

# Lecture 21

## CH-4114

### Molecular Simulation

**"Everything that living things do can be understood in terms of the jiggings and wiggings of atoms."**

**- Richard P. Feynman**

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# Pressure Calculation, Coupling and Concept of Barostat:

## The NPT Ensemble and Volume Fluctuation

- **NPT Ensemble (Isothermal-Isobaric):** This ensemble maintains a constant number of particles (**N**), constant pressure (**P**), and constant temperature (**T**). To keep the pressure constant, the **volume (V) of the simulation box is allowed to fluctuate**.
- **Isothermal Compressibility ( $\beta_T$ ):** The extent to which the volume fluctuates is an inherent property of the system, mathematically defined by the isothermal compressibility:

$$\beta_T = -\frac{1}{V} \left( \frac{\partial V}{\partial P} \right)_T$$

## Weak Coupling to a Pressure Bath

The core idea of the Berendsen method is to "**weakly couple**" the system's pressure,  $P$ , to an external reference pressure,  $P_0$ . This coupling is designed to make the pressure relax exponentially towards the target pressure.

- **Relaxation Term:** The text uses the equation:

$$\frac{dP}{dt} \propto \frac{P_0 - P}{\tau_P}$$

This means the rate of change of pressure ( $\frac{dP}{dt}$ ) is proportional to the difference between the target pressure ( $P_0$ ) and the current pressure ( $P$ ), scaled by a **pressure relaxation time constant** ( $\tau_P$ ). The term  $\tau_P$  dictates how quickly the system pressure reacts to the bath.

## Berendsen Pressure Coupling

### The Target: Exponential Pressure Relaxation

The Berendsen method postulates that the system's instantaneous pressure ( $P$ ) should relax toward the target pressure ( $P_0$ ) exponentially with a time constant  $\tau_P$ :

$$\left. \frac{dP}{dt} \right|_{\text{bath}} = \frac{P_0 - P}{\tau_P} \quad (1)$$

**Recall how this is analogous to Berendsen Temperature Coupling**

$$\frac{dT}{dt} = \frac{1}{\tau_T} (T_0 - T)$$

### Solve for $T(t)$

This differential equation has an exponential relaxation solution:

$$T(t) = T_0 + [T(0) - T_0]e^{-t/\tau_T}$$

So the system temperature relaxes exponentially toward the bath temperature  $T_0$ .

### Velocity rescaling form (used in MD)

Since  $E_k \propto v^2 \propto T$ , controlling  $T$  means **rescaling velocities** at each time step:

Let  $\lambda$  be the scaling factor for velocities at time  $t$ :

$$\mathbf{v}'_i = \lambda \mathbf{v}_i$$

Then  $T' = \lambda^2 T$ .

# Berendsen Pressure Coupling

## The Mechanism: Coordinate and Volume Scaling

In Molecular Dynamics (MD), the pressure is changed by continuously scaling the simulation box volume ( $V$ ).

The Berendsen barostat applies an extra velocity term to the equation of motion for each coordinate  $\mathbf{x}$  (where  $\mathbf{x}$  is  $x, y, z$  for a given particle), proportional to a scaling factor  $\alpha$ :

$$\dot{\mathbf{x}} = \mathbf{v} + \alpha\mathbf{x} \quad (2)$$

The rate of change of a linear dimension (e.g., box length  $L$ ) is  $\dot{L} = \alpha L$ .

For a 3D isotropic system, the rate of change of the **volume** ( $V = L^3$ ) is:

$$\frac{dV}{dt} = \frac{d(L^3)}{dt} = 3L^2 \frac{dL}{dt} = 3L^2(\alpha L) = 3\alpha L^3$$

$$\frac{dV}{dt} = 3\alpha V \quad (3)$$

### 3. The Constraint: Isothermal Compressibility

Thermodynamically, the change in volume is related to the change in pressure via the **isothermal compressibility** ( $\beta_T$ ):

$$\beta_T = -\frac{1}{V} \left( \frac{\partial V}{\partial P} \right)_T$$

Rearranging this to relate the time derivatives  $\frac{dP}{dt}$  and  $\frac{dV}{dt}$ :

$$\frac{dV}{dt} = -\beta_T V \frac{dP}{dt} \quad (4)$$

*(Note: This relationship assumes the system is close to equilibrium, which is the underlying approximation of the Berendsen method).*

## Combining the Terms (Derivation of $\alpha$ )

Now, we equate the two expressions for the rate of change of pressure ( $\frac{dP}{dt}$ ):

- Substitute the coupling requirement (1) for  $\frac{dP}{dt}$  into the thermodynamic relation (4):

$$\frac{dV}{dt} = -\beta_T V \left( \frac{P_0 - P}{\tau_P} \right)$$

$$\frac{dV}{dt} = \frac{\beta_T V}{\tau_P} (P - P_0) \quad (5)$$

- Finally, equate the volume change derived from the scaling (3) with the combined expression (5):

$$3\alpha V = \frac{\beta_T V}{\tau_P} (P - P_0)$$

- Solving for  $\alpha$  by dividing both sides by  $3V$ :

$$\alpha = \frac{\beta_T}{3\tau_P} (P - P_0)$$

## Key Insight

$$\alpha = \frac{\beta_T}{3\tau_P}(\mathbf{P} - \mathbf{P}_0)$$

- This factor  $\alpha$  (alpha) is a simple scalar that dictates the velocity by which all coordinates and box vectors are scaled at each time step to drive the pressure toward the target value  $P_0$ .
- Interpretation of the Terms:  $(P - P_0)$ : The sign of  $\alpha$  (alpha) depends on the pressure difference. If  $P > P_0$ ,  $\alpha$  is positive, leading to an expansion ( $\dot{x} > 0$ ), which reduces the pressure.
- $\tau_p$ : The coupling time constant. A smaller  $\tau_p$  results in a larger  $\alpha$ , meaning stronger, faster coupling.
- $\beta_T$ : The isothermal compressibility. A system that is highly compressible (large  $\beta_T$ ) requires a larger  $\alpha$  for a given pressure change.

## Computing Pressure: The Virial Theorem

The final expression for pressure in an MD simulation that relates macroscopic pressure ( $P$ ) to microscopic properties (particle positions  $\mathbf{r}$  and forces  $\mathbf{F}$ ) is:

$$P = \frac{NkT}{V} + \frac{1}{3V} \sum_{i < j} \langle \mathbf{r}_{ij} \cdot \mathbf{F}_{ij} \rangle$$

Origin: Statistical  
Mechanics/Mayer's  
Cluster Expansion

This formula effectively breaks the pressure into two intuitive components:

### 1. The Ideal Gas Component ( $P_{\text{Ideal}}$ )

This component accounts for the kinetic energy of the particles, representing the pressure that would exist even if the particles had no interactions (an ideal gas).

- **Starting Point:** The ideal gas law is  $PV = NkT$ .

## 2. The Interaction Component ( $P_{\text{Interaction}}$ )

This component accounts for the effects of inter-particle forces ( $\mathbf{F}_{ij}$ ) on the system pressure. This is where the **Virial Theorem** comes in.

- **The Virial ( $\mathbf{W}$ ):** In classical mechanics, the Virial is defined as the sum of the product of coordinates and forces. The Virial Theorem connects the system's kinetic energy ( $\sum \frac{1}{2} m_i \mathbf{v}_i^2 = \frac{3}{2} N k T$ ) to the system's Virial  $W$ :

$$\sum_i \langle \mathbf{r}_i \cdot \mathbf{F}_i \rangle = -3PV$$

*This is the core result that links microscopic forces to macroscopic pressure,  $P$ .*

- **Breaking Down the Force ( $\mathbf{F}_i$ ):** The total force  $\mathbf{F}_i$  on any particle  $i$  has two origins:
  1. **External Force ( $\mathbf{F}_i^{\text{ext}}$ ):** The force exerted by the walls of the container, which is directly related to the system pressure  $P$ .
  2. **Internal Force ( $\mathbf{F}_i^{\text{int}}$ ):** The forces exerted on particle  $i$  by all other particles  $j$  (i.e.,  $\mathbf{F}_i^{\text{int}} = \sum_{j \neq i} \mathbf{F}_{ij}$ ).

## Virial Contribution to Pressure in Molecular Dynamics or Monte Carlo Simulations Using Truncated Lennard-Jones (LJ) potentials.

- In simulations, we can't sum over all infinite pairs of particles — so we **truncate** the potential at a cutoff  $r_c$ .  
However, there are still weak attractive interactions beyond  $r_c$  that **contribute to energy and pressure**.
- In our early classes we have shown energy correction.
- To estimate and fix pressure now, we compute their **mean-field (analytical)** contribution assuming that Beyond  $r_c$ , the radial distribution function  $g(r) \approx 1$  i.e., particles are uniformly distributed).
- This approximation gives rise to “tail corrections” for **energy** and **pressure**.

### Virial Expression for Pressure:

In statistical mechanics, the pressure  $P$  of a system can be related to the virial as:

$$P = \rho k_B T + \frac{1}{3V} \left\langle \sum_{i < j} \mathbf{r}_{ij} \cdot \mathbf{F}_{ij} \right\rangle$$

When using a Lennard-Jones potential:

$$U(r) = 4\epsilon \left[ \left( \frac{\sigma}{r} \right)^{12} - \left( \frac{\sigma}{r} \right)^6 \right]$$

And the corresponding force  $F(r) = -\frac{dU}{dr}$ , the virial integral for  $r > r_c$  can be evaluated analytically under the assumption of a uniform particle distribution beyond  $r_c$ .

## Plug into the Pressure Tail Integral

$$P_{\text{tail}} = -\frac{2\pi}{3}\rho^2 \int_{r_c}^{\infty} r^3 \frac{dU(r)}{dr} dr$$

Substitute and simplify:

$$P_{\text{tail}} = \frac{8\pi\epsilon}{3}\rho^2 \int_{r_c}^{\infty} \left[ 12\frac{\sigma^{12}}{r^{10}} - 6\frac{\sigma^6}{r^4} \right] dr$$

Integrate term by term:

$$\int_{r_c}^{\infty} r^{-10} dr = \frac{1}{9r_c^9}, \quad \int_{r_c}^{\infty} r^{-4} dr = \frac{1}{3r_c^3}$$

$$P_{\text{tail}} = \frac{8\pi\epsilon}{3}\rho^2 \left[ 12\frac{\sigma^{12}}{9r_c^9} - 6\frac{\sigma^6}{3r_c^3} \right]$$

### ■ Tail Correction to Pressure (WLRC):

$$P_{\text{tail}} = \frac{16}{9} \pi \rho^2 \sigma^3 \epsilon \left( 2 \left( \frac{\sigma}{r_c} \right)^9 - 3 \left( \frac{\sigma}{r_c} \right)^3 \right)$$

This corresponds to the formula:

$$\text{WLRC} = \frac{16}{9} \pi \rho N \sigma^3 \left( 2 \left( \frac{\sigma}{r_c} \right)^9 - 3 \left( \frac{\sigma}{r_c} \right)^3 \right)$$

Where:

- $\rho = N/V$  is the number density
- $\sigma$ : LJ parameter
- $r_c$ : cutoff radius
- $\left( \frac{\sigma}{r_c} \right)^3 = SR3$ , and  $\left( \frac{\sigma}{r_c} \right)^9 = SR9$  **Code Implementation**
- $\epsilon$  is absorbed in unit conversion or set to 1 (reduced units)

This is derived by integrating the virial component of the force over the spherical shell beyond  $r_c$  assuming a constant density (mean-field approximation).

# Dimensions and Units 1. Magnitudes

- Important extensive quantities small in magnitude
  - when expressed in macroscopic units
- Small numbers are inconvenient
- Two ways to magnify them
  - work with atomic-scale units
    - ps, amu, nm or Å
  - make dimensionless with characteristic values
    - model values of size, energy, mass

Symbol	Definition	Value
<i>1. Constants</i>		
$k$	Boltzmann's constant	$1.3806 \times 10^{-23} \text{ J/(molec}\cdot\text{K)}$
$N_0$	Avagadro's number	$6.022 \times 10^{23}$
<i>2. Simulation Variables</i>		
$N$	Number of molecules	$\sim 10^3$
$V$	Simulation cell volume	$\sim 10^{-24} \text{ m}^3$
$m$	Molecular mass	$\sim 10^{-25} \text{ kg/molec}$
$\rho$	Number density	$\sim 10^{27} \text{ molec/m}^3$
$E$	Energy (total)	$\sim 10^{-20} \text{ J/molec}$
$t$	time	$\sim 10^{-12} \text{ s}$
<i>3. Model Variables</i>		
$\sigma$	Size variable	$\sim 5 \times 10^{-10} \text{ m}$
$\epsilon$	Energy variable	$\sim 10^{-21} \text{ J/molec}$
$r_b$	Bond distance	$\sim 10^{-10} \text{ m}$
$k_v$	Vibrational spring constant	$\sim 10^3 \text{ J/m}^2$

## Dimensions and Units 2. Scaling

- In simulations it is often convenient to express quantities such as temperature, density, pressure and the like in reduced units. This means that we choose a convenient unit of energy, length and mass and then express all the other quantities in terms of these basic units. A natural choice of our basic units is the following

- size  $\sigma$
- energy  $\epsilon$
- mass  $m$

In terms of these  
basic units, all  
other units follow

Symbol	Meaning	Definition
$r^*$	dimensionless distance	$r/\sigma$
$E^*$	dimensionless energy	$E/\epsilon$
$T^*$	dimensionless temperature	$kT/\epsilon$
$U^*$	dimensionless internal energy	$U/\epsilon$
$t^*$	dimensionless time	$t/[\sigma(m/\epsilon)^{0.5}]$
$v^*$	dimensionless velocity	$v/(\epsilon/m)^{0.5}$
$F^*$	dimensionless force	$F\sigma/\epsilon$
$P^*$	dimensionless pressure	$P\sigma^3/\epsilon$
$D^*$	dimensionless self diffusion coefficient	$D/[\sigma(\epsilon/m)^{0.5}]$

## Why reduced Units?

- Many combinations of  $\rho$ ,  $T$ ,  $\epsilon$  and  $\sigma$  all correspond to the same state in reduced units. This is the law of corresponding states: the same simulation can of a LJ model can be used to study the Argon at 60 K and density 840 kg/m<sup>3</sup> and Xe at 112 K and a density at 1617 kg/m<sup>3</sup>. In reduced unit both simulations corresponds to the state point  $\rho = 0.5$  and  $T = 0.5$ . Scaling by model parameters
- In reduced units almost all quantities of interest are of order 1 (say between 10<sup>-3</sup> and 10<sup>3</sup>). Hence if we suddenly find very large (or very small) number in our simulations, suspect some error somewhere.
- Simulation results obtained in reduced units can be translated back into real units.

See the following table

Conversion of reduced Units to real Units for LJ argon system:  $\epsilon/k_B = 119.8 \text{ K}$ ,  
 $\sigma = 3.405 \times 10^{-10} \text{ m}$ ,  $m = 0.03994 \text{ kg/mol}$

Quantity	Reduced Units	Real Units
Temperature	$T^* = 1$	$T = \epsilon/k_B = 119.8 \text{ K}$
Density	$\rho^* = 1$	$\rho = 1680 \text{ kg/m}^3$
Time	$\delta t^* = 0.005$	$\delta t = 1.09 \times 10^{-14} \text{ s}$
Pressure	$P^* = 1$	$P = 41.9 \text{ MPa}$

Thank you for joining CH4114!

May you look back on this class and recognize the **valuable transformation** that took place here.

---- Susmita Roy