

# Report for Worksheet 1: Integrators

Markus Baur and David Beyer

November 8, 2019

## Contents

<b>1</b>	<b>Cannonball</b>	<b>1</b>
1.1	Simulating a cannonball . . . . .	1
1.2	Influence of friction and wind . . . . .	1
<b>2</b>	<b>Solar system</b>	<b>3</b>
2.1	Simulating the solar system with the Euler scheme . . . . .	3
2.2	Integrators . . . . .	5
2.2.1	Velocity Verlet algorithm . . . . .	5
2.2.2	Verlet algorithm . . . . .	6
2.2.3	Implementation of the symplectic Euler algorithm . . . . .	6
2.2.4	Implementation of the Velocity Verlet algorithm . . . . .	6
2.3	Long-term stability . . . . .	6

## 1 Cannonball

### 1.1 Simulating a cannonball

The function for the (constant) gravitational force can be implemented in the following way in Python, it returns the force as a vector:

```
def force(mass, gravity):
    return np.array([0.0, -mass * gravity])
```

The Euler scheme is implemented in this fashion:

```
def step_euler(x, v, dt, mass, gravity, f):
    x += v * dt
    v += f / mass * dt
    return x, v
```

It is crucial to first update the positions  $\mathbf{x}$  and then update the velocities  $\mathbf{v}$ . To simulate the cannonball until it hits the ground, a while-loop is used.

A plot of the simulated trajectory  $y(x)$  is shown in Figure 1. As we would expect from the analytical solution

$$y(x) = x - \frac{gx^2}{2v_0^2}, \quad (1)$$

the trajectory looks like a parabola.

### 1.2 Influence of friction and wind

To account for friction, the function which calculates the force is modified in the following way:

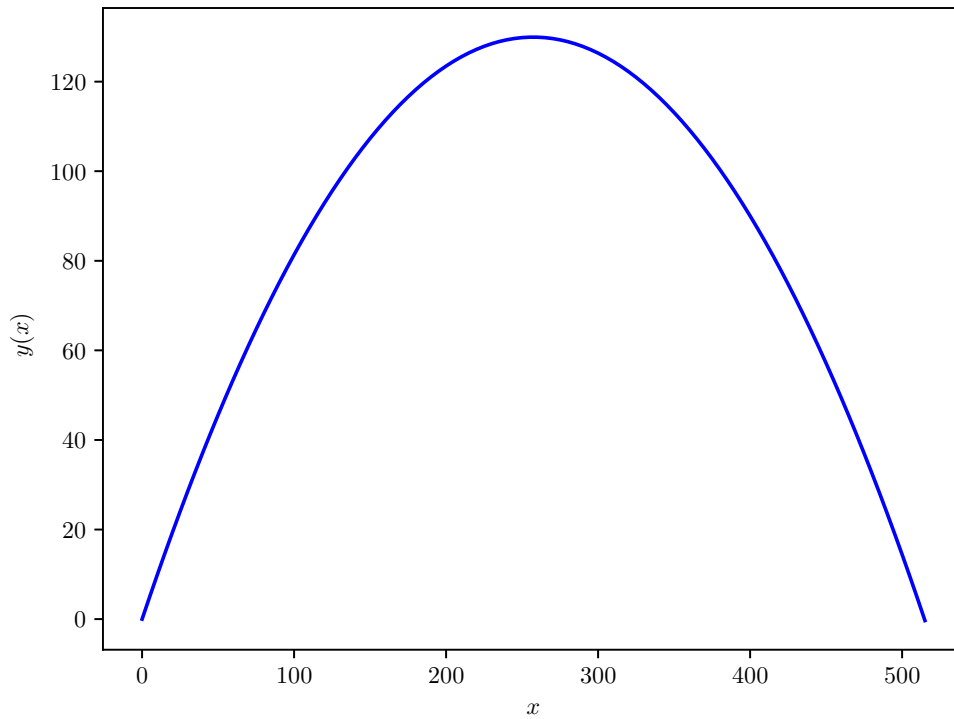


Figure 1: Simulated trajectory  $y(x)$  for the system without friction. The used integrator is the Euler scheme.

```
def force(mass, gravity, v, gamma, v_0):
    ret = np.array([0.0, -mass * gravity])
    ret -= gamma * (v - v_0)
    return ret
```

Because the force is now not constant anymore along the trajectory (it varies with the velocity  $\mathbf{v}$ ), the function for the Euler step has to be modified as well:

```
def step(x, v, dt, mass, gravity, gamma, v_0):
    f = force(mass, gravity, v, gamma, v_0)
    x += v * dt
    v += f / mass * dt
    return x, v
```

In contrast to the previous task, the force is now evaluated every time the Euler step is called.

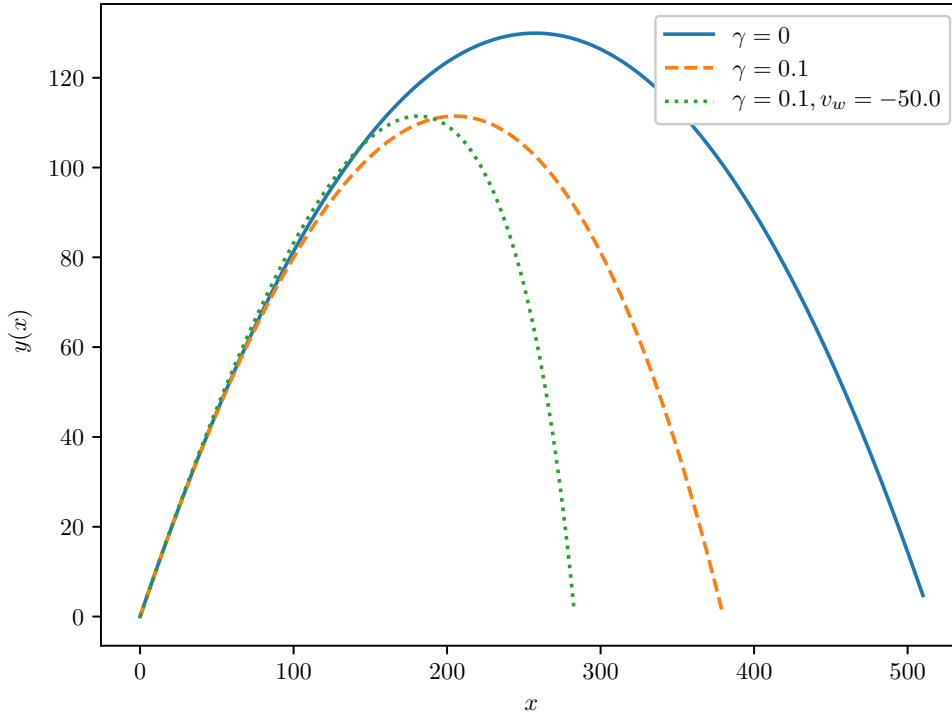


Figure 2: Simulated trajectories  $y(x)$  for a friction coefficient  $\gamma = 0.1$  and different values of the wind speed  $v_w$ . The used integrator is the Euler scheme.

Figure 2 shows simulated trajectories for the system without friction as well as for a system with friction coefficient  $\gamma = 0.1$  and wind speeds of  $v_w = 0, -50.0$ . Comparing the trajectory of the cannonball without friction to the ones with friction, we can easily see that the friction leads to both a decreased maximum height  $y_{\max}$  and a decreased range  $x_{\max}$ . This is caused by the dissipation of energy through the non-conservative friction force. We can also identify that the negative wind speed  $v_w = -50.0$  leads to an even larger decrease of the range  $x_{\max}$ , this is also expected, because the friction force is proportional to the relative velocity of the air and the cannonball. Because the wind blows in the  $x$ -direction only, the maximum height  $y_{\max}$  is the same for both cases.

In Figure 3 we see multiple trajectories for  $\gamma = 0.1$  and different wind speeds. As the wind speed becomes more negative, the range of the cannonball becomes smaller because the friction increases. For  $v_w \approx 200$  it hits the ground at its starting point.

## 2 Solar system

### 2.1 Simulating the solar system with the Euler scheme

The gravitational force between two particles is calculated using this function:

```
def force(r_ij, m_i, m_j, g):
    return - g * m_i * m_j * r_ij / np.linalg.norm(r_ij) ** 3
```

To calculate all the forces on all the particles, the following function is used.

```
def forces(x, masses, g):
    ret = np.zeros(x.shape)

    for i in range(len(x)):
        for j in range(i + 1, len(x)):
```

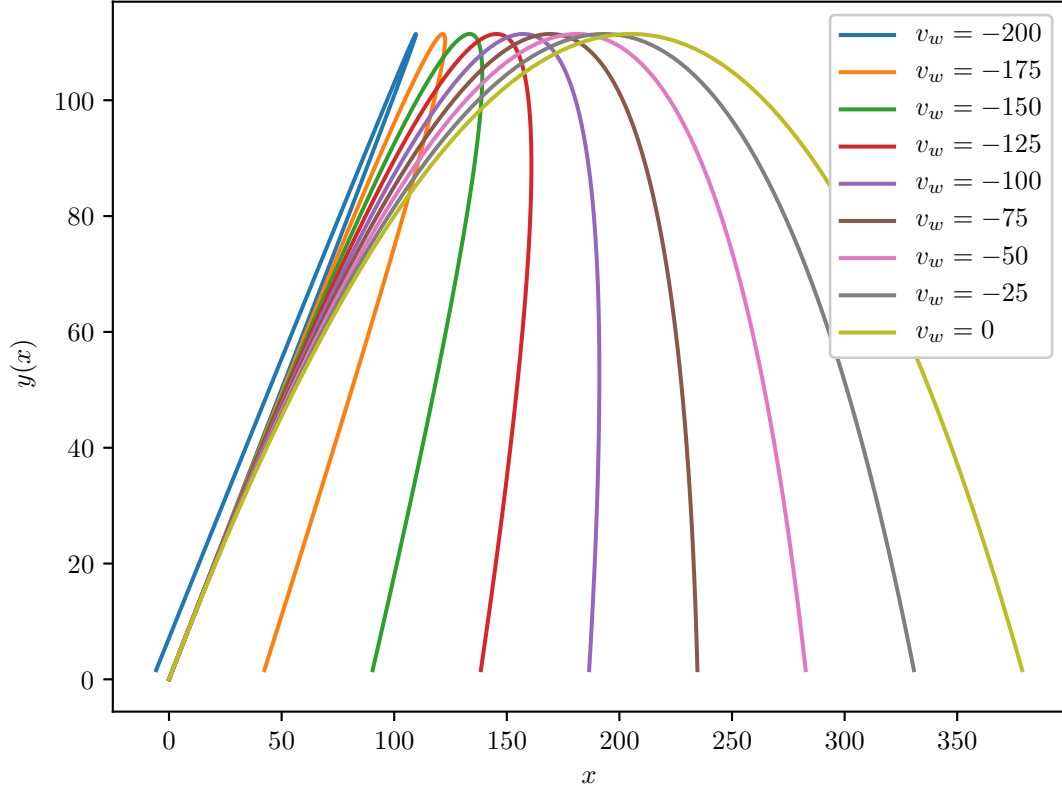


Figure 3: Simulated trajectories  $y(x)$  for different values of the friction coefficient  $\gamma$  and the wind speed  $v_w$ . For  $v_w = -200$ , the cannonball hits the ground closely to the starting point. The used integrator is the Euler scheme.

```

        f = force(x[j] - x[i], masses[i], masses[j], g)
        ret[i] -= f
        ret[j] += f

    return ret

```

The function returns an array which contains all forces.

A small modification of the Euler step is necessary because the particles have different masses:

```

def step_euler(x, v, dt, mass, g):
    x += v * dt
    f = forces(x, mass, g)
    # calculate acceleration per coordinate dimension
    f[:, 0] /= mass
    f[:, 1] /= mass
    v += f * dt
    return x, v

```

In simulations with many particles, the computationally most expensive step is the evaluation of the forces  $\mathbf{F}_{ij}$ . For a system of  $n$  particles, the complexity of this task scales like  $\mathcal{O}(n^2)$  because the pairwise forces  $\mathbf{F}_{ij}$  have to be evaluated for every of the  $n(n-1)$  possible combination of  $i, j$  (or at least half of them using Newton's third law).

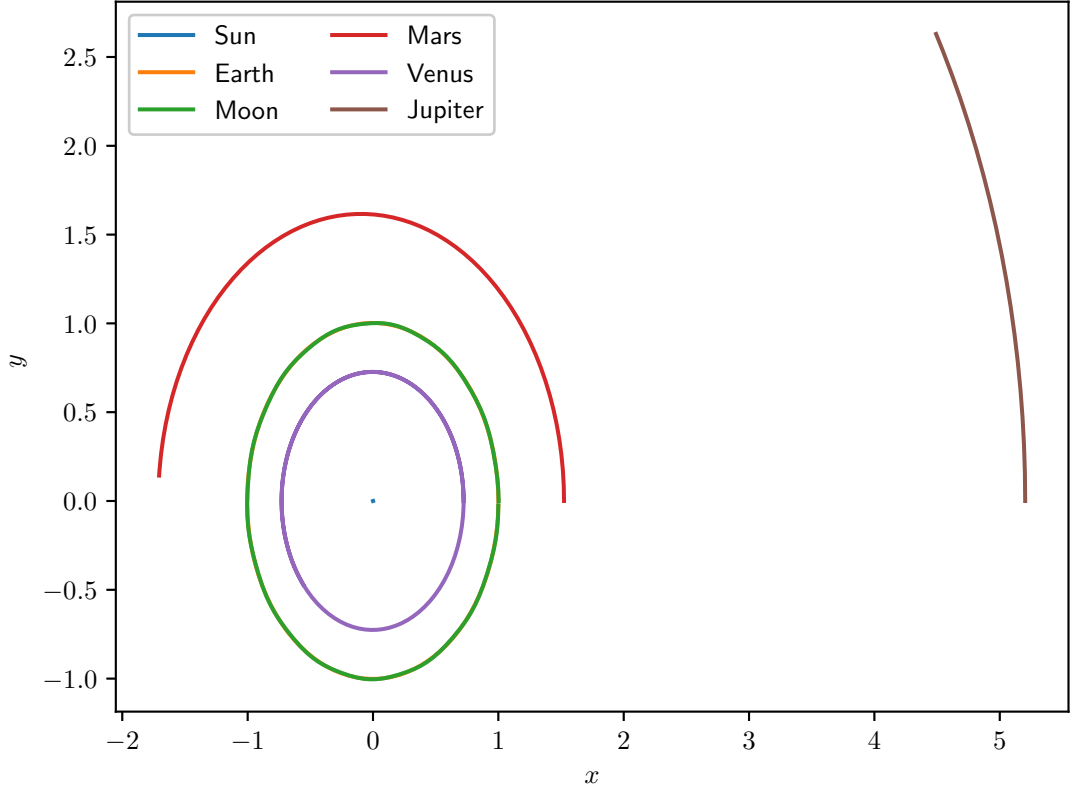


Figure 4:

## 2.2 Integrators

### 2.2.1 Velocity Verlet algorithm

The Velocity Verlet algorithm can be derived in the following way: For the positions  $\mathbf{x}$ , we perform a Taylor expansion up to the second order in  $\Delta t$ :

$$\begin{aligned}\mathbf{x}(t + \Delta t) &= \mathbf{x}(t) + \dot{\mathbf{x}}(t) \cdot \Delta t + \frac{\ddot{\mathbf{x}}(t)}{2} \cdot \Delta t^2 + \mathcal{O}(\Delta t^3) \\ &= \mathbf{x}(t) + \mathbf{v}(t) \cdot \Delta t + \frac{\mathbf{a}(t)}{2} \cdot \Delta t^2 + \mathcal{O}(\Delta t^3)\end{aligned}\tag{2}$$

For the velocities  $\mathbf{v}$ , we also perform a Taylor expansion up to the second order:

$$\begin{aligned}\mathbf{v}(t + \Delta t) &= \mathbf{v}(t) + \dot{\mathbf{v}}(t) \cdot \Delta t + \frac{\ddot{\mathbf{v}}(t)}{2} \cdot \Delta t^2 + \mathcal{O}(\Delta t^3) \\ &= \mathbf{v}(t) + \mathbf{a}(t) \cdot \Delta t + \frac{\dot{\mathbf{a}}(t)}{2} \cdot \Delta t^2 + \mathcal{O}(\Delta t^3)\end{aligned}\tag{3}$$

To get an expression for  $\ddot{\mathbf{v}}(t)$ , we perform a Taylor expansion of  $\dot{\mathbf{v}}(t + \Delta t)$ :

$$\dot{\mathbf{v}}(t + \Delta t) = \dot{\mathbf{v}}(t) + \ddot{\mathbf{v}}(t) \cdot \Delta t + \mathcal{O}(\Delta t^2)\tag{4}$$

We can solve this expression for  $\ddot{\mathbf{v}}(t)$ :

$$\ddot{\mathbf{v}}(t) = \frac{\dot{\mathbf{v}}(t + \Delta t) - \dot{\mathbf{v}}(t)}{\Delta t} + \mathcal{O}(\Delta t) = \frac{\mathbf{a}(t + \Delta t) - \mathbf{a}(t)}{\Delta t} + \mathcal{O}(\Delta t)\tag{5}$$

Plugging  $\ddot{\mathbf{v}}(t)$  into Equation 3, we get

$$\mathbf{v}(t + \Delta t) = \mathbf{v}(t) + \mathbf{a}(t) \cdot \Delta t + \frac{\mathbf{a}(t + \Delta t) + \mathbf{a}(t)}{2} \cdot \Delta t^2 + \mathcal{O}(\Delta t^3)\tag{6}$$

Equation 2 and Equation 6 define the Velocity Verlet algorithm.

### 2.2.2 Verlet algorithm

The Verlet algorithm which was presented in the lecture can also be derived from the Velocity Verlet algorithm. Using Equation 2, we can write  $\mathbf{x}(t)$  as

$$\dots \tag{7}$$

As we can see in ..., the Verlet algorithm needs the positions  $\mathbf{x}(t)$  and  $\mathbf{x}(t - \Delta t)$  to calculate the position  $\mathbf{x}(t + \Delta t)$ . However, the initial conditions only include the position at exactly one point in time  $t_0$ , this means that the initial conditions are not sufficient to solve the problem with the Verlet algorithm. To get the value of  $\mathbf{x}(t_0 - \Delta t)$  we have to use another integrator like the Euler scheme, this results in a bigger error.

### 2.2.3 Implementation of the symplectic Euler algorithm

### 2.2.4 Implementation of the Velocity Verlet algorithm

## 2.3 Long-term stability