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K. Bock Name > Major Atmospheric Gamma Imaging Cherenkov Telescope project (MAGIC) Site > http://wwwmagic.mppmu.mpg.de Contact > rkb@mail.cern.ch **OBJECTIVE** \* The objective of the MAGIC gamma telescope is to develop an classification prediction system for distinguishing between gamma rays and hadrons in the field of astrophysics. \* The classification of Gamma and Hadron rays is crucial for unraveling the mysteries of the universe, and the MAGIC Telescope dataset provides valuable information for this purpose. \* This study utilizes machine learning techniques such as Logistic Regression, Decision Tree, Random Ensemble Techniques to predict and differentiate between these high-energy particles. \* By applying these advanced algorithms, our aim is to create a reliable and accurate prediction model that can effectively classify Gamma and Hadron rays. INTRODUCTION **GAMMA & HADRON RAYS** \* Gamma Rays: Gamma rays are the highest energy form of electromagnetic radiation. They originate from various astrophysical sources such as supernovae, pulsars, and black holes. Gamma rays carry valuable information about these extreme cosmic events, allowing scientists to study the most energetic processes in the universe. \* Hadron Rays: Hadron rays are particles composed of quarks, such as protons and neutrons. They are abundant in cosmic ray showers that result from interactions between high-energy particles and the Earth's atmosphere. Hadrons can produce signals that mimic gamma rays, making their distinction crucial for accurate astrophysical observations. MOTIVATION BEHIND CLASSIFICATION \* Distinguishing between gamma and hadron rays helps us gain insights into the most energetic and exotic phenomena occurring in the universe. \* By separating these two types of particles, scientists can study the processes associated with gamma rays, such as the birth and death of stars, cosmic explosions, and the behavior of matter in extreme conditions. \* Discriminating between gamma and hadron rays enables scientists to filter out background noise and focus on the genuine gamma-ray signals. \* This distinction ensures that observations and measurements are not contaminated by hadronic signals that can mimic gamma-ray signatures. IMPORTING LIBRARIES In [1]: | import warnings import numpy as np import pandas as pd import sweetviz as sv import catboost as cb import seaborn as sns from scipy import stats import matplotlib.pyplot as plt warnings.filterwarnings("ignore") from xgboost import XGBClassifier from catboost import CatBoostClassifier from sklearn.impute import SimpleImputer from imblearn.over\_sampling import SMOTE pd.set option("display.max columns", None) from sklearn.tree import DecisionTreeClassifier from sklearn.feature\_selection import SelectKBest from sklearn.neighbors import KNeighborsClassifier from sklearn.linear\_model import LogisticRegression from imblearn.under sampling import RandomUnderSampler from sklearn.preprocessing import LabelEncoder , StandardScaler from statsmodels.stats.outliers\_influence import variance\_inflation\_factor from sklearn.ensemble import RandomForestClassifier , AdaBoostClassifier , GradientBoostingClassifier from sklearn.model\_selection import train\_test\_split, KFold, cross\_validate, cross\_val\_score , GridSearchCV , K from sklearn.metrics import accuracy\_score, precision\_score, recall\_score, roc\_auc\_score, roc\_curve, r2\_score , IMPORTING DATA In [2]: DF = pd.read\_csv ( 'MagicTelescope.csv' ) Out[2]: ID fLength: fWidth: fSize: fConc: fConc1: fAsym: fM3Long: fM3Trans: fAlpha: fDist: class: 28.7967 16.0021 2.6449 0.3918 0.1982 27.7004 22.0110 -8.2027 40.0920 81.8828 1 g 31.6036 11.7235 2.5185 0.5303 0.3773 26.2722 23.8238 -9.9574 6.3609 205.2610 2 3 3 162.0520 136.0310 4.0612 0.0374 0.0187 -64.8580 -45.2160 76.9600 256.7880 116.7410 g 23.8172 9.5728 2.3385 0.6147 0.3922 27.2107 -6.4633 -7.1513 10.4490 116.7370 30.9205 3.1611 0.3168 0.1832 75.1362 -5.5277 28.5525 21.8393 4.6480 356.4620 g **19015** 19016 19016 21.3846 10.9170 2.6161 0.5857 0.3934 15.2618 11.5245 2.8766 2.4229 106.8258 h **19016** 19017 19017 28.9452 6.7020 2.2672 0.5351 0.2784 37.0816 13.1853 -2.9632 86.7975 247.4560 **19017** 19018 19018 75.4455 47.5305 3.4483 0.1417 0.0549 -9.3561 41.0562 -9.4662 30.2987 256.5166 h **19018** 19019 19019 120.5135 76.9018 3.9939 0.0944 0.0683 5.8043 -93.5224 -63.8389 84.6874 408.3166 **19019** 19020 19020 187.1814 53.0014 3.2093 0.2876 0.1539 -167.3125 -168.4558 31.4755 52.7310 272.3174 h 19020 rows × 13 columns **DESCRIPTION OF VARIABLES** 1. fLength: It represents the major axis of an ellipse in millimeters. The major axis is the longest diameter of the elliptical shape formed by the detected particle. Unit of measurement = mm 2. fWidth: This variable indicates the minor axis of the ellipse in millimeters. The minor axis is the shortest diameter of the elliptical shape. Unit of measurement = millimeters , mm 3. fSize: It represents the logarithm (base 10) of the sum of the content of all the pixels in the image. It provides information about the overall magnitude or intensity of the detected signal. Unit of measurement = millimeters , mm 4. fConc: This variable measures the ratio of the sum of the two highest pixels' values over fSize. It provides insight into the concentration or distribution of pixel values within the image. Unit of measurement = dimensionless ratio 5. fAsym: This variable measures the distance from the highest pixel to the center of the ellipse, projected onto the major axis in millimeters. It provides information about the asymmetry or displacement of the signal with respect to the center. Unit of measurement = millimeters , mm 6. fConc1: It represents the ratio of the value of the highest pixel over fSize. This variable provides information about the dominance or intensity of the highest pixel in the image. Unit of measurement = dimensionless ratio 7. fM3Long: It represents the third root of the third moment along the major axis of the ellipse in millimeters. The third moment is a statistical measure that provides information about the shape or pixel values along the major axis. Unit of measurement = millimeters , mm 8. fM3Trans: This variable indicates the third root of the third moment along the minor axis of the ellipse. It provides information about the shape or distribution of pixel values along the minor axis. Unit of measurement = millimeters , mm 9. fAlpha: It represents the angle (in degrees) between the major axis of the ellipse and a vector connecting the origin (coordinate [0, 0]) to the ellipse's center. It provides information about the orientation or alignment of the detected signal. Unit of measurement = degrees (°) 10. fDist: This variable represents the distance from the origin to the center of the ellipse in millimeters. It provides information about the spatial position or radial distance of the detected signal. Unit of measurement = millimeters , mm 11. class: This variable indicates the class or type of the detected signal. It has two categories: "g" for gamma rays (signal) and "h" for hadron rays (background). This variable is used for classification purposes, distinguishing between gamma rays and hadron rays. TARGET VARIABLE > Class DATA DESCRIPTION DROPPING THE UNIQUE COLUMNS In [3]: DF = DF.iloc [ : , 2 : ] fWidth: fSize: fConc: fConc1: Out[3]: fLength: fAsym: fM3Long: fM3Trans: fAlpha: fDist: class: **0** 28.7967 16.0021 2.6449 0.3918 0.1982 27.7004 22.0110 -8.2027 40.0920 81.8828 g 31.6036 11.7235 2.5185 0.5303 0.3773 26.2722 23.8238 -9.9574 6.3609 205.2610 -45.2160 76.9600 256.7880 **2** 162.0520 136.0310 4.0612 0.0374 0.0187 116.7410 -64.8580 23.8172 9.5728 2.3385 0.6147 0.3922 27.2107 -6.4633 -7.1513 10.4490 116.7370 30.9205 3.1611 0.3168 21.8393 4.6480 356.4620 75.1362 0.1832 -5.5277 28.5525 g 21.3846 10.9170 2.6161 0.5857 0.3934 15.2618 11.5245 2.8766 2.4229 106.8258 **19016** 28.9452 6.7020 2.2672 0.5351 0.2784 37.0816 13.1853 -2.9632 86.7975 247.4560 **19017** 75.4455 47.5305 3.4483 0.1417 0.0549 -9.3561 41.0562 -9.4662 30.2987 256.5166 **19018** 120.5135 76.9018 3.9939 0.0944 0.0683 5.8043 -93.5224 -63.8389 84.6874 408.3166 **19019** 187.1814 53.0014 3.2093 0.2876 0.1539 -167.3125 -168.4558 31.4755 52.7310 272.3174 19020 rows × 11 columns **INFORMATION** Information = pd.DataFrame ( { 'Columns' : DF.columns.tolist ( ) , 'Null Count' : DF.count().tolist ( ) , 'Data Type' : DF.dtypes.tolist ( ) } ).reset index ( ).drop ( columns = [ 'index' ] ) Information Out[4]: **Columns Null Count Data Type** fLength: 19020 float64 fWidth: 19020 float64 fSize: 19020 float64 19020 fConc: float64 19020 fConc1: float64 fAsym: 19020 float64 fM3Long: 19020 float64 7 fM3Trans: 19020 float64 19020 fAlpha: float64 fDist: 19020 float64 10 class: 19020 object STATISTICAL SUMMARY In [5]: Description = pd.DataFrame ( DF.describe ( ) ).transpose ( ) Description Out[5]: count mean min 25% **50% 75%** max **fLength:** 19020.0 53.250154 42.364855 4.2835 24.336000 37.14770 70.122175 334.1770 **fWidth:** 19020.0 22.180966 18.346056 0.0000 11.863800 17.13990 24.739475 256.3820 **fSize:** 19020.0 2.825017 0.472599 1.9413 2.477100 2.73960 3.101600 5.3233 **fConc:** 19020.0 0.380327 0.182813 0.0131 0.235800 0.35415 0.503700 0.8930 **fConc1:** 19020.0 0.214657 0.110511 0.0003 0.128475 0.19650 0.285225 0.6752 **fAsym:** 19020.0 -4.331745 59.206062 -457.9161 -20.586550 4.01305 24.063700 575.2407 **fM3Long:** 19020.0 10.545545 51.000118 -331.7800 -12.842775 15.31410 35.837800 238.3210 **fM3Trans:** 19020.0 0.249726 20.827439 -205.8947 -10.849375 0.66620 10.946425 179.8510 **fAlpha:** 19020.0 27.645707 26.103621 0.0000 5.547925 17.67950 45.883550 90.0000 **fDist:** 19020.0 193.818026 74.731787 1.2826 142.492250 191.85145 240.563825 495.5610 **EXPLORATORY DATA ANALYSIS** VISUAL REPORT In [6]: report = sv.analyze ( DF ) report.show\_notebook ( ) | [ 0%] 00:00 ->... **DataFrame DUPLICATES** 115 2.6 MB RAM **FEATURES** 11 Get updates, docs & report issues here CATEGORICAL Created & maintained by Francois Bertrand **ASSOCIATIONS** NUMERICAL 10 Graphic design by <u>Jean-Francois Hains</u> 0 TEXT DataFrame fLength: VALUES: 19,020 (100%) MAX 334 **RANGE** 330 MISSING: 45.8 95% 140 IOR STD 42.4 70 03 DISTINCT: **18,643** (98%) AVG 53 VAR 1,795 MEDIAN 37 20% 4.97 ZEROES: 01 24 KURT. SKEW 2.01 5% 16 MIN SUM 1.0M 4 -100 VALUES: 19,020 (100%) MAX 256 **RANGE** 256 75% MISSING: 95% 58 IOR 12.9 Q3 25 STD 18.3 DISTINCT: **18,200** (96%) AVG 22 VAR 337 MFDIAN 17 ZEROES: KURT. 98 (<1%) Q1 12 16.8 5% SKEW 3.37 SUM 422k MIN 0 0% -50 0 50 100 ↑ fSize: 30% VALUES: 19,020 (100%) MAX 5.32 **RANGE** 3.38 0.624 MISSING: 95% 3.72 IOR 0.473 03 STD 3.10 20% DISTINCT: **7,228** (38%) AVG 2.83 VAR 0.223 2.74 MEDIAN KURT. ZEROES: 0.727 01 2.48 2.19 SKEW 0.876 5% SUM 53,732 1.94 1.00 2.00 3.00 fConc: VALUES: 19,020 (100%) MAX 0.893 **RANGE** 0.880 0.268 MISSING: 95% 0.734 IOR STD 0.183 0.504 DATA PREPROCESSING DF.columns = DF.columns.str.replace ( ':', '' ) In [7]: DF Out[7]: fLength **fWidth** fConc fConc1 fM3Long fM3Trans fAlpha fDist class fSize fAsym 28.7967 16.0021 2.6449 0.3918 0.1982 27.7004 22.0110 -8.2027 40.0920 81.8828 g 23.8238 -9.9574 31.6036 11.7235 2.5185 0.5303 0.3773 26.2722 6.3609 205.2610 g 162.0520 136.0310 4.0612 0.0374 0.0187 116.7410 -64.8580 -45.2160 76.9600 256.7880 g 23.8172 9.5728 2.3385 0.6147 0.3922 27.2107 -6.4633 -7.1513 10.4490 116.7370 g 75.1362 30.9205 3.1611 0.3168 0.1832 -5.5277 28.5525 21.8393 4.6480 356.4620 g 21.3846 19015 10.9170 2.6161 0.5857 0.3934 15.2618 11.5245 2.8766 2.4229 106.8258 h 19016 28.9452 6.7020 2.2672 0.5351 0.2784 37.0816 13.1853 -2.9632 86.7975 247.4560 47.5305 3.4483 0.1417 75.4455 0.0549 41.0562 -9.4662 30.2987 256.5166 **19018** 120.5135 76.9018 3.9939 0.0944 0.0683 -63.8389 84.6874 408.3166 5.8043 -93.5224 **19019** 187.1814 53.0014 3.2093 0.2876 0.1539 -167.3125 -168.4558 31.4755 52.7310 272.3174 19020 rows × 11 columns In [8]: DF['class'].replace ( { 'g' : 0 , 'h' : 1 } , inplace = True ) DF['class'].value counts ( ) 12332 Out[8]: 6688 Name: class, dtype: int64 DATA VISUALIZATION COMPARISON OF FLENGTH OF HADRON AND GAMMA RAYS \* 'fLength' specifically refers to the major axis of an ellipse that is fitted to the shower image detected by the telescope. \* This feature is measured in millimeters and provides information about the spatial extent of the shower image. \* A "shower image" refers to the pattern of Cherenkov light produced by the electromagnetic shower generated when high-energy gamma rays or cosmic rays interact with the Earth's atmosphere. In [9]: plt.hist ( DF [ DF [ 'class' ] == 1 ] [ 'fLength' ] , color = '#7A6174' , bins = 10 , label = 'Hadron Rays' ); plt.hist ( DF [ DF [ 'class' ] == 0 ] [ 'fLength' ] , color = '#DB5375' , bins = 10 , label = 'Gamma Rays' , al plt.xlabel ( 'Major Axis of an Ellipse in millimeters' ) plt.legend ( ) plt.show ( ) 7000 Hadron Rays Gamma Rays 6000 5000 4000 3000 2000 1000 150 200 50 250 0 300 350 Major Axis of an Ellipse in millimeters \* The 'FLength' (Major Axis of an Ellipse in millimeters) is identified as a distinguishing feature rays and Hadron rays. The significant difference of nearly 4000 mm suggests that this feature could be critical in discriminating between the two types of particles. \* It suggests that the spatial extent of the shower image produced by gamma rays is much larger than that of Hadron rays. This could be due to different interactions of gamma rays and Hadron rays with the Earth's atmosphere. \* Gamma rays are high-energy photons, whereas Hadron rays consist of particles like protons and neutrons. \* When gamma rays enter the Earth's atmosphere, they interact with air nuclei, leading to the creation of an electromagnetic cascade. \* On the other hand, Hadron rays, being composed of charged particles, undergo strong interactions with atomic nuclei, leading to a hadronic cascade. COMPARISON OF FWIDTH OF HADRON AND GAMMA RAYS \* This feature represents the minor axis length of an ellipse that is fitted to the shower image detected by the MAGIC Telescope during the observation of high-energy gamma-ray and background (non-gamma-ray) events. In [10]: plt.hist ( DF [ DF [ 'class' ] == 1 ] [ 'fWidth' ] , color = '#8DAA91' , bins = 10 , label = 'Hadron Rays' ); plt.hist ( DF [ DF [ 'class' ] == 0 ] [ 'fWidth' ] , color = '#453643' , bins = 10 , label = 'Gamma Rays' , alp plt.xlabel ('Minor Axis of an Ellipse in millimeters') plt.legend ( ) plt.show ( ) Hadron Rays Gamma Rays 6000 5000 4000 3000 2000 1000 0 150 250 Minor Axis of an Ellipse in millimeters **INFERENCE** \* The difference in 'Fwidth' is a result of the distinct energy deposition and interactions of gamma rays and Hadron rays with the Earth's atmosphere. \* The wider 'Fwidth' for gamma rays suggests a more extensive and energetic electromagnetic cascade compared to the more confined hadronic cascade for Hadron rays. \* When gamma rays enter the Earth's atmosphere, they interact with air nuclei, initiating an electromagnetic cascade. \* This cascade involves the production of high-energy electrons and positrons, which travel at nearly the speed of light.

**CONTENTS** 

2 INTRODUCTION

• 3 IMPORTING LIBRARIES

4 IMPORTING DATA

5 DATA DESCRIPTION

2.1 GAMMA & HADRON RAYS

4.1 DESCRIPTION OF VARIABLES

2.2 MOTIVATION BEHIND CLASSIFICATION

1 OBJECTIVE

\* These charged particles emit Cherenkov radiation and create an extensive and widespread shower.

\* The Cherenkov radiation emitted by high-energy electrons and positrons in the electromagnetic

\* As a result, the fitted ellipse to this shower image has a larger minor axis ('Fwidth').

\* Each pixel in the camera records the intensity of the Cherenkov light it receives.

\* It represents the concentration of the Cherenkov light within the shower image.

\* The higher the intensity, the higher the energy of the particles detected in that region.

\* 'FSize' is another feature in the dataset and represents the "size of the fitted ellipse in the

\* It is related to the spatial extent of the shower image, typically measured in number of pixels.

\* 'Fconc' is calculated by taking the sum of the two highest pixel values in the shower image and then

\* A higher value of 'Fconc' indicates that a significant portion of the Cherenkov light is concentrated

\* Conversely, a lower value of 'Fconc' suggests a more spread-out distribution of light intensity.

\* It represents the proportion of the Cherenkov light concentration contributed by the pixel with the

\* A higher value of 'Fconc1' suggests that a significant portion of the Cherenkov light is concentrated

with the highest intensity, indicating a highly localized and intense light emission in the shower

\* 'Fconc1' can be important in discriminating between different types of particle events. It may help

plt.hist ( DF [ DF [ 'class' ] == 1 ] [ 'fConc' ] , color = '#DB5375' , bins = 10 , label = 'FConc Hadron Rays'

plt.hist ( DF [ DF [ 'class' ] == 0 ] [ 'fConc' ] , color = '#0C1B33' , bins = 10 , label = 'FConc Gamma Rays'

plt.hist ( DF [ DF [ 'class' ] == 1 ] [ 'fConc1' ] , color = '#407899' , bins = 10 , label = 'Fconc1 Hadron Ray

plt.hist ( DF [ DF [ 'class' ] == 0 ] [ 'fConc1' ] , color = '#D57A66' , bins = 10 , label = 'Fconc1 Gamma Rays

3000

2500

1500

1000

500

\* This implies that the Cherenkov light emitted by gamma rays tends to be more concentrated in a

\* The Cherenkov light emitted by the high-energy electrons and positrons in the electromagnetic

\* The higher energy of gamma rays compared to most Hadron rays contributes to more extensive and

electromagnetic cascades. The higher energy deposition in the atmosphere leads to a more

\* Gamma rays have no rest mass, whereas Hadron rays (e.g., protons, neutrons) have rest masses

\* This mass difference affects the energy deposition and interactions of the particles in the

\* The higher 'FConc1' values for gamma rays suggest that gamma-ray showers tend to have a more

and focused light emission, with the highest pixel being a major contributor to the total

\* The lower 'FConc1' values for Hadron rays indicate that the distribution of Cherenkov light in

\* The difference in 'FConc1' can be attributed to the varying energy and particle interactions of

\* Higher-energy gamma rays might produce more intense showers, leading to higher 'FConc1' values.

images is more evenly spread out, with no single pixel dominating the emission.

sns.heatmap ( Correlation Matrix , annot = True , fmt = '0.2f' , cmap = 'pink' );

influencing the characteristics of the generated showers and, consequently, the 'Fconc'

more concentrated within a smaller spatial area. This concentration of Cherenkov light results in

within the shower image, while Hadron rays exhibit a more spread-out distribution.

Fconc1 Hadron Rays

Fconc1 Gamma Rays

The ratio of the value of the highest pixel over fSize

\* Gamma-ray showers, being more focused and intense, might have higher 'Fconc1' values, while

\* 'Fconc1' is calculated by dividing the value of the highest pixel intensity by 'FSize.'

between gamma-ray events and background events (non-gamma-ray events).

plt.xlabel ( "The ratio of the sum of the two highest pixels' values over fSize" )

FConc Hadron Rays

FConc Gamma Rays

plt.xlabel ( "The ratio of the value of the highest pixel over fSize" )

The ratio of the sum of the two highest pixels' values over fSize

'Fconc' value for gamma rays.

gamma rays, thus increasing 'Fconc.'

are more diffuse, might have lower 'Fconc1' values.

cascade produces a

**FConc** 

shower image."

dividing that

within a

FConc1

highest

intensity.

in the pixel

differentiate

background showers, which

In [11]: plt.figure ( figsize = ( 15 , 5 ) )

plt.subplot(1, 2, 1)

plt.legend ( )

plt.legend ( )

plt.show ( )

1000

500

**INFERENCE** 

smaller area

a higher

energetic

measurements.

intense, localized,

observed by the telescope.

Hadron rays with the Earth's atmosphere.

Cherenkov light

Hadron shower

gamma rays and

**CORRELATION HEAT MAP** 

plt.figure ( figsize = ( 20 , 20 ) )

In [12]: Correlation Matrix = DF.corr ()

FCONC1

cascade tends to be

concentrated shower for

**FCONC** 

plt.subplot(1, 2, 2)

sum by 'FSize.'

smaller area of the shower image.

wide and elongated shower image.

COMPARISON OF FConc (PIXEL RATIO) OF HADRON AND GAMMA RAYS



	<pre><axessubplot:> The performance parameters of the logistic regression model :  Values (%) Accuracy 93.56 Precision 92.16 Recall 95.66 F1_Score 93.88 FPR 0.09</axessubplot:></pre>
	- 3000 - 2838 270 - 2000
	- 144 3175 - 1000 - 500
In [27]:	<pre>ENSEMBLE TECHNIQUES : BOOSTING  GRADIENT BOOSTING  HYPER PARAMETER TUNING  GB = GradientBoostingClassifier() Parameters = {</pre>
	<pre>'learning_rate': [0.1 , 0.01 ],  'n_estimators': [ 150 , 200 ],  'max_depth': [ 7 , 9 ]  }  GS = GridSearchCV ( estimator = GB , param_grid = Parameters , cv = 5 )  GS.fit ( Train_X, Train_Y )  print("Best Hyperparameters : ", GS.best_params_ )</pre>
	Best_Model = GS.best_estimator_  Validation_Accuracy = Best_Model.score ( Test_X, Test_Y )  print( "\nValidation Set Accuracy with Best Model : ", Validation_Accuracy)  Best Hyperparameters : {'learning_rate': 0.1, 'max_depth': 9, 'n_estimators': 200}  Validation Set Accuracy with Best Model : 0.9349618795705616  PREDICTION
In [28]:	<pre>GB = GradientBoostingClassifier ( **GS.best_params_ ) GB.fit ( Train_X , Train_Y ) Y_GB = GB.predict ( Test_X ) Accuracy_GB = accuracy_score ( Test_Y , Y_GB ) * 100 Precision_GB = precision_score ( Test_Y , Y_GB ) * 100 Recall_GB = recall_score ( Test_Y , Y_GB ) * 100</pre>
	<pre>F1_Score_GB = f1_score ( Test_Y , Y_GB ) * 100 Confusion_GB = confusion_matrix ( Test_Y , Y_GB ) FPR_GB = Confusion_GB[0][1] / ( Confusion_GB [0][0] + Confusion_GB [0][1] ) GB_Performance_Metrics = { 'Accuracy ' : Accuracy_GB , 'Precision ' : Precision_GB , 'Recall ' : Recall_GB ,</pre>
	<pre>display ( sns.heatmap ( pd.DataFrame ( Confusion_GB ) , annot = True , fmt = '0.0f' , cmap = 'RdGy' ) ); print ( "The performance parameters of the logistic regression model : \n" ) display ( GB_Performance_Metrics )  <axessubplot:> The performance parameters of the logistic regression model :  Values(%)  Accuracy 93.70  Decided the performance of the logistic regression model :</axessubplot:></pre>
	Precision 92.58  Recall 95.45  F1_Score 93.99  FPR 0.08  - 3000  - 2500
	- 2854 254 - 2000 - 2000 - 1500 - 1000
	EXTREME GRADIENT BOOSTING  HYPER PARAMETER TUNING
In [29]:	<pre>XGB = XGBClassifier ( ) Parameters = {     'learning_rate': [0.1, 0.01],     'max_depth': [3, 5, 7],     'n_estimators': [100, 200, 300],     'gamma': [0, 0.1, 0.2], }</pre>
	<pre>GS = GridSearchCV ( estimator = XGB , param_grid = Parameters , cv = 5 )  GS.fit ( Train_X, Train_Y )  print("Best Hyperparameters : ", GS.best_params_ )  Best_Model = GS.best_estimator_  Validation_Accuracy = Best_Model.score ( Test_X, Test_Y )  print( "\nValidation Set Accuracy with Best Model : ", Validation_Accuracy)  Best Hyperparameters : {'gamma': 0.2, 'learning_rate': 0.1, 'max_depth': 7, 'n_estimators': 300}</pre>
In [30]:	<pre>Validation Set Accuracy with Best Model : 0.9190913334370624  PREDICTION  XGB = XGBClassifier ( **GS.best_params_ )  XGB.fit ( Train_X , Train_Y )  Y_XGB = XGB.predict ( Test_X )  Accuracy_XGB = accuracy_score ( Test_Y , Y_XGB ) * 100</pre>
	<pre>Precision_XGB = precision_score ( Test_Y , Y_XGB ) * 100  Recall_XGB = recall_score ( Test_Y , Y_XGB ) * 100  F1_Score_XGB = f1_score ( Test_Y , Y_XGB ) * 100  Confusion_XGB = confusion_matrix ( Test_Y , Y_XGB )  FPR_XGB = Confusion_XGB[0][1] / ( Confusion_XGB [0][0] + Confusion_XGB [0][1] )  XGB_Performance_Metrics = { 'Accuracy ' : Accuracy_XGB , 'Precision ' : Precision_XGB , 'Recall ' : Recall_XGB</pre>
	<pre>XGB_Performance_Metrics = pd.DataFrame ( np.round ( list(XGB_Performance_Metrics.values()), 2 ) , index = list() display ( sns.heatmap ( pd.DataFrame ( Confusion_XGB ) , annot = True , fmt = '0.0f' , cmap = 'RdGy' ) ); print ( "The performance parameters of the logistic regression model : \n" ) display ( XGB_Performance_Metrics ) </pre> <pre><axessubplot:> The performance parameters of the logistic regression model :</axessubplot:></pre> <pre>Values(%)</pre>
	Accuracy       91.91         Precision       92.40         Recall       91.90         F1_Score       92.15         FPR       0.08
	- 2500 - 2000 - 1500
	269 3050 - 1000 - 500 COMPARING THE MODELS
In [33]:	<pre># Classification algorithm names algorithm_names = ['Logistic Regression', 'Decision Tree', 'Random Forest', 'Gradient Boosting', 'Extreme Grad # Accuracy scores accuracy_scores = [ Accuracy_LR , Accuracy_DT , Accuracy_RF , Accuracy_GB , Accuracy_XGB ] # Precision scores</pre>
	<pre>precision_scores = [ Precision_LR , Precision_DT , Precision_RF , Precision_GB , Precision_XGB ]  # Recall scores  recall_scores = [ Recall_LR , Recall_DT , Recall_RF , Recall_GB , Recall_XGB ]  # Plotting the scores  plt.figure( figsize = ( 15 , 10 ) )</pre>
	<pre># Accuracy scores plot  plt.plot(algorithm_names, accuracy_scores, label = 'Accuracy', marker = 'o', color = '#42253B')  for i in range(len(algorithm_names)):     plt.text(algorithm_names[i], accuracy_scores[i], f'{accuracy_scores[i]:.2f}', ha ='center', va = 'bottom')  # Precision scores plot  plt.plot(algorithm_names, precision_scores, label='Precision', marker='o', color = '#9E4770')  for i in range (langerithm_names));</pre>
	<pre>for i in range ( len( algorithm_names ) ):     plt.text( algorithm_names [ i ] , precision_scores[ i ], f'{precision_scores[i]:.2f}', ha='center', va='bot  # Recall scores plot  plt.plot(algorithm_names, recall_scores, label='Recall', marker='o', color='#003459')  for i in range(len(algorithm_names)):     plt.text(algorithm_names[i], recall_scores[i], f'{recall_scores[i]:.2f}', ha='center', va='bottom')</pre>
	<pre># Labeling the axes and title plt.xlabel( 'Classification Algorithms' ) plt.ylabel( 'Scores' ) plt.title( 'Performance Comparison' )</pre>
	<pre># Adding a legend plt.legend()  # Rotating the x-axis labels for better visibility plt.xticks( rotation = 45 )  # Displaying the plot plt.show ( )</pre>
	95 66 95 45 93 56 93 70 93 16 92 16  Performance Comparison   Accuracy Precision Recall  92 16  93 70 92 16
	87,02 88 - 79,97
	76.71
In [34]:	Classification Algorithms  Classification Algorithms  Performance = pd.DataFrame ( index = algorithm_names )  Performance [ "Accuracy" ] = accuracy_scores  Performance [ "Precision" ] = precision_scores  Performance [ "Recall" ] = recall_scores
Out[34]:	Performance           Accuracy         Precision         Recall           Logistic Regression         76.707640         79.967105         73.244953           Decision Tree         89.466314         87.023543         93.552275           Random Forest         93.558425         92.162554         95.661344           Gradient Boosting         93.698460         92.577440         95.450437           Extreme Gradient Boosting         91.909133         92.396244         91.895149