

6G Midband Technology and Spectrum

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Current allocations and uses in 7.125 - 24 GHz in the U.S.

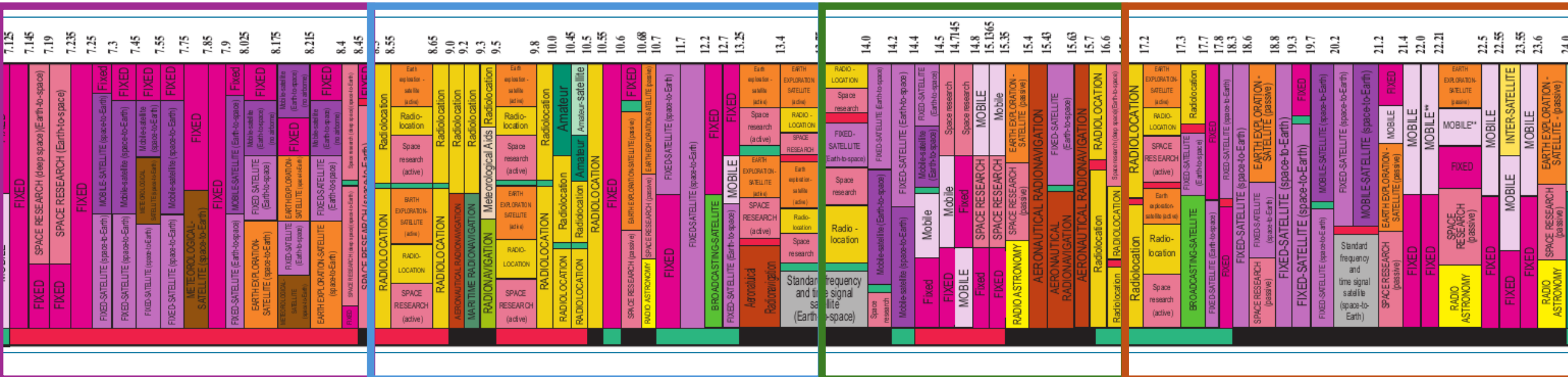
- 7.125 - 24 GHz band is split into four sub-categories for ease of presentation and analysis

7.125 - 8.5 GHz

8.5-13.75 GHz

13.75-17.1 GHz

17.1 - 24 GHz



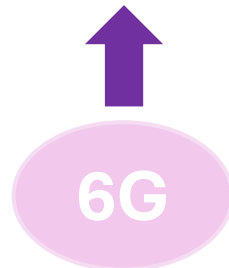
7.125 - 8.5 GHz

	7125	7145	7190	7235	7250	7300	7375	7450	7550	7750	7900	8025	8175	8215	8400	8450
F	Fixed (FS)															
F			Earth Exploration Satellite(EESS)				Mobile (MS)						Earth Exploration Satellite(EESS)			
N													EESS			
F					Fixed satellite (FSS)							Fixed satellite (FSS)				
F			MetSat					MetSat		MetSat		MetSat				
F					Mobile-Satellite (MSS)							Mobile-Satellite (MSS)				
F		Space Research (SRS)					Maritime Mobile Satellite (MMSS)									Space Research (SRS)
N		SRS														SRS

- FS, FSS and MSS are the largest current allocations
- Approximately 20% of FS use is by the Department of Defense (DoD),
- The 8400 – 8500 MHz band is allocated for SRS (DL).
- Other uses may not be ubiquitous and hence perhaps more amenable to sharing.

7.125 - 8.5 GHz

	7125	7145	7190	7235	7250	7300	7375	7450	7550	7750	7900	8025	8175	8215	8400	8450
F						Fixed (FS)										
F			Earth Exploration Satellite(EESS) ↑				Mobile (MS)					Earth Exploration Satellite(EESS) ↓				
N												EESS ↓				
F				Fixed satellite (FSS) ↓						Fixed satellite (FSS) ↑						
F			MetSat					MetSat		MetSat		MetSat				
F				Mobile-Satellite (MSS) ↓						Mobile-Satellite (MSS) ↑						
F		Space Research (SRS)					Maritime Mobile Satellite (MMSS)								Space Research (SRS)	
N		SRS													SRS	



8.5-13.75 GHz

	8.5		8.65		8.85		9.2		9.5		9.9		10		10.4		10.5		10.6		10.7	11.2	11.7	12.1	12.5	12.75	13.4					
N																		FS		Fixed (FS)					Fixed (FS)							
N							EESS											EESS							BSS	Mobile (MS)	EESS					
	Radiolocation (RLS)																(RAS)															RLS
																				FSS							FSS					
							SRS			MetSat									SRS										SRS			
				A R		A R																						A R				

- RLS is the single largest allocation.
- ~2500 MHz in 10.7 – 13.25 GHz is allocated for non-federal use, of which 12.2 - 13.25 GHz is already under consideration by the FCC for sharing with other services.
- 13.25 - 13.75 may also be available for sharing with limited restrictions.

8.5-1375 GHz

	8.5		8.65		8.85		9.2		9.5		9.9		10		10.4		10.5		10.6		10.7		11.2		11.7		12.1		12.5		12.75		13.4				
N																			FS		Fixed (FS)						Fixed (FS)										
N							EESS												EESS								BSS		Mobile (MS)		EESS						
	Radiolocation (RLS)																		(RAS)																		
																					FSS								FSS								
							SRS				MetSat									SRS													SRS				
				A R		A R																										A R					

- Passive EESS in the nearby 10.6–10.7 GHz band

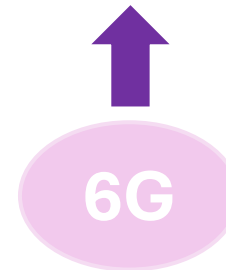
13.75-17.1 GHz

	13.75	14	14.2	14.3	14.4	14.47	14.5	14.71	14.75	14.8	15.13	15.35	15.4	15.43	15.63	15.7	16.6
					Fixed (FS)												
					Mobile(MS)												
F											EESS	EESS					
												RAS					
	RLS												Radiolocation (RLS)				
								Space Researc (SRS)					Aeronautical Radionavigation (AR)			SRS	
			Mobile-Satelite (MSS)														
	Fixed Satellite (FSS)																

- RLS, SRS, FS and MS are the largest allocations in this part of the spectrum.
- 200 MHz between 14.2 – 14.4 GHz is not allocated for federal use.

13.75-17.1 GHz

	13.75	14	14.2	14.3	14.4	14.47	14.5	14.71	14.75	14.8	15.13	15.35	15.4	15.43	15.63	15.7	16.6
					Fixed (FS)												
					Mobile(MS)												
F											EESS	EESS					
												RAS					
	RLS												Radiolocation (RLS)				
							Space Research (SRS)						Aeronautical Radionavigation (AR)				SRS
			Mobile-Satellite (MSS)														
	Fixed Satellite (FSS)																



17.1-24 GHz

	17.1	17.3	17.8	18.1	18.4	18.8	19.7	20.2	21.4	22.21	22.55	23.55
					Fixed (FS)			MS S	Fixed (FS)			
			Fixed (FS)				MSS		Fixed (FS)			
									Mobile(MS)			
										RAS	RAS	RAS
	RLS			MetSat							ISS	
		SRS				SRS			SRS		SRS	SRS
		EES S				EES S			EES S		EES S	EES S
		FSS	Fixed Satellite (FSS)									
			Fixed Satellite (FSS)									

- 100 MHz spectrum in 17.7 – 17.8 GHz is not allocated for federal use.
- 2,200 MHz of spectrum in 17.8 – 18.6 GHz and 18.8 – 20.2 GHz may be practical for coexistence if used for earth stations rather than user devices.

Potential Shared Spectrum Bands

7.125-8.5 GHz

- Federal Fixed
- Federal Fixed satellite
- Federal Mobile Satellite

10.7-13.25 GHz

- Non-Federal Satellite Services

14.0-14.2 GHz

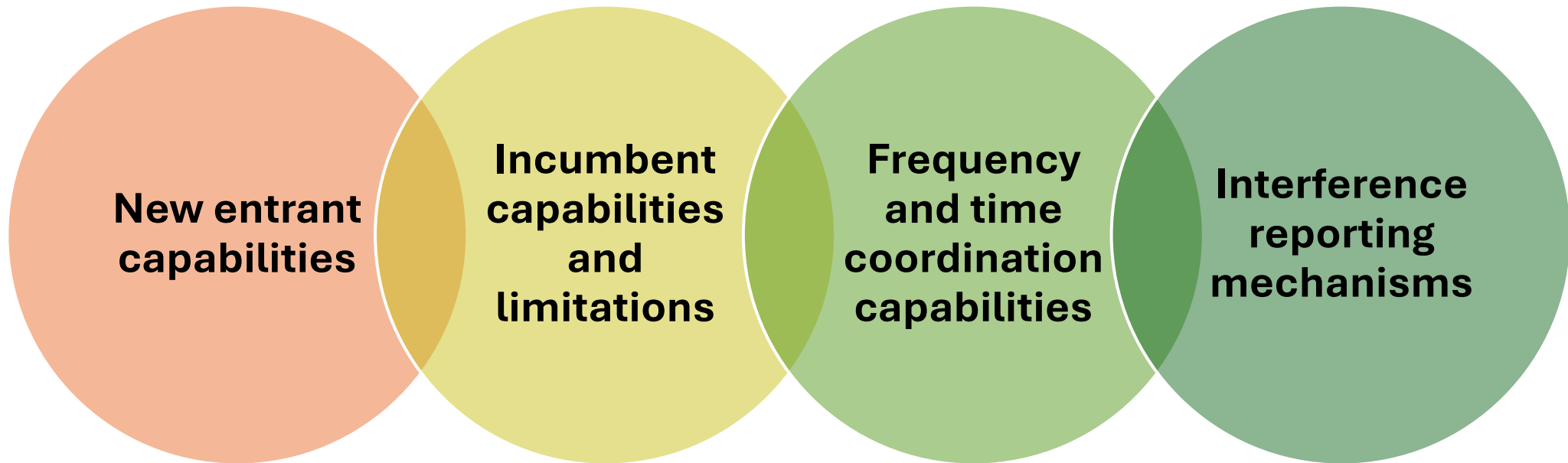
- Space Research

17.8-18.6 GHz and
18.8-20.2 GHz

- Federal Satellite
- Additional analysis need to be done with regard to commercial satellite

Fundamental sharing framework

Success of spectrum sharing depends upon a few basic components:



They have to work in harmony with each other for the overall benefit

6G Spectrum

6G demand : **Peak data rate** 200 Gbit/s

$$c = \text{Bandwidth} \cdot \text{Layer} \cdot \log_2(1 + \text{SNR})$$

**More
Bandwidth?**

5G

Current 5G candidates:

- 3.3-3.8 GHz (400 MHz)
- 6.4-7.1 GHz (700 MHz)

6G

Upper Mid

WRC-23 candidates:

- 4.4-4.8 GHz (400 MHz)
- 7.1-8.4 GHz (600-1275 MHz)
- 14.8-15.35 GHz (550 MHz)

6G will only have slightly more bandwidth than 5G

The need for 6G

6G demand : **Peak data rate** 200 Gbit/s

6G will only have slightly more bandwidth than 5G

$$c = \textit{Bandwidth} . \textit{Layer} . \log_2(1 + \textit{SNR})$$



**More
Bandwidth?**



**Bigger
MIMO**



Increase the
Frequency and
number of the
antenna

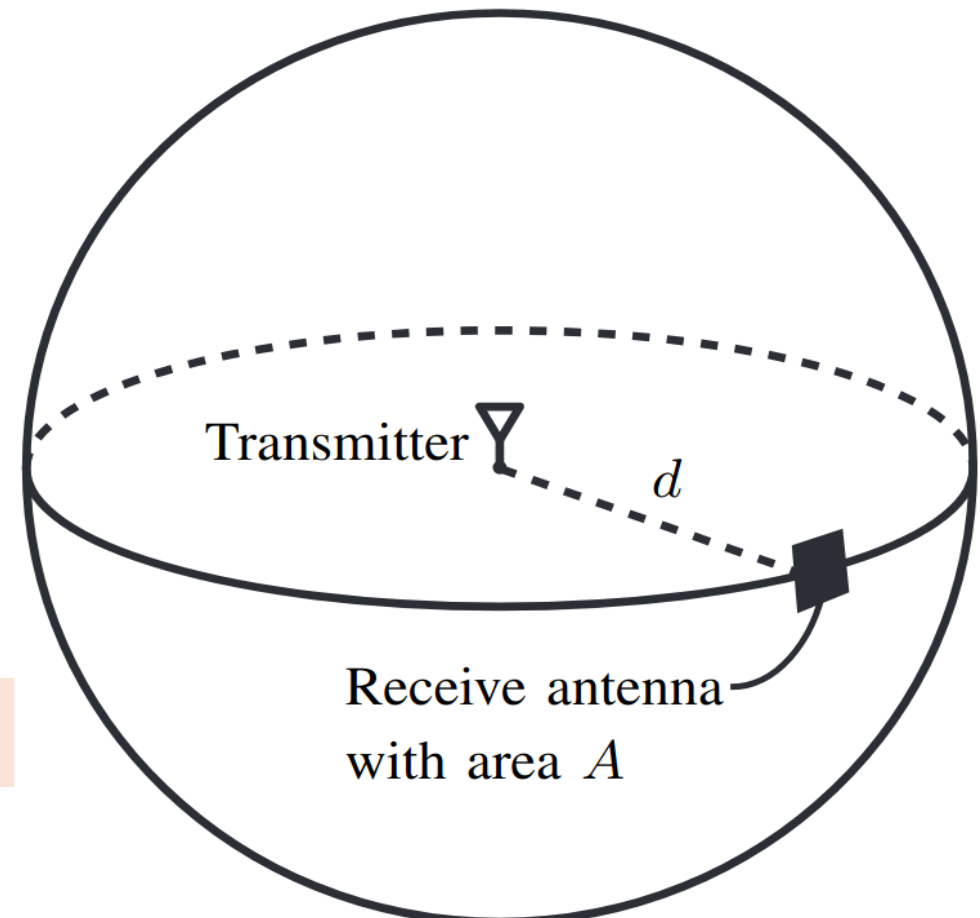
What Happens to Path Loss at High Frequencies?

- Friis' formula: the received power from an isotropic transmitter antenna

$$P_{rx} = \frac{A}{4\pi d^2} P_{tx} \quad A = \frac{\lambda^2}{4\pi}$$

$$p_r = \frac{\lambda^2}{(4\pi d)^2} \cdot p_t$$

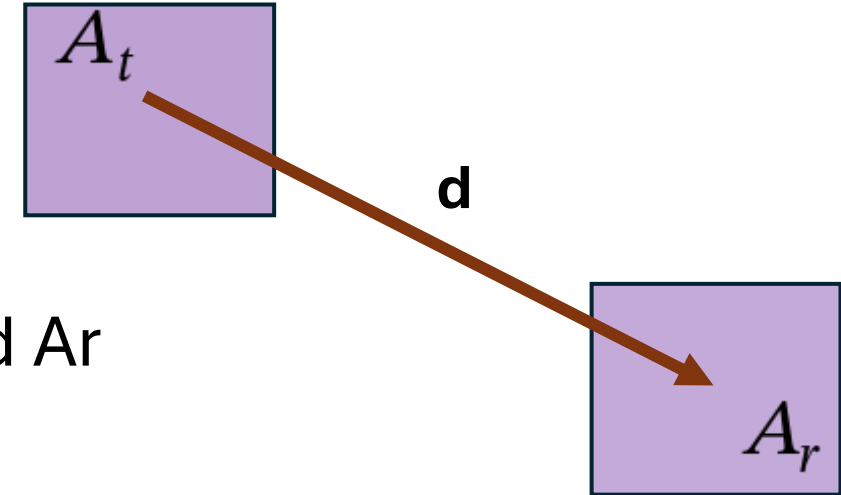
- Higher frequency Less power



What Happens to Path Loss at High Frequencies?

- A reasonable design goal is to keep the free-space path-loss constant when frequency increased

$$P_{rx} = \frac{A}{4\pi d^2} P_{tx}$$



- Received power with array areas A_t and A_r

$$p_r = \frac{A_r A_t}{d^2 \lambda^2} \cdot p_t$$

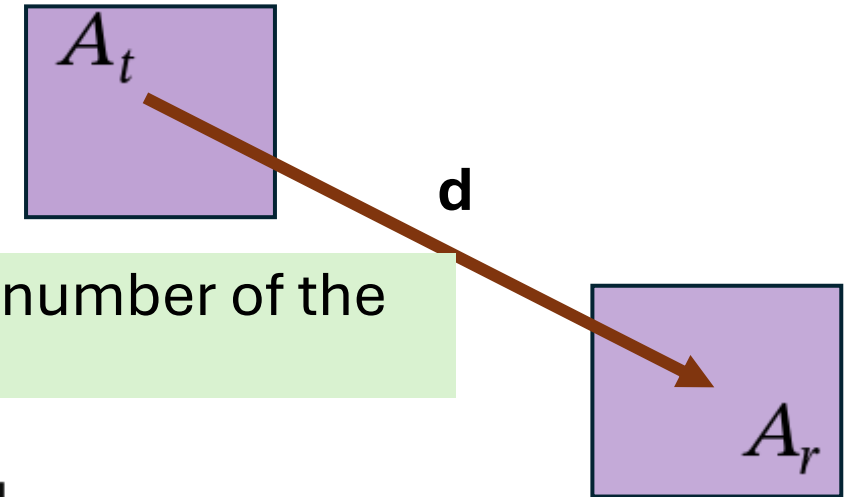
• Higher frequency More power!!

• We need more antenna to fill the area

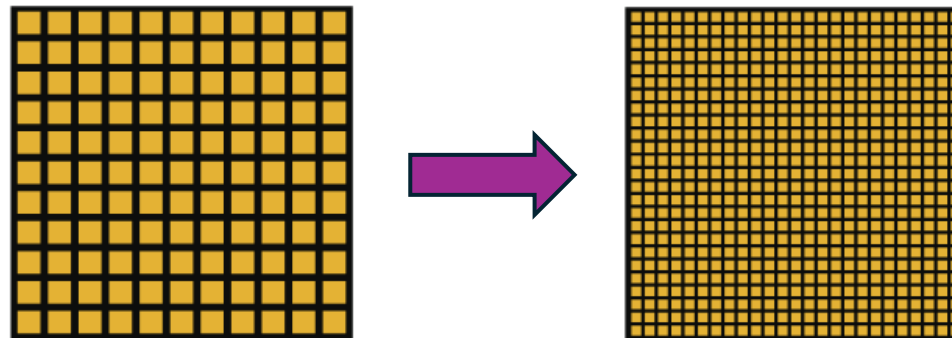
What happens to pathloss in high frequency?

$$p_r = \frac{A_r A_t}{d^2 \lambda^2} \cdot p_t$$

- Higher frequency More power!!



- Improve the free space pathloss by increasing the number of the antenna but the same aperture area



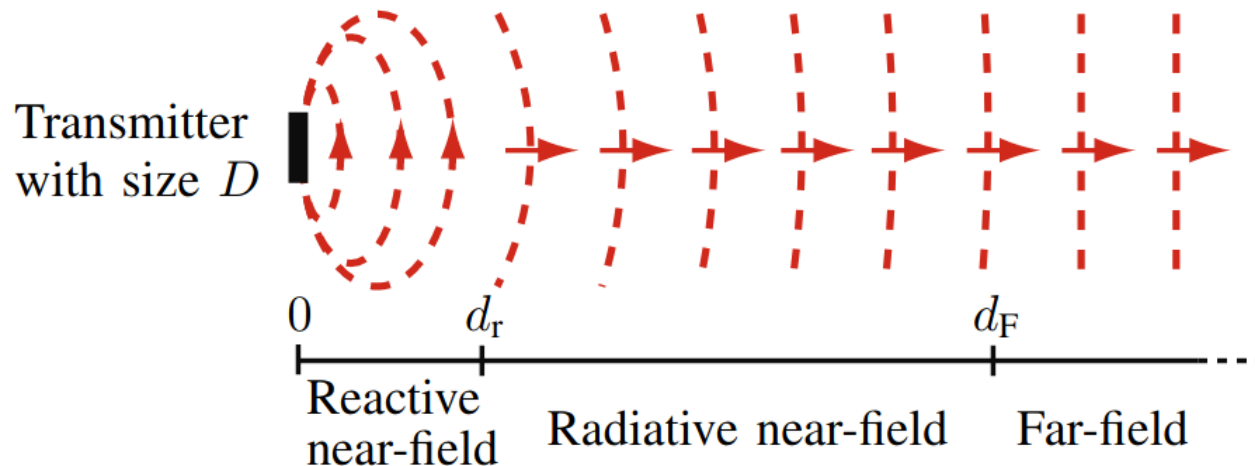
- So what is the problem with mmWave?

What happen if we increase the number of the antennas?

- Just more links between the transmitter and receiver?
- Just a bigger channel matrix?
- Just more complexity?

What happen if we increase the number of the antennas?

- Just more links between the transmitter and receiver?
- Just a bigger channel matrix?
- Just more complexity?



**Nearfield
Communication**

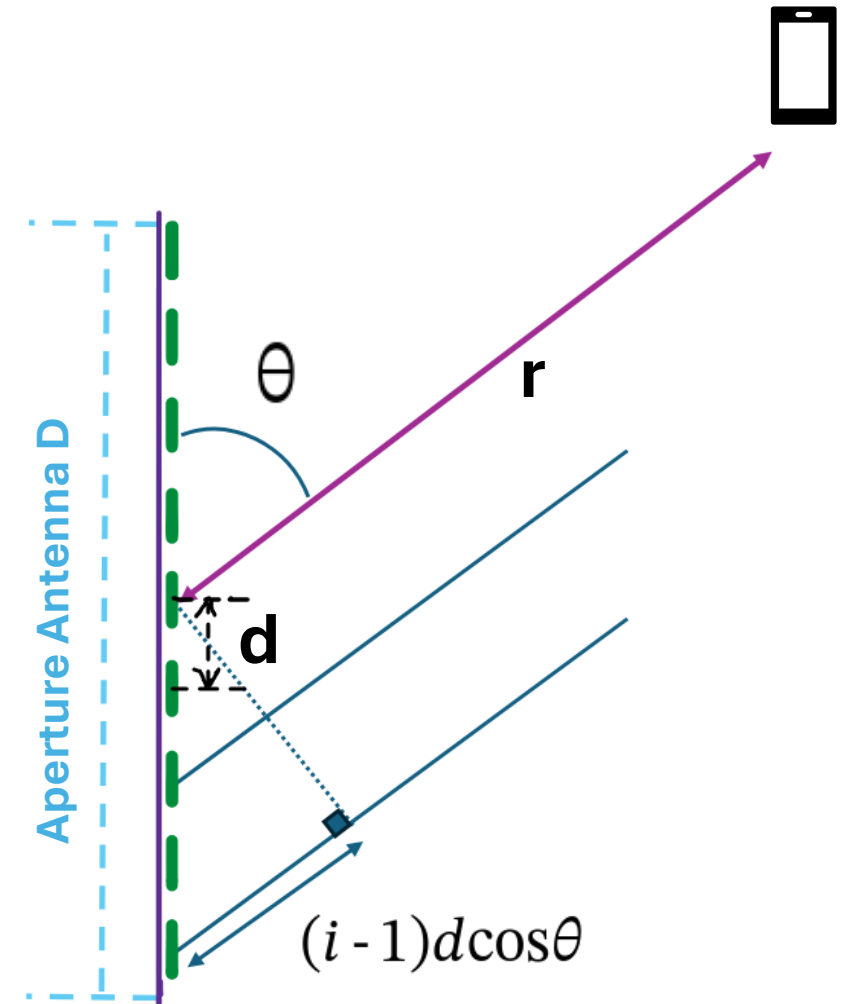
SIMO Channel Model

- Considering a simple SIMO channel with LOS

$$h_1 = \frac{\lambda\sqrt{G}}{4\pi r} e^{(-j\frac{2\pi}{\lambda}r)}$$

$$h_i = \frac{\lambda\sqrt{G}}{4\pi r} e^{-(j\frac{2\pi}{\lambda}(r + (i-1)d\cos\theta))}$$

$$H = h_1 \left[1, e^{-j\frac{2\pi}{\lambda}d\cos(\theta)}, \dots, e^{-j\frac{2\pi}{\lambda}(N-1)d\cos\theta} \right]^T$$



SIMO Channel Model in Near-field

- Considering a simple SIMO channel with LOS

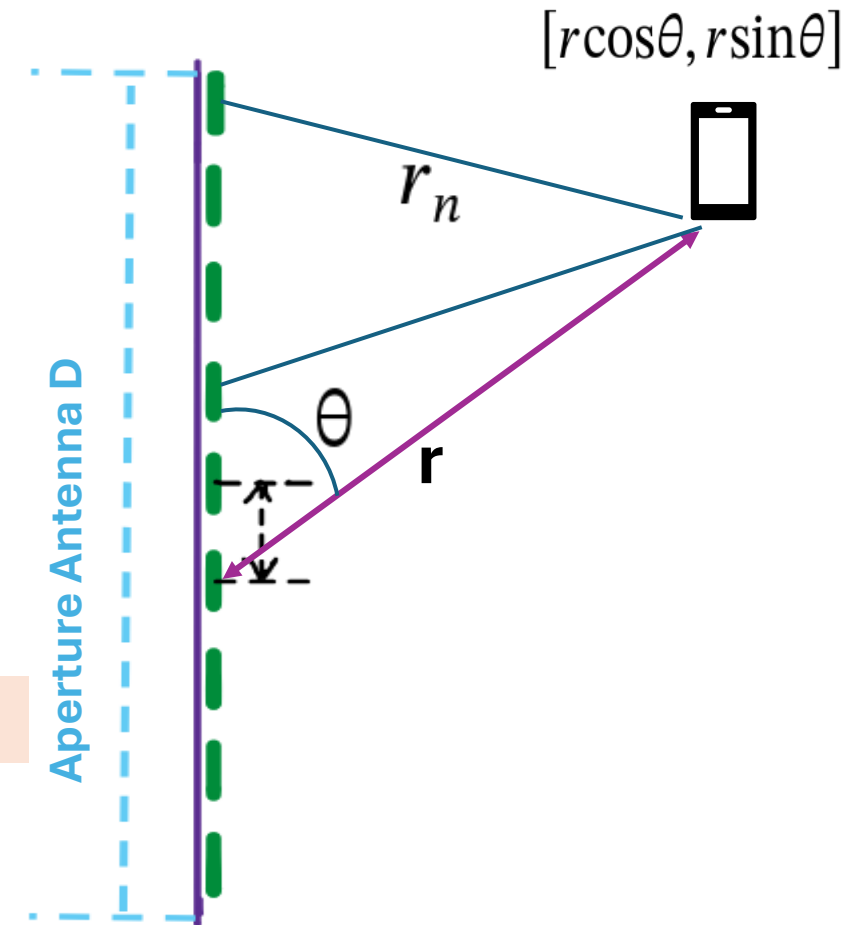
$$h_1 = \frac{\lambda\sqrt{G}}{4\pi r} e^{(-j\frac{2\pi}{\lambda}r)}$$

$$h_n = \frac{\lambda\sqrt{G}}{4\pi r_n} e^{-j\frac{2\pi}{\lambda}r_n}$$

$$r_n = \sqrt{r^2 + n^2 d^2 - 2r n d \cos\theta}$$

$$r_n = r \left(1 + \frac{n^2 d^2}{2r^2} - \frac{n d \cos\theta}{r} \right)$$

- Taylor approximation



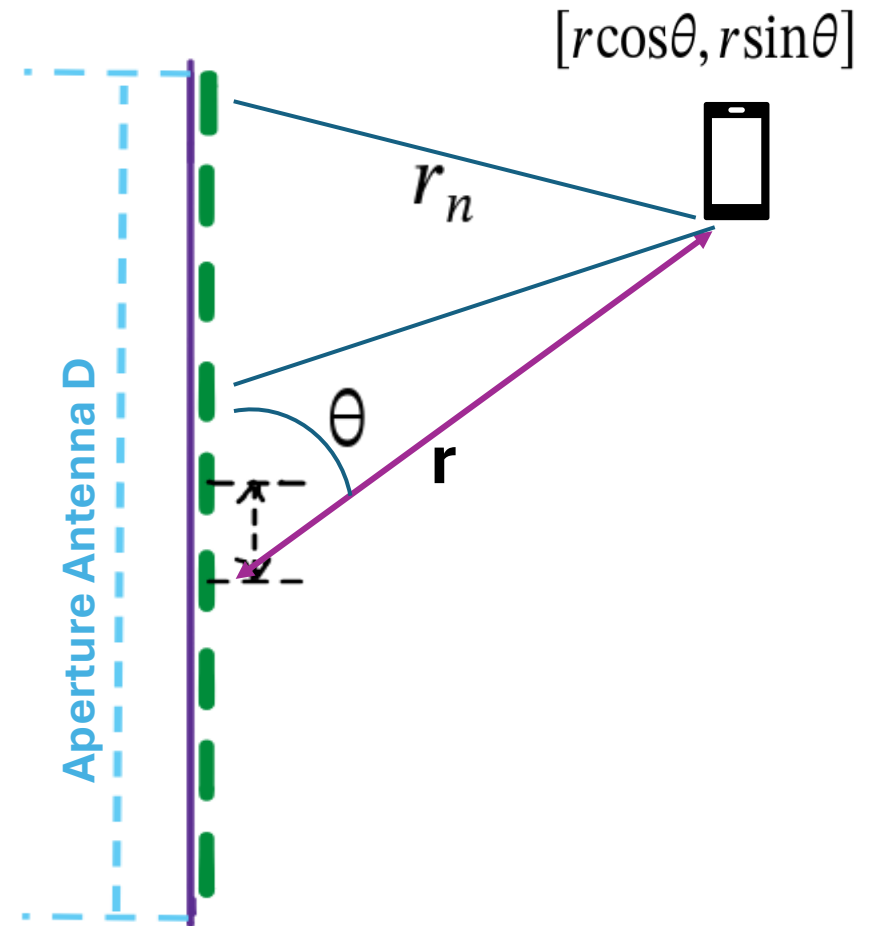
SIMO Channel Model in Near-field

$$r_n = r \left(1 + \frac{n^2 d^2}{2r^2} - \frac{nd \cos \theta}{r} \right)$$

$$h_n = h_1 e^{-j \frac{2\pi}{\lambda} \left(\frac{n^2 d^2}{2r} - nd \cos \theta \right)}$$

$$\frac{2\pi}{\lambda} \left(\frac{n^2 d^2}{2r} \right)$$

For the enough large r (far-field)
equal to **0**



SIMO Channel Model in Nearfield

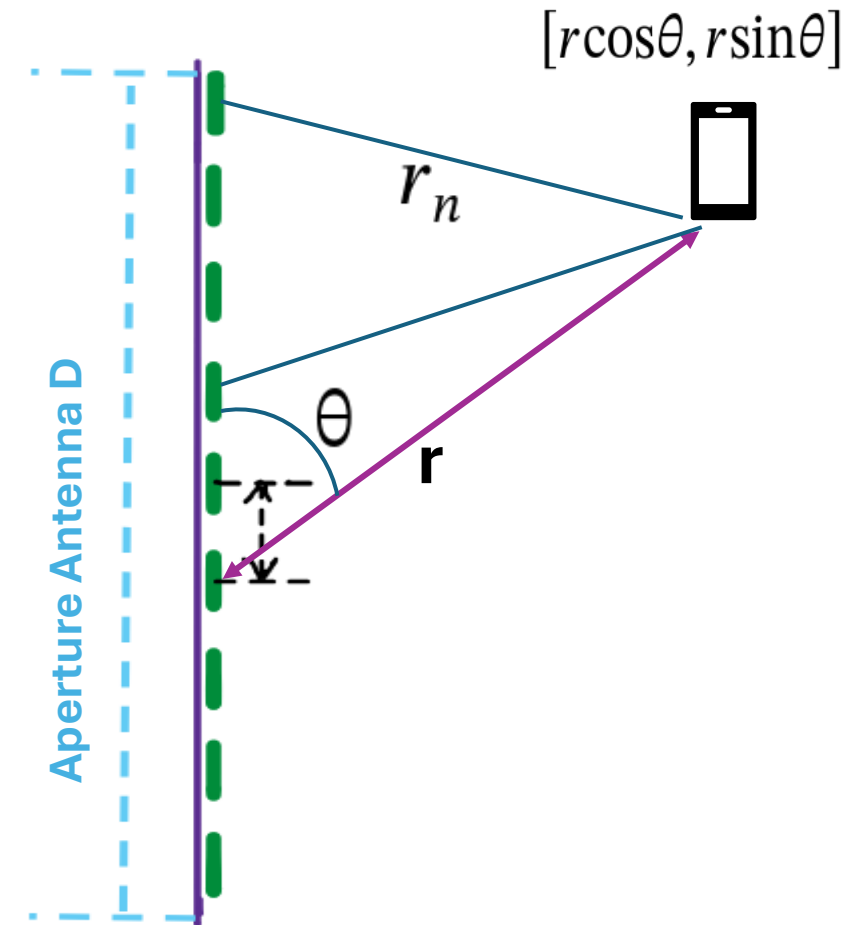
$$r_n = r \left(1 + \frac{n^2 d^2}{2r^2} - \frac{nd \cos \theta}{r} \right)$$

$$h_n = h_1 e^{-j \frac{2\pi}{\lambda} \left(\frac{n^2 d^2}{2r} - nd \cos \theta \right)}$$

$$\frac{2\pi}{\lambda} \left(\frac{n^2 d^2}{2r} \right)$$

For the enough large r (far-field)
equal to **0**

But how much large?



Phase Variation in the Nearfield of the Antenna Array

$$\frac{2\pi}{\lambda} \left(\frac{n^2 d^2}{2r} \right)$$

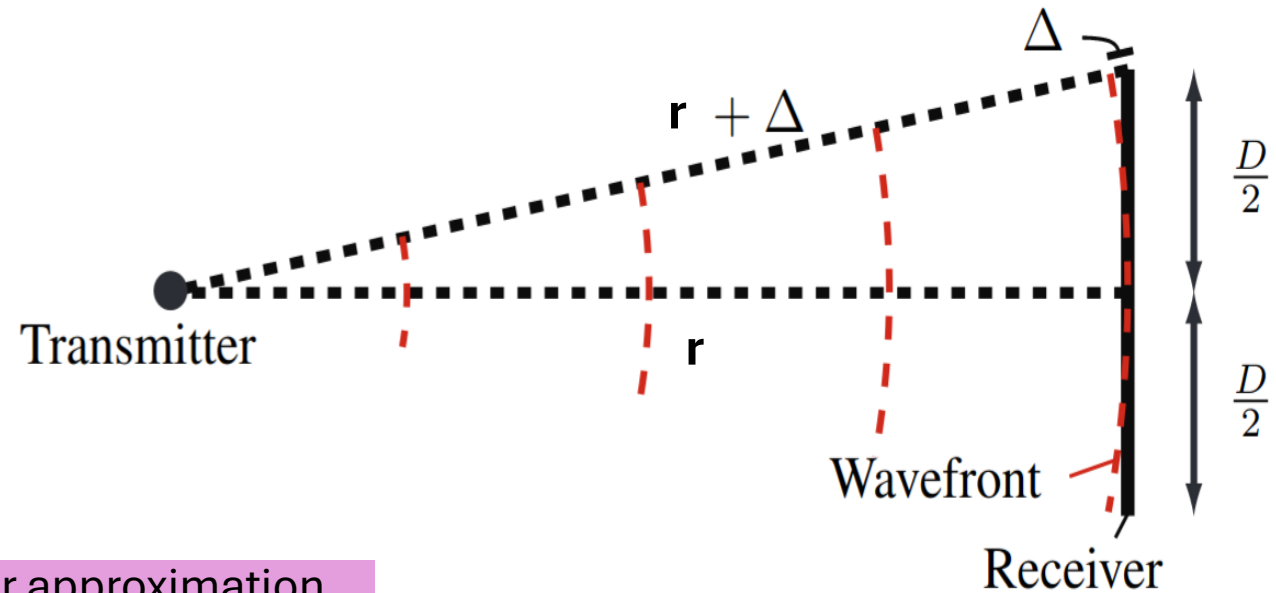
What is the largest phase shift for a given r ?

• Taylor approximation

$$\Delta = \sqrt{r^2 + \frac{D^2}{4}} - r = r \sqrt{1 + \frac{D^2}{4r^2}} - r \approx \frac{D^2}{8r}$$

$$d_F = \frac{2D^2}{\lambda}$$

Fraunhofer array distance



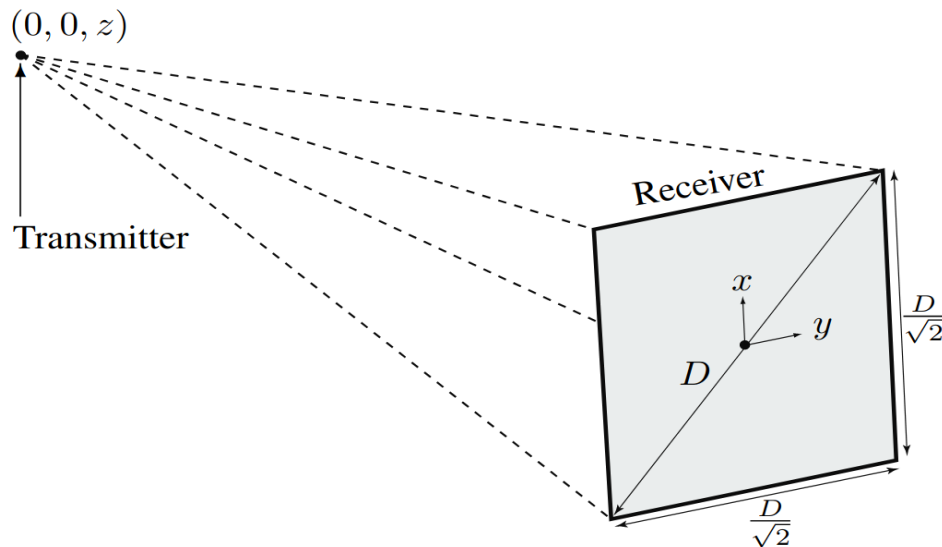
$$\frac{2\pi}{\lambda} \frac{D^2}{8r} = \frac{\pi}{8}$$

$$\cos\left(\frac{\pi}{8}\right) \approx .92 \approx 1$$

Phase Variation in the Nearfield of the Antenna Array

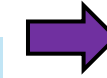
Fraunhofer array distance

$$d_F = \frac{2D^2}{\lambda}$$



Planar wavefront

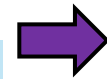
phase variation
depends on **the incident angle.**



Far-field

Curved wavefront

phase variation
depends on **the incident angle and distance.**



Nearfield

Power Variation in Nearfield of the Antenna Array

- The spherical curvature also gives rise to power variations over the receive antenna.

The power differences

$$\frac{r^2}{(r + \Delta)^2} \approx \frac{r^2}{\left(r + \frac{D^2}{8r}\right)^2}$$

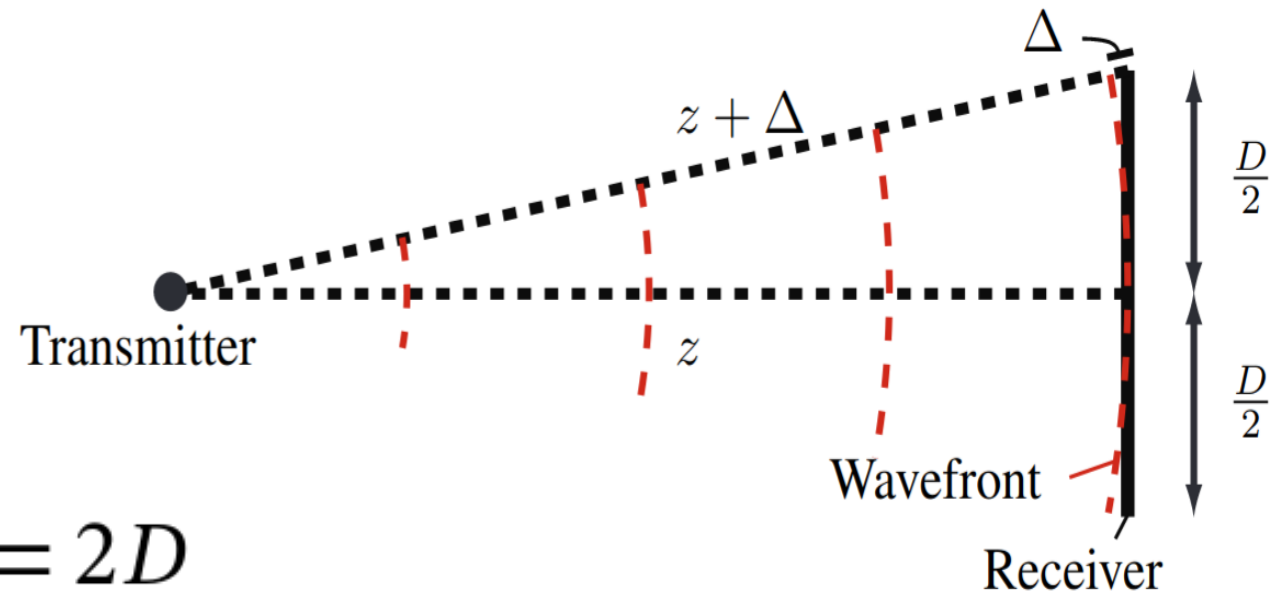
Björnson distance

$$r = d_B = 2D$$

$$d_F = d_B \frac{D}{\lambda}$$



Phase variations are more prevalent



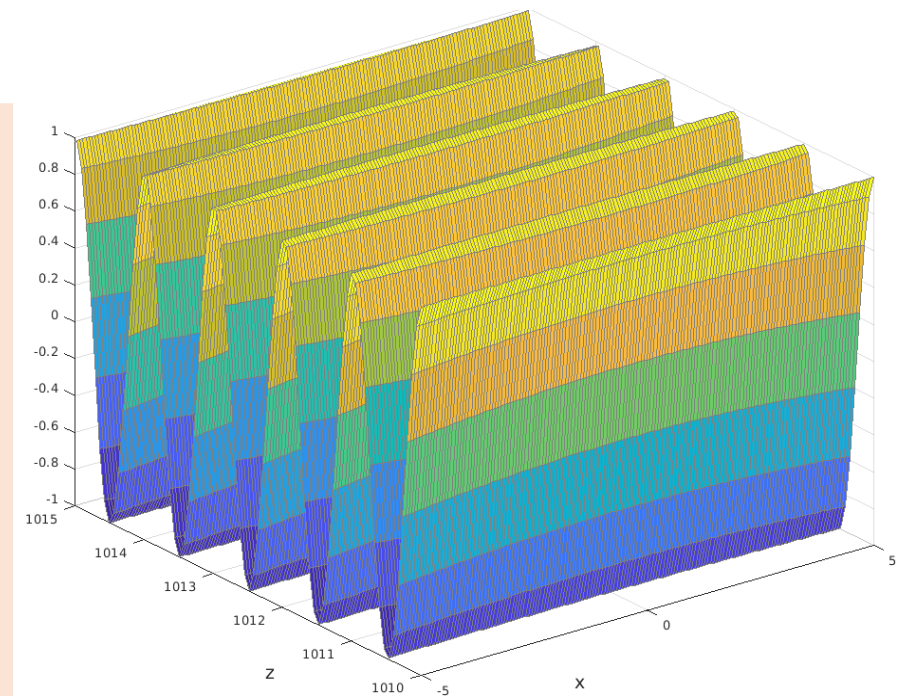
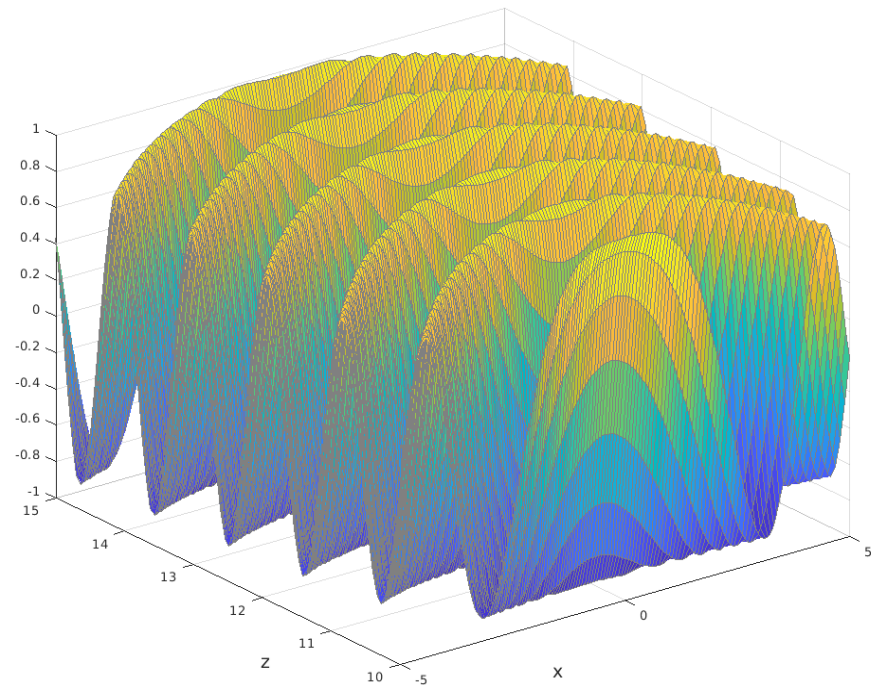
Nearfield of the Antenna Array

- Fresnel region and is characterized by the fact that the amplitude variations can be neglected, but not the phase variations.

$$d_B < r < d_F$$

Example : Fraunhofer distance

0.5×0.5 , 3 GHz = 10 m
 1×1 m, 15 GHz = 200m
 1×1 m, 30 GHz = 400 m

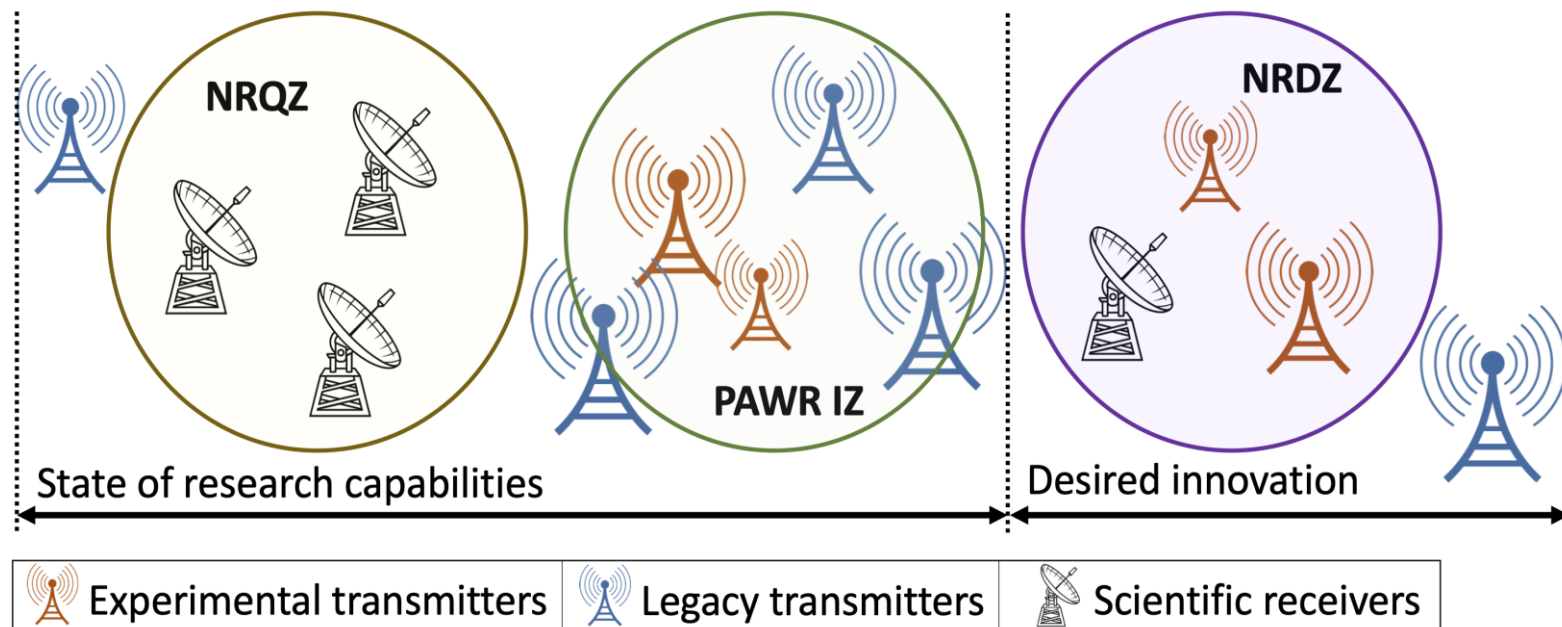


National Radio Dynamic Zones (SII-NRDZ)

- What is the Dynamic Zone?

- Why It Matters

- Allows radio waves to transmit freely.
- Advanced technologies for **spectrum sharing** and coexistence.
- Spectrum is a limited resource.
- Traditional methods are insufficient.



<https://www.cs.albany.edu/nrdz-ra/>

How to detect aliens' signals?

NASA's **Project Cyclops** in 1971



Detect signals traveling up to
1000 light-years

Absorption
Noise
Distinguishable from natural
sources

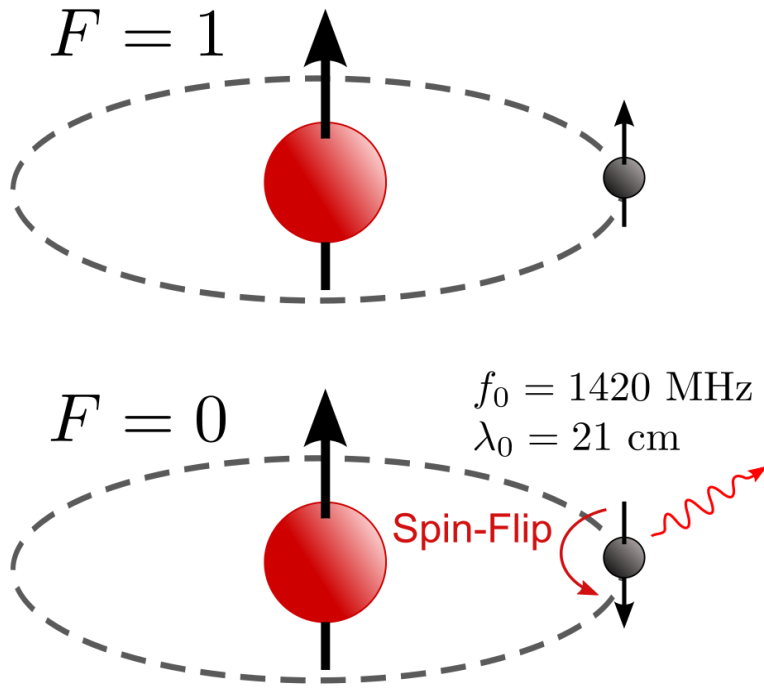


Microwaves
1 mm – 1 m

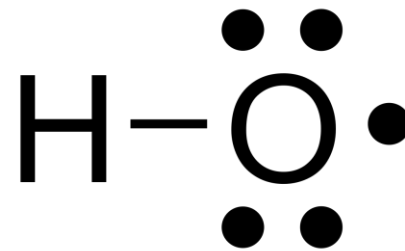
Waterhole

There is a magical narrow gate in Microwaves

https://en.wikipedia.org/wiki/Hydrogen_line



https://en.wikipedia.org/wiki/Hydroxyl_radical

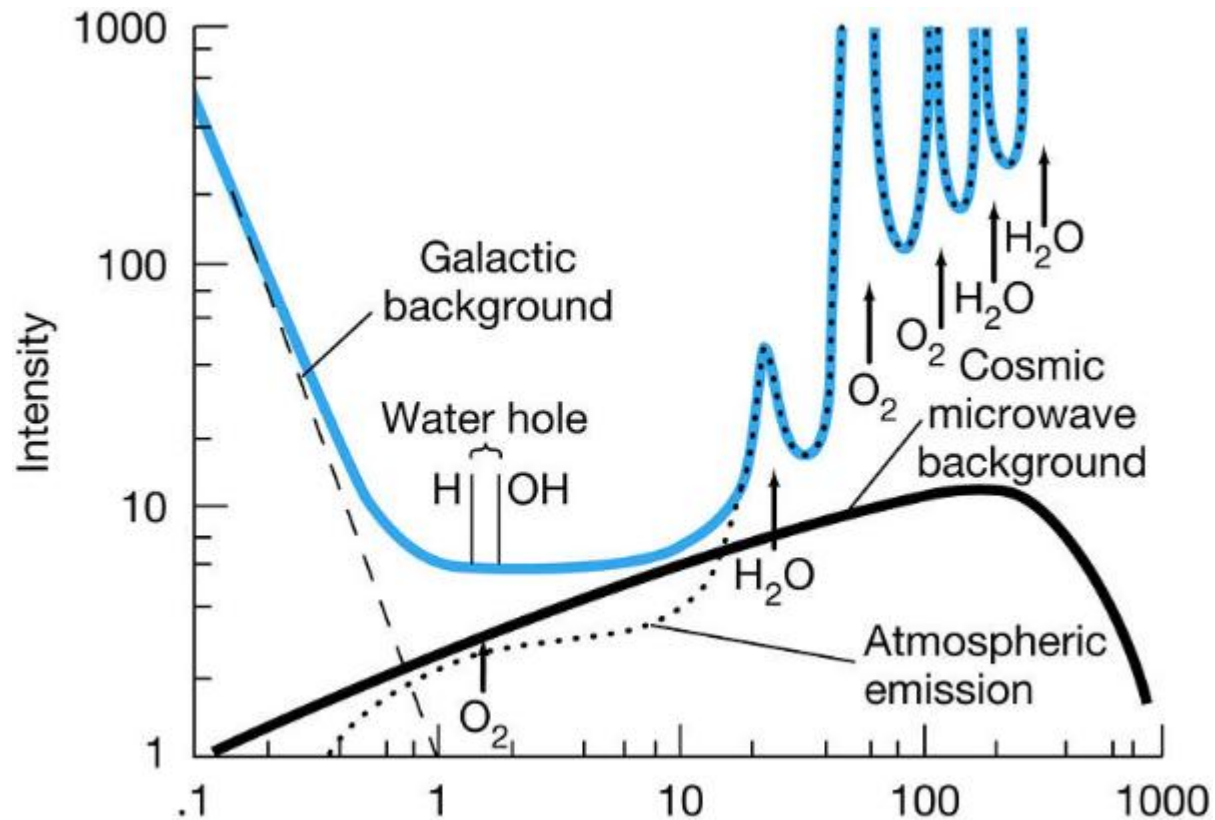


$f = 1662 \text{ MHz}$
 $\lambda = 18 \text{ cm}$

Water hole
18cm-21cm

The 242 MHz area, between the hydrogen line (1420 MHz) and the hydroxyl lines (1662 MHz), are the quietest part of the spectrum

Waterhole



Waterhole
18cm-21cm

1420-1662 MHz

Oliver:

“What more **poetic** place could there be for **water-based life** to seek its kind than the age-old meeting place for all species: the water hole”