

driftsim: Simulation of radial drift in PPDs

Omkar H. Ramachandran

November 11, 2017

Mathematical Formulation

Lagrangian formulation without dissipation

For a dust particle with mass m at a position (r, θ) from the central star of mass M encountering no dissipation, the Lagrangian can be written as follows:

$$L = T - U = \frac{1}{2}m\dot{r}^2 + \frac{1}{2}mr^2\dot{\theta}^2 + \frac{GMm}{r} \quad (1)$$

We know from the Euler-Lagrange equation that

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_i} \right) = \frac{\partial L}{\partial q_i} \quad (2)$$

where q_i refers to the i th generalized coordinate. Substituting equation 1 into 2, we get the following:

$$m\ddot{r} = -\frac{GMm}{r^2} + mr\dot{\theta}^2 \quad (3)$$

$$mr^2\ddot{\theta} = 0 \quad (4)$$

Naive implementation of dissipation due to gas

Assuming that no energy dissipation occurs along r , the only damping factor is caused by the difference in relative velocities between gas and dust. This leaves equation 3 unchanged while altering 4 as follows:

$$mr^2\ddot{\theta} = -br\dot{\theta}_{rel}$$

which implies that

$$\ddot{\theta} = -\frac{\alpha}{mr}\dot{\theta} \quad (5)$$

where $\alpha = b(1 - \beta)$ where $\beta \in [0, 1]$ is the ratio of velocities between gas and dust. Note that equation 5 depends on the mass of the dust object, thus implying that more massive dust particles will drift slower than less massive ones at the same r .