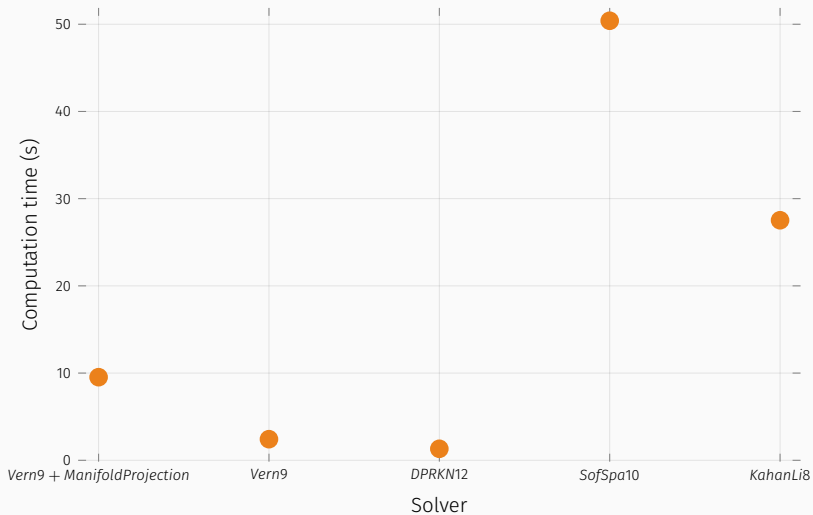


Figure 6: Computational time benchmark for long integration time



- The maximal Lyapunov exponent is defined as

$$\lambda = \lim_{t \rightarrow \infty} \frac{1}{t} \ln \frac{dx(t)}{dx(0)}$$

- One of the signature characteristics of chaos is the sensitivity to the initial conditions ( $\lambda > 0$ ).
- We can get some intuitive insight for the sensitivity to the initial conditions by following the evolution of two nearby initial conditions.

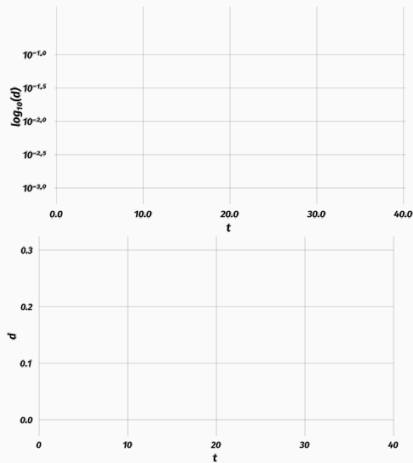
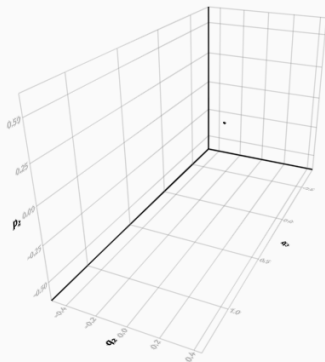


Figure 7: The distance between nearby trajectories for  $B = 0.5$  and  $E = 0.3$ .

- For a given set of parameters ( $B$  and  $E$ ) we have a set of compatible initial conditions.
- Poincaré sections give us a global picture of the dynamics.
- For a better (visual) understanding of the dynamics we can show the Lyapunov exponents on the Poincaré map.

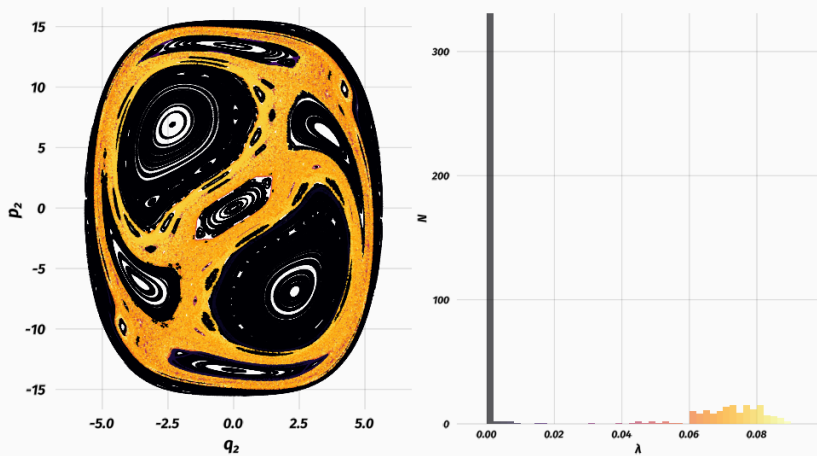


Figure 8: A Poincaré section at  $B = 0.5, E = 120$

## A note regarding the $\lambda$ histogram

- Theoretically the Lyapunov exponent histogram should have two sharp peaks: one for the regular part and one for the chaotic one.
- The spread in the chaotic part is given by finite time effects.
- To better understand this we will take a look at how the integration time affects the results.

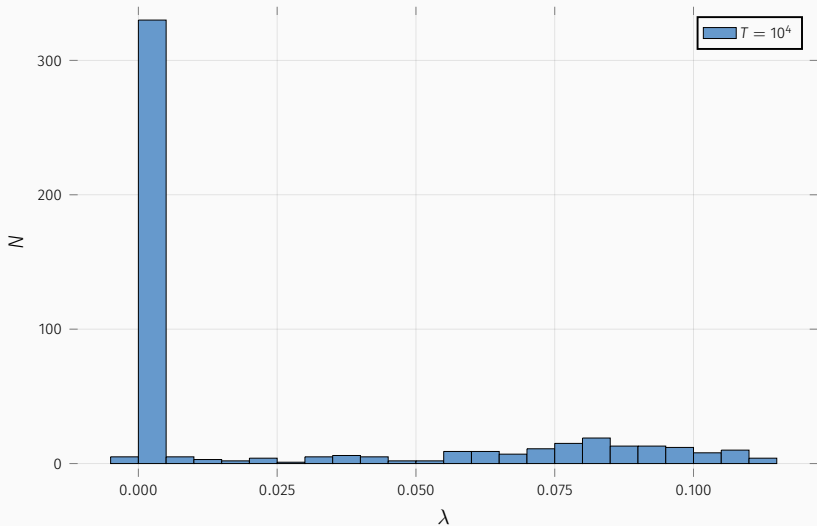


Figure 9: Maximal Lyapunov coefficient histogram for  $B = 0.5$  and  $E = 120$ .

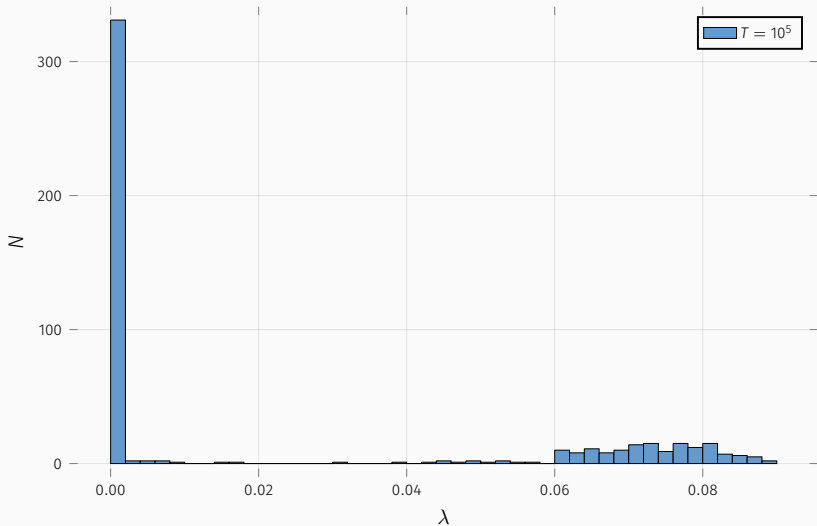


Figure 10: Maximal Lyapunov coefficient histogram for  $B = 0.5$  and  $E = 120$ .

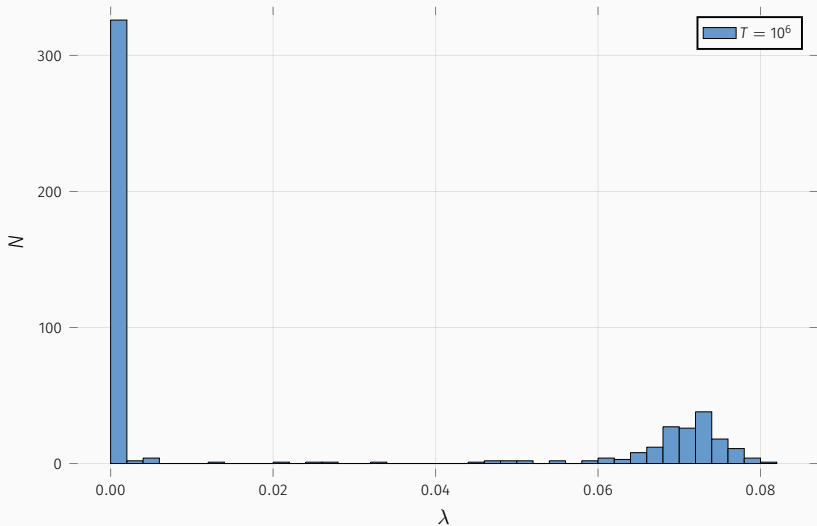


Figure 11: Maximal Lyapunov coefficient histogram for  $B = 0.5$  and  $E = 120$ .



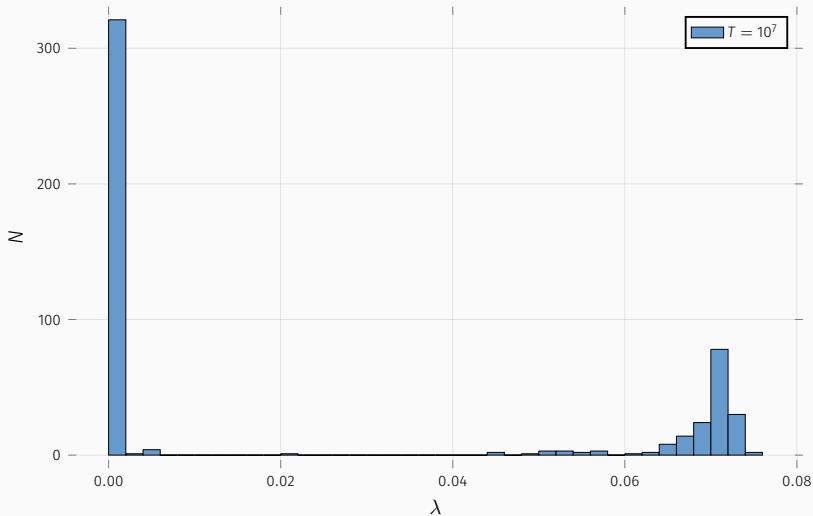


Figure 12: Maximal Lyapunov coefficient histogram for  $B = 0.5$  and  $E = 120$ .

## Averaging $\lambda$ over the initial conditions

- We define the averaged Lyapunov coefficient as the mean of the maximal Lyapunov exponents in the chaotic region.
- We need a sufficiently robust method of selecting the  $\lambda$ s in the chaotic region.
- We consider as chaotic everything after the first local maxima in the histogram.

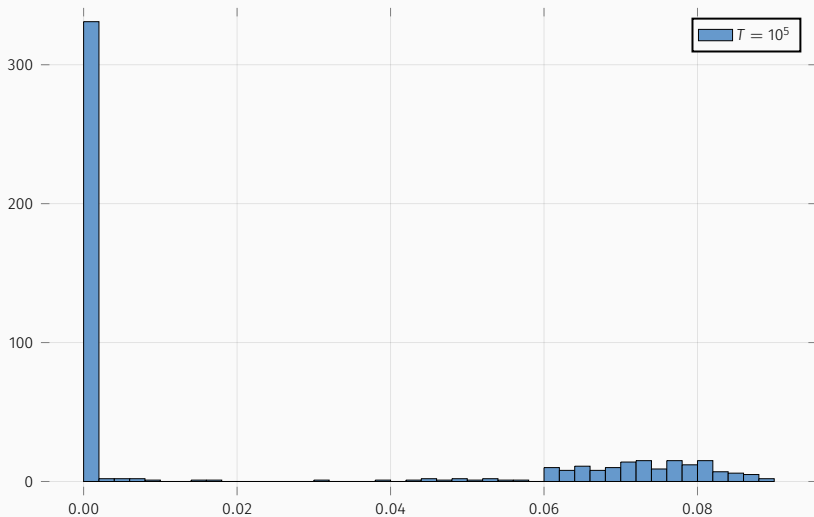


Figure 13: Selecting the chaotic trajectories for  $B = 0.5$  and  $E = 120$ .

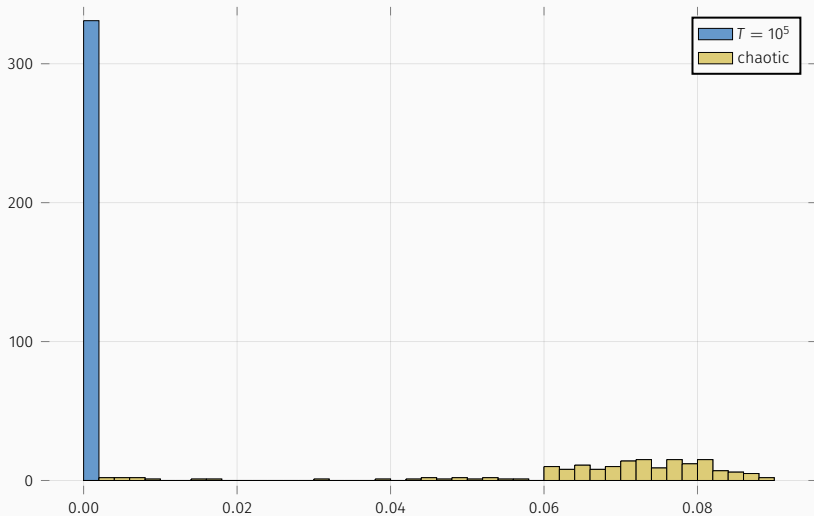


Figure 14: Selecting the chaotic trajectories for  $B = 0.5$  and  $E = 120$ .

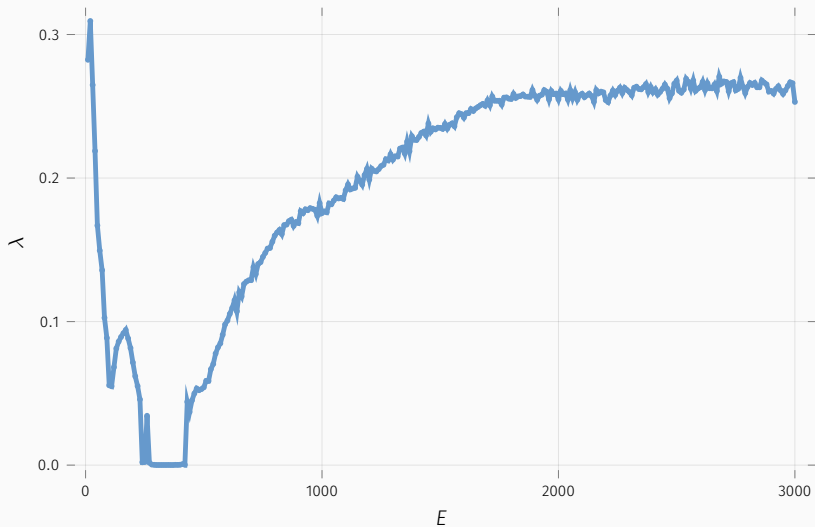


Figure 15: Averaged  $\lambda$  for  $B = 0.5$  and  $E \in (10, 3000)$ .

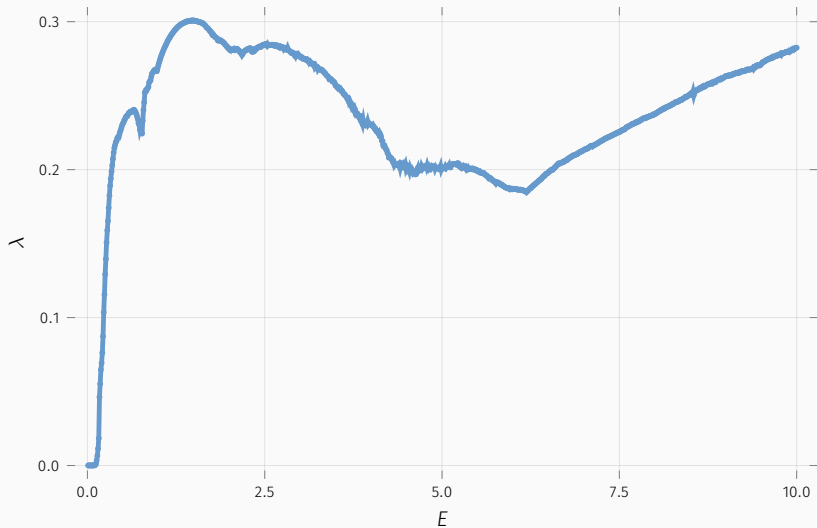


Figure 16: Averaged  $\lambda$  for  $B = 0.5$  and  $E \in (0.01, 10)$ .

## Other indicators

- We can look at the distance between two nearby trajectories in the limit of  $T \rightarrow \infty$  in order to get an estimate of the phase space volume.
- In a finite phase space volume we cannot have only an exponential divergence of trajectories, so there must be some something that folds the trajectories back after the initial divergence.
- We define  $\Gamma$  as a measure of folding

$$\Gamma = \frac{e^\lambda - 1}{d_\infty}$$

- We can apply similar averaging techniques as for  $\lambda$ .

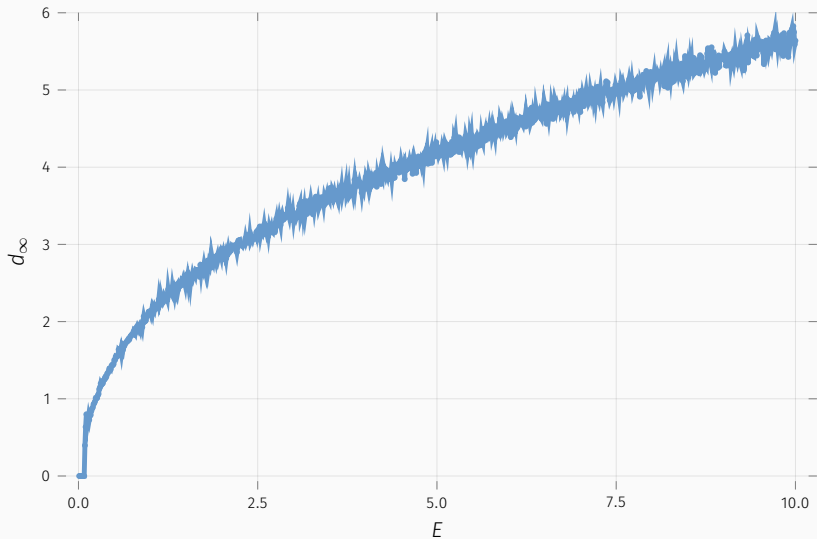


Figure 17: Averaged  $d_\infty$  for  $B = 0.5$  and  $E \in (0.01, 10)$ .



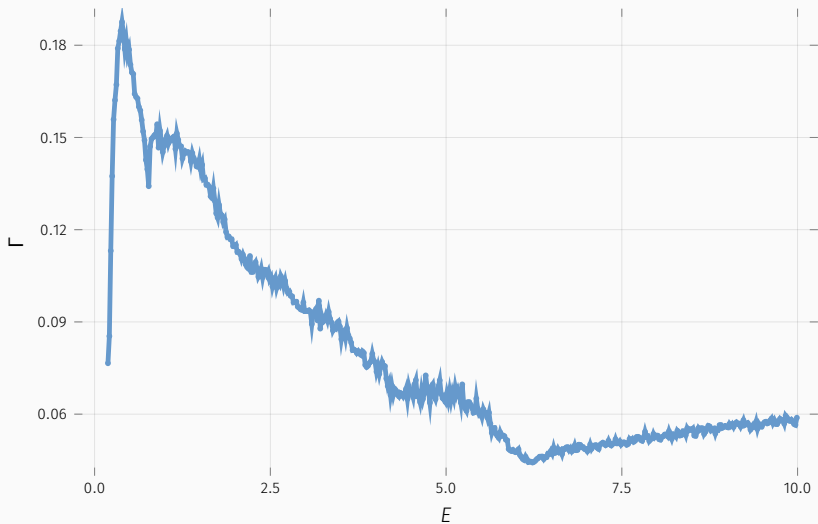


Figure 18: Averaged  $\Gamma$  for  $B = 0.5$  and  $E \in (0.01, 10)$ .

## Averaging $\lambda$ over the energy

- We can get a global picture of the system by integrating the averaged Lyapunov coefficient over an energy interval.
- For a small energy interval in the low energy limit we get a monotonously increasing dependence.
- For a large energy interval we have a non-trivial dependence of the averaged Lyapunov exponent with respect to the non-integrability parameter  $B$ .

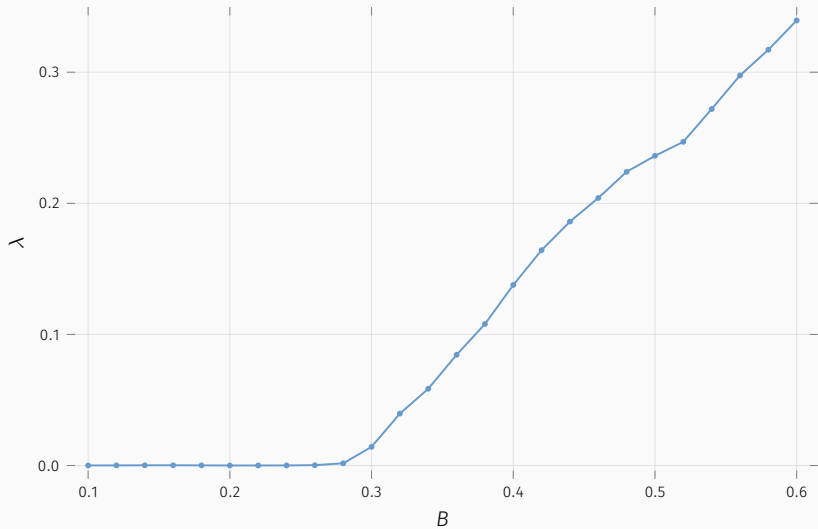


Figure 19: Averaged  $\lambda$  for  $B = 0.5, E \in (10, 3000)$  and  $B \in (0.1, 0.6)$ .

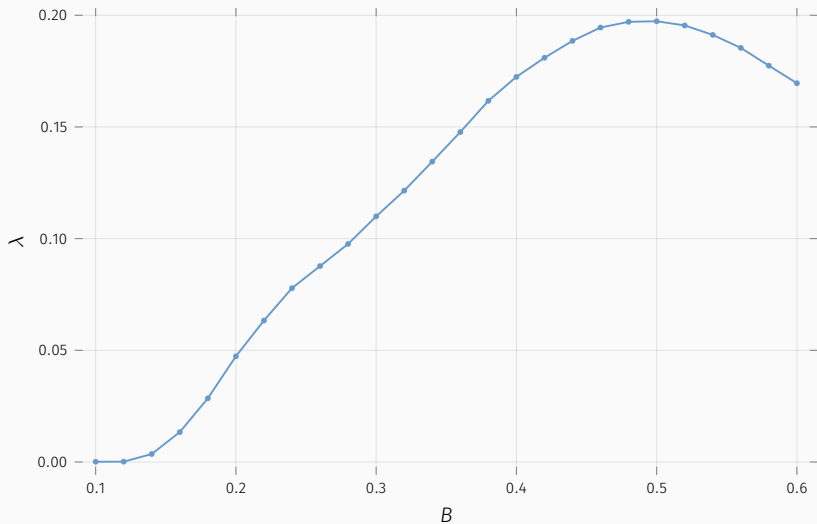


Figure 20: Averaged  $\lambda$  for  $B = 0.5, E \in (0.01, 10)$  and  $B \in (0.1, 0.6)$ .

# Conclusions

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# Conclusions

- The **Julia** ecosystem provides performant and flexible tools for exploring dynamical systems.
- We investigated in detail the classical dynamics of a non-integrable system and found a series of interesting phenomena with respect to its phase-space structure as function of energy and the non-integrability parameter.
- In order to propote open and reproducible[3, 5] science the presentation and the dataset needed for the visualizations are freely available online at **`https://github.com/SebastianM-C/DS19Presentation`**.



J. Bezanson, A. Edelman, S. Karpinski, and V. B. Shah.  
**Julia: A fresh approach to numerical computing.**  
*SIAM Review*, 59(1):65–98, 2017.



G. Datseris.  
**Dynamicalsystems.jl: A julia software library for chaos and nonlinear dynamics.**  
*Journal of Open Source Software*, 3(23):598, mar 2018.



G. M. Kurtzer, V. Sochat, and M. W. Bauer.  
**Singularity: Scientific containers for mobility of compute.**  
*PLOS ONE*, 12(5):1–20, 05 2017.



C. Rackauckas and Q. Nie.

Differentialequations.jl – a performant and feature-rich ecosystem for solving differential equations in julia.

*Journal of Open Research Software*, 5(1):15, 2017.



L. White, R. Togneri, W. Liu, and M. Bennamoun.

Datadeps.jl: Repeatable data setup for replicable data science, 2018.



Thank you!

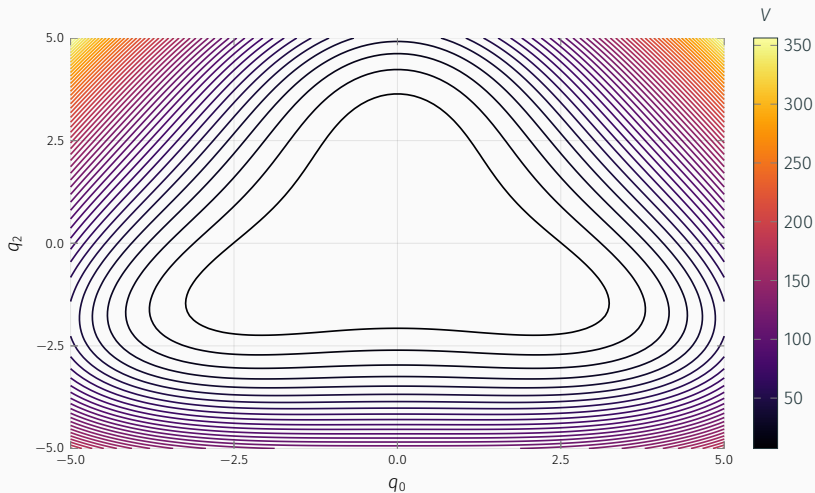


Figure 21: The equipotential lines for  $V$  at  $B = 0.5$ .

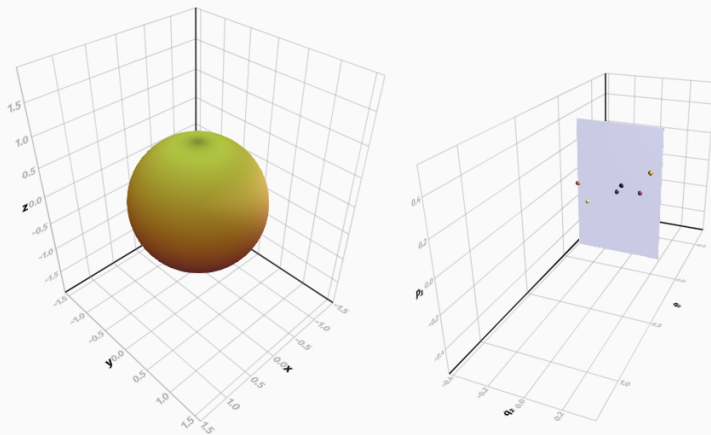
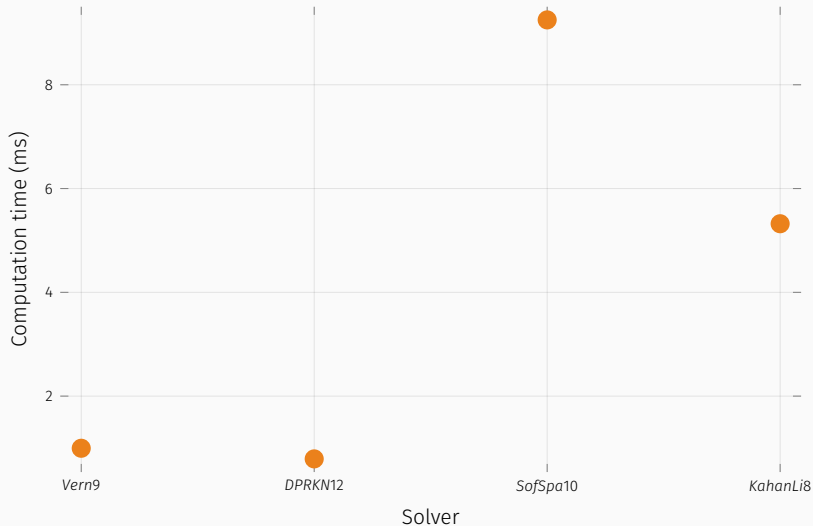


Figure 22: The nucleus and the corresponding trajectory in the phase space for a regular trajectory with  $B = 0.5, E = 0.3$



**Figure 23:** Energy error benchmark for short integration time with rescaling



**Figure 24:** Computational time benchmark for short integration time with rescaling

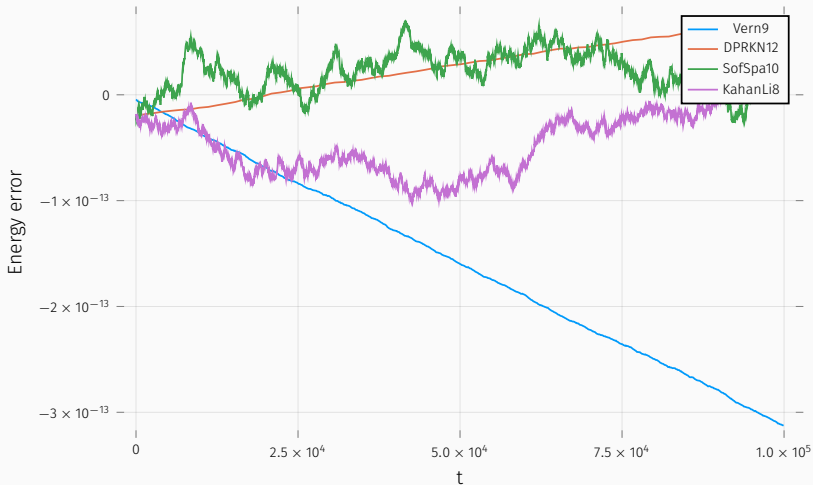


Figure 25: Energy error benchmark for long integration time with rescaling

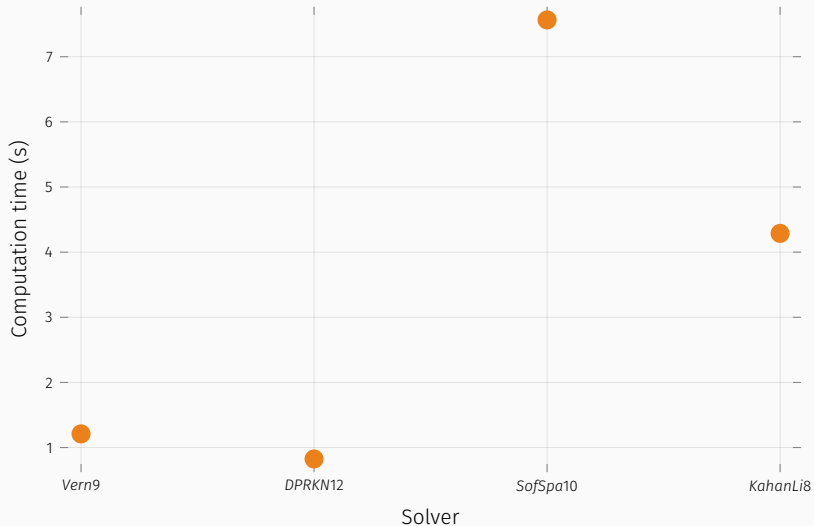


Figure 26: Computational time benchmark for long integration time with rescaling

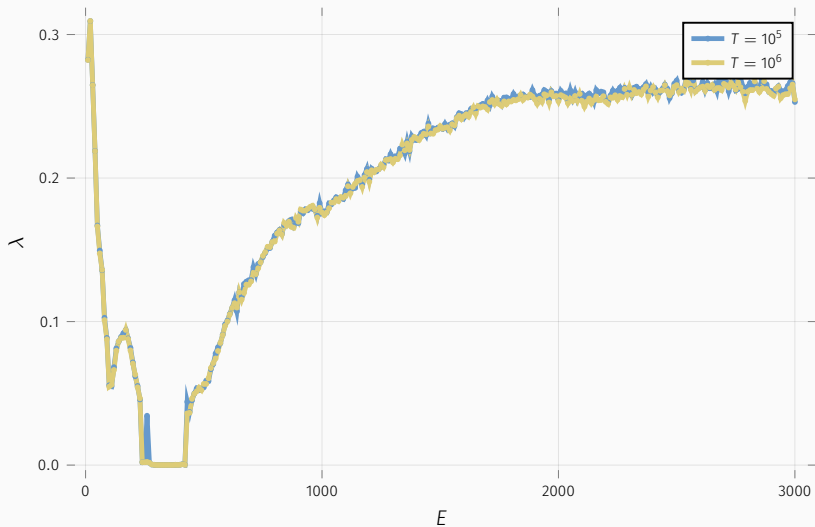


Figure 27: Averaged  $\lambda$  for  $B = 0.5$  and  $E \in (10, 3000)$ .



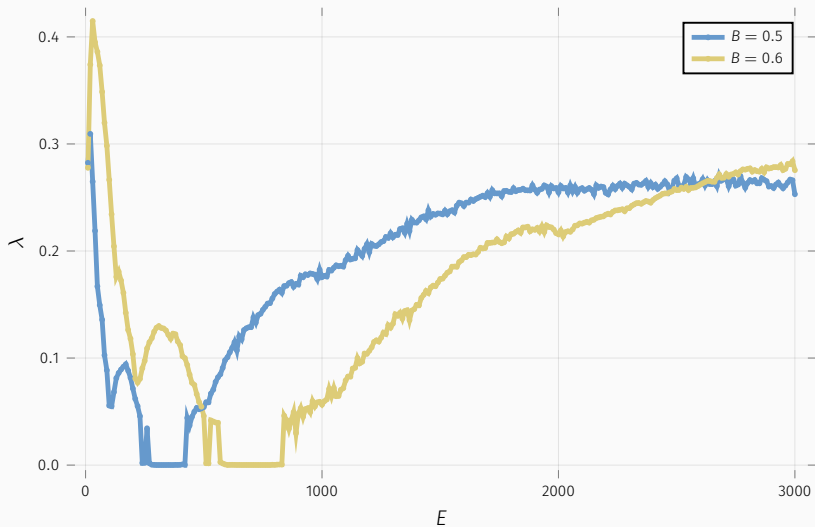


Figure 28: Averaged  $\lambda$  for  $B \in 0.5, 0.6$  and  $E \in (10, 3000)$ .