

320.

$$(a) v_{circ} = \sqrt{\frac{GM}{r}} = \sqrt{\frac{6.67E-11 \text{ mks} \times 2E30 \text{ kg}}{40 \times 1.5e11 \text{ m}}} = 4.72 \text{ km/s}$$

(b) similar calculation: $v_{circ} = 29.7 \text{ km/s}$

(c) In one day, the comet moves $4.72 \text{ km/s} \times 86400 \text{ s/day} = 4.08 \times 10^5 \text{ km}$. At a distance of 40 AU, this angle is $4.08E5 \text{ km} / (40 \times 1.5E8 \text{ km}) \text{ rad} = 6.8e-5 \text{ rad} = 6.8e-5 \text{ rad} \times 206265''/\text{rad} = 14.0''$. So the comet moves $14''/\text{day}$ forwards.

[You could have just used the formula for angular velocity to do this: $\omega = v / r$.]

(d) Similar calculation, except you use the earth's motion of 29.7 km/s .

A simple diagram will convince you that this motion is in the opposite direction to the answer in (c) - or see me - just like retrograde motion that we discussed in class. The angular velocity is 88.3 arcsec/day .

(e) The net motion of the comet is $88.3 - 14.0 = 74.3''/\text{day}$, in a direction opposite to the comet's motion. The earth's orbital motion dominates by far!

501.

$$\text{parallax["]} = 1 / d [\text{pc}]$$

Therefor $\text{parallax} = 1/250'' = 0.004''$. This is marginal. I said in the notes that you could measure parallax to $0.001''$; the text says $0.01''$. So either answer is correct. For reference I think the text meant to say "parallax can be measured reasonably accurately to 0.01 arcsec ".

502.

The simple way: Pluto is 39 times further away than the earth is from the sun, so parallax has to be 39 times bigger. Therefore the parallax angle is $39 \times 0.004'' = 0.156$ arcsec.

The more detailed way:

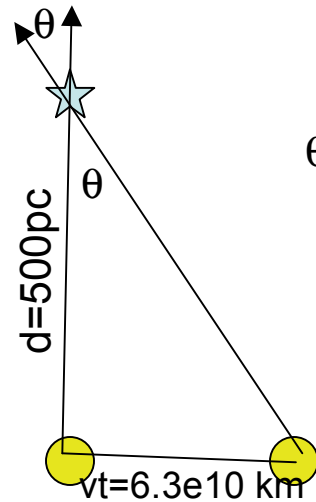
$$\Delta\theta = 39 AU / d_{AU} \text{ rad} = 39 \times 206265 / d_{AU} \text{ arcsec}$$

But $d_{AU} = 206265 d_{pc}$. So substituting above, the 206265's cancel out and we're left with:

$$\Delta\theta = 39 / d_{pc} \text{ arcsec, which for } d=250\text{pc works out to } 0.156 \text{ arcsec as above.}$$

504.

The sun moves $vt = 20 \text{ km/s} \times 100 \text{ yr} \times (3.16 \times 10^7 \text{ s/yr}) = 6.32 \times 10^{10} \text{ km}$ in 100 yr. You can ignore the earth's motion around the sun and the earth's rotation in all this. So the basic diagram looks like:



$$\begin{aligned}\theta &= vt/d \text{ rad} = 6.3 \times 10^{10} \text{ km} / (500 \text{ pc} \times 3.1 \times 10^{13} \text{ km/pc}) \\ &= 4.065 \times 10^{-6} \text{ rad} \\ &\times 206265 \text{ ''/rad} = 0.84 \text{ arcsec}\end{aligned}$$

[I realize that "change in position" was a bit ambiguous; what I meant to say was "change in angular position on the sky". So I will accept the answer $6.3 \times 10^{10} \text{ km}$.]